ROSETTA-ALICE TO PLANETARY SCIENCE ARCHIVE INTERFACE CONTROL DOCUMENT

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Contract JPL 960549

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REVISION NOTICE

Initial Issue: January 2004

Revision 1: January 2007 Major revision to reflect new format of Alice data products

Revision 2: April 2007 Multiple revisions to correct inaccuracies in the description of the Alice data products. Removed all TBD items from document

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TBD LIST

Section	Description

1. INTRODUCTION

1.1 Purpose and Scope

The purpose of this EAICD (Experimenter to (Science) Archive Interface Control Document) is twofold. First it provides users of the *Alice* instrument on the *Rosetta* mission with detailed description of the product and a description of how it was generated, including data sources and destinations. Secondly, it is the official interface between the *Alice* instrument team and the archiving authority.

1.2 Archiving Authorities

The Planetary Data System Standard is used as archiving standard by

- NASA for U.S. planetary missions, implemented by PDS
- ESA for European planetary missions, implemented by the Research and Scientific Support Department (RSSD) of ESA

1.2.1 ESA's Planetary Science Archive (PSA)

ESA implements an online science archive, the PSA,

- to support and ease data ingestion
 - to offer additional services to the scientific user community and science operations teams as e.g.
 - o search queries that allow searches across instruments, missions and scientific disciplines
 - o several data delivery options as
 - direct download of data products, linked files and data sets
 - ftp download of data products, linked files and data sets

The PSA aims for online ingestion of logical archive volumes and will offer the creation of physical archive volumes on request.

1.3 Contents

This document describes the data flow of the *Alice* instrument on *Rosetta* from the s/c until the insertion into the PDS-SBN for NASA and PSA for ESA. It includes information on how data were processed, formatted, labeled and uniquely identified. The document discusses general naming schemes for data volumes, data sets, data and label files. Standards used to generate the product are explained. Software that may be used to access the product are explained.

The design of the data set structure and the data product is given. Examples of these are given in the appendix.

1.4 Intended Readership

The staff of archiving authority (Planetary Data System for NASA, Planetary Science Archive for ESA) design team and any potential user of the *Rosetta-Alice* data.

1.5 Applicable Documents

AD1. Planetary Data System Data Preparation Workbook, 1995 February 17, Version 3.1, JPL, D-7669, Part 1

- AD2. Planetary Data System Standards Reference, 2003 August 1, Version 3.6, JPL, D-7669, Part 2
- AD3. Rosetta Archive Generation, Validation and Transfer Plan, 2003 December 22, Issue 2, Revision 1, RO-EST-PL-5011
- AD4. Rosetta-Alice Data Processing Software, 2003 April 22

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- AD6. Data Delivery Interface Document (DDID), 2003 October 23, Issue B6, RO-ESC-IF-5003
- AD7. Rosetta Time Handling, 2004 Januaryu 9, Draft 2, RO-EST-TN-3165

1.6 Relationships to Other Interfaces

Any change in this ICD may need to be reflected in the data processing software (LIMA, MIKE, and NOVEMBER) and the document that describes the software (AD4). Correspondingly, this ICD may need to be changed to reflect changes in the software or changes in the understanding of requirements for archiving imposed either by NASA or by NASA's Planetary DataSystem and ESA's Planetary Science Archive.

1.7 Acronyms and Abbreviations

APID	Application Process ID
DDS	Data Distribution System
FITS	Flexible Image Transport System
FOV	field-of-view
HK	house keeping
ICD	interface control document
MCP	microchannel plate
OBT	Onboard Time
PDS	Planetary Data System
PSA	Planetary Science Archive
SFDU	Standard Format Data Unit
SCET	Spacecraft Elapsed Time
TC	telecommand
ТМ	telemetry
UV	ultraviolet

1.8 Contact names and Addresses

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- Maarten Versteeg, Southwest Research Institute, mversteeg@swri.org, +1 210-522-5029

2. OVERVIEW OF INSTRUMENT DESIGN, DATA HANDLING, AND PRODUCT GENERATION

2.1 Instrument Design

2.1.1 Hardware and Characteristics

The *Alice* UV spectrometer is a very simple instrument. An opto-mechanical layout of the instrument is shown in Figure 2-1. Light enters the telescope section through a 40 x 40 mm² entrance aperture and is collected and focused by an f/3 off-axis paraboloidal (OAP) primary mirror onto the approximately 0.1° x 6° spectrograph entrance slit. After passing through entrance slit, the light falls onto a toroidal holographic grating, where it is dispersed onto a microchannel plate (MCP) detector that uses a double-

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delay line (DDL) readout scheme. The 2-D (1024 x 32)-pixel format, MCP detector uses dual, side-byside, solar-blind photocathodes of potassium bromide (KBr) and cesium iodide (CsI). The predicted spectral resolving power ($\lambda/\Delta\lambda$) of *Alice* is in the range of 105–330 for an extended source that fills the instantaneous field-of-view (IFOV) defined by the size of the entrance slit. *Alice* is controlled by an SA 3865 microprocessor, and utilizes lightweight, compact, surface mount electronics to support the science detector, as well as the instrument support and interface electronics. Figure 2-22 shows a 3D external view of *Alice*. The resulting design is highly systems-engineered to minimize mass and complexity, and enjoys strong parts-level heritage from previous UV spectrometers.



Figure 2-1: The opto-mechanical layout of Alice.

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Figure 2-2: External view of Alice.

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Table 2-1: Pre-flight <i>Alice</i> Characteristics & Performance Overview			
Bandpass	700-2050 Å		
Spectral Resolution (Δλ FWHM)	10 Å (700 Å) - 13 Å (2050 Å) for extended sources 3–5 Å for point sources		
Spatial Resolution	0.1° x 0.5°		
Nominal Sensitivity	$0.5 (1900 \text{ Å}) - 7.8 (1150 \text{ Å}) \text{ counts s}^{-1} \text{ R}^{-1}$		
Field of View	0.1° x 6.0°		
Pointing	Boresight with OSIRIS, VIRTIS		
	Nucleus Imaging and Spectroscopy; Coma Spectroscopy		
Observation Types	Jet and Grain Spectrophotometry; Stellar Occultations (secondary observations)		
Telescope/Spectrograph	Off-axis telescope, 0.15-m diameter Rowland circle spectrograph		
Detector Type	2-D Microchannel Plate with delay-line anode readout		

2.1.2 Science Modes

The instrument has three modes of taking data: **histogram**, **pixel list**, and **count rate** ("photometer"). For the first two, the time of arrival of a charge at both ends of the orthogonal delay lines is used to determine the pixel on which a photon landed.

- In **histogram** mode, the instrument uses a histogram memory, with memory locations that map oneto-one to pixels on the detector, to store the total number of counts in each pixel. This mode integrates an image and is used at high count rates, and will likely be the most commonly-used mode. Images are 1024x32 pixels (the long axis corresponds to wavelength, the short axis is the spatial dimension), but the first 100 columns of the image are and "unexposed" part of the detector. This region contains the pulse height array, and also will provide monitoring of background/cosmic ray and dark counts. The histogram values (counts) saturate at 65535 counts.
- In **pixel list** mode, the instrument records the x-y location of each photon and inserts time tags in this list at regular (user-definable) intervals to provide the timing information for the photons. This mode is used at low count rates. The maximum number of entries is 32678 (including photons and time tags), but may be less depending on the brightness of the target and the exposure time of the particular observation. Once the memory is filled to this limit, no more data are recorded.
- In **count rate** mode (also called "photometer" mode), the x-y information for each photon is ignored and a series of counts is summed with time tags inserted at uniform intervals to provide the timing information. This mode is likely to not be widely used but is particularly useful for observing very bright ultraviolet stars, e.g. at the limb of the cometary nucleus. The maximum number of entries is 32678, but may be less depending on the time interval and total exposure time of the particular

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observation. (Note that the number of entries is independent of the the brightness of the target.) Values (summed counts per time interval) saturate at 65535 counts.

In all modes, data from portions of the array can be "windowed". E.g., in histogram mode, transmission of data to Earth can be of subarrays rather than of the complete array. This allows for minimizing bandwidth by reading subsets (windows) of data stored in memory. The data processing software will consider each of these windows to be separate data sets, so that multiple reads of separate windows of a single observation will be saved into separate files. For this reason, the file naming convention (Section 3.1.4) includes a "dump number" for each file/image.

2.2 Scientific Objectives

See *Alice* EID-B, § 1.1.1.

The scientific objectives of the *Alice* investigation are to characterize the composition of the nucleus and coma, and the coma/nucleus coupling of comet 67P/Churyumov-Gerasimenko. This will be accomplished through the observation of spectral features in the 700-2050 Å EUV/FUV region. *Alice* will provide measurements of noble gas abundances in the coma, the atomic budget in the coma, major ion abundances in the tail and in the region where solar wind particles interact with the ionosphere of the comet, determine the production rates, variability, the structure of H₂O, CO and CO₂ gas surrounding the nucleus, and the far-UV properties of solid grains in the coma. *Alice* will also map the cometary nucleus in the FUV, and study Mars and the *Rosetta* asteroid flyby targets while en route to 67P/Churyumov-Gerasimenko.

Ultraviolet spectroscopy is a powerful tool for studying astrophysical objects, and has been applied with dramatic success to the study of comets. *Alice* will provide unprecedented improvements in sensitivity and spatial resolution over previous cometary UV observations. For example, *Alice* will move the sensitivity threshold from the ~1 Rayleigh level achievable with the Hubble Space Telescope to the milliRayleigh level in deep integrations. In addition, *Alice* will (by virtue of its location at the comet) move the spatial exploration of nucleus UV surface properties from the present-day state-of-the-art (i.e., no data available on any comet) to complete nuclear maps at Nyquist-sampled resolutions of a few hundred meters. Stars occulted by the absorbing coma will also be observed and used to map the water molecule spatial distribution, giving us hints at how the production regions are located on the nuclear surface.

Through its remote-sensing nature, Alice will be able to:

- Obtain compositional and morphological information on the comet prior to the rendezvous, thereby providing planning observations for in situ instruments prior to entering orbit about the comet.
- Map the spatial distribution of key species in the coma, and small coma dust grains, as a function of time as the comet responds to the changing solar radiation field during its approach to the Sun.
- Obtain compositional and production rate measurements of nuclear jets and other inner coma features even when the Orbiter is not in the vicinity of these structures.
- Obtain certain ion abundance measurements around perihelion in order to connect nucleus activity to changes in tail morphology and structure, and coupling to the solar wind.

The primary scientific themes of the *Alice* investigation are the following (not prioritized):

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- Determine the rare gas content of the nucleus to provide information on the temperature of formation and the thermal history of the comet since its formation. Ar and Ne will be prime targets of the *Alice* investigations.
- Determine the production rates and spatial distributions of the key parent molecule species, H₂O, CO and CO₂, thereby allowing the nucleus/coma coupling to be directly observed and measured on many time-scales in order to study the chemical heterogeneity of the nucleus and its coupling to the coma.
- Obtain an unambiguous budget of the cosmogonically most important atoms C, H, O, N, and S through the detection of their emissions far from the nucleus. This is required to understand their production processes and to derive the elemental composition of the volatile fraction of the nucleus. Coupled to the measure of the major molecule abundances of the nucleus, this will give us the total contribution of the secondary parent species to the compositional makeup of the nucleus.
- Study the onset of nuclear activity and nucleus output variations related to changing solar aspect and nuclear rotation with unprecedented sensitivity.

Additional scientific themes Alice will address include the following:

- Spectral mapping of the complete nucleus at far-UV wavelengths to characterize the distribution of UV absorbers on the surface, in particular icy patches and organics.
- Photometric properties and ice/rock ratio of small grains in the coma as an aid to understanding the size distribution of cometary grains and how they vary in time. Also, studying the grain coma to establish the relative contributions of the nucleus and coma grains to the observed gases.
- Mapping the time variability of O⁺, N⁺, and possibly S⁺ and C⁺ emissions in the coma and ion tail in order to connect nuclear activity to changes in tail morphology and structure, and tail interaction/coupling to the solar wind.

Alice will take data during cruise (typically checkout data during periods defined by ESA) and at every encounter with a planetary and asteroid body (within activity limitations imposed by ESA and the *Rosetta* project). Extensive observations will be made during the various phases at the comet. At each asteroid flyby, for example, *Alice* may spatially scan the region of the asteroid by moving the slit across the field thought to contain the asteroid. If the S/N is high, this could result in a complete image cube. However, we may also, for example, carry out a staring observation at the limb which would ultimately result in a single, one-dimensional spectrum. Observations at the encounters with Earth and Mars will be aimed at measuring the airglow from those planets. Observations during the first, closest Earth-Moon flyby will use the Moon as a calibration source.

At the comet, a variety of observing scenarios will be used. These include spatial scans with spatially resolved spectra at each point, 1-D spectra, and time series as stars are viewed through the coma and approaching occultation by the nucleus. This last can involve recording complete 2-D spectra as a function of time, but then collapsing the data in transmission to provide 1-D spectra, possibly integrating over significant wavelength regions, before transmission to Earth.

2.3 Data Handling Process

2.3.1 Responsible Individuals

The following individuals are responsible for the data preparation, monitoring and maintenance of the processing and calibration pipeline, and submission of the product delivery to the PSA:

- Andrew Steffl, Southwest Research Institute, steffl@boulder.swri.edu, +1 303-546-9670
- Joel Parker, Southwest Research Institute, joel@boulder.swri.edu, +1 303-546-9670

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• Maarten Versteeg, Southwest Research Institute, mversteeg@swri.org, +1 210-522-5029

2.3.2 Data Flow

The Alice instrument typically sends HK data on a regular time interval. When exposures are taken, the HK data available at the end of the exposure are included as a header to the science data to provide instrument context for the exposure. Events (command acknowledgements and error reports) are sent in response to TC. Dumps of memory and context/parameters are sent when requested by specific TCs. These different types of data from the *Alice* instrument are associated with the following APIDs:

- 1473 (acknowledgements)
- 1476 (HK packets)
- 1479 (events)
- 1481 (memory dump)
- 1483 (context dump)
- 1484 (science {includes diagnostic and parameter dump} and telecommands)
- -4 (the "super APID", that includes all *Alice* APIDs, so all data can be obtained from DDS in a single request)

Additional spacecraft APIDs are available, and used by *Alice* software, but will not be included in the archived *Alice* data products.

The *Alice* data are sent to the *Rosetta* spacecraft, relayed to ground, and stored in the DDS. These data are sent ("pushed") to computers at the *Alice* team institution in Boulder, Colorado at regular, pre-defined intervals (e.g., daily), and also by direct request to the DDS via FTP or web interface (see AD6). The data are passed through instrument team software for calibration and analysis. These pipeline programs can be run interactively, but typically will be called by a UNIX shell script whenever the DDS regularly pushes *Alice* data.

Auxiliary programs will be run on the data to create the appropriate PDS labels and format for archiving.

2.3.3 Processing Software

1. The *Rosetta-Alice* data processing software is used internally by the instrument team, and is included in the DOCUMENT/CODE/MIKE directory only for reference (i.e., it is *not* included in the archive *as supported software*). The software consists of programs to process and calibrate the science data. Steps include flat fielding, deadtime correction, and applying wavelength and flux calibration. This is performed by a set of IDL programs called "Mike".

2.3.4 Archive Data Levels

There are two types of notation for "processing levels" in use in the *Rosetta* mission: those defined by the ESA PSA, and those defined by the National Research Council (NRC) Committee on Data Management and Computation (CODMAC). The PDS uses the CODMAC data processing levels as defined in Section 6.3.2 in AD2. For consistency, we use the CODEMAC levels (unless explicitly datted otherwise) in the text of this EAICD and in the descriptions of the PDS data labels in Section 4. The relationship of the levels in question are given in Appendix B of AD3, but are repeated and expanded here:

Table 2-2: Data Processing Levels					
CODMAC level	ESA/PSA level		PRODUCT_TYPE PDS keyword value	Data Record (DR) Type	<i>Alice</i> File Type

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L	1	1a		UDR	Unprocessed (Raw)	1
	2	1b	0	EDR	Experiment (Edited)	engineering (eng)
	3	2	1A	RDR or REDR	Reduced (Calibrated)	science (sci)
	4		1B	REFDR	Reformatted (Resampled)	
	5	3	2-5	DDR	Derived	
	6			ANCDR	Ancillary	

The *Alice* data archive will contain:

- Level 2 data, which are derived from the (Level 1?) data sent by the DDS to the PI's institution in Boulder, Colorado, USA (to the computer alice.boulder.swri.edu), and decommutated by the *Rosetta-Alice* data collection program Lima.
- Level 3 data. These calibrated data are created by the data processing program Mike from the Level 2 data (as output by Lima).

No Level 3 or higher data are expected to be produced for the *Alice* archive.

2.4 Overview of Data Products

2.4.1 Pre-Flight Data Products

N/A

2.4.2 Instrument Calibrations

Instrument calibration files are included in the CALIB/ directory.

2.4.3 In-Flight Data Products

As described in Section 2.1.2, *Alice* has three science modes: histogram, pixel list, and count rate. Each mode corresponds to one type of data product that will be archived as FITS files, and the archive will include CODMAC Level 2 ("engineering") and Level 3 ("science") data for each mode.

The data from each observation also include a full set of HK data, providing a snapshot of the instrument status *at the end of the given observation* (though some parameters, such as the MCP voltage and strip current also include the value at the beginning of the observation). In the FITS files output by the processing software, those HK parameters are included as keyword=value header fields as shown in the example in Section 4.8 (but only a few are propagated into the PDS label files, since many of them are very instrument-specific and not of general use to PDS users). These header fields include raw *and* calibrated parameter values where appropriate, distinguished by the last character of the FITS header keyword: "R" for raw and "C" for calibrated. For example, T_MIRR1R is the "raw" (ADC counts) temperature value for mirror sensor #1, and T_MIRR1C is the corresponding "calibrated" (degrees Celsius) temperature; MCPVR is the "raw" (ADC counts) MCP voltage, and MCPVC is the "calibrated" (volts) voltage.

We also include *all* HK data sent in the TM stream as an in-flight data product. These data are archived as flat ASCII files (tables), and provide a larger context of the instrument health and behavior over the course of the mission, and are also useful for calibration of the instrument as a function of any parameters that may affect instrument performance.

Below we describe all these data products in considerable detail. Examples of the PDS labels are given in Section 4. Detailed descriptions of the FITS headers are given in Section 4.8.

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2.4.3.1 Level 2 ("Engineering") data

2.4.3.1.1 Histogram Mode

Each dump of histogram data from the *Alice* instrument also contains pulse-height data (16 integers) and count rate data (up to 100 integers). These arrays will be stored as additional blocks ("extensions") in the FITS file. So, the FITS file for histogram data will contain:

- Extension = 0 : The 1024x32 histogram image (or a subsection if windowing was selected in the dump)
- Extension = 1 : The 16-value pulse-height distribution table. The pulse height distribution is a histogram of the amplitude of the photon events that were detected during the histogram acquisition. The histogram values (counts) saturate at 65535 counts. For *Rosetta-Alice* the resolution is rather coarse: the quantization is four bits, but the upper and lower 3 bins are hardly used.
- Extension = 2 : Up to 100-element count rate series. These count rate data give the count rate measurements as reconstructed from the observation, providing a crude time-resolved "photometer" measurement of the observation. This is possible because even in Histogram mode because the photons are time-tagged.

2.4.3.1.2 Pixel List Mode

Pixel list data will be stored as a 1-dimensional array of size equal to the number of values in the pixel list (number of entries may be less than the full 32678 limit depending on the particular observation) in extension 1. For ease of quick-look, the extension 0 in the FITS file will contain a "ground-software-reconstructed" histogram (image) representation of the data, so any FITS viewer can display the image. In addition, a count rate representation (using the pixel list time hack period) will be created and stored in extension 2. So, the FITS file for pixel list data will contain:

- Extension = 0: The 1024x32 reconstructed histogram-formatted image
- Extension = 1 : The list of all the pixel list values
- Extension = 2 : reconstructed count rate data with a time interval equal to the selected time hack period

The value of each pixel list event is the encoded X,Y positions of the detected event. In addition to the detector events, a time-mark value is also inserted into the list on a periodic basis. The pixel list data are encoded as follows:

Value (for events) = X * 1024 + Y

Value (for timehacks) = 65535

Another way to describe this is by representing the bits in the pixel list events (16-bit words) as:

haaaaaeeeeeeeee

where each letter represents a single bit with the following meaning:

- h (most significant bit), set to 1 for a timehack, and 0 for a normal photon event. For timehacks, the other 15 bits have no meaning and are set to 1.
- a the five bits of spatial information of a photon event.

e - the ten bits of spectral information of a photon event.

2.4.3.1.3 Count Rate Mode

Data will be stored as a 1-dimensional FITS image with the number of elements equal to the number of intervals for which a count rate was obtained (the number of elements may be less than the full 32678 limit depending on the particular observation).

2.4.3.1.4 Housekeeping Telemetry

HK data are sent from *Alice* at regular, user-defined time intervals, and provide values for all instrument parameters of interest, e.g., temperatures, currents, voltages, limit values, software flags, door status, etc. When these data are received from the DDS and processed through the data pipeline (by the "Lima" program), they are added to a running ASCII file. Note that since one HK TM file can contain data covering a wide range of time, the time field used in the filename identification of the files will be the UTC-format SCET of the first data packet that appears in the file.

The structure of the ASCII files will be a header section followed by a data section (the PDS label will point past this header to the data field). For use by other file-reading programs, the beginning of the data section is identified by the explicit text "START DATA" alone on a line.

- Description of Header block:
 - Each line in the header block defines one column of data in the data block. The columns in the header are:
 - Column Number: column number of value in data block,
 - Key: the FITS-compatible keyword name,
 - Unit: the physical unit (if any) of the values in the column, a dash when no unit is applicable,
 - Width: width of column, including padding spaces,
 - Format: format for data ("I"=integer, "F"=floating, "A"=string),
 - Info: additional explanatory info about the entry (this field may be empty and starts with the pound (#) character), can contain the name of the parameter in the *Alice* code
 - All fields except the "Info" field *must* contain at least one character/value with *no* whitespace separating values. So, e.g., if there are no Units or Name for a given entry, the field may have "-" or "n/a" or some other indicator of no value.
 - The header section also may contain comment lines, identified as having the pound character (#) as the first character on the line.
 - The names of the calibration files used to translate from "raw" TM values to "engineering" values (e.g., from ADC counts to temperature) will be included as comments in the header.
- Description of Data block:
 - A single line (record) is added to the file for each received TM packet (i.e., from a single dump)
 - The order of the values/columns are as described in the header block.
 - No removal of doubles or sorting is to be performed. Such sorting and trimming can be done easily by external programs (e.g., the UNIX or IDL commands "sort" and "uniq")
 - The data will include both the raw TM value and a calibrated value.

2.4.3.2 Level 3 ("Science") data

2.4.3.2.1 Histogram Mode

After Histogram-mode data are processed through the processing and calibration pipeline (the "Mike" program), the result is a 2-D array of the same size as the original: 1024x32 pixels (note that the histogram images will *not* be trimmed, i.e., they will retain the first 100 columns of the "unexposed"

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detector). The axes still correspond to wavelength vs. spatial, and value of each pixel is the calculated flux. However, since the wavelength calibration may not be linear and may vary with Y position, it cannot be easily parameterized and stored in the FITS header; therefore, the wavelength calibration is stored as a second array in the FITS file. The pulse-height and count rate data that are stored as additional extensions in the "engineering" files (Level 2 data), will be retained and are *not* processed:

- Extension = 0 : The 1024x32 histogram image (processed and flux calibrated); [photons- $cm^{-2}-s^{-1}$] implicitly per pixel in wavelength
- Extension = 1 : 1024x32 element image containing the statistical uncertainty associated with extension 0.
- Extension = 2 : The 1024x32 histogram wavelength image (calibrated wavelength at each pixel); [Å]
- Extension = 3 : The 16-value pulse-height array (*not* processed; same as in Level 2 data)
- Extension = 4 : The 100-value count rate array (*not* processed; same as in Level 2 data)

2.4.3.2.2 Pixel List Mode

The pixel list data will be a 2-dimensional array (effectively a table) with the following columns: [X, Y, Wavelength, Time Step]; this will be in the third extension of the FITS file. For quick-look purposes, the primary extension in the FITS file is a histogram-formatted version of the data as is done for the Level 2 data.

- Extension = 0 : The 1024x32 constructed histogram-formatted image (processed and flux calibrated); [photons-cm⁻²-s⁻¹] implicitly per pixel in wavelength
- Extension = 1 : 1024x32 element image containing the statistical uncertainty associated with extension 0.
- Extension = 2 : The 1024x32 histogram wavelength image (calibrated wavelength at each pixel); [Å]
- Extension = 3 : Table of pixel list data containing the x-position of the event, y-position, wavelength, and time step number.
- Extension = 4 : Reconstructed count rate data with interval equal to selected hack period (*not* processed, i.e. identical to Level 2 data)

2.4.3.2.3 Count Rate Mode

As with the Level 2 data, this is a 1-D array stored in extension 1 as a FITS binary table. Values are still in counts (since there is no wavelength information, the counts cannot be converted to flux), but some calibrations have been applied, e.g., deadtime and dark current corrections (since there is no wavelength or position information, a flat field sensitivity correction cannot be applied to these data).

2.4.4 Software

As listed in Section 2.3.3, the *Alice* data calibration software "Mike" is used internally by the instrument team and is included as IDL procedures (source code) in the /DOCUMENT/CODE/ directory for reference only (i.e., it is not included in the archive *as supported software*). Here we describe the software in more detail:

2.4.4.1 Mike

- Reads science-mode (pixel list, histogram, count rate) data files.
- Apply corrections as appropriate: deadtime, bias, dark and flat.

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- Apply wavelength and flux calibrations.
- Write calibrated data to multi-extension FITS-format file, where additional extensions may include wavelength values, effective count rate array, etc.

2.4.5 Documentation

In addition to the required ASCII formats for documents, the *Alice* team will provide formatted documents in MS Word, and /or PDF. Documents to be included will describe the instrument and its calibration, data processing, and data archiving and may also include scientific results. Documents will be located in the /DOCUMENT/ directory.

2.4.6 Derived and other Data Products

No derived/other data products are expected from the Rosetta-Alice project.

2.4.7 Ancillary Data Usage

The only ancillary data used by the *Rosetta-Alice* pipeline software are SPICE kernels, which are used to calculate observation geometry information (e.g., target distance, attitude, etc.). These can be obtained from the NAIF at JPL.

3. ARCHIVE FORMAT AND CONTENT

3.1 Format and Conventions

3.1.1 Deliveries and Archive Volume Format

Deliveries of Rosetta-Alice data will occur according to the schedule.

3.1.2 Data Set ID Formation

All the *Rosetta-Alice* datasets have the naming convention generally following the definition in Section 6 in the PDS Standards Reference [AD2], and limited to 40 characters in length. This will be the same value that will appear in the DATA_SET_ID element in the PDS label. The format is:

DATA_SET_ID = RO-<target>-ALI-<L>-<phase>-<text>-V<#.#>

where:

- <target> is the abbreviated name of the target: "E" (Earth), "M" (Mars), "A" (asteroid), "C" (comet), "CAL" (calibration), "X" (other, e.g., checkout), "Y" (sky), or other targets as described in Section 6.3.2 in [AD2].
- <L> is the **CODMAC** processing level: "2" or "3" (corresponding to PSA levels "1b" and "2", respectively). The CODMAC level "6" may also be used to describe ancilliary data.
- <phase> is the abbreviated name of the phase (e.g., CVP1, EAR1, GMP, COP etc. as defined in Appendix F, Table 2 in AD3)
- <text> is additional descriptive text if needed. Since the DATA_SET_ID is limited to 40 characters, this text field is then limited to typically no more than about 20 characters.
- V<#. #> is a version number

The expanded DATA_SET_NAME element will have the format (limited to 60 characters):

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```
"ROSETTA-ORBITER <full target> ALICE <L> <full phase> <text> V<#.#>"
```

where:

- <full_target> is the *full* target name (Earth, Mars, Asteroid <name>, Comet 67P/Churyumov-Gerasimenko, Calibration).
- <full_phase>: is the *full* name of the phase (e.g., Commissioning, Earth Swing-by 1, Global Mapping Phase, Close Observation Phase, etc.)
- <L>, <text>, V<#.#>: are as described above in the DATA_SET_ID.

Note that if the DATA_SET_NAME would exceed 60 characters, then the target and/or phase names will be abbreviated using the abbreviations in DATA_SET_ID. Also note that if multiple targets or phases are included in a given dataset, their identifiers are concatenated with a forward slash, "/", such as:

DATA SET ID = "RO-E/CAL-ALI-2-EAR1-V1.0"

to describe a dataset that includes Earth and calibration observations during the first Earth flyby.

3.1.3 Data Directory Naming Convention

Note that since a submitted archive data set cannot contain multiple data processing levels (see the last paragraph in section 6.1 in the PDS Standards Reference [AD2]), we expect to deliver two data sets per phase. In one data set, the data directories will contain Level 2 data files (also referred to as "engineering files"), which are identified with "ENG" in the filename as shown in Section 3.1.4.1. In the other data set, the data directories will contain Level 3 data files (also referred to as "science files"), which are identified with "ENG" in the filename as shown in Section 3.1.4.1. In the other data set, the data directories will contain Level 3 data files (also referred to as "science files"), which are identified with "SCI" in the filename as shown in Section 3.1.4.2.

The Alice archive data directory structure has the following naming convention:

/DATA/<yyyy>/<mm>/<filename>

where

- \circ <yyyy> is the year of the observation
- o <mm> is the month of the observation e.g., 01 for January

3.1.4 Filenaming Convention

All the *Rosetta-Alice* files are prefixed by "RA_" and are ISO 9660/level 2 compliant.

3.1.4.1 Engineering Files

These Level 2, "ENG" type files are output by the *Alice* data collection program Lima. They have the filename convention:

RA_<YYMMDDhhmmss>_<TYPE>_ENG.<EXT>

where:

- <YYMMDDhhmmss> is the 12-character UTC date and time of the start of the observation (the start of collecting photons).
- <type> is one of the following 4-character values:
 - PIX# (pixel list mode data; "#" is a 1-digit version number to distinguish the different dumps that might result from a single observation)

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- HIS# (histogram mode data)
- CNT# (count rate mode data)
- HKTM (housekeeping telemetry data; these files contain a header describing the contents and columns, but such a header is not PDS-compliant, so these files are used only internally by the *Alice* team, and should not appear in the archive)
- HKTMX (housekeeping telemetry data; the "x" indicates that the header has been stripped off as required for archive format, and so these are the files that will appear in the archive)
- <EXT> is one of the following 3-character extensions:
- FIT (for FITS files: PIX, HIS, CNT)
- TAB (for ASCII table files: HKTM)

3.1.4.2 Science Files

Level 3, "SCI" files (also sometimes referred to as "processed" or "calibrated" files) are output by the processing and calibration program Mike. *Only the science data (types: PIX, HIST, CNT) are processed and calibrated.* They have the filename convention:

RA_<YYMMDDhhmmss>_<type>_SCI_<v>.FIT

where:

- <YYMMDDhhmmss> is the 12-character UTC date and time of the start of the observation (the start of collecting photons). This will be the same value as in the filename of the input engineering file.
- <type> is one of the following 4-character values (the same value as in the filename of the input engineering file):
 - PIX# (pixel list mode data; "#" is a 1-digit version number to distinguish the different dumps that might result from a single observation)
 - HIS# (histogram mode data)
 - CNT# (count rate mode data)
- <v> is a running version number to allow for different processing runs of the data (e.g., with different, improved calibrations)
- The extension is always "FIT" since these are all FITS files.

3.1.4.3 Science Calibration Files

These are files used by the processing and calibration program Mike (e.g., darks, flats, wavelength, parameter files). They have the filename convention:

where:

- <type> is one of the following 4-character values:
 - o AEFF (effective area calibration vectors)
 - O DARK (dark image)
 - FLAT (flatfield image)
 - o WAVE (wavelength calibration file)
- <vvv> is a running 3-digit version number.
- <EXT> is one of the following extensions:
 - FIT (for FITS files: DARK, FLAT, WAVE)
 - TAB (for ASCII table files: AEFF)

3.2 Standards Used in Data Product Generation

3.2.1 PDS Standards

The PDS Standards Reference version 3.7 was used for the design of the Rosetta-Alice archive

3.2.2 Time Standards

The timing information used in all *Rosetta-Alice* data is derived from the *Rosetta* onboard clock. The *Alice* clock resolution is 1/256 seconds (~3.9 millisecond), although *Rosetta* provides sub-millisecond time resolution. The OBT is a counter related to the UTC by the Time Correlation Packets (TCP) distributed via the DDS (APID 1966, type 190, subtype 40); note that the SCET can be written in either OBT or UTC formats. The various time definitions, formats, and relationships are described in AD6 and AD7.

In the *Alice* data (e.g., the headers of the FITS files), "OBT" is the time that *Alice* includes in its TM packets once the instrument has become synchronized to the spacecraft (which is done after a time update TC is sent to *Alice*); before receiving this synchronization, the time provided by *Alice* is the time since instrument startup. The OBT is the number of seconds since the spacecraft clock was last reset (though it is expected that the clock will not ever need to be reset during the mission). The format of OBT in the PDS labels for *Rosetta* data is: "<r>/<f.f.>", where <r> is the reset number, and <f.f> is the floating-point OBT value. A reset value of <r>=1 indicates a zero point of January 1st, 2003 (2003-01-01T00:00:00), so the associated OBT is the number of seconds (not corrected for drift, etc.) since that zero point.

We use the term "SCET" for the ground-corrected OBT, modified by known deviations of the onboard clock as described in AD6 and AD7; the offset is given as the value for the *Alice*-specific keyword TOFFSET. As defined those documents, the time format used is the Sun MJT, as standard on Sun Solaris UNIX platforms. The format is two 32-bit integers. The first contains the number of seconds since the beginning of January 1st, 1970 (1970-01-01T00:00:00) and the second integer contains the number of microseconds.

In the *Alice* data pipeline, the OBT and UTC/SCET used to calculate TOFFSET are the two values that appear in the *Alice* data TM header (see Figure 1 in AD7). The OBT is the time of "measurement", placed in the TM packet header by *Alice*. The UTC/SCET is the corresponding time of the same event, calculated using the TCP and placed in the TM header by the DDS. The calculated TOFFSET is then applied in the *Alice* data pipeline to calculate the UTC for all other OBT values that appear in the *Alice* data (e.g., exposure start/stop times, dust alert times, etc. Note that although this is not precisely valid since using a single offset does not take into account any time drift (or other effects) that are more fully corrected by the functional form of the TCP, it should be sufficiently accurate based on the assumption that any drift is very small over the range of time covered by a given observation (represented by a single *Alice* data file).

The format used to represent "raw" time values (OBT and SCET) is a real number representing the number of seconds since time zero. The format used to represent "calibrated" time values (SCET only) is the correlated UTC represented in the ISO 8601 compliant standard time string:

<YYYY>-<MM>-<DD>T<hh>:<mm>:<ss.ss>

Examples of the time format are given by the *Alice*-specific keywords parameters: SCETR, SCETC, INSTTIME, PRESSCET, and the PDS label keywords: SPACECRAFT_CLOCK_START_COUNT and SPACECRAFT_CLOCK_STOP_COUNT (which are OBT, "raw" times) and START_TIME and STOP_TIME (which are SCET "calibrated" UTC times).

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N.B. During the original commissioning period (March—May 2004), a drift of the *Alice* internal timekeeping was discovered. This amounted to -1.8 seconds/hour relative to the spacecraft clock (i.e., the *Alice* timekeeping ran slow). This was not due to the actual clock drifting, but a hardware miscalculation (dividing the *Alice* clock ticks by 4000 rather than 4002). The drift was corrected each time *Alice* use a time synchronization signal which are nominally sent from the spacecraft every 30 minutes; however, by design, *Alice* does not accept time synchronization signals when an acquisition is active, so if there are many observations in a row, the time error due to the drift can be significant (several seconds). This problem was corrected by a software patch fix that was uploaded to the flight instrument in September 2004. In all data files that were created before that date, the times listed in the data files will have uncertainties up to several seconds.

3.2.3 Reference Systems

- ISO 9660, Level 2: file naming
- ISO 8601: time format

3.2.4 Other Applicable Standards

Observation geometry parameters provided in the data product labels and the FITS headers are computed for the start time of a given observation. The Earth Mean Equator reference frame of J2000 is used as the inertial reference frame for geometry items in the data headers.

Latitude and longitude (IAU2000 model) are given in the PLANETOGRAPHIC ("geodetic") coordinate system. Light travel-time from target to observer is taken into account. Latitude is the angle between the equatorial plane and the line perpendicular to the reference ellipsoid of the body (e.g., reflects body oblateness). Positive longitude is to the west. Units: DEGREES

3.3 Data Validation

The *Alice* team will perform scientific validation of the data initially with a "quick-look" of the data soon after receipt, and more thoroughly by in-depth analysis and publication of the data. If an error in the pipeline is discovered, the software will be corrected, and all affected data will be reprocessed through the pipeline. If the affected data includes data already previously submitted to the archive, the subsequent archive submission will include a full incremental dataset containing the corrected data.

Validation of archive format compliance will be made by use of PDS validation tools:

- readpds Software tool supplied by the PDS Small Bodies Node used in the verification of PDS Labels and data products.
- pvv The PSA Validation and Verification Tool from ESA.

The readpds tool is available from the PDS Small Bodies Node at:

• http://pdssbn.astro.umd.edu/nodehtml/software.shtml

3.4 Content

3.4.1 Volume Set

Each submission of *Alice* data to ESA will be separate volumes, where each volume is a single data set. Data sets are described below.

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3.4.2 Data Set

The *Rosetta-Alice* data set delivery phases and approximate sizes are given in in Appendix E in AD3, and given in more detail here. We only list here the mission phases that are likely to generate data (e.g., *Alice* will not obtain data during the Deep Space Maneuver phases). "All" data types means both processing levels (2 and 3) and possibly all three *Alice* data types (histogram, pixel list, count rate, housekeeping). Also note that at least two, possibly three datasets will be delivered for each phase, corresponding to the different data processing levels, so that "<L>" in the table below will be CODMAC data processing level numbers "2" or "3".

	Table 3-1: Planned Alice Data Sets				
Phase	DATA_SET_ID	Delivery Date	Size	Data Types	
Commissioning 1	RO-CAL-ALICE- <l>-CVP1-V1.0</l>	April 2007	76 MB	All	
Cruise 1	RO-CAL-ALICE- <l>-CR1-V1.0</l>	April 2007	19 MB	All	
Commissioning 2	RO-CAL-ALICE- <l>-CVP2-V1.0</l>	April 2007	395 MB	All	
Earth Swing-by 1	RO-CAL-ALICE- <l>-EAR1-V1.0</l>	April 2007	34MB	All	
Cruise 2	RO-CAL-ALICE- <l>-CR2-V1.0</l>	April 2007	536 MB	All	
Mars	RO-CAL-ALICE- <l>-MARS-V1.0</l>	November 2007	100 MB?	All	
Cruise 3	RO-CAL-ALICE- <l>-CR3-V1.0</l>	March 2008	TBD	All	
Earth Swing-by 2	RO-CAL-ALICE- <l>-EAR2-V1.0</l>	June 2008	TBD	All	
Earth Swing-by 3	RO-CAL-ALICE- <l>-EAR3-V1.0</l>	June 2010	TBD	All	
Asteroid Fly-by 1	RO-CAL-ALICE- <l>-AST1-V1.0</l>	TBD	TBD	All	
Asteroid Fly-by 2	RO-CAL-ALICE- <l>-AST2-V1.0</l>	TBD	TBD	All	
Comet phases	RO-CAL-ALICE- <l>-<ph>-V1.0</ph></l>	6 months after the	500 MB	All	
(each phase)	(<ph> = NCD, FAT, CAT,)</ph>	end of each phase			

3.4.3 Directories

The sections below illustrate the structure of the final archive of data from *Alice* to be delivered to the ESA-PDS team for inclusion in the *Rosetta* archive. Intermediate (incremental) deliveries during the mission will likely not include all types of data products. The details of what components are included in the archive and how they are grouped into data sets, subdirectories, and files comprises the design of the archive.

The *Rosetta* archive will be an online archive, where the data are delivered electronically. Thus, there is no need to bundle several *data sets* into one *volume*, and one data set will correspond to one volume. A data set will include the data products, secondary data, software and documentation that completely document and support the use of the data products from one or more mission phases.

3.4.3.1 Root Directory

The top-level structure of the root (/) directory of a typical data archive volume (= data set) corresponds to chapter 19 of the PDS Standards Reference (AD2) and is summarised below. It will contain the following files:

• **AAREADME.TXT** : This file describes the volume (= data set) as a whole. It gives an overview of the contents and organisation of the data set, general instructions for its use and contact information.

• **VOLDESC.CAT** : This file contains the VOLUME object, which gives a high-level description of the contents of the volume (= data set).

The following directory structure is found under the root (/) directory.

3.4.3.2 CALIB Directory

The calibration data will be archive in the /CALIB/ directory. All calibration files have a version number in their filename (Section 1.1.1.1 and 3.1.4.3), which is also time-ordered. Files in this directory include:

- **CALINFO.TXT** : Description of the contents of the CALIB directory.
- Flat field, wavelength, and dark calibration files (PDS-labelled FITS format) sorted by version number.
- Effective area calibration table (2-column ASCII table)

3.4.3.3 CATALOG Directory

This directory contains the catalog object files for the entire volume (= data set).

- **CATINFO.TXT** : Description of the contents of the CATALOG directory.
- **MISSION.CAT** : PDS mission catalog information about the *Rosetta* mission. This file will be provided by ESA.
- **INSTHOST.CAT** : PDS instrument host catalog information about the *Rosetta* spacecraft and the mounting relationship of the instruments within the spacecraft. This file will be provided by ESA.
- **INST.CAT** : PDS instrument catalog information about the *Alice* instrument (likely to be the same in all deliveries, unless updates are needed).
- **DATASET.CAT** : PDS data set catalog information about the data set submitted.
- **<TARGET_NAME>.CAT** : PDS target catalog information about the observation target, i.e. comet, asteroid, Earth or Mars, will be provided by ESA.
- **PERSON.CAT** : PDS personnel catalog information about the members of the *Alice* instrument team responsible for generating the data products.
- **REF.CAT** : PDS reference catalog information about every journal article, book, or other published reference mentioned in the above catalog objects or their components.

3.4.3.4 Index Directory

This directory contains the index files summarising all data products in the volume (= data set) by mode, key instrument parameters or mission phase, and organised to facilitate finding the data of interest for a particular scientific question. Information about the observation geometry of the data products is also included here, i.e. spacecraft position and attitude, illumination conditions etc. Information that is not accurately known at the time of delivery and thus will probably be updated later should be stored in the index files rather than in the data product labels.

- **INDXINFO.TXT** : Description of the contents of the INDEX directory.
- **INDEX.LBL** : Detached label for the index table INDEX.TAB. The INDEX_TABLE specific object should be used to identify and describe the columns of the index table.
- **INDEX.TAB** : Index of the data set in tabular format.

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3.4.3.5 BROWSE Directory and BROWSE Files

The *Rosetta-Alice* team does not foresee using the BROWSE directory. The *Alice* data files are relatively small, so that reduced resolution data are not necessary.

3.4.3.6 GEOMETRY Directory

The *Rosetta-Alice* team does not foresee providing any geometric information beyond that provided by ESOC. This directory, therefore, will be blank until the composite archive is produced by the EPT, at which time the geometry files from ESOC will be included.

3.4.3.7 SOFTWARE Directory

The *Rosetta-Alice* team does not foresee providing any software that is intended for regular updates and multiple platform capability. The source code of software used internally by the *Rosetta-Alice* team will be provided as reference documents in the DOCUMENT directory. This software is for documentation only, and is not supported.

3.4.3.8 GAZETTEER Directory

The GAZETTEER directory will not be used for *Rosetta-Alice*.

3.4.3.9 LABEL Directory

The label directory will not be used for Rosetta-Alice.

3.4.3.10 DOCUMENT Directory

This directory provides documentation and supplementary and ancillary information to assist in understanding and using the data products in the volume (= data set). All documents will be present in ASCII format to ensure long-term readability. Other formats (Word, PDF, TeX etc.) will be included as well as appropriate.

Files will include:

- **DOCINFO.TXT** : Description of the contents of the DOCUMENT directory.
- *Alice* Experiment User Manual (AD5)
- This EAICD.
- The *Rosetta-Alice* Data Processing Software Manual (AD4). This provides details of the behavior, use, and methods of the Lima, Mike, and November programs.
- The IDL calibration programs (Mike) used for pipeline processing. Note that the Mike software includes the IDL programs for wavelength calibration and deadtime correction mentioned in Section 3.4.3.2.
- Calibration reports if available.

3.4.3.11 EXTRAS Directory

Use of this directory is TBD. No specific items have yet been identified for this directory.

3.4.3.12 DATA Directory

This directory contains the actual *Alice* data (images and tables). Data files with detached PDS label files are arranged in a logical subdirectory structure with subdirectories corresponding to the year and month the data were obtained, e.g. DATA/2004/04/.

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- Level 2 data (also referred to in this document as "engineering" data).
- Level 3 data (also referred to in this document as "science", "processed", or "calibrated" data).

4. DETAILED INTERFACE SPECIFICATIONS: DATA PRODUCT DESIGN

All our data products use detached labels. In the subsections below, we provide explicit examples of the label files for all data products.

4.1 HIS, Level 2 (engineering histogram data)

PDS VERSION ID = PDS3 /* FILE CHARACTERISTICS DATA ELEMENTS */ = FIXED_LENGTH
= 2880 /* FITS standard record length */
= 33
= "RA_040419231832_HIS0_ENG.LBL" RECORD TYPE RECORD_BYTES FILE_RECORDS FILE_NAME FILE NAME DATA FORMAT = FITS /* DATA OBJECT POINTERS IDENTIFICATION DATA ELEMENTS */ = ("RA_040419231832_HIS0_ENG.FIT",1) ^HEADER

 ^IMAGE
 = ("RA_040419231832_HIS0_ENG.FIT",7)

 ^PULSE_HEIGHT_HEADER
 = ("RA_040419231832_HIS0_ENG.FIT",30)

 ^PULSE_HEIGHT_TABLE
 = ("RA_040419231832_HIS0_ENG.FIT",31)

 ^COUNT_RATE_HEADER
 = ("RA_040419231832_HIS0_ENG.FIT",32)

 ^COUNT_RATE_SERIES
 = ("RA_040419231832_HIS0_ENG.FIT",33)

 /* OBSERVATION AND EXPOSURE DATA ELEMENTS */ INSTRUMENT_MODE_ID = "HIS" /* HIS, PIX, CNT */ INSTRUMENT_MODE_DESC = "HISTOGRAM" START_TIME = 2004-04-19T23:18:31.633 /* UTC */ STOP_TIME = 2004-04-19T23:18:51.782 /* UTC */ SPACECRAFT_CLOCK_START_COUNT = "1/41037497.246" SPACECRAFT_CLOCK_STOP_COUNT = "1/41037517.395" EXPOSURE_DURATION = 20.1480 /* SECONDS */ DATA_QUALITY_ID = 0 DATA_QUALITY_DESC = "Data quality is nominal" MISSION_PHASE_NAME = "COMMISSIONING 1" /* see Phase Definition table in RSOC Design Specification, */ /* RO-EST-PL-2010 and Archive GVT Plan, RO-EST-PL-5011 */ TARGET_NAME = CHECKOUT = "N/A" TARGET TYPE

/* INSTRUMENT AND MISSION DESCRIPTIVE DATA ELEMENTS */

INSTRUMENT_ID = "ALICE"

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INSTRUMENT_NAME = "ALICE" INSTRUMENT_TYPE = "ULTRAVIOLET SPECTROMETER" INSTRUMENT_HOST_ID = "RO" INSTRUMENT_HOST_NAME = "ROSETTA-ORBITER" MISSION_ID = "ROSETTA" MISSION_NAME = "INTERNATIONAL ROSETTA MISS = "INTERNATIONAL ROSETTA MISSION" /* PRODUCT, DATA SET, AND VOLUME DATA ELEMENTS */ FRODUCT_ID= "RA_040419231832_HIS0_ENG.FIT"PRODUCT_TYPE= EDR /* CODMAC_E PRODUCT_TYPE= EDR /* CODMAC Data Level 2 */PRODUCT_CREATION_TIME= 2004-09-20T22:37:07DATA_SET_ID= "RO-C/CAL/X-ALICE-2-CVP1-V1.0"DATA_SET_NAME= " ROSETTA-ORBITER 2002T7/CAL/CHECK ALICE 2 CVP1 V1.0" PRODUCER_ID = "ROSETTA_ALICE" PRODUCER_FULL_NAME = "ANDREW J. STEFFL" PRODUCER INSTITUTION NAME = "SOUTHWEST RESEARCH INSTITUTE" /* DATA OBJECT DEFINITIONS */ = HEADER OBJECT = 17280 /* RECORD BYTES x RECORDS */ BYTES BYTES HEADER_TYPE INTERCHANGE FORMAT = FITS = ASCII = 6 RECORDS = " DESCRIPTION FITS header for Rosetta-Alice uncalibrated (CODMAC Data Level 2) histogram observation." END OBJECT = HEADER DBJECT= IMAGELINE_SAMPLES= 1024 /* FITS NAXIS1 keyword */LINES= 32 /* FITS NAXIS2 keyword */SAMPLE_BITS= 16SAMPLE_TYPE= MSB_INTEGERINTERCHANGE_FORMAT= BINARYOFFSET= 32768 /* FITS BZERO keyword */SCALING_FACTOR= 1.00000 /* FITS BSCALE keyword */AXIS_ORDER_TYPE= FIRST_INDEX_FASTESTLINE_DISPLAY_DIRECTION= WPSAMPLE_DISPLAY_DIRECTION= RTGHT OBJECT SAMPLE DISPLAY DIRECTION = RIGHT = " DESCRIPTION FITS image for Rosetta-Alice uncalibrated (CODMAC Data Level 2) histogram observation. Units are raw data numbers." END OBJECT = IMAGE

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= 2880 /* RECORD BYTES x RECORDS */ BYTES HEADER TYPE = FITS INTERCHANGE FORMAT = ASCII RECORDS = 1 DESCRIPTION = "FITS extension header" END OBJECT = PULSE HEIGHT HEADER OBJECT = PULSE HEIGHT TABLE = "PULSE HEIGHT DISTRIBUTION" NAME COLUMNS = 1 = 16 /* FITS NAXIS2 keyword */ ROWS ROW BYTES = 2 INTERCHANGE FORMAT = BINARY = " DESCRIPTION The pulse height distribution (PHD) is a one-dimensional histogram, that bins each of the detected events (pulses) based on their 'amplitude'. Detected events are those that are larger than the discriminator setting, which is an analog (voltage) value. Of the 16 PHD bins, both the top and bottom 3 bins don't get any events (this is caused by the digitization of this amplitude information), so only 10 bins contain non-zero values. The sampling parameter interval is effectively unknown and somewhat arbitrary anyway since only the relative shape and stability of the distribution is of interest." OBJECT = COLUMN " NAME = "PHD DATA TYPE = MSB INTEGER /* 16-bit signed integers */ = 2 BYTES = 1 START BYTE = 32768 /* FITS TZERO1 keyword */ OFFSET END OBJECT = COLUMN END OBJECT = PULSE HEIGHT TABLE = COUNT RATE HEADER OBJECT = 2880 /* RECORD BYTES x RECORDS */ BYTES HEADER_TYPE = FITS INTERCHANGE FORMAT = ASCII RECORDS = 1 = "FITS extension header" DESCRIPTION = COUNT RATE HEADER END OBJECT OBJECT = COUNT RATE SERIES = 1 COLUMNS ROWS = 100 /* FITS NAXIS1 keyword */ = 2 ROW BYTES = BINARY INTERCHANGE FORMAT SAMPLING_PARAMETER_INTERVAL= 1.90000 SAMPLING PARAMETER NAME = TIME

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```
SAMPLING PARAMETER UNIT
                               = SECONDS
                               = "
  DESCRIPTION
    These count rate data are a table that gives the measured count
    rate from the observation. Units are counts per time hack interval"
  OBJECT
                               = COLUMN
                               = "COUNT RATE"
    NAME
    DATA TYPE
                              = MSB INTEGER
                              = 2
    BYTES
                         = 2 /* 16-bit signed integer */
= 32768 /* FITS TZERO1 keyword */
= COLUMN
= COLUMN
    START BYTE
    ITEM BYTES
    OFFSET
 END OBJECT
END OBJECT
                             = COUNT RATE SERIES
```

END

4.2 HIS, Level 3 (science histogram data)

```
PDS_VERSION_ID = PDS3
/* FILE CHARACTERISTICS DATA ELEMENTS */
RECORD_TYPE = FIXED_LENGTH
RECORDBYTES = 2880 /* FITS standard record length */
FILE_RECORDS = 109
FILE_NAME = "RA_040419231832_HIS0_SCI.LBL"
DATA_FORMAT = FITS
/* DATA OBJECT POINTERS IDENTIFICATION DATA ELEMENTS */
^HEADER = ("RA_040419231832_HIS0_SCI.FIT",1)
^IMAGE = ("RA_040419231832_HIS0_SCI.FIT",13)
^WAVELENGTH_HEADER = ("RA_040419231832_HIS0_SCI.FIT",59)
^WAVELENGTH_HEADER = ("RA_040419231832_HIS0_SCI.FIT",60)
^PULSE_HEIGHT_TABLE = ("RA_040419231832_HIS0_SCI.FIT",106)
^PULSE_HEIGHT_TABLE = ("RA_040419231832_HIS0_SCI.FIT",107)
^CCOUNT_RATE_HEADER = ("RA_040419231832_HIS0_SCI.FIT",108)
^COUNT_RATE_SERIES = ("RA_040419231832_HIS0_SCI.FIT",109)
```

N.B.: The formats of the following sets of data elements are identical to	
those in Section 4.1, so their descriptions are omitted here:	
/* OBSERVATION AND EXPOSURE DATA ELEMENTS */	
/* INSTRUMENT AND MISSION DESCRIPTIVE DATA ELEMENTS */	
/* PRODUCT AND DATA SET DATA ELEMENTS */	
/* STRUCTURE DEFINITION OF INSTRUMENT PARAMETER OBJECTS */	
/* DESCRIPTION OF INSTRUMENT */	

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/* PARAMETERS INDEX FILE DEFINITION */
/* MISSION SPECIFIC KEYWORDS */

/* OBSERVATION GEOMETRY ELEMENTS */ /* S/C TO SUN ELEMENTS */ SC SUN POSITION VECTOR = (1.1170497e+08, 71725860., 30237734.) SC_SUN_VELOCITY VECTOR = (-20.288503, 23.269792, 9.9525278) SUB_SOLAR_LONGITUDE= 160.16964SUB_SOLAR_LATITUDE= -4.4102611 SUB SOLAR LATITUDE /* S/C TO TARGET ELEMENTS */ SC TARGET POSITION VECTOR = (-1.0000000e+32, -1.0000000e+32, -1.0000000e+32)SC TARGET VELOCITY VECTOR = (-1.0000000e+32, -1.0000000e+32, -1.0000000e+32) SUB_SPACECRAFT_LONGITUDE = -1.0000000e+32 SUB SPACECRAFT LATITUDE = -1.0000000e+32SPACECRAFT ALTITUDE = -1.0000000e+32CENTRAL_BODY DISTANCE = -1.0000000e+32/* INSTRUMENT BORESIGHT LINE OF SIGHT ELEMENTS */ RIGHT ASCENSION = 46.574495 DECLINATION PHASE ANGLE = -80.222574 = 0 INCIDENCE ANGLE = -1.0000000e+32EMISSION ANGLE = -1.0000000e+32= {"naif0008.tls" SPICE FILE NAME "ROS 060511 STEP.TSC", "ROS V06.TF", "EARTHFIXEDIAU.TF", "DE403-MASSES.TPC", "PCK00008.TPC", "DE405S.BSP", "C2002T7.bsp", "TEMPEL1-9P-DI-P.BSP", "mar033 2000-2025.bsp", "sat164.bsp", "2867 Steins 2004 2016.bsp", "21 Lutetia 2004 2016.bsp", "67P CHURY GERAS 2004 2016.BSP", "ORER _____00031.BSP", "ORFR 00047.BSP", ____00047.BSP", "ORGR _____00047.BSP", "ORHR "ORMR 00047.BSP", "ATNR P040302093352 00046.BC"}

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```
/* DATA OBJECT DEFINITIONS */
OBJECT
                                    = HEADER
                                    = 34560 /* RECORD BYTES x RECORDS */
  BYTES
  HEADER TYPE
                                   = FITS
  INTERCHANGE FORMAT
                                   = ASCII
  RECORDS
                                   = 12
                                     = "
  DESCRIPTION
    FITS header for Rosetta-Alice calibrated (CODMAC Data Level 3)
    histogram observation."
                                     = HEADER
END OBJECT
OBJECT
                                   = IMAGE
  LINE SAMPLES
                                   = 1024 /* FITS NAXIS1 keyword */
                                   = 32 /* FITS NAXIS2 keyword */
  LINES
                                    = 32
  SAMPLE BITS
 SAMPLE_TYPE = IEEE_REAL

OFFSET = 0.00000 /* FITS BZERO keyword */

SCALING_FACTOR = 1.00000 /* FITS BSCALE keyword */

AXIS_ORDER_TYPE = FIRST_INDEX_FASTEST

LINE_DISPLAY_DIRECTION = UP

SAMPLE_DISPLAY_DIRECTION = UP
  SAMPLE_DISPLAY_DIRECTION = RIGHT
                                     = "
  DESCRIPTION
     FITS image for Rosetta-Alice calibrated (CODMAC Data Level 3)
    histogram observation. Units are photons cm**-2 s**-1 Data are in
    units of photons per square centimeter per second"
END OBJECT
                                     = IMAGE
OBJECT
                                     = WAVELENGTH HEADER
                                     = 2880 /* RECORD BYTES x RECORDS */
  BYTES
  HEADER TYPE
                                    = FITS
 INTERCHANGE_FORMAT
                                   = ASCII
                                   = 1
 RECORDS
  DESCRIPTION
                                   = "FITS extension header"
END OBJECT
                                   = WAVELENGTH HEADER
                              = WAVELENGTH_IMAGE
= 1024 /* FITS NAXIS1 keyword */
= 32 /* FITS NAXIS2 keyword */
OBJECT
  LINE SAMPLES
  LINES
 LINES = 32 /* FITS NAXIS2 keyword */

SAMPLE_BITS = 32

SAMPLE_TYPE = IEEE_REAL

INTERCHANGE_FORMAT = BINARY

OFFSET = 0.00000 /* FITS BZERO keyword */

SCALING_FACTOR = 1.00000 /* FITS BSCALE keyword */

AXIS_ORDER_TYPE = FIRST_INDEX_FASTEST

LINE_DISPLAY_DIRECTION = UP
  SAMPLE DISPLAY DIRECTION = RIGHT
                                     = "
  DESCRIPTION
    FITS extension containing calibrated wavelengths for each detector
     pixel. Units are Angstroms. Wavelengths are in Angstroms."
END OBJECT
                                     = WAVELENGTH IMAGE
```

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OBJECT BYTES HEADER_TYPE INTERCHANGE_FORMAT RECORDS DESCRIPTION END_OBJECT	<pre>= PULSE_HEIGHT_HEADER = 2880 /* RECORD_BYTES x RECORDS */ = FITS = ASCII = 1 = "FITS extension header" = PULSE_HEIGHT_HEADER</pre>
histogram, that bins eac their 'amplitude'. Detec the discriminator settin the 16 PHD bins, both th events (this is caused b information), so only 10 sampling parameter inter	<pre>= PULSE_HEIGHT_TABLE = "PULSE_HEIGHT_DISTRIBUTION" = 1 = 16 /* FITS NAXIS2 keyword */ = 2 = BINARY = " oution (PHD) is a one-dimensional th of the detected events (pulses) based on ted events are those that are larger than g, which is an analog (voltage) value. Of the top and bottom 3 bins don't get any by the digitization of this amplitude bins contain non-zero values. The val is effectively unknown and somewhat nly the relative shape and stability of nterest."</pre>
OBJECT NAME DATA_TYPE BYTES START_BYTE OFFSET END_OBJECT END_OBJECT	<pre>= COLUMN = "PHD " = MSB_INTEGER /* 16-bit signed integers */ = 2 = 1 = 32768 /* FITS TZER01 keyword */ = COLUMN = PULSE_HEIGHT_TABLE</pre>
OBJECT BYTES HEADER_TYPE INTERCHANGE_FORMAT RECORDS DESCRIPTION END_OBJECT	<pre>= COUNT_RATE_HEADER = 2880 /* RECORD_BYTES x RECORDS */ = FITS = ASCII = 1 = "FITS extension header" = COUNT_RATE_HEADER</pre>
OBJECT COLUMNS ROWS ROW_BYTES INTERCHANGE_FORMAT SAMPLING_PARAMETER_INTERVA SAMPLING_PARAMETER_NAME SAMPLING_PARAMETER_UNIT DESCRIPTION	L= 1.90000 = TIME

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These count rate data are a table that gives the measured count rate from the observation. Units are counts per time hack interval" OBJECT = COLUMN = "COUNT RATE" NAME DATA TYPE = MSB INTEGER BYTES = 2 START BYTE = 1 = 2 /* 16-bit signed integer */ ITEM BYTES = 32768 /* FITS TZER01 keyword */ OFFSET = COLUMN END OBJECT END OBJECT = COUNT RATE SERIES

END

4.3 PIX, Level 2 (engineering pixel list data)

PDS VERSION ID = PDS3 /* FILE CHARACTERISTICS DATA ELEMENTS */ RECORD TYPE = FIXED LENGTH = 2880 $\overline{/}$ * FITS standard record length */ RECORD BYTES FILE_RECORDS = 63 FILE NAME = "RA 040323225136 PIX0 ENG.LBL" DATA FORMAT = FITS /* DATA OBJECT POINTERS IDENTIFICATION DATA ELEMENTS */ ^HEADER = ("RA 040323225136 PIX0 ENG.FIT",1)

 ^IMAGE
 = ("RA_040323225136_PIX0_ENG.FIT",7)

 ^PIXEL_LIST_HEADER
 = ("RA_040323225136_PIX0_ENG.FIT",30)

 ^PIXEL_LIST_TABLE
 = ("RA_040323225136_PIX0_ENG.FIT",31)

 ^COUNT_RATE_HEADER
 = ("RA_040323225136_PIX0_ENG.FIT",49)

 ^COUNT_RATE_SERIES
 = ("RA_040323225136_PIX0_ENG.FIT",50)

 N.B.: The formats of the following sets of data elements are identical to those in Section 4.1, so their descriptions are omitted here: /* OBSERVATION AND EXPOSURE DATA ELEMENTS */ /* INSTRUMENT AND MISSION DESCRIPTIVE DATA ELEMENTS */ /* PRODUCT AND DATA SET DATA ELEMENTS */ /* STRUCTURE DEFINITION OF INSTRUMENT PARAMETER OBJECTS */ /* DESCRIPTION OF INSTRUMENT */ /* PARAMETERS INDEX FILE DEFINITION */ /* MISSION SPECIFIC KEYWORDS */

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```
/* DATA OBJECT DEFINITIONS */
OBJECT
                             = HEADER
                            = 17280 /* RECORD BYTES x RECORDS */
 BYTES
 HEADER TYPE
                            = FITS
 INTERCHANGE FORMAT
                            = ASCII
                             = 6
 RECORDS
                              = "
 DESCRIPTION
    FITS header for Rosetta-Alice uncalibrated (CODMAC Data Level
    2) pixel list observation."
END OBJECT
                              = HEADER
OBJECT
                             = IMAGE
                             = 1024 /* FITS NAXIS1 keyword */
 LINE SAMPLES
                            = 32 /* FITS NAXIS2 keyword */
 LINES
 SAMPLE BITS
                            =
                                          16
 SAMPLE_TYPE
 SAMPLE_TYPE = HOL_
INTERCHANGE_FORMAT = BINARY
= 32768 /* FITS BZERO keyword */
 OFFSET= 52760 /* FITS B2ERO Reyword */SCALING_FACTOR= 1.00000 /* FITS BSCALE keyword */AXIS_ORDER_TYPE= FIRST_INDEX_FASTESTLINE_DISPLAY_DIRECTION= UP
  SAMPLE DISPLAY DIRECTION = RIGHT
                             = "
  DESCRIPTION
   Reconstructed histogram image for Rosetta-Alice uncalibrated
    (CODMAC Data Level 2) pixel list observation. Units are raw data
    numbers."
END OBJECT
                             = IMAGE
OBJECT
                             = PIXEL LIST HEADER
 BYTES
                            = 2880 /* RECORD BYTES x RECORDS */
 HEADER TYPE
                            = FITS
 INTERCHANGE FORMAT
                            = ASCII
 RECORDS
                             = 1
                            = "FITS extension header"
 DESCRIPTION
                            = PIXEL LIST HEADER
END OBJECT
OBJECT
                            = PIXEL LIST TABLE
 COLUMNS
                             = 1
 ROWS
                             = 24491 /* FITS extension NAXIS1 */
 ROW BYTES
                             = 2
 INTERCHANGE FORMAT
                            = BINARY
  DESCRIPTION
                              = "
    These pixel list data are an array up to 32767 elements that gives
    the X,Y positions of every detected event. In pixel list mode, the
    event addresses that are generated by the detector are stored
    sequentially in the memory array, as in a list. In addition to the
    detector events, a time-mark value is also inserted into the list
```

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on a periodic basis. The data are encoded as follows: Value (for events) = $X \times 1024 + Y$ Value (for timehacks) > 32768 Another way to describe this is by representing the bits in the pixel list events (16-bit words) as: h a a a a a e e e e e e e e e where each letter represents a single bit with the following meaning: h - (MSB) set to 1 for a timehack, and 0 for a normal photon event. For timehacks, the other 15 bits have no meaning. a - the five bits of spatial information of a photon event. e - the ten bits of spectral information of a photon event." OBJECT = COLUMN = "PIXEL LIST" NAME = MSB INTEGER DATA TYPE BYTES = 2 START BYTE = 1 = 2 /* Each item is signed 16 bit integer */ ITEM BYTES OFFSET = 32768 /* BZERO keyword in FITS extension header END OBJECT = COLUMN END OBJECT = PIXEL LIST TABLE OBJECT = COUNT RATE HEADER BYTES = 2880 /* RECORD BYTES x RECORDS */ HEADER TYPE = FITS = ASCII INTERCHANGE FORMAT = 1 RECORDS = "FITS extension header" DESCRIPTION END OBJECT = COUNT RATE HEADER OBJECT = COUNT RATE SERIES COLUMNS = 1 ROWS = 19221 /* FITS NAXIS1 keyword */ ROW BYTES = 2 = BINARY INTERCHANGE FORMAT SAMPLING PARAMETER INTERVAL= 0.016000000 SAMPLING PARAMETER NAME = TIME SAMPLING PARAMETER UNIT = SECONDS DESCRIPTION = " These count rate data are a table that gives the count rate measurements as reconstructed from the pixel list observation. Units are raw counts per time hack interval" OBJECT = COLUMN = "COUNT RATE" NAME DATA TYPE = MSB INTEGER BYTES = 2
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```
START_BYTE= 1ITEM_BYTES= 2 /* 16-bit signed integer */OFFSET= 32768 /* FITS TZER01 keyword */END_OBJECT= COLUMNEND_OBJECT= COUNT_RATE_SERIES
```

END

4.4 PIX, Level 3 (science pixel list data)

PDS VERSION ID = PDS3 /* FILE CHARACTERISTICS DATA ELEMENTS */ RECORD_TYPE = FIXED_LENGTH = 2880 $\overline{/*}$ FITS standard record length */ RECORD_BYTES FILE_RECORDS = 138 FILE NAME = "RA 040323225136 PIX0 SCI.LBL" DATA FORMAT = FITS /* DATA OBJECT POINTERS IDENTIFICATION DATA ELEMENTS */ = ("RA 040323225136 PIX0 SCI.FIT",1) ^HEADER = ("RA 040323225136 PIX0 SCI.FIT",13) ^IMAGE

 *IMAGE
 = ("RA_040323225136_PIX0_SCI.FIT",13)

 *WAVELENGTH_HEADER
 = ("RA_040323225136_PIX0_SCI.FIT",59)

 *WAVELENGTH_IMAGE
 = ("RA_040323225136_PIX0_SCI.FIT",60)

 *PIXEL_LIST_HEADER
 = ("RA_040323225136_PIX0_SCI.FIT",106)

 *PIXEL_LIST_TABLE
 = ("RA_040323225136_PIX0_SCI.FIT",107)

 *COUNT_RATE_HEADER
 = ("RA_040323225136_PIX0_SCI.FIT",124)

 *COUNT_RATE_SERIES
 = ("RA_040323225136_PIX0_SCI.FIT",125)

N.B.: The formats of the following sets of data elements are identical to those in Section 4.2, so their descriptions are omitted here: /* OBSERVATION GEOMETRY ELEMENTS */ /* S/C TO TARGET ELEMENTS */ /* INSTRUMENT BORESIGHT LINE OF SIGHT ELEMENTS */

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/* DATA OBJECT DEFINITIONS */ OBJECT = HEADER = 34560 /* RECORD_BYTES x RECORDS */ BYTES HEADER TYPE = FITS = ASCII INTERCHANGE FORMAT RECORDS = 12 = " DESCRIPTION FITS header for Rosetta-Alice calibrated (CODMAC Data Level 3) pixel list observation." END OBJECT = HEADER OBJECT = IMAGE LINE SAMPLES = 1024 /* FITS NAXIS1 keyword */ = 32 /* FITS NAXIS2 keyword */ = 32 LINES SAMPLE_BITS SAMPLE_TYPE SAMPLE_BITS= 52SAMPLE_TYPE= IEEE_REALOFFSET= 0.00000 /* FITS BZERO keyword */SCALING_FACTOR= 1.00000 /* FITS BSCALE keyword */AXIS_ORDER_TYPE= FIRST_INDEX_FASTESTLINE_DISPLAY_DIRECTION= UP SAMPLE DISPLAY DIRECTION = RIGHT = " DESCRIPTION FITS image for reconstructed Rosetta-Alice calibrated (CODMAC Data Level 3 histogram observation. Units are photons cm**-2 s**-1 Data are in units of photons per square centimeter per second" END OBJECT = IMAGE OBJECT = WAVELENGTH HEADER BYTES = 2880 /* RECORD BYTES x RECORDS */ BYTES HEADER_TYPE = FITS INTERCHANGE_FORMAT = ASCII - 1 RECORDS = 1 DESCRIPTION = "FITS extension header" END OBJECT = WAVELENGTH HEADER = WAVELENGTH_IMAGE
= 1024 /* FITS NAXIS1 keyword */
= 32 /* FITS NAXIS2 keyword */ OBJECT LINE SAMPLES LINES= 32 /* FITS NAXIS2 keyword */SAMPLE_BITS= 32SAMPLE_TYPE= IEEE_REALINTERCHANGE_FORMAT= BINARYOFFSET= 0.00000 /* FITS BZERO keyword */SCALING_FACTOR= 1.00000 /* FITS BSCALE keyword */AXIS_ORDER_TYPE= FIRST_INDEX_FASTESTLINE_DISPLAY_DIRECTION= UPSAMPLE_DISPLAY_DIRECTION= DECUT SAMPLE_DISPLAY DIRECTION = RIGHT = " DESCRIPTION FITS extension containing calibrated wavelengths for each detector pixel. Units are Angstroms. Wavelengths are in Angstroms." END OBJECT = WAVELENGTH IMAGE

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OBJECT = PIXEL LIST HEADER = 2880 /* RECORD BYTES x RECORDS */ BYTES HEADER TYPE = FITS INTERCHANGE FORMAT = ASCII RECORDS = 1 = "FITS extension header" DESCRIPTION = PIXEL LIST HEADER END OBJECT OBJECT = PIXEL LIST TABLE COLUMNS = 4 = 5270 /* NAXIS2 in FITS extension header ROWS ROW BYTES = 9 INTERCHANGE FORMAT = BINARY = " DESCRIPTION These pixel list data are a 4-column table that lists: X, Y, Wavelength, Time Step. Any necessary calibrations (flatfield, dark) are applied to that pixel location. The deadtime correction is applied to the total counts in a timestep. Photons appearing at different X,Y positions during a particular timestep are listed, in the order detected, as records with the same Time Step value." OBJECT = COLUMN NAME = "X POSITION" DATA TYPE = MSB INTEGER BYTES = 2 START BYTE = 1 = 2 /* Each item is Integer*2 (short integer) */ = 0 ITEM BYTES OFFSET END OBJECT = COLUMN OBJECT = COLUMN = "Y POSITION" NAME DATA TYPE = MSB INTEGER BYTES = 1 START BYTE = 3 = 1 /* Each item is Integer*1 (byte) */ ITEM BYTES = 0 OFFSET = COLUMN END OBJECT OBJECT = COLUMN NAME = "PIXEL LIST WAVELENGTH" = "Calibrated wavelength of pixel list event" DESCRIPTION DATA TYPE = IEEE REAL BYTES = 4 START BYTE = 4 ITEM BYTES = 4 /* Each item is an IEEE floating point */ = 0 OFFSET = COLUMN END OBJECT OBJECT = COLUMN NAME = "PIXEL LIST TIME STEP"

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```
DATA TYPE
                            = MSB INTEGER
                            = 2
   BYTES
                            = 8
    START BYTE
   ITEM BYTES
                           = 2 /* Each item is Integer*2 (short integer) */
   OFFSET
                           = 0
 END OBJECT
                            = COLUMN
                           = PIXEL LIST TABLE
END OBJECT
OBJECT
                         = COUNT RATE HEADER
                           = 2880 /* RECORD BYTES x RECORDS */
 BYTES
 HEADER_TYPE
                           = FITS
 INTERCHANGE FORMAT
                           = ASCII
                           = 1
 RECORDS
                           = "FITS extension header"
 DESCRIPTION
END OBJECT
                            = COUNT RATE HEADER
OBJECT
                            = COUNT RATE SERIES
 COLUMNS
                            = 1
                           = 19221 /* FITS NAXIS1 keyword */
 ROWS
                           = 2
 ROW BYTES
 INTERCHANGE FORMAT = BINARY
  SAMPLING PARAMETER INTERVAL= 0.016000000
  SAMPLING PARAMETER NAME = TIME
  SAMPLING PARAMETER UNIT
                           = SECONDS
                           = "
  DESCRIPTION
   These count rate data are a table that gives the measured count
   rate from the observation. Units are counts per time hack interval"
 OBJECT
                            = COLUMN
   NAME
                            = "COUNT RATE"
                            = MSB INTEGER
   DATA TYPE
   BYTES
                           = 2
   START BYTE
                           = 1
                         = 2 /* 16-bit signed integer */
= 32768 /* FITS TZERO1 keyword */
   ITEM BYTES
   OFFSET
 END OBJECT
                           = COLUMN
END OBJECT
                           = COUNT RATE SERIES
```

END

4.5 CNT, Level 2 (engineering count rate data)

PDS_VERSION_ID = PDS3
/* FILE CHARACTERISTICS DATA ELEMENTS */
RECORD_TYPE = FIXED_LENGTH
RECORD_BYTES = 2880 /* FITS standard record length */

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```
FILE RECORDS
                               = 8
                               = "RA 040419231322 CNT0_ENG.LBL"
FILE NAME
DATA FORMAT
                               = FITS
/* DATA OBJECT POINTERS IDENTIFICATION DATA ELEMENTS */

      ^HEADER
      = ("RA_040419231322_CNT0_ENG.FIT",1)

      ^COUNT_RATE_HEADER
      = ("RA_040419231322_CNT0_ENG.FIT",7)

      ^COUNT_RATE_SERIES
      = ("RA_040419231322_CNT0_ENG.FIT",8)

N.B.: The formats of the following sets of data elements are identical to
those in Section 4.1, so their descriptions are omitted here:
      /* OBSERVATION AND EXPOSURE DATA ELEMENTS */
      /* INSTRUMENT AND MISSION DESCRIPTIVE DATA ELEMENTS */
      /* PRODUCT AND DATA SET DATA ELEMENTS */
      /* STRUCTURE DEFINITION OF INSTRUMENT PARAMETER OBJECTS */
      /* DESCRIPTION OF INSTRUMENT */
      /* PARAMETERS INDEX FILE DEFINITION */
      /* MISSION SPECIFIC KEYWORDS */
/* DATA OBJECT DEFINITIONS */
OBJECT
                               = COUNT RATE HEADER
                               = 17280 /* RECORD BYTES x RECORDS */
  BYTES
  HEADER TYPE
                              = FITS
  INTERCHANGE FORMAT
                             = ASCII
  RECORDS
                              = 6
  DESCRIPTION
                              = "FITS extension header"
END OBJECT
                              = COUNT RATE HEADER
                               = COUNT RATE SERIES
OBJECT
 COLUMNS
                               = 1
  ROWS
                             = 1000 /* FITS NAXIS1 keyword */
  ROW BYTES
                              = 2
                       = BINARY
  INTERCHANGE FORMAT
  SAMPLING PARAMETER_INTERVAL= 0.09000000
  SAMPLING PARAMETER NAME = TIME
  SAMPLING PARAMETER UNIT
                              = SECONDS
                               = "
  DESCRIPTION
    These count rate data are a table that gives the count rate
    measurements as reconstructed from the pixel list observation.
    Units are raw counts per time hack interval"
  OBJECT
                               = COLUMN
                               = "COUNT RATE"
    NAME
                              = MSB INTEGER
    DATA TYPE
    BYTES
                              = 2
    START BYTE
                              = 1
   ITEM BYTES
                             = 2 /* 16-bit signed integer */
    OFFSET
                               = 32768 /* FITS TZERO1 keyword */
```

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END_OBJECT	=	COLUMN	I	
END_OBJECT	=	COUNT	_RATE_	SERIES

END

4.6 CNT, Level 3 (science count rate data)

PDS_VERSION_ID = PDS3

/* FILE CHARACTERISTICS DATA ELEMENTS */

RECORD_TYPE	= FIXED_LENGTH
RECORD BYTES	= 2880 /* FITS standard record length */
FILE RECORDS	= 15
FILE NAME	= "RA 040419231322 CNT0 SCI.LBL"
DATA_FORMAT	= FITS
—	

/* DATA OBJECT POINTERS	IDENTIFICATION DATA ELEMENTS */
^HEADER	= ("RA_040419231322_CNT0_SCI.FIT",1)
^COUNT RATE HEADER	= ("RA 040419231322 CNTO SCI.FIT",13)
^COUNT_RATE_SERIES	= ("RA_040419231322_CNT0_SCI.FIT",14)

N.B.: The formats of the following sets of data elements are identical to	,
those in Section 4.1, so their descriptions are omitted here:	
/* OBSERVATION AND EXPOSURE DATA ELEMENTS */	
/* INSTRUMENT AND MISSION DESCRIPTIVE DATA ELEMENTS */	
/* PRODUCT AND DATA SET DATA ELEMENTS */	
/* STRUCTURE DEFINITION OF INSTRUMENT PARAMETER OBJECTS */	
/* DESCRIPTION OF INSTRUMENT */	
/* PARAMETERS INDEX FILE DEFINITION */	
/* MISSION SPECIFIC KEYWORDS */	

/* DATA OBJECT DEFINITIONS */
OBJECT = COUNT_RATE_HEADER
BYTES = 34560 /* RECORD_BYTES x RECORDS */
HEADER_TYPE = FITS
INTERCHANGE_FORMAT = ASCII
RECORDS = 12
DESCRIPTION = "FITS extension header"

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= COUNT RATE_HEADER END OBJECT = COUNT RATE SERIES OBJECT COLUMNS = 1 ROWS = 1000 /* FITS NAXIS1 keyword */ ROW BYTES = 4 = BINARY INTERCHANGE FORMAT SAMPLING PARAMETER INTERVAL= 0.09000000 SAMPLING PARAMETER NAME = TIME SAMPLING PARAMETER_UNIT = SECONDS = " DESCRIPTION These count rate data are a table that gives the count rate measurements as reconstructed from the pixel list observation. Units are raw counts per time hack interval" OBJECT = COLUMN = "COUNT RATE" NAME DATA TYPE = MSB INTEGER = 4 BYTES START BYTE = 1 = 4 /* 32-bit floating point */ = 32768 /* FITS TZERO1 keyword */ ITEM BYTES OFFSET = COLUMN END OBJECT END OBJECT = COUNT RATE SERIES

END

4.7 HKTM (housekeeping telemetry data)

PDS VERSION ID = PDS3/* FILE CHARACTERISTICS DATA ELEMENTS */ RECORD TYPE = FIXED LENGTH RECORD BYTES = 449 FILE RECORDS = 3210 = "RA 040322224947 HKTM ENG.LBL" /* this label FILE NAME /* DATA OBJECT POINTERS IDENTIFICATION DATA ELEMENTS */ ^HOUSEKEEPING_TABLE = ("RA_040322224947_HKTM_ENG.TAB",1) /* TIME RANGE DATA ELEMENTS */ START TIME = 2004-03-22T22:49:47.995 /* UTC */ STOP TIME = 2004-03-23T06:50:49.512 /* UTC */ SPACECRAFT CLOCK START COUNT = "1/41.511" /* OBT */ SPACECRAFT CLOCK STOP COUNT = "1/38645435.812" /* OBT */

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/* OBSERVATION AND EXPOSURE DATA ELEMENTS */ = "N/A" = "N/A" = "N/A" INSTRUMENT MODE ID INSTRUMENT MODE DESC EXPOSURE DURATION DATA QUALITY ID = 0 = "Data quality is nominal" = "COMMISSIONING 1" DATA QUALITY DESC MISSION_PHASE NAME /* see Phase Definition table in RSOC Design Specification, */ /* RO-EST-PL-2010 and Archive GVT Plan, RO-EST-PL-5011 */ TARGET NAME = "N/A" TARGET TYPE = HOUSEKEEPING

```
N.B.: The formats of the following sets of data elements are identical to
those in Section 4.1, so their descriptions are omitted here:
    /* INSTRUMENT AND MISSION DESCRIPTIVE DATA ELEMENTS */
    /* PRODUCT AND DATA SET DATA ELEMENTS */
    /* STRUCTURE DEFINITION OF INSTRUMENT PARAMETER OBJECTS */
    /* DESCRIPTION OF INSTRUMENT */
    /* PARAMETERS INDEX FILE DEFINITION */
    /* MISSION SPECIFIC KEYWORDS */
```

/* DATA OBJECT DEFINITION */ OBJECT = HOUSEKEEPING HEADER BYTES = 3903 HEADER TYPE = PLAIN TEXT END OBJECT = HOUSEKEEPING_HEADER OBJECT = HOUSEKEEPING TABLE INTERCHANGE FORMAT = ASCII ROWS = 865 COLUMNS = 73 ROW BYTES = 442 DESCRIPTION = "Parameter values from Alice housekeeping telemetry (HKTM) data." OBJECT = COLUMN NAME = "ScetR" DATA TYPE = IEEE REAL STARTBYTE = 1 BYTES = 15 DESCRIPTION = "SpaceCraft Event Time (secs since 1970)" UNIT = counts END OBJECT = COLUMN OBJECT = COLUMN

```
NAME
           = "ScetC"
 DATA_TYPE = CHARACTER
STARTBYTE = 15
 BYTES = 24
 DESCRIPTION = "SpaceCraft Event Time (UTC)"
 UNIT = "N/A"
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "InstTime"
DATA_TYPE = IEEE_REAL
 STARTBYTE = 39
 BYTES = 15
 DESCRIPTION = "Alice instrument time (OBT, seconds)"
 UNIT = counts
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "TOffset"
 DATA TYPE = IEEE REAL
 STARTBYTE = 54
 BYTES = 15
 DESCRIPTION = "Alice OBT offset from SCET (seconds)"
 UNIT = "N/A"
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "InstSync"
 DATA_TYPE = MSB_INTEGER
STARTBYTE = 69
 BYTES = 2
 DESCRIPTION = "Alice instrument time synchronized"
 UNIT = "N/A"
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "HvOn"
DATA_TYPE = CHARACTER
 STARTBYTE = 71
 BYTES = 4
 DESCRIPTION = "HkHvOn"
 UNIT = "N/A"
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "DetStim"
 DATA TYPE = CHARACTER
 STARTBYTE = 75
 BYTES = 5
 DESCRIPTION = "HkDetStim"
 UNIT = "N/A"
END OBJECT = COLUMN
```

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OBJECT = COLUMN NAME = "AcqMode" DATA_TYPE = CHARACTER STARTBYTE = 80 BYTES = 10 DESCRIPTION = "HkAcqCountRate/HkAcqMode" UNIT = "N/A" END OBJECT = COLUMN OBJECT = COLUMN NAME = "AcqActiv" DATA TYPE = CHARACTER STARTBYTE = 90 BYTES = 4 DESCRIPTION = "HkAcqActive" UNIT = "N/A" END OBJECT = COLUMN OBJECT = COLUMN NAME = "DetDoor" DATA TYPE = CHARACTER STARTBYTE = 94 BYTES = 7 DESCRIPTION = "HkDetDoor" UNIT = "N/A"END OBJECT = COLUMN OBJECT = COLUMN NAME = "ApDoor" DATA_TYPE = CHARACTER STARTBYTE = 101 BYTES = 8 DESCRIPTION = "HkApDoor" UNIT = "N/A"END OBJECT = COLUMN OBJECT = COLUMN NAME = "Code" DATA TYPE = CHARACTER STARTBYTE = 109 BYTES = 8 DESCRIPTION = "EEPromCode/EEPromPage" UNIT = "N/A"END OBJECT = COLUMN OBJECT = COLUMN NAME = "State" DATA TYPE = CHARACTER STARTBYTE = 117 BYTES = 9 DESCRIPTION = "HkState"

```
UNIT = "N/A"
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "MirrHtr"
 DATA TYPE = CHARACTER
 STARTBYTE = 126
 BYTES = 4
 DESCRIPTION = "HkMirrorHtrOn"
 UNIT = "N/A"
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "GratHtr"
 DATA TYPE = CHARACTER
 STARTBYTE = 130
 BYTES = 4
 DESCRIPTION = "HkGratingHtrOn"
 UNIT = "N/A"
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "CntRate"
DATA_TYPE = MSB_INTEGER
 STARTBYTE = 134
 BYTES = 6
 DESCRIPTION = "HkCountRate"
 UNIT = Hz
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "CntRateL"
 DATA TYPE = MSB INTEGER
 STARTBYTE = 140
 BYTES = 6
 DESCRIPTION = "HkCountRateLo"
       = Hz
 UNIT
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "CntRateH"
 DATA TYPE = MSB INTEGER
 STARTBYTE = 146
 BYTES
           = 6
 DESCRIPTION = "HkCountRateHi"
 UNIT = Hz
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "HvSetR"
DATA_TYPE = MSB_INTEGER
 STARTBYTE = 152
```

```
BYTES
            = 4
  DESCRIPTION = "HkHvSet"
  UNIT = counts
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "HvSetC"
DATA_TYPE = MSB_INTEGER
 STARTBYTE = 156
BYTES = 5
 DESCRIPTION = ""
 UNIT = Volt
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "McpVR"
DATA_TYPE = MSB_INTEGER
  STARTBYTE = 161
 BYTES = 4
 DESCRIPTION = "HkMcpV"
 UNIT = counts
END_OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "McpVC"
  DATA TYPE = MSB INTEGER
  STARTBYTE = 165
 BYTES = 5
  DESCRIPTION = ""
 UNIT = Volt
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "McpVLoR"
DATA_TYPE = MSB_INTEGER
 STARTBYTE = 170
BYTES = 4
 DESCRIPTION = "HkMcpVlo"
 UNIT = counts
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "McpVLoC"
DATA_TYPE = MSB_INTEGER
STARTBYTE = 174
 BYTES = 5
 DESCRIPTION = ""
  UNIT = Volt
END OBJECT = COLUMN
OBJECT = COLUMN
  NAME = "McpVHiR"
```

```
DATA_TYPE = MSB_INTEGER
STARTBYTE = 179
BYTES = 4
  DESCRIPTION = "HkMcpVhi"
 UNIT = counts
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "McpVHiC"
DATA_TYPE = MSB_INTEGER
 STARTBYTE = 183
 BYTES = 5
 DESCRIPTION = ""
 UNIT = Volt
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "Anovr"
 DATA TYPE = MSB INTEGER
 STARTBYTE = 188
 BYTES = 4
 DESCRIPTION = "HkAnodeV"
 UNIT = counts
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "AnoVC"
  DATA TYPE = MSB INTEGER
 STARTBYTE = 192
 BYTES = 4
 DESCRIPTION = ""
 UNIT = Volt
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "AnoVLoR"
DATA_TYPE = MSB_INTEGER
  STARTBYTE = 196
 BYTES = 4
 DESCRIPTION = "HkAnodeVlo"
 UNIT = counts
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "AnoVLoC"
 DATA TYPE = MSB INTEGER
 STARTBYTE = 200
 BYTES = 4
 DESCRIPTION = ""
 UNIT = Volt
END OBJECT = COLUMN
```

```
OBJECT = COLUMN
 NAME = "AnoVHiR"
DATA_TYPE = MSB_INTEGER
STARTBYTE = 204
 BYTES = 4
 DESCRIPTION = "HkAnodeVhi"
 UNIT = counts
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "AnoVHiC"
 DATA TYPE = MSB INTEGER
 STARTBYTE = 208
 BYTES
           = 4
 DESCRIPTION = ""
 UNIT = Volt
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "StrIR"
 DATA TYPE = MSB INTEGER
 STARTBYTE = 212
 BYTES = 4
 DESCRIPTION = "HkStripI"
 UNIT = counts
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "StrIC"
 DATA_TYPE = IEEE_REAL
STARTBYTE = 216
 BYTES = 6
 DESCRIPTION = ""
 UNIT = microAmp
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "StrILOR"
 DATA TYPE = MSB INTEGER
 STARTBYTE = 222
 BYTES = 4
 DESCRIPTION = "HkStripIlo"
 UNIT = counts
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "StrILoC"
 DATA TYPE = IEEE REAL
 STARTBYTE = 226
 BYTES = 6
 DESCRIPTION = ""
 UNIT = microAmp
```

```
END_OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "StrIHiR"
 DATA TYPE = MSB INTEGER
 STARTBYTE = 232
 BYTES = 4
 DESCRIPTION = "HkStripIhi"
 UNIT = counts
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "StrIHiC"
 DATA TYPE = IEEE REAL
 STARTBYTE = 236
 BYTES
            = 6
 DESCRIPTION = ""
 UNIT = microAmp
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "T_Mirr1R"
DATA_TYPE = MSB_INTEGER
 STARTBYTE = 242
 BYTES = 4
 DESCRIPTION = "HkMirrorTempPri"
 UNIT = counts
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "T_Mirr1C"
 DATA TYPE = IEEE REAL
 STARTBYTE = 246
 BYTES = 6
 DESCRIPTION = ""
 UNIT = deqC
END_OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "T Mirr2R"
 DATA TYPE = MSB INTEGER
 STARTBYTE = 252
           = 4
 BYTES
 DESCRIPTION = "HkMirrorTempSec"
 UNIT = counts
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "T Mirr2C"
 DATA_TYPE = IEEE_REAL
STARTBYTE = 256
 BYTES = 6
```

```
DESCRIPTION = ""
 UNIT = degC
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "TS MirrR"
 DATA TYPE = MSB INTEGER
 STARTBYTE = 262
 BYTES = 4
 DESCRIPTION = "HkMirrorHtrSet"
 UNIT = counts
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "TS MirrC"
 DATA_TYPE = IEEE_REAL
STARTBYTE = 266
 BYTES = 6
 DESCRIPTION = ""
 UNIT = degC
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "T_Grat1R"
DATA_TYPE = MSB_INTEGER
 STARTBYTE = 272
 BYTES = 4
 DESCRIPTION = "HkGratingTempPri"
 UNIT = counts
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "T Grat1C"
 DATA TYPE = IEEE REAL
 STARTBYTE = 276
 BYTES = 6
 DESCRIPTION = ""
 UNIT = degC
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "T Grat2R"
 DATA_TYPE = MSB_INTEGER
STARTBYTE = 282
 BYTES = 4
 DESCRIPTION = "HkGratingTempSec"
 UNIT = counts
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "T_Grat2C"
 DATA TYPE = IEEE REAL
```

```
STARTBYTE = 286
 BYTES = 6
 DESCRIPTION = ""
 UNIT = degC
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "TS_GratR"
DATA TYPE = MSB INTEGER
 STARTBYTE = 292
 BYTES = 4
 DESCRIPTION = "HkGratingHtrSet"
 UNIT = counts
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "TS_GratC"
 DATA TYPE = IEEE REAL
 STARTBYTE = 296
 BYTES = 6
 DESCRIPTION = ""
 UNIT = degC
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "DetDiscR"
 DATA TYPE = MSB INTEGER
 STARTBYTE = 302
 BYTES = 4
 DESCRIPTION = "HkDetDiscr"
 UNIT = counts
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "DetDiscC"
DATA_TYPE = IEEE_REAL
 STARTBYTE = 306
 BYTES = 5
 DESCRIPTION = ""
 UNIT = Volt
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "CycRem"
 DATA TYPE = MSB INTEGER
 STARTBYTE = 311
 BYTES = 4
 DESCRIPTION = "HkCyclesRemaining"
 UNIT = counts
END OBJECT = COLUMN
OBJECT = COLUMN
```

```
= "LastPrsR"
 NAME
 DATA_TYPE = MSB_INTEGER
STARTBYTE = 315
 BYTES = 4
 DESCRIPTION = "HkLastPressure"
 UNIT = counts
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "LastPrsC"
DATA_TYPE = IEEE_REAL
 STARTBYTE = 319
 BYTES = 6
 DESCRIPTION = ""
 UNIT = mbar
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "LastPRaR"
 DATA TYPE = MSB INTEGER
 STARTBYTE = 325
 BYTES = 4
 DESCRIPTION = "HkLastPrsRate"
 UNIT = counts
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "LastPRaC"
 DATA_TYPE = IEEE_REAL
STARTBYTE = 329
 BYTES = 6
 DESCRIPTION = ""
 UNIT = mbar/s
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "PresTime"
DATA_TYPE = IEEE_REAL
 STARTBYTE = 335
 BYTES = 15
 DESCRIPTION = "HkPressureTime (OBT, seconds)"
 UNIT = counts
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "PresScet"
 DATA TYPE = CHARACTER
 STARTBYTE = 350
 BYTES = 24
 DESCRIPTION = "HkPressureTime (SCET, UTC)"
 UNIT = "N/A"
END OBJECT = COLUMN
```

```
OBJECT = COLUMN
 NAME = "Unsafe"
DATA_TYPE = MSB_INTEGER
 STARTBYTE = 374
 BYTES = 6
 DESCRIPTION = "HkUnsafeTimer"
 UNIT = sec
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "Safety"
 DATA TYPE = MSB INTEGER
 STARTBYTE = 380
 BYTES = 4
 DESCRIPTION = "HkSafety"
 UNIT = "N/A"
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "OverRide"
 DATA TYPE = CHARACTER
 STARTBYTE = 384
BYTES = 4
 DESCRIPTION = "HkSafetyOverride"
 UNIT = "N/A"
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "SafeMask"
 DATA_TYPE = MSB_INTEGER
 STARTBYTE = 388
 BYTES = 4
 DESCRIPTION = "HkSafetyMask"
 UNIT = "N/A"
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "T DElecR"
 DATA TYPE = MSB INTEGER
 STARTBYTE = 392
 BYTES = 4
 DESCRIPTION = "HkDetElecTemp"
 UNIT = counts
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "T DElecC"
 DATA TYPE = IEEE REAL
 STARTBYTE = 396
 BYTES = 6
 DESCRIPTION = ""
```

```
UNIT = deqC
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "T DHousR"
 DATA TYPE = MSB INTEGER
 STARTBYTE = 402
 BYTES = 4
 DESCRIPTION = "HkDetHousingTemp"
 UNIT = counts
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "T DHousC"
 DATA_TYPE = IEEE_REAL
STARTBYTE = 406
 BYTES = 6
 DESCRIPTION = ""
 UNIT = degC
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "T_CNDHR"
DATA_TYPE = MSB_INTEGER
 STARTBYTE = 412
 BYTES = 4
 DESCRIPTION = "HkCdhElecTemp"
 UNIT = counts
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "T CNDHC"
 DATA TYPE = IEEE REAL
 STARTBYTE = 416
 BYTES = 6
 DESCRIPTION = ""
 UNIT = degC
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "T_IHousR"
 DATA TYPE = MSB INTEGER
 STARTBYTE = 422
 BYTES = 4
 DESCRIPTION = "HkInstrHousingTemp"
 UNIT = counts
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "T_IHousC"
DATA_TYPE = IEEE_REAL
 STARTBYTE = 426
```

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```
BYTES
           = 6
 DESCRIPTION = ""
 UNIT = deqC
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "TestStat"
 DATA TYPE = MSB INTEGER
 STARTBYTE = 432
 BYTES = 4
 DESCRIPTION = "HkTestStatus"
 UNIT = "N/A"
END OBJECT = COLUMN
OBJECT = COLUMN
 NAME = "SwVers"
 DATA TYPE = IEEE REAL
 STARTBYTE = 436
 BYTES = 6
 DESCRIPTION = "HkAfsVersion"
 UNIT = "N/A"
END OBJECT = COLUMN
```

```
END OBJECT = HOUSEKEEPING TABLE
```

4.8 Sample FITS header

Below is a sample of a FITS header from an *Alice* histogram calibrated science data file. Note this is the *main* header, and that the headers for *extensions* in the FITS files are "minimal" headers that do not contain much descriptive information about the contents of the data block. Some of the comments (the text following the forward slash in a line in a FITS header record) are truncated because the line length is limited to 80 characters.

```
SIMPLE =
                           T /image conforms to FITS standard
BITPIX =
                         -32 /bits per data value
                           2 /number of axes
NAXIS =
NAXIS1 =
                        1024 /
NAXIS2 =
                          32 /
EXTEND =
                           T /file may contain extensions
                     'SwRI' / Southwest Research Institute (R)
ORIGIN =
DATE = '2004-09-20T22:37:16' / file creation/processing date/time
LIMA VER=
                       '0.1' / LIMA software version number
LIMA DAT=
               ' 2004-09-03' / LIMA software version date
MISSION =
                   'Rosetta' / ESA comet exploration mission
                    'R-Alice' / Ultra-Violet Spectrometer
INSTRUME=
COMMENT Alice Histogram data acquisition: Received Histogram data
BUNIT = 'PHOTONS CM-2 S-1' /Data units are photons per square cm per
second
                      0.00000 /
BZERO =
```

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1.00000 / BSCALE = COMMENT Science TM Packet re-assembly results TOTLPKTS= 19 / expected number of packets RECVPKTS= 19 / received number of packets LOSTPKTS= 0 / missing packets (bitmask) / COMMENT Acquisition common: used calibration files L CALDIR= '../../data/cal/lima/' / directory of LIMA calibration data 'ra hvs 000.cal' / high voltage set calibration file CALHVSET= 'ra mcp 000.cal' / microchannel plate voltage calibration file CALMCPV = 'ra ano 000.cal' / anode voltage calibration file CALANOV = 'ra stp 000.cal' / strip current calibration file CALSTPI = CALDISC = 'ra_dis_000.cal' / discriminator voltage calibration file CALTEMP = 'ra_tmp_000.cal' / temperature calibration file (all channels) COMMENT Acquisition common section 1 SCETR = 1082680611.696 / SpaceCraft Event Time (seconds since 1970) SCETC = '2004-04-23T00:36:51.696' / SpaceCraft Event Time (UTC) INSTTIME= 41301397.234 / Alice Instrument time (OBT) 1041379214.462 / Alice OBT offset from SCET (seconds) TOFFSET = 132 / Mcp high level during acquisition (ADC MCPMAXVR= counts) MCPMAXVC= -3727 / McpMaxV (Volt) 132 / Mcp low level during acquisition (ADC MCPMINVR= counts) MCPMINVC= -3727 / McpMinV (Volt) MCPSTRVR= 132 / Mcp level at start of acquisition (ADC counts) -3727 / McpStartV (Volt) MCPSTRVC= 132 / Mcp level at end of acquisition (ADC counts) MCPSTPVR= MCPSTPVC= -3727 / McpStopV (Volt) 51 / strip current high level during acquisition STRMAXIR= (AD STRMAXIC= 5.26 / StripMaxI (microAmp) 45 / strip current low level during acquisition STRMINIR= (ADC STRMINIC= 4.97 / StripMinI (microAmp) STRSTRIR= 48 / strip current at start of acquisition (ADC coun STRSTRIC= 5.11 / StripStartI (microAmp) 47 / strip current at end of acquisition (ADC STRSTPIR= counts STRSTPIC= 5.07 / StripStopI (microAmp) / COMMENT Common Housekeeping section; Hk acquired at start of data dump HVON = 'on' / HkHvOn (on/off) DETSTIM = 'stim' / HkDetStim (stim/off) = 'Histogram' / AcquisitionMode ACQMODE (PixelList/Histogram/CountRate) ACOACTIV= 'off' / HkAcqActive (on/off) DETDOOR = 'open' / HkDetDoor (open/closed)

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APDOOR = 'open' / HkApDoor (error/closed/open/between) 'PROM' / HkEEPromCode/HkEEPromPage (PROM/EEPROM0-3) = CODE 'dump' / HkState STATE (checkout/acquire/dump/safe/hold/safedu 'off' / HkMirrorHtrOn (on/off) MIRRHTR = 'off' / HkGratingHtrOn (on/off) GRATHTR = 204 / HkCountRate (Hz) CNTRATE = 204 / HkCountRateLo (Hz) CNTRATEL= 250 / HkCountRateHi (Hz) CNTRATEH= 136 / High Voltage Set Point (DAC counts) HVSETR = -3821 / HkHvSet (Volt) 132 / Mcp Voltage (A -3819 / HkMcpV (Volt) HVSETC = 132 / Mcp Voltage (ADC counts) MCPVR = MCPVC = MCPVLOR = 132 / Mcp Voltage low level during Hk period (ADC cou -3819 / HkMcpVlo (Volt) MCPVLOC = 132 / Mcp Voltage high level during Hk period (ADC MCPVHIR = со -3819 / HkMcpVhi (Volt) MCPVHIC = 225 / Anode Voltage (ADC counts) ANOVR = ANOVC = -596 / HkAnodeV (Volt) 225 / Anode Voltage low level during Hk period ANOVLOR = (ADC C -596 / HkAnodeVlo (Volt) ANOVLOC = 226 / Anode Voltage high level during Hk period ANOVHIR = (ADC ANOVHIC = -599 / HkAnodeVhi (Volt) STRPIR = 47 / Strip Current (ADC counts) STRPIC = 5.07 / HkStripI (microAmp) STRPILOR= 46 / Strip Current low level during Hk period (ADC C STRPILOC= 5.02 / HkStripIlo (microAmp) 50 / Strip Current high level during Hk period STRPIHIR= (ADC STRPIHIC= 5.21 / HkStripIhi (microAmp) T MIRR1R= 49 / Mirror Temperature 1 (ADC counts) T MIRR1C= 4.1 / HkMirrorTempPri (degC) 48 / Mirror Temperature 2 (ADC counts) T MIRR2R= T MIRR2C= 3.7 / HkMirrorTempSec (degC) TS MIRRR= 0 / Mirror Heater Set Point (ADC counts) -33.3 / HkMirrorHtrSet (degC) TS MIRRC= T GRAT1R= 40 / Grating Temperature 1 (ADC counts) T GRAT1C= 0.0 / HkGratingTempPri (degC) T GRAT2R= 40 / Grating Temperature 2 (ADC counts) T GRAT2C= 0.0 / HkGratingTempSec (degC) TS GRATR= 0 / Grating Heater Set Point (ADC counts) -33.3 / HkGratingHtrSet (degC) TS GRATC= 8 / Detector Dicriminator readback (ADC counts) DETDISCR= 0.09 / HkDetDiscr (Volt) DETDISCC= CYCLSREM= 3 / HkCyclesRemaining (in current MAD command) 90 / Last reported ROSINA pressure (RosinaReal) LASTPRSR= 90 / Last reported Rubing 5E-10 / HkLastPressure (mbar) LASTPRSC=

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LASTPRAR= 255 / Last reported ROSINA pressure trend (RosinaReal -1.0 / HkLastPrsRate (mbar/s) LASTPRAC= PRESTIME= 41301349.449 / HkPressureTime (instrument time) PRESSCET= '2004-04-23T00:36:03.912' / HkPressureTime (SCET) SAFETIME= 0 / HkUnsafeTimer (sec, timeout remaining) 0 / HkSafety (bitmask) SAFETY = 'off' / HkSafetyOverride (on/off) SAFEOVER= 32 / HkSafetyMask (bitmask) SAFEMASK= 60 / Detector Electronics Temperature (ADC T DELECR= counts) 8.9 / HkDetElecTemp (degC) T DELECC= T DHOUSR= 46 / Detector Housing temperature (ADC counts) 70 / C&DH Electronics Tem 13.3 / HkCdhElecTemp (degC) 43 / Instrument Housing Te 1.5 / HkInstrue T DHOUSC= 2.8 / HkDetHousingTemp (degC) 70 / C&DH Electronics Temperature (ADC counts) T CDHR = T CDHC = 43 / Instrument Housing Temperature (ADC counts) T IHOUSR= T IHOUSC= 1.5 / HkInstrHousingTemp (degC) TESTSTAT= 0 / HkTestStatus (bitmask) '2.03' / HkAfsVersion (version number of instrument SOFTVERS= SW) COMMENT Common Instrument Parameter section HSNSGRAT= 'pri' / PmHtrSenseGrating (pri/sec) 'pri' / PmHtrSenseMirror (pri/sec) HSNSMIRR= 'enabled' / PmDoorEnable (enabled/disabled) DOORENA = DACADCFR= 231 / PmDacAdcFactor (/240) DACADCFC= 0.963 / conversion factor from HvSet to McpV 'ground' / PmEventSafety (ground/on-board) EVNTSAFT= DISCSETR= 10 / Discriminator acquisition setpoint (DAC counts) 0.12 / PmDiscriminator (Volt) '0-0000' / PmHotSeg1: spatial-spectral '0-0000' / PmHotSeg2: spatial-spectral '0-0000' / PmHotSeg3: spatial-spectral '0-0000' / PmHotSeg5: spatial-spectral '0-0000' / PmHotSeg6: spatial-spectral '0-0000' / PmHotSeg7: spatial-spectral '0-0000' / PmHotSeg8: spatial-spectral '0-0000' / PmHotSeg8: spatial-spectral 0.12 / PmDiscriminator (Volt) DISCSETC= HOTSEG1 = HOTSEG2 = HOTSEG3 = HOTSEG4 = HOTSEG5 = HOTSEG6 = HOTSEG7 =HOTSEG8 = 136 / HV acquisition setpoint (DAC counts) HVLEVELR= -3821 / PmHvLevel (Volt) HVLEVELC= 4 / HV step size (DAC counts) HVSTEPR = HVSTEPC = -92 / PmHvStepSize (Volt) HVRTIMR = 2 / HV runtime (kernel ticks) 0.2 / PmHvRunTime (sec) HVRTIMC = 168 / Maximum Strip Current (ADC counts) MAXSTRIR= 10.96 / PmHvMaxStripI (microAmp) MAXSTRIC= 55 / Minimum Anode Voltage (ADC counts) MINANOVR= -150 / PmHvMinAnodeV (Volt) MINANOVC= MAXANOVR= 245 / Maximum Anode Voltage (ADC counts) MAXANOVC= -649 / PmHvMaxAnodeV (Volt)

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MCPTOLR = 8 / HV-Mcp tolerance (delta ADC counts) 184 / PmHvMcpTol (Volt) MCPTOLC = 4 / PmHvFailAnode (count) ANOFAIL = 4 / PmHvFailMcp (count) MCPFAIL = STPFAIL = 4 / PmHvFailStrip (count) 'enabled' / Door sensor enabled (top bit PmDoorTimeout) DOORSENS= DOORTIME= 2 / Aperture Door control time, PmDoorTimeout (sec) 10 / Detector event counter interval (kernel CNTINTVR= ticks) CNTINTVC= 0.96 / PmCounterInterval (sec) 100 / PmCountSlots (fixed at afs compile time) CNTSLOTS= 21 / Pressure Limit (Rosina Real) MAXPRESR= 1E-05 / PmMaxPressure (mbar) MAXPRESC= 60 / Pressure trend time horizon (sec*10) PTIMMARR= 600 / PmPressTimeMargin (sec) PTIMMARC= 57.7 / PmMaxTemp (degC) 255 / PmTempWin 192 / Maximum temperature (ADC counts) T MAXR = T MAXC = 255 / PmTempMask (bitmask) T MASK = 8 / Grating heater temperature tolerance (ADC TOLGRATR= cnts) 3.0 / PmHtrTolGrating (degC) TOLGRATC= TOLMIRRR= 8 / Mirror heater temperature tolerance (ADC counts 3.0 / PmHtrTolMirror (degC) TOLMIRRC= TSTMASK = 0 / PmTestMask (bitmask) 5 / PmHkRate (sec, nk your 'diag' / PmDiagEnable (diag/off) 40000 / PmMaxCountRate (Hz, saf HKRATE = 5 / PmHkRate (sec, Hk generation rate) DIAGENA = 40000 / PmMaxCountRate (Hz, safety limit) MAXCOUNT= SAFTIME = 600 / PmSafetyTimeout (sec) MAXDUST = 30 / PmMaxDust (part/dm2/sec?, safety limit) COMMENT Acquisition common section 2 PRSMAXR = 90 / high ROSINA pressure reported during acquisitio 5E-10 / PressMax (mbar) PRSMAXC = 90 / low ROSINA pressure reported during PRSMINR = acquisition PRSMINC = 5E-10 / PressMin (mbar) STRTTIME= 41300188.230 / exposure start time (instrument time) 5E-10 / PressMin (mbar) STRTSCET= '2004-04-23T00:16:42.693' / exposure start time (SCET) DATE-OBS= '2004-04-23T00:16:42.693' / exposure start time (SCET) 41301397.227 / exposure stop time (instrument time) STOPTIME= STOPSCET= '2004-04-23T00:36:51.689' / exposure stop time (SCET) EXPTIME = 1208.996 / Calculated actual exposure time (seconds) DUMPNO = 0 / DumpNo, dump specification number (0-7) DETCNTS = 257113 / DetectorCounts, total counts from detector even / COMMENT Histogram acquisition specific keywords 1200 / AcqTime, commanded integration time (sec) ACOTIME = WILOSPEC= 0 / WinLowSpectral, spectral window lower

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(inclusiv 1023 / WinHighSpectral, spectral window upper WIHISPEC= (inclusi WICOSPEC= 1 / CollapseSpectral, spectral collapse factor 0 / WinLowSpatial, spatial window lower WILOSPAT= (inclusive) WIHISPAT= 31 / WinHighSpatial, spatial window upper (inclusive 1 / CollapseSpatial, spatial collapse factor WICOSPAT= / COMMENT Calculated event statistics COMMENT Flat field is already normalized to average of active non-gap region 123993 / Events counted in Histogram EVENTS = PHDEVNTS= 124046 / Events counted in PHD 32768.0 /Original Data is unsigned Integer O BZERO = _____ MIKE BEG= 'Jan 22 13:28:17 2007' / START MIKE KEYWORD BLOCK MIKE VER= '3.0 [2006 Jan 24]' /Version of Mike pipeline code K_MODE = 'ACQMODE ' / Keyword containing the mode name K_ETIME = 'EXPTIME ' / Keyword for the effective exposure time FILE_IN = './2004/04/ra_040423001643_his0_eng.fit' / Input file for processing FILE OUT= 'ra 040423001643 his0 sci.fit' / Output file after processing DATA DIR= '/Users/steffl/ralice/data/' / Top-level R-Alice data directory BADFILE = ' / FITS file of bad pixel mask array BADFLAG = -1 / Bad pixel mask flag -666 / Bad pixel value BADVALUE= DEADFILE= 'deadtime/ra_dead_003.txt' / Deadtime correction file 1 / DEADFLAG= DEADTYPE= 'FUN ' / Correct using FUNction or lookup table (LUT)? DEADCORR= 'TOTAL ' / Correct by TOTAL or each PIXEL count rate? DARKFILE= 'dark/RA DARK 003.FIT' / Dark image filename 1 / DARKFLAG= FLATFILE= 'flat/RA_FLAT_003.FIT' / Flag field image filename FLATFLAG= 1 / FLATNORM= 'NONE ' / How to normalize flatfield WCALFLAG= 1 / WCALPRO = 'ralice wavecal' / IDL program to perform wavelength calibration WCALPARS= 'T_DELECC' / keywords for parameters to use for wave cal -1 / Wavelength linear rectify flag WRECFLAG= WRECPRO = 'alice_waverec' / IDL program to perform wavelength rectify 1 / AEFFFLAG= AEFFPRO = 'ralice_aeff' / IDL program to get effective area AEFFPARS= 'T DELECC' / keywords for parameters to get effective area KERNFILE= '/Users/steffl/kernels/rosetta/rosetta kernels.txt' / filename of the LOG FILE= ' / Filename to save log file (default = append

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to LOG MAIL= ' / address (if any) to e-mail log file MIKE ERR= 0 / MIKE END= 'Mon Jan 22 13:28:18 2007' / END MIKE KEYWORD BLOCK _____ COMMENT BEGIN GEOMETRY KEYWORD BLOCK COMMENT Values calulated for midpoint of observation 46.5509243021 / Right Ascension of ALICE Boresight [deg.] RA = -80.2268061330 / Declination of ALICE Boresight [deg.] DEC = COMMENT S/C -> Sun Geometry SCSUNX = 106162996.345 / J2000 EME S/C -> Sun X vector [km] 77717972.0051 / J2000 EME S/C -> Sun Y vector [km] SCSUNY = 32801372.5140 / J2000 EME S/C -> Sun Z vector [km] SCSUNZ = SCSUNVX = SCSUNVY = -21.8059492212 / X vel. of Sun wrt S/C in [km/s] (J2000 EME) 22.2283091967 / Y vel. of Sun wrt S/C in [km/s] (J2000 EME) 9.51321161942 / Z vel. of Sun wrt S/C in [km/s] (J2000 EME) SCSUNVZ = SCSUNLON=120.536950103 / Sub solar longitude [deg.] (heliographic)SCSUNLAT=-4.04206437239 / Sub solar latitude [deg.] COMMENT S/C -> Target Geometry TARGNAME= 'CALIBRATION' / Name of observation target TARGTYPE= 'CALIBRATION' / Type of observation target SCTARGX = -1.00000E+32 / J2000 EME S/C -> target X vector [km] SCTARGY = -1.00000E+32 / J2000 EME S/C -> target Y vector [km] -1.00000E+32 / J2000 EME S/C -> target Z vector [km] SCTARGZ = -1.00000E+32 / Distance to target center [km] TARGDIST= -1.00000E+32 / X vel. of target wrt S/C in [km/s] (J2000 SCTARGVX= EME) SCTARGVY= -1.00000E+32 / Y vel. of target wrt S/C in [km/s] (J2000 EME) SCTARGVZ= -1.00000E+32 / Z vel. of target wrt S/C in [km/s] (J2000 EME) -1.00000E+32 / Sub Spacecraft planetographic longitude SCLON = [deg.] -1.00000E+32 / Sub Spacecraft latitude [deg.] SCLAT = SCALT = -1.00000E+32 / Spacecraft altitude [km] -1.00000E+32 / Phase angle (Sun-Target-S/C) [deg] TARGPHAS= COMMENT Row-specific Geometry Information COMMENT The ALICE slit is imaged onto detector rows 5-23 (inclusive) ROW5RA = 44.1262677608 / Right Ascension [deg.] -77.1829689075 / Declination [deg.] ROW5DEC = ROW5LON = -1.00000E+32 / Planetographic Lon. of target along LOS [deg.] ROW5LAT = -1.00000E+32 / Lat. of target along LOS [deg.] -1.00000E+32 / Incidence angle at LOS intercept [deg.] ROW5INC = -1.00000E+32 / Emission angle at LOS intercept [deg.] ROW5EMI = 44.3162846028 / Right Ascension [deg.] ROW6RA =

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ROW6DEC = -77.4881388348 / Declination [deg.] -1.00000E+32 / Planetographic Lon. of target along LOS ROW6LON = [deq.] -1.00000E+32 / Lat. of target along LOS [deg.] ROW6LAT = -1.00000E+32 / Incidence angle at LOS intercept [deg.] ROW6INC = -1.00000E+32 / Emission angle at LOS intercept [deg.] ROW6EMI = 44.5156451580 / Right Ascension [deg.] ROW7RA = ROW7DEC = -77.7931691390 / Declination [deg.] -1.00000E+32 / Planetographic Lon. of target along LOS ROW7LON = [deg.] -1.00000E+32 / Lat. of target along LOS [deg.] ROW7LAT = ROW7INC = -1.00000E+32 / Incidence angle at LOS intercept [deg.] ROW7EMI = -1.00000E+32 / Emission angle at LOS intercept [deg.] 44.7250630416 / Right Ascension [deg.] ROW8RA = ROW8DEC = -78.0980488551 / Declination [deg.] ROW8LON = -1.0000025102 -1.00000E+32 / Planetographic Lon. of target along LOS [deq.] ROW8LAT = -1.00000E+32 / Lat. of target along LOS [deg.] -1.00000E+32 / Incidence angle at LOS intercept [deg.] ROW8INC = ROW8EMI = -1.00000E+32 / Emission angle at LOS intercept [deg.] 44.9453257058 / Right Ascension [deg.] ROW9RA = ROW9DEC = -78.4027656823 / Declination [deg.] ROW9LON = -1.00000E+32 / Planetographic Lon. of target along LOS [deq.] -1.00000E+32 / Lat. of target along LOS [deg.] ROW9LAT = ROW9INC = -1.00000E+32 / Incidence angle at LOS intercept [deg.] ROW9EMI = -1.00000E+32 / Emission angle at LOS intercept [deg.] ROW10RA = 45.1773053188 / Right Ascension [deg.] -78.7073073112 / Declination [deg.] ROW10DEC= ROW10LON= -1.00000E+32 / Planetographic Lon. of target along LOS [deq.] ROW10LAT= -1.00000E+32 / Lat. of target along LOS [deg.] ROW10INC= -1.00000E+32 / Incidence angle at LOS intercept [deg.] ROW10EMI= -1.00000E+32 / Emission angle at LOS intercept [deg.] ROW11RA = 45.4219681971 / Right Ascension [deg.] ROW11DEC= -79.0116586098 / Declination [deg.] -1.00000E+32 / Planetographic Lon. of target along LOS ROW11LON= [deg.] -1.00000E+32 / Lat. of target along LOS [deg.] ROW11LAT= ROW11INC= -1.00000E+32 / Incidence angle at LOS intercept [deg.] -1.00000E+32 / Emission angle at LOS intercept [deg.] ROW11EMI= 45.6803893947 / Right Ascension [deg.] ROW12RA = -79.3158033199 / Declination [deg.] ROW12DEC= -1.00000E+32 / Planetographic Lon. of target along LOS ROW12LON= [deq.] -1.00000E+32 / Lat. of target along LOS [deg.] ROW12LAT= ROW12INC= -1.00000E+32 / Incidence angle at LOS intercept [deg.] -1.00000E+32 / Emission angle at LOS intercept [deg.] ROW12EMI= 45.9537680482 / Right Ascension [deg.] ROW13RA = ROW13DEC= -79.6197235526 / Declination [deg.] ROW13LON= -1.00000E+32 / Planetographic Lon. of target along LOS [deg.]

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ROW13LAT= -1.00000E+32 / Lat. of target along LOS [deg.] -1.00000E+32 / Incidence angle at LOS intercept [deg.] ROW13INC= -1.00000E+32 / Emission angle at LOS intercept [deg.] ROW13EMI= 46.2434448721 / Right Ascension [deg.] ROW14RA = ROW14DEC= -79.9233987429 / Declination [deg.] ROW14LON= -1.00000E+32 / Planetographic Lon. of target along LOS [deg.] -1.00000E+32 / Lat. of target along LOS [deg.] ROW14LAT= -1.00000E+32 / Incidence angle at LOS intercept [deg.] ROW14INC= -1.00000E+32 / Emission angle at LOS intercept [deg.] ROW14EMI= ROW15RA = 46.5509243021 / Right Ascension [deg.] -80.2268061330 / Declination [deg.] ROW15DEC= ROW15LON= -1.00000E+32 / Planetographic Lon. of target along LOS [deg.] -1.00000E+32 / Lat. of target along LOS [deg.] ROW15LAT= -1.00000E+32 / Incidence angle at LOS intercept [deg.] ROW15INC= ROW15EMI= -1.00000E+32 / Emission angle at LOS intercept [deg.]

 ROW16RA =
 46.8778995329 / Right Ascension [d

 ROW16DEC=
 -80.5299199809 / Declination [deg.]

 46.8778995329 / Right Ascension [deg.] -1.00000E+32 / Planetographic Lon. of target along LOS ROW16LON= [deg.] ROW16LAT= -1.00000E+32 / Lat. of target along LOS [deg.] -1.00000E+32 / Incidence angle at LOS intercept [deg.] ROW16INC= -1.00000E+32 / Emission angle at LOS intercept [deg.] ROW16EMI= ROW17RA = 47.2262827887 / Right Ascension [deg.] 47.2262827887 / Right Ascension [d -80.8327111667 / Declination [deg.] ROW17DEC= ROW17LON= -1.00000E+32 / Planetographic Lon. of target along LOS [deg.] ROW17LAT= -1.00000E+32 / Lat. of target along LOS [deg.] ROW17INC= -1.00000E+32 / Incidence angle at LOS intercept [deg.] ROW17EMI= -1.00000E+32 / Emission angle at LOS intercept [deg.] 47.5982415262 / Right Ascension [deg.] ROW18RA = -81.1351467038 / Declination [deg.] ROW18DEC= ROW18LON= -1.00000E+32 / Planetographic Lon. of target along LOS [deq.] -1.00000E+32 / Lat. of target along LOS [deg.] ROW18LAT= -1.00000E+32 / Incidence angle at LOS intercept [deg.] ROW18INC= -1.00000E+32 / Emission angle at LOS intercept [deg.] ROW18EMI= ROW19RA = 47.9962413080 / Right Ascension [deg.] ROW19DEC= -81.4371886618 / Declination [deg.] -1.00000E+32 / Planetographic Lon. of target along LOS ROW19LON= [deq.] -1.00000E+32 / Lat. of target along LOS [deg.] ROW19LAT= ROW19INC= -1.00000E+32 / Incidence angle at LOS intercept [deg.] ROW19EMI= -1.00000E+32 / Emission angle at LOS intercept [deg.] 48.4230992177 / Right Ascension [deg.] ROW20RA = ROW20DEC= -81.7387941479 / Declination [deg.] ROW20LON= -1.00000E+32 / Planetographic Lon. of target along LOS [deg.] ROW20LAT= -1.00000E+32 / Lat. of target along LOS [deg.] ROW20INC= -1.00000E+32 / Incidence angle at LOS intercept [deg.] ROW20EMI= -1.00000E+32 / Emission angle at LOS intercept [deg.]

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ROW21RA = 48.8820465135 / Right Ascension [deg.] -82.0399135339 / Declination [deg.] ROW21DEC= -1.00000E+32 / Planetographic Lon. of target along LOS ROW21LON= [deg.] ROW21LAT= -1.00000E+32 / Lat. of target along LOS [deg.] ROW21INC= -1.00000E+32 / Incidence angle at LOS intercept [deg.] -1.00000E+32 / Emission angle at LOS intercept [deg.] ROW21EMI= ROW22RA = ROW22DEC= POW22LON= ROW22RA = 49.3768057093 / Right Ascension [deg.] -82.3404889540 / Declination [deg.] ROW22LON= -1.00000E+32 / Planetographic Lon. of target along LOS [deg.] -1.00000E+32 / Lat. of target along LOS [deg.] ROW22LAT= ROW22INC= -1.00000E+32 / Incidence angle at LOS intercept [deg.] ROW22EMI= -1.00000E+32 / Emission angle at LOS intercept [deg.] ROW23RA = 49.9116890755 / Right Ascension [deg.] ROW23DEC= -82.6404546406 / Declination [deg.] -1.00000E+32 / Planetographic Lon. of target along LOS ROW23LON= [deg.] ROW23LAT= -1.00000E+32 / Lat. of target along LOS [deg.] -1.00000E+32 / Incidence angle at LOS intercept [deg.] ROW23INC= -1.00000E+32 / Emission angle at LOS intercept [deg.] ROW23EMI= COMMENT SPICE files used in geometry calculations NUMBSPCK= 20 / Number of loaded SPICE kernels / SPICE kernel 0001 SPCK0001= 'naif0008.tls' SPCK0002= 'ROS 060511 STEP.TSC' / SPICE kernel 0002 SPCK0003= 'ROS V06.TF' / SPICE kernel 0003 SPCK0004= 'EARTHFIXEDIAU.TF' / SPICE kernel 0004 SPCK0004-EARINFIXEDIA0.1F/ SFICE kernel 0004SPCK0005='DE403-MASSES.TPC'/ SPICE kernel 0005SPCK0006='PCK00008.TPC'/ SPICE kernel 0006SPCK0007='DE405S.BSP'/ SPICE kernel 0007SPCK0008='C2002T7.bsp'/ SPICE kernel 0008 SPCK0009= 'TEMPEL1-9P-DI-P.BSP' / SPICE kernel 0009 SPCK0010= 'mar033 2000-2025.bsp' / SPICE kernel 0010 SPCK0011= 'sat164.bsp' / SPICE kernel 0011 SPCK0012= '2867_Steins_2004_2016.bsp' / SPICE kernel 0012 SPCK0013= '21_Lutetia_2004_2016.bsp' / SPICE kernel 0013 SPCK0014= '67P_CHURY_GERAS_2004_2016.BSP' / SPICE kernel 0014 SPCK0015= 'ORER_____00031.BSP' / SPICE kernel 0015 _____00047.BSP' / SPICE kernel 0016 SPCK0016= 'ORFR ____00047.BSP' / SPICE kernel 0017 SPCK0017= 'ORGR 00047.BSP' / SPICE kernel 0018 00047.BSP' / SPICE kernel 0019 SPCK0018= 'ORHR SPCK0019= 'ORMR SPCK0020= 'ATNR P040302093352 00046.BC' / SPICE kernel 0020 END