



Institut für Planetenforschung

**HRSC on MarsExpress**

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<b>Doc.-No.</b>	:	<b>HRSC-DLR-TN-4200-008</b>
<b>Issue</b>	:	<b>005</b>
<b>Date</b>	:	<b>16-Mar-2015</b>

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**MEX-HRSC To Planetary  
Science Archive  
Interface Control Document  
(EAICD)**

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## 1 INTRODUCTION

### 1.1 Purpose and Scope

The purpose of this EAICD (Experimenter to (Science) Archive Interface Control Document) is two fold. First it provides users of the HRSC instrument with detailed description of the product and a description of how it was generated, including data sources and destinations. Secondly, it is the official interface between the HRSC team and the Planetary Science Archive.

### 1.2 Archiving Authorities

The Planetary Data System Standard is used as archiving standard by

- NASA for U.S. planetary missions, implemented by PDS
- ESA for European planetary missions, implemented by the Research and Scientific Support Department (RSSD) of ESA

#### ESA's Planetary Science Archive (PSA)

ESA implements an online science archive, the PSA,

- to support and ease data ingestion
- to offer additional services to the scientific user community and science operations teams as e.g.
  - search queries that allow searches across instruments, missions and scientific disciplines
  - several data delivery options as
    - direct download of data products, linked files and data sets
    - ftp download of data products, linked files and data sets

The PSA aims for online ingestion of logical archive volumes and will offer the creation of physical archive volumes on request.

### 1.3 Contents

This document describes the data flow of the HRSC instrument on MarsExpress from the s/c until the insertion into the PSA for ESA. It includes information on how data were processed, formatted, labeled and uniquely identified. The document discusses general naming schemes for data volumes, data sets,

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data and label files. Standards used to generate the product are explained. Software that may be used to access the product is explained furtheron.

The design of the data set structure and the data product is given. Examples of these are given in the appendix.

## 1.4 Intended Readership

The staff of the archiving authority (Planetary Science Archive, ESA, RSSD, design team) and any potential user of the HRSC data.

## 1.5 Applicable Documents

1. Planetary Data System Preparation Workbook, February 1, 1995, Version 3.1, JPL, D-7669, Part1
2. Planetary Data System Standards Reference, June 1, 1999, Version 3.3, JPL, D-7669, Part 2
3. MarsExpress Archive Generation, Validation and Transfer Plan, J. Zender, , ESA-MEX-TN-4009, Rev. 1.0, 21-Jun-2001
4. HRSC: the High Resolution Stereo Camera of Mars Express, . G. Neukum1, R. Jaumann and the HRSC Co-Investigator and Experiment Team, ESA SP-1240, 2004.
5. Mars Express - HRSC Data Products Naming Convention, T. Roatsch, HRSC-DLR-TN-4200-001, Issue 005, 6-August-2003.
6. PDS Standards Reference , <http://pds.jpl.nasa.gov/documents/sr/index.html>
7. Planetary Science Data Archive Technical Note – Geometry and Position Information, Issue 3, Revision 1, J. Diaz del Rio, ESA RSSD Planetary Missions Division, SOP-RSSD-TN-010, 20-September-2004.
8. MarsExpress - HRSC Level-1 Product Description, T. Roatsch, HRSC-DLR-TN-4200-004, Issue 001, 10-November-2000.
9. MarsExpress - HRSC Level-2 Product Description, T. Roatsch, HRSC-DLR-TN-4200-006, Issue 001, 10-November-2000.
10. MarsExpress – HRSC VICAR Label Description Document, T. Roatsch, HRSC-DLR-TN-4200-002, Issue 004, 24-August-2004.
11. Flight User Manual for HRSC on MarsExpress, R. Pischel, HRSC-DLR-MA-4100-001/4D, 24-July-2002
12. The VICAR Image Processing System, <http://www-mipl.jpl.nasa.gov/external/vicar.html>
13. Navigation and Ancillary Information Facility (NAIF), <http://pds-naif.jpl.nasa.gov/>
14. Roatsch, T., Oberst, J., Acton, C., Zender, J.: SPICE Usage on the Mars Express Orbiter. XXVI. General Assembly of the European Geophysical Society, Nice, France, March 25-30, 2001, European Geophysical Society, (2001)
15. Zender, J., Bachman, N., Semenov, B., Acton, C., Roatsch, T.: Implementation Concept for Exchanging Ancillary Spacecraft Data for ESA's Mission to Mars, Mars Express. XXVI. General Assembly of the European Geophysical Society, Nice, France, March 25-30, 2001, European Geophysical Society, (2001)
16. Planetary Science Archive PVV User Manual, J. Zender and D. Heather, SOP-RSSD-UM-004, Issue 1, 06-November-2003
17. Mars Express Agreement on the Long-term Preservation of HRSC Camera Data, ME-EST-TN-11420, Issue 1.09 May 2003

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## 1.6 Relationships to Other Interfaces

This document is in close relationship to

- Mars Express - HRSC Data Products Naming Convention [5]
- MarsExpress - HRSC Level-2 Product Description [9]
- MarsExpress – HRSC VICAR Label Description Document [10]

The contents of these documents is summarized in this document for easier use.

## 1.7 Acronyms and Abbreviations

CoI	Co-Investigator
DLR	German Aerospace Center
ESOC	European Space Operation Center
ESTEC	European Space Research and Technology Center
FUB	Free University of Berlin
HRSC	High Resolution Stereo Scanner
JPL	NASA Jet Propulsion Laboratory
PI	Principal Investigator
PSA	Planetary Science Archive
PVV	Planetary Data Science Volume Verifier
SRC	Super Resolution Camera
VICAR	Video Image Communication and Retrieval

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## 1.8 Contact Names and Addresses

Archive generation software	Klaus-Dieter Matz, DLR
Archive distribution	Thomas Roatsch, DLR
Levels 1 generation software cognizant engineer	Thomas Roatsch, DLR
Map projection and generation	Frank Scholten, DLR
Calibration software and procedures	Klaus-Deter Matz, DLR
Operations	Reñe Pischel, DLR
VICAR software cognizant Engineer	Robert G. Deen, JPL
PDS Central Node MEX Data Engineer	Betty Sword, JPL
PDS Imaging Node contact	Rafael Alanis, JPL

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## 2 OVERVIEW OF INSTRUMENT DESIGN, DATA HANDLING PROCESS AND PRODUCT GENERATION

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### 2.1 Instrument Design Overview

The High Resolution Stereo Camera (HRSC), originally developed for the Russian-led Mars-96 mission, was selected as part of the Orbiter payload for ESA's Mars Express mission. The HRSC is a pushbroom scanning instrument with nine CCD line detectors mounted in parallel in the focal plane. Its unique feature is the ability to obtain near-simultaneous imaging data of a specific site at high resolution, with along-track triple stereo, four colours and five different phase angles, thus avoiding any time-dependent variations of the observational conditions. An additional Super-Resolution Channel (SRC) a framing device will yield nested images in the metre-resolution range for detailed photogeologic studies. The spatial resolution from the nominal periapsis altitude of 250 km will be 10 m px<sup>-1</sup>, with an image swath of 53 km, for the HRSC and 2.3 m px<sup>-1</sup> for the SRC. During the mission's nominal operational lifetime of 1 martian year (2 Earth years) and assuming an average HRSC data transfer share of 40%, it will be possible to cover at least 50% of the martian surface at a spatial resolution of  $\leq 15$  m px<sup>-1</sup>. More than 70% of the surface can be observed at a spatial resolution of  $\leq 30$  m px<sup>-1</sup>, while more than 1% will be imaged at better than 2.5 m px<sup>-1</sup>. The HRSC will thus close the gap between the medium- to low-resolution coverage and the very high-resolution images of the Mars Observer Camera on the Mars Global Surveyor mission and the in situ observations and measurements by landers. The HRSC will make a major contribution to the study of martian geosciences, with special emphasis on the evolution of the surface in general, the evolution of volcanism, and the role of water throughout martian history. The instrument will obtain images containing morphologic and topographic information at high spatial and vertical resolution, allowing the improvement of the cartographic base down to scales of 1:50 000. The experiment will also address atmospheric phenomena and atmosphere-surface interactions, and will provide urgently needed support for current and future lander missions as well as for exobiological studies. The goals of HRSC on Mars Express will not be met by any other planned mission or instrument.

Further information about the instrument and its operation can be found in [4] and [11].

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### 2.2 Data Handling Process

All HRSC data are processed at DLR in Berlin, Germany.

The data processing consists of the following steps:

- transfer of data from ESOC to DLR
- remove all transmission headers to get the original camera data
- sort camera data by sensor and combine them with the housekeeping data
- decompress the data to get Level 1
- radiometric calibration of the data
- calculate footprints of every image file and get Level 2

The formats of the data levels 1 and 2 are described in [8] and [9], respectively.

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All data processing steps are performed in the VICAR environment [12], a software package developed and maintained by JPL and used for the data processing of many planetary missions.

DLR developed specific VICAR modules for every processing step.

The cognizant persons for the specific task are listed in chapter 1.8. Please, adress all questions and comments through the Data Processing Manager ([thomas.roatsch@dlr.de](mailto:thomas.roatsch@dlr.de)).

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## 2.3 Product Generation

The starting point for the product generation are the bested data which are delivered by ESOC to the PI teams on CD-ROMs. Therefore, it is very important to get this data set right in time for the product generation. The following steps are the same as for the usual data product generation. The only addition is the final step of conversion from the VICAR format to PDS format and the generation of the complete data sets. This step is also performed at DLR in Berlin.

The final products will be sent to the PI and the CoIs who are in charge for the data validation. The data will be send from DLR to PSA after succesful validation and PI approval.

## 3 ARCHIVE FORMAT AND CONTENT

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### 3.1 Format and Conventions

#### 3.1.1 Deliveries and Archive Volume Format

The HRSC data will be delivered to PSA every 6 months. Every delivery contains the data taken during a time period of 6 months. The delivery will be performed only via file transfer, no storage media like CD or DVD will be used. The data will be split in a couple of sub-datasets to avoid file transfer problems with very huge files. All sub-datasets will have a size < 10 GByte. Thesub- datasets will also be compressed (using bzip2) to minimize the file transfer time.

The following file naming scheme combining country and instrument name (as suggested by PSA) will be used for the sub-dataset names:

DE\_DLR\_PF\_MEXHRSC\_0000

where 0000 will be the four digit number of the first orbit in a sub-dataset,

All sub-datasets will be packed to one single file (only for the file transfer from DLR to PSA) using the UNIX tar command with the following options

tar cfvj and the extension bz2 will be added to the filename.



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### 3.1.2 Data Set ID Formation

All HRSC data belong to one data set. Therefore, the DATA\_SET\_ID for all data is constant and has the following value:

MEX-M-HRSC/SRC-3-RDR-0010-V3.0

This name follows the standard PDS rules and contains the mission name, the instrument name, describes the level of processing (RDR), the start orbit number (e.g. 0010), and the version number. Version 3 was delivered based on an improved radiometric calibration. The DATA\_SET\_ID must be changed whenever it will become necessary to deliver different versions.

### 3.1.3 Data Directory Naming Convention

The HRSC data are sorted by orbit in the DATA directory, each sub-directory will have the name oooo where oooo is the number of the orbit in which the data were taken.

### 3.1.4 Filenaming Convention

The file naming convention is described in detail in [5]. The image files in the DATA directories follow this convention:

HOOOO\_MMM\_DD2.IMG

where

O O O O        4 digit orbit number

M M M        number of the image in this orbit

D D         sensor name (can be ND,S1,S2,P1,P2,BL,GR,IR,RE,SR)

The '2' indicates the level of processing which is archived in PSA/PDS.

Please, note that all line sensor data which were taken together will get the same number.

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## 3.2 Standards Used in Data Product Generation



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### 3.2.1 PDS Standards

All data apply to version 3.6 of the PDS Standards Reference, please see [1], [6] for details.

### 3.2.2 Time Standards

All time information in the data follows the SPICE time standards. Please, see [13], [14], and [15] for details. Within the data products themselves, the time standard used is ET (Ephemeris Time), which is a double precision number of seconds. The starting point for this time is the J2000 epoch. This epoch is Greenwich noon on January 1, 2000 Barycentric Dynamical Time. This ephemeris time is calculated from the Spacecraft Onboard Time using the appropriate SPICE routines and the time correlation packages which are provided by ESTEC as a SPICE Clock Kernel. Outside of the products themselves, there are a few instances in the HRSC data sets where time flags are provided. The main time values are provided in the data product labels, which provide a start and stop time for the measurement, and a corresponding clock count from the spacecraft. Below, the standards used to define these values are described.

#### 3.2.2.1 START\_TIME and STOP\_TIME Formation

The PDS formation rule for dates and time in UTC is:

YYYY-MM-DDThh:mm:ss.fff

YYYY	year (0000-9999)
MM	month (01-12)
DD	day of month (01-31)
T	date/time separator
hh	hour (00-23)
mm	minute (00-59)
ss	second (00-59)
fff	fractions of second (000-999) (restricted to 3 digits)

This standard is followed for all START\_TIME and STOP\_TIME values in the products included in the HRSC data sets.

#### 3.2.2.2 SPACECRAFT\_CLOCK\_START\_COUNT and SPACECRAFT\_CLOCK\_STOP\_COUNT

The SPACECRAFT\_CLOCK\_START\_COUNT and SPACECRAFT\_CLOCK\_STOP\_COUNT values represent the on-board time counters (OBT) of the spacecraft and instrument computers. This OBT counter is given in the headers of the experiment telemetry source packets. It contains the data acquisition start time as 32-bit of unit seconds followed by 16-bit of fractional seconds. The time resolution of the fractional part is  $2^{-16} = 1.52 \times 10^{-5}$  seconds. Thus, the OBT is represented as a decimal real number in floating-point notation with 5 digits after the decimal point.

A reset of the spacecraft clock is represented by an integer number followed by a slash, e.g. "1/" or "2/".



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Example:

```
SPACECRAFT_CLOCK_START_COUNT = "1/21983325.39258"
```

### 3.2.3 Reference Systems

The reference systems used for orbit, attitude, and target body follow the SPICE standards and are defined in the different SPICE kernels. Please, see [3], [13], [14], and [15] for details.

All latitudes and longitudes are given in degrees, latitudes are planetocentric.

### 3.2.4 Other Applicable Standards

No other standards are used.

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## 3.3 Data Validation

The validation of these volumes is divided into two processes:

The first process is to check that the volumes are technically correct:

- Insure that the volume is complete, and has correct structure as defined in this document.
- Insure that dynamically generated file, such as index and catalog files are correct and complete.
- Insure that structure of each generated volume is PDS compliant

These steps will be performed using PVV, the PSA Validation and Verification Tool developed by ESTEC [16].

The second process is to check that the image data contained in the data volumes are correct. This will be done by visual inspection by the PI and three CoIs. Specific tools for automated checks may be developed by the teams in charge for this step.

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## 3.4 Content

### 3.4.1 Volume Set

There are no volume sets since the data will be delivered electronically. It is not planned to generate archives on any medium (like CD-ROM or DVD).

### 3.4.2 Data Set

The data set name is defined in chapter 3.1.2.

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### 3.4.3 Directories

#### 3.4.3.1 Root Directory

The Root Directory contains the following standard PDS files:

AAREADME.TXT  
VOLDESC.CAT

#### 3.4.3.2 Calibration Directory

There is no calibration directory, the data are already radiometrically calibrated.

#### 3.4.3.3 Catalog Directory

The Catalog Directory contains the following standard PDS files:

CATINFO.TXT  
DATASET.CAT  
INST.CAT  
INSTHOST.CAT  
MISSION.CAT  
PERSON.CAT  
REF.CAT  
SOFT.CAT

#### 3.4.3.4 Index Directory

##### 3.4.3.4.1 Dataset Index File, `index.lbl` and `index.tab`

The Index Directory contains the required PDS index files which are generated by PVV [16].

##### 3.4.3.4.2 Geometric Index File, `geindex.lbl` and `geindex.tab`

The Index Directory also contains the Geometric Index File as defined in [7].

These geometric index files currently do not exist for Phobos and Deimos data. The reason is that it is not possible to calculate the information based on the current SPICE kernels. Both orbit and pointing information must be improved, ESTEC is investigating this problem.

##### 3.4.3.4.3 other Index Files

The data set also contains browse index files which are generated using PVV [16].

#### 3.4.3.5 Browse Directory and Browse Files

The browse directory has sub-directories for every orbit which belongs to the data set, the directory name is the four digits orbit number. The browse images are generated from the original data using the following steps:

- reducing the size both in line and sampl direction by a factor of 8
- conversion to unsigned 8 bit data



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- conversion to raw data
- conversion to jpeg using the UNIX program cjpeg with the highest compression quality (100)

The filenames of the browse files are the same as the original image files, the extension is changed from IMG for the image files to JPG for the browse files.

The browse directory also contains a PDS labels for every browse image.

Please, note that huge images (number of lines > 240,00) are reduced by a higher factor to avoid browse images with a number of lines > 30,000 which can not be displayed by some image display programs (like Adobes Photoshop).

### 3.4.3.6 Geometry Directory

There is no Geometry Directory, all external geometry information (orbit data, attitude data, etc.) can be found in the Mars Express SPICE data set. The SPICE data are released as separate data sets and are necessary for the generation of higher data levels. All other files necessary for the generation of higher data levels are included in the software directory (cahper 3.7).

SPICE kernels will not be available as separate release at the time of the first HRSC data release, unfortunately.

The user of the HRSC data has to get the SPICE kernels from the ftp server at ESTEC

<ftp://gorilla.estec.esa.int/pub/projects/mex/data/spice/kernels/>

or from the ftp mirror at NAIF

<ftp://naif.jpl.nasa.gov/pub/naif/MEX/kernels/>

### 3.4.3.7 Software Directory

The Software Directory contains the following files:

GEOCAL.LBL

GEOCAL.ZIP

MINIVICAR.LBL

MINIVICAR.ZIP

SOFTINFO.TXT

SOFTINFO.TXT is the standard PDS info file for this directory.

GEOCAL.ZIP is a zip-file which contains the geometric calibration files (one file per HRSC line sensor) which are necessary to run the geometric correction software (contained in MINIVICAR.ZIP), GEOCAL.LBL is the standard PDS label file for this zip-file.

MINIVICAR.ZIP is a zip-file which contains the source code and necessary utilities to build the programs (both for Linux-RedHat-Enterprise and Sun-Solaris 9) listed below, MINIVICAR.LBL is the standard PDS label file for this zip-file.

List of programs contained in MINIVICAR.ZIP:

label	a program to list the label of the files contained in the data directory
xvd	a program to display these data
hrpref2asc	a program to list the binary prefix of these data
hrfill	a program to interpolate over bad pixels (e.g. blemishes, saturated pixels)
flform_lut	a program to generate a look-up table for the 16-bit to 8-bit conversion of the image data
dlr12to8	a program to perform this conversion (both without and with look-up table)



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hrortho a program for the geometric correction and map projection of HRSC line sensor data  
frameortho a program for the geometric correction and map projection of SRC images

MINIVICAR.ZIP also contains a converter (written in JAVA) to convert the images into a couple of image formats.

The installation and usage of the software is described in HRSC\_MINIVICAR.TXT which can be found in the document directory.

### 3.4.3.8 Gazetter Directory

There is no Gazetter Directory.

### 3.4.3.9 Label Directory

There is no Label directory.

### 3.4.3.10 Document Directory

This directory contains the documentation for the HRSC data sets.  
The content is:

- DOCINFO.TXT the standard PDS info file.
- HRSC\_EAICD.LBL the label for the Experimenter to Archive ICD
- HRSC\_EAICD.PDF the Adobe PDF file of the experimenter to Archive ICD
- HRSC\_EAICD.TXT the Text file of the experimenter to Archive ICD
- HRSC\_ESA\_SP.LBL the label for the HRSC Instrument Description published in the ESA SP-1240
- HRSC\_ESA\_SP.PDF the Adobe PDF file of the ESA SP-1240
- HRSC\_LABEL.LBL the label for the HRSC VICAR and PDS label table
- HRSC\_LABEL.PDF the Adobe PDF file of the HRSC VICAR and PDS label table
- HRSC\_LABEL\_HEADER.LBL the label for the HRSC VICAR and PDS label description
- HRSC\_LABEL\_HEADER.PDF the Adobe PDF file of the HRSC VICAR and PDS label description
- HRSC\_MINIVICAR.LBL the label for the MINIVICAR documentation
- HRSC\_MINIVICAR.TXT the Text file describing the how to install and run MINIVICAR
- VICAR2.LBL the Label for the description of the VICAR labels
- VICAR2.TXT Text file describing the VICAR label.

### 3.4.3.11 Extras Directory

There is no Extras Directory.

### 3.4.3.12 Data Directory

The Data Directory contains sub-directories for every orbit which is part of the data set, the directory names are the four digits orbit number. The contents of these sub-directories is described in 3.4.8.

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### 3.4.4 Other Data Products

No Pre-Flight Data Products, Sub-System test data, and instrument calibration data will be delivered to PSA/PDS.

### 3.4.5 In-Flight Data Products

The HRSC data archive contains all data which were taken in Mars orbit from Mars and its satellites Phobos and Deimos.

The data are radiometrically calibrated, the HRSC team calls them level-2 data. The HRSC team will not archive the uncalibrated data in the PSA as agreed in [17].

These data can be used both for cross-instrument calibration (e.g. with the spectrometer OMEGA) or with instruments from other missions (e.g. from NASA's MER lander mission).

### 3.4.6 Software

The HRSC processing software was developed in the VICAR environment [12]. VICAR was developed by NASA/JPL and was used for the processing of camera data from many planetary missions (e.g. Viking, Galeileo). The data processing team at DLR in Berlin developed specific modules to process the HRSC data.

These modules perform the following steps:

- remove all telemetry headers from the data
- sort the data by sensor and combine the image data with the housekeeping data
- decompression of the data
- radiometric calibration of the data
- calculation of the footprints for every image

### 3.4.7 Documentation

The contents of the documentation directory is described in 3.4.3.10.

### 3.4.8 Derived and other Data Products

There are currently no plans to deliver derived and other data products. Also, no data based on the cooperation with other Mars Express teams will be delivered.

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## 4. DETAILED INTERFACE SPECIFICATIONS

### 4.1. Data Product Structure

The data structure consists of an ASCII PDS label, followed by an embedded ASCII VICAR label, followed by a  $n \times m$  block of binary image data. Inherent to the VICAR label is the possibility of an ASCII EOL label being appended after the binary data in order to handle label modifications. This EOL label is simply a continuation field for the main VICAR label, when there is no more space for expansion before the image data.

### 4.2 Label and Header Descriptions

#### 4.2.1 PDS Label

HRSC data have an attached PDS label. A PDS label is object-oriented and describes the objects in the data file. The PDS label contains keywords for product identification. The label also contains descriptive information needed to interpret or process the data in the file.

PDS labels are written in Object Description Language (ODL [1]). PDS label statements have the form of "keyword = value". Each label statement is terminated with a carriage return character (ASCII 13) and a line feed character (ASCII 10) sequence to allow the label to be read by many operating systems. Pointer statements with the following format are used to indicate the location of data objects in the file:

^object = location

where the carat character (^, also called a pointer) is followed by the name of the specific data object. The location is the 1-based starting record number for the data object within the file.

#### 4.2.2 PDS Image Object

An IMAGE object is a two-dimensional array of values, all of the same type, each of which is referred to as a *sample*. IMAGE objects are normally processed with special display tools to produce a visual representation of the samples by assigning brightness levels or display colors to the values. An IMAGE consists of a series of lines, each containing the same number of samples.

The required IMAGE keywords define the parameters for simple IMAGE objects:

- LINES is the number of lines in the image.
- LINE\_SAMPLES is the number of samples in each line.
- SAMPLE\_BITS is the number of bits in each individual sample.
- SAMPLE\_TYPE defines the sample data type.



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The IMAGE object has a number of keywords relating to image statistics. These keywords will be present in all data, the statistics keywords are:

- MEAN
- MEDIAN
- MAXIMUM
- MINIMUM
- STANDARD\_DEVIATION

Many variations on the basic IMAGE object are possible with the addition of optional keywords and/or objects. The “^IMAGE” keyword identifies the start of the image data and will skip over the VICAR label.

#### 4.2.3 Keyword Length Limits

All PDS keywords are limited to 30 characters in length (Section 12.7.3 in PDS Standards Reference). Therefore, software that reads HRSC PDS labels must be able to ingest keywords up to 30 characters in length.

#### 4.2.4 Data Type Restrictions

In order to accommodate VICAR dual-labeled files, 16-bit data must be stored as signed data. Unsigned 16-bit data is not supported.

#### 4.2.5 Interpretation of N/A, UNK, and NULL

During the completion of data product labels or catalog files, one or more values may not be available for some set of required data elements. In this case PDS provides the symbolic literals “N/A”, “UNK”, and “NULL”, each of which is appropriate under different circumstances.

- “N/A” (“Not Applicable”) indicates that the values within the domain of this data element are not applicable in this instance.
- “UNK” (“Unknown”) indicates that the value for the data element is not known and never will be.

“NULL” is used to flag values that are *temporarily* unknown. It indicates that the data preparer recognizes that a specific value should be applied, but that the true value was not readily available. “NULL” is a placeholder

#### 4.2.6 VICAR Label

For all data products, an embedded VICAR label follows the PDS label and is pointed to by the PDS pointer “^IMAGE\_HEADER”. The VICAR label is also organized in an ASCII, “keyword = value” format, although there are only spaces between keywords (no carriage return/line feeds as in PDS). The information in the VICAR label is an exact copy of the information in the PDS label as defined in the next section.



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#### 4.2.7 VICAR Format

The reader is referred to the VICAR File Format document for details of the format, which is available at the URL "[http://www-mipl.jpl.nasa.gov/vicar/vic\\_file\\_fmt.html](http://www-mipl.jpl.nasa.gov/vicar/vic_file_fmt.html)". The following text is an excerpt which describes the basic structure:

A VICAR file consists of two major parts: the labels, which describe what the file is, and the image area, which contains the actual image. The labels are potentially split into two parts, one at the beginning of the file, and one at the end. Normally, only the labels at the front of the file will be present. However, if the EOL keyword in the system label (described below) is equal to 1, then the EOL labels (End Of file Labels) are present. This happens if the labels expand beyond the space allocated for them. The VICAR file is treated as a series of fixed-length records, of size RECSIZE (see below). The image area always starts at a record boundary, so there may be unused space at the end of the label, before the actual image data starts.

The label consists of a sequence of "keyword=value" pairs that describe the image, and is made up entirely of ASCII characters. Each keyword-value pair is separated by spaces. Keywords are strings, up to 32 characters in length, and consist of uppercase characters, underscores ("\_"), and numbers (but should start with a letter). Values may be integer, real, or strings, and may be multiple (e.g. an array of 5 integers, but types cannot be mixed in a single value). Spaces may appear on either side of the equals character (=), but are not normally present. The first keyword is always LBLSIZE, which specifies the size of the label area in bytes. LBLSIZE is always a multiple of RECSIZE, even if the labels don't fill up the record. If the labels end before LBLSIZE is reached (the normal case), then a 0 byte terminates the label string. If the labels are exactly LBLSIZE bytes long, a null terminator is not necessarily present. The size of the label string is determined by the occurrence of the first 0 byte, or LBLSIZE bytes, whichever is smaller. If the system keyword EOL has the value 1, then End-Of-file Labels exist at the end of the image area (see above). The EOL labels, if present, start with another LBLSIZE keyword, which is treated exactly the same as the main LBLSIZE keyword. The length of the EOL labels is the smaller of the length to the first 0 byte or the EOL's LBLSIZE. Note that the main LBLSIZE does not include the size of the EOL labels. In order to read in the full label string, simply read in the EOL labels, strip off the LBLSIZE keyword, and append the rest to the end of the main label string.

### 4.3 Binary Data Storage Conventions

HRSC data are stored as binary data. The data are stored in signed 16-bit integers. The PDS and VICAR labels are stored as ASCII text.

The ordering of bits and bytes is only significant for pixel data; all other labeling information is in ASCII. All data are stored as Most Significant Byte first ("big-endian", as used by e.g. Sun computers and Java)

### 4.4 PDS keyword table

The same keywords are used for all data. These keywords are described in the following table:

FILE_NAME	usual default name of the output file; this entry allows the user to check for accidental renaming of files, filename without path
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DATA_SET_ID	The data_set_id element is a unique alphanumeric identifier for a data set or a data product.
DETECTOR_ID	identifies which of the ten CCD detectors was used for this particular image.
EVENT_TYPE	identifies the classification of an event, HRSC specific, to be defined by HRSC planning group
INSTRUMENT_HOST_ID	The instrument_host_id element provides a unique identifier for the host where an instrument is located.
INSTRUMENT_HOST_NAME	full name of the spacecraft
INSTRUMENT_ID	The instrument_id element provides an abbreviated name or acronym which identifies an instrument.
INSTRUMENT_NAME	full name of an instrument
MISSION_NAME	full name of mission
MISSION_PHASE_NAME	The mission_phase_name element provides the commonly-used identifier of a mission phase.
PROCESSING_LEVEL_ID	identifies the processing level of a data set; parameter must be updated after each processing step according to the program specification , DLR-Levels
PRODUCT_CREATION_TIME	The product_creation_time element defines the UTC system format time when a product was created.
PRODUCT_ID	The product_id data element represents a permanent, unique identifier assigned to a data product by its producer.
SPACECRAFT_CLOCK_START_COUNT	provides the value of the spacecraft clock at the beginning of a time period of interest. This is the same for all line sensors and SRC images during one imaging sequence.
SPACECRAFT_CLOCK_STOP_COUNT	provides the value of the spacecraft clock at the end of a time period of interest. This is the same for all line sensors and SRC images during one imaging sequence.
START_TIME	date and time of recording of the first image line in UTC format "YYYY-MM-DDTHH:MM:SS.MMMZ" (corresponds to the ephemeris time prefix entry of that line )
STOP_TIME	date and time of recording of the last image line in UTC format "YYYY-MM-DDTHH:MM:SS.MMMZ" (corresponds to the ephemeris time prefix entry of that line )



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ASCENDING_NODE_LONGITUDE	value of the angle of the xy-plane of the J2000 coordinate system to the ascending node computed from the spacecraft's position- and velocity vector at periapsis (not to be used during test and cruise)	deg
MAXIMUM_RESOLUTION	highest resolution in an image	m/
FOOTPRINT_POINT_LATITUDE	The footprint_point_latitude element provides the latitude of a point within an array of points along the border of a footprint, described as a polygon, outlining an imaged area on the planet's surface. Latitude values are planetocentric.	deg
FOOTPRINT_POINT_LONGITUDE	The footprint_point_longitude element provides the longitude of a point within an array of points along the border of a footprint, described as a polygon, outlining an imaged area on the planet's surface. Longitude values are planetocentric.	deg
ORBIT_NUMBER	number of the orbital revolution of the s/c around the target body (not to be used during test and cruise)	
ORBITAL_ECCENTRICITY	value of orbit eccentricity computed from the spacecraft's position- and velocity vector at periapsis (not to be used during test and cruise)	
ORBITAL_INCLINATION	value of the angle of inclination with respect to the xy-plane computed from the spacecraft's position- and velocity vector at periapsis	
ORBITAL_SEMIMAJOR_AXIS	value of orbit semi-major axis computed from spacecraft 's position - and velocity vector at periapsis (not to be used during test and cruise)	km
PERIAPSIS_ALTITUDE	The PERIAPSIS_ALTITUDE element provides the distance between the spacecraft and the target body at periapsis. Periapsis is the closest approach point of the spacecraft to the target body in its orbit around the target body. Note: For MARS EXPRESS, the al	km
PERIAPSIS_ARGUMENT_ANGLE	angle in the xy-plane of the J2000 coordinate system from the ascending node to periapsis (not to be used during test and cruise)	deg
PERIAPSIS_TIME	The PERIAPSIS_TIME element is the time, in UTC format "YYYY-MM-DDThh:mm:ss[.fff]Z", when the spacecraft passes through periapsis. Periapsis is the closest approach point of the spacecraft to the target body in its orbit around the target body. (not to be	time
SPACECRAFT_ORIENTATION	The spacecraft orientation element provides the orientation of a spacecraft in orbit or cruise in respect to a given frame. E.g. a non-spinning spacecraft might be flown in +Y or -Y direction in respect to the spacecraft mechanical build frame. This element shall be used in combination with the keyword spacecraft_orientation_desc that describes the convention used to describe the spacecraft orientation. The spacecraft orientation shall be given as a 3-tuple, one value for the x,y and z axes	
SPACECRAFT_POINTING_MODE	The spacecraft pointing element provides information on the pointing mode of the spacecraft. The definition of the modes and the standard values are given in the s/c pointing mode description element, that shall always accompany the keyword	
RIGHT_ASCENSION	The right_ascension element provides the right ascension value. Right_ascension is defined as the arc of the celestial equator between the vernal equinox and the point where the hour circle through the given body intersects the Earth's mean equator (reckoned eastward).	deg



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DECLINATION	The declination element provides the value of an angle, corresponding to latitude, used to fix position on the celestial sphere. Declination is measured positive north and negative south of the celestial equator, and is defined relative to a specified reference period or epoch.	deg
OFFSET_ANGLE	offset from nadir looking during ACROSS_TRACKING or ALONG_TRACKING	deg
SPACECRAFT_SOLAR_DISTANCE	the spacecraft's distance to the Sun measured from its position vector at periapsis (not to be used during test and cruise)	km
TARGET_NAME	name of the target body	
DETECTOR_TEMPERATURE	TEMPERATURE SPL_F (Dornier HKD doc.) for sensor P2,RE,S2, TEMPERATURE SPL_N (Dornier HKD doc.) for sensor BL,ND,GR, TEMPERATURE SPL_A (Dornier HKD doc.) for sensor P1,IR,S1; , temp_fpm in hrhk23, SRC first level is 2	Cels
FOCAL_PLANE_TEMPERATURE	TEMPERATURE OPTICS in Dornier HKD document, HRSC only, temp_co in hrhk23	Cels
INST_CMPRS_NAME	flag indicating whether spacecraft on-board compression has been bypassed, in which case, the received data were uncompressed; HRSC: config. byte 1/2, bit 2 = 1 ==> BYPASS_FLAG = "YES"	
INST_CMPRS_QUALITY	The compression index parameter in the table of scale factors (TABF). It is in the range from 0 to 15. A higher value means more compression	
INST_CMPRS_QUANTZ_TBL_ID	Number of the quantization matrix in the PMEM file, TB*2 + Malgo	
INST_CMPRS_RATIO	mean compression rate for the entire image data represented in the file, this number is =1 for data collected in the bypass mode.	
INSTRUMENT_TEMPERATURE	TEMPERATURE FEE in Dornier HKD document, temp_fee in hrhk23, SRC first level is 2	Cels
LENS_TEMPERATURE	TEMPERATURE OPTICAL BENCH in Dornier HKD document, HRSC only, temp_ob in hrhk23	Cels
MACROPIXEL_SIZE	macropixel format	
MISSING_FRAMES	The MISSING_FRAMES element is the total number of frames that are missing from a file. (Cf. ERROR_FRAMES and OVERFLOW_FRAMES). Note: For MARS EXPRESS, a frame, which is also called a "row", is eight lines of data. Each line, in turn, is composed of a s	
PIXEL_SUBSAMPLING_FLAG	The PIXEL_SUBSAMPLING_FLAG element indicates whether this product is the result of subsampling of the data, HRSC only.	
SAMPLE_FIRST_PIXEL	position of the first pixel on the CCD line that contributes to the first VICAR macropixel (may still be dark or dummy pixels for level 1 images); config bytes 3,4; bits 0-12; note: FIRST_PIXEL = start_pixel_number + 1	
SIGNAL_CHAIN_ID	The SIGNAL_CHAIN_ID element identifies the signal chain (electronic signal path) number selected for charge-coupled device (CCD) output. Note: For MARS EXPRESS the High-Resolution Stereo Colour Imager (HRSC) is composed of 10 channels, each consisting of	
BANDWIDTH	The bandwidth element provides a measure of the spectral width of a filter or channel. For a root-mean-square detector this is the effective bandwidth of the filter i.e., the full width having a flat response over the bandwidth and zero response elsewhere. For HRSC this value is for the whole sensor (CCD+Optics).	nm

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CENTER_FILTER_WAVELENGTH	The center_filter_wavelength element provides the mid_point wavelength value between the minimum and maximum instrument filter wavelength values. For HRSC this value is for the whole sensor (CCD+Optics).	nm
RADIANCE_OFFSET	The radiance_offset element provides the constant value by which a stored radiance is added. Note: Expressed as an equation: $true\_radiance\_value = radiance\_offset + radiance\_scaling\_factor * stored\_radiance\_value$ .	W/n n
RADIANCE_SCALING_FACTOR	The radiance_scaling_factor element provides the constant value by which a stored radiance is multiplied. Note: Expressed as an equation: $true\_radiance\_value = radiance\_offset + radiance\_scaling\_factor * st$	W/n n
REFLECTANCE_SCALING_FACTOR	The reflectance_scaling_factor element identifies the conversion factor from DN to reflectance.	

## 4.5. Example PDS Product Label

```

PDS_VERSION_ID                = PDS3

/* FILE DATA ELEMENTS */

RECORD_TYPE                    = FIXED_LENGTH

RECORD_BYTES                   = 10420

FILE_RECORDS                   = 251387

LABEL_RECORDS                  = 2

/* POINTERS TO DATA OBJECTS */

^IMAGE_HEADER                  = 3

^IMAGE                          = 4

/* IDENTIFICATION DATA ELEMENTS */

FILE_NAME                      = "H0024_0000_ND2.IMG"

DATA_SET_ID                    = "MEX-M-HRSC/SRC-3-RDR-
0010-V1.0"

DETECTOR_ID                    = MEX_HRSC_NADIR

EVENT_TYPE                     = "MARS-LOCAL-CARTOGRAPHY-

```



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Im-Lc-Tc"

INSTRUMENT\_HOST\_ID = MEX

INSTRUMENT\_HOST\_NAME = "MARS EXPRESS"

INSTRUMENT\_ID = HRSC

INSTRUMENT\_NAME = "HIGH RESOLUTION STEREO CAMERA"

MISSION\_NAME = "MARS EXPRESS"

MISSION\_PHASE\_NAME = MC\_Phase\_1

PROCESSING\_LEVEL\_ID = 2

PRODUCT\_CREATION\_TIME = 2004-11-24T19:53:14.000Z

PRODUCT\_ID = "H0024\_0000\_ND2.IMG"

/\* TIME DATA ELEMENTS \*/

SPACECRAFT\_CLOCK\_START\_COUNT = "1/0022332827.31245"

SPACECRAFT\_CLOCK\_STOP\_COUNT = "1/0022333524.07540"

START\_TIME = 2004-01-16T11:35:55.639Z

STOP\_TIME = 2004-01-16T11:52:36.574Z

/\* ORBITAL DATA ELEMENTS \*/

ASCENDING\_NODE\_LONGITUDE = 228.47

MAXIMUM\_RESOLUTION = 11.7 <m/pixel>

FOOTPRINT\_POINT\_LATITUDE = (-51.592,-51.3204,-51.3182,-  
50.2127,  
  
-49.1343,-  
47.0532,-45.0716,-44.1141,  
  
-43.1777,-42.2627,-  
40.4919,-38.7932,



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34.0794, -31.9082, -30.5201, -37.1618, -

29.1757, -27.2292, -26.5962, -29.8429, -

25.3606, -24.7522, -22.3965, -25.9749, -

21.2573, -20.691, -20.1401, -21.8232, -

17.9641, -17.4256, -16.345, -18.502, -

15.2604, -13.6279, -13.0823, -15.8032, -

11.9887, -10.3444, -9.79539, -12.5358, -

8.69579, -7.59391, -7.04251, -9.24578, -

5.38784, -4.28411, -3.73212, -6.4911, -

2.07651, -1.52482, -0.973341, -3.18019, -

0.679678, 3.42738, 4.52295, 5.61604,

7.79353, 9.41786, 12.1229, 12.1305, 12.1303

, 11.5946, 11.0577, 10.5197, 9.98082,

9.44085, 8.89994, 6.18221, -3.71932,

6.48098, -7.03285, -7.58452, -5.92888, -

8.68704, -9.78828, -10.3382, -8.13596, -



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11.9848,-12.5326,-13.0799, -11.4364,-  
15.2612,-15.8047,-16.3473, -13.6262,-  
17.9697,-18.5084,-19.0461, -16.8891,-  
22.9913,-23.5757,-26.6196, -21.8354,-  
29.2072,-34.1319,-34.8845, -28.5455,-  
48.2611,-51.592) -45.2103,-  
FOOTPRINT\_POINT\_LONGITUDE = (175.791,171.024,171.001,171.293,  
171.557,172.005,172.39,172.557,172.701,  
172.843,173.092,173.309,173.491,173.8,  
173.991,174.097,174.144,174.195,174.321  
,174.357,174.389,174.422,174.449,  
174.566,174.588,174.613,174.635,174.656  
,174.702,174.713,174.736,174.764,  
174.772,174.775,174.811,174.827,174.834  
,174.839,174.883,174.904,174.916,  
174.931,174.913,174.901,174.901,174.909



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<b>MEX-HRSC to Planetary Science Archive Interface Control Document (EAICD)</b>	<b>Doc.-No.</b> : <b>HRSC-DLR-TN-4200-008</b>
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,174.914,174.912,174.914,174.915,  
  
174.913,174.913,174.907,174.888,174.878  
  
,174.866,174.836,174.809,174.756,  
  
176.143,176.149,176.143,176.139,176.138  
  
,176.133,176.13,176.127,176.115,176.094  
  
,176.094,176.093,176.09,176.089,176.089  
  
,176.088,176.092,176.091,176.084,  
  
176.083,176.082,176.087,176.081,176.077  
  
,176.08,176.081,176.084,176.085,176.082  
  
,176.081,176.072,176.067,176.063,  
  
176.051,176.048,176.042,176.014,176.006  
  
,175.906,175.858,175.791)

ORBIT\_NUMBER = 24  
  
ORBITAL\_ECCENTRICITY = 0.667  
  
ORBITAL\_INCLINATION = 86.66  
  
ORBITAL\_SEMIMAJOR\_AXIS = 11020.2  
  
PERIAPSIS\_ALTITUDE = 274.22  
  
PERIAPSIS\_ARGUMENT\_ANGLE = 355.69



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<b>MEX-HRSC to Planetary Science Archive Interface Control Document (EAICD)</b>	<b>Doc.-No.</b> : <b>HRSC-DLR-TN-4200-008</b>
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PERIAPSIS\_TIME = 2004-01-16T11:48:35.000Z

SPACECRAFT\_ORIENTATION = (0.0,1.0,0.0)

^MEX\_ORIENTATION\_DESC = "MEX\_ORIENTATION\_DESC.TXT"

SPACECRAFT\_POINTING\_MODE = ACROSSTRACK

^MEX\_POINTING\_DESC = "MEX\_POINTING\_DESC.TXT"

RIGHT\_ASCENSION = -1e+32

DECLINATION = -1e+32

OFFSET\_ANGLE = 2.04

SPACECRAFT\_SOLAR\_DISTANCE = 2.23543e+08

TARGET\_NAME = MARS

/\* CAMERA DATA ELEMENTS \*/

DETECTOR\_TEMPERATURE = 18.3705 <degC>

FOCAL\_PLANE\_TEMPERATURE = 9.8067 <degC>

INST\_CMPRS\_NAME = "DISCRETE COSINE TRANSFORMATION  
(DCT) "

INST\_CMPRS\_RATIO = 7.43431

INST\_CMPRS\_QUALITY = 0

INST\_CMPRS\_QUANTZ\_TBL\_ID = 0

INSTRUMENT\_TEMPERATURE = 11.4379 <degC>

LENS\_TEMPERATURE = 10.0106 <degC>

MACROPIXEL\_SIZE = 1

MISSING\_FRAMES = 0

PIXEL\_SUBSAMPLING\_FLAG = N



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<b>MEX-HRSC to Planetary Science Archive Interface Control Document (EAICD)</b>	<b>Doc.-No.</b> : <b>HRSC-DLR-TN-4200-008</b>
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```
SAMPLE_FIRST_PIXEL          = 80

SIGNAL_CHAIN_ID              = 0

/* RADIOMETRIC DATA ELEMENTS */

BANDWIDTH                    = 177.0 <nm>

CENTER_FILTER_WAVELENGTH    = 677.5 <nm>

RADIANCE_OFFSET              = 0.0 <W*m**-2*sr**-1>

RADIANCE_SCALING_FACTOR     = 0.0695439 <W*m**-2*sr**-1>

REFLECTANCE_SCALING_FACTOR  = 0.00184611

/* DATA OBJECT DEFINITIONS */

OBJECT                       = IMAGE

  INTERCHANGE_FORMAT         = BINARY

  LINES                      = 251384

  LINE_PREFIX_BYTES         = 68

  LINE_SAMPLES               = 5176

  SAMPLE_TYPE                = MSB_INTEGER

  SAMPLE_BITS                = 16

  BANDS                      = 1

  BAND_STORAGE_TYPE         = BAND_SEQUENTIAL

  MAXIMUM                    = 209

  MEAN                       = 127.59

  MINIMUM                    = 62

  STANDARD_DEVIATION        = 23.3313
```



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<b>MEX-HRSC to Planetary Science Archive Interface Control Document (EAICD)</b>	<b>Doc.-No.</b> : <b>HRSC-DLR-TN-4200-008</b>
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```
END_OBJECT = IMAGE

/* IMAGE HEADER DATA ELEMENTS */

OBJECT = IMAGE_HEADER

HEADER_TYPE = VICAR2

INTERCHANGE_FORMAT = ASCII

BYTES = 10420

^DESCRIPTION = "VICAR2.TXT"

END_OBJECT = IMAGE_HEADER

END
```

## Appendix: Available Software to read PDS files

The software delivery which is part of the data sets contains the display program xvd which can display images both in PDS and VICAR format. This programs uses the standard PDS library to open and read the data in PDS format. This library can not handle images which are larger than 2 GBytes, unfortunately. The same problem occurs also for other PDS programs (e.g. nasaview) which can not read large files. The only solution currently is to display the large files as VICAR files using xvd.