

MARS EXPRESS

PLANETARY FOURIER SPECTROMETER

INSTRUMENT CALIBRATION REPORT

VOLUME I

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TABLE OF CONTENT:

 The Series of Measurements Performed 1.1 -27 Dec 2001 1.2 -28 Dec 2001 1.3 -19 Jan 2002 1.4 -29 Dec 2001 1.5 -02 Jan 2002 1.6 -03 Jan 2002 1.7 -04 Jan 2002 1.8 -05 Jan 2002 1.9 -Fov Measurements 1.10 -Scanner Transfer Function 1.11 -Out of the Window Measurements 	page 2 page 3 page 8 page 11 page 14 page 18 page 19 page 22 page 23 page 23 page 25 page 30 page 34
 1.12 -Thermovacuum. 08 Jan 2002 1.13 -10 Jan 2002 1.14 -21 Jan 2002 1.15 -Oriel Lamp 1.16 -Minerals 1.17 -Data Taken in Alenia in Turin 	page 35 page 41 page 42 page 48 page 48 page 51
 2Typical Raw Spectra Obtained 2.1 - Thermal Lw Channel 2.2 - Thermal Sw Channel 2.3 - Russian Black Body 2.4 - Gosts Features 2.5 - Oriel and Ceramic Element Lamp 2.6 - Forward and Reverse Motion 2.7 - Integrating Sphere 2.8 - Internal Lamp 2.9 - Internal Black Body 2.10 - Mercury Lamp 2.11 - Autotest Data 2.12 - Minerals 2.13 - Earth's Atmosphere 2.14 - Interferograms Statistics 	page 53 page 53 page 53 page 75 page 75 page 91 page 104 page 104 page 132 page 132 page 132 page 132 page 136 page 139 page 147
2.15 - Stability of the Experiment3 - Summary and Conclusions on the Instrumental Effects	page 148 page 157



1. THE SERIES OF MEASUREMENTS PERFORMED

Data collected have been ordered by time : we started measurements with the lw channel and the black bodies.

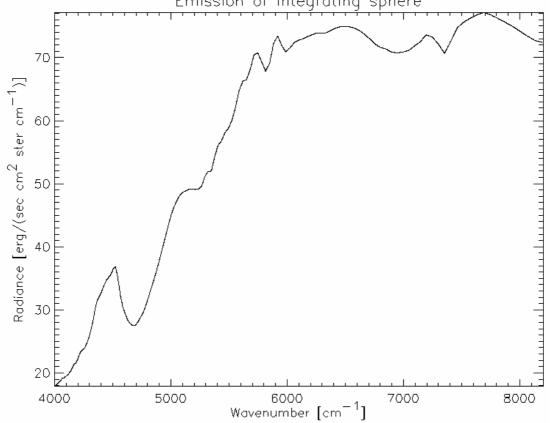
Reference sources used are:

- 1. "IFSI" black body (MIKRON M345X4 UDC): it has a certified emission of 0.97±0.005 in the range [8; 15] μm and [3; 5] μm
- 2. Integrating sphere (Optronic Lab., inc., OL 455 12 2): it provides the radiance spectrum of fig. 1 when the current is set to 5440 mA, to which 5403 fL correspond. The spectrum can be scaled proportionally to luminance when acting on internal slit.

Other non-certified sources are:

- 1. IKI black bodies: a value of emissivity of $1. \pm 0.1$ has been estimated in Lecce laboratories
- 2. Ceramic element: stable source, able to provide intense signal in the whole PFS spectral range, but decreasing very much in the range 5000 8000 cm-1.
- 3. Mercury lamp: (Oriel 6035) very stable against temperature variations, presents lines at 9862.13 8859.44 7369.09 7313.42 6537.74 5923.64 5910.10 5902.49 5857.27 5844.57 5770.54 5515.60 and 5076.09 cm⁻¹. The line at 6537.74 is the strongest and the only actually reliable for our purposes. The width of these lines are declared as "negligible" by the manufacturer but no precise figure in this sense is provided.

Moreover, the PFS scanner hosts a calibration lamp and internal black bodies. The characterization of these sources is one of the most important tasks for the whole calibration program. A first order evaluation of black body paint emissivity and bare lamp intensity are provided by the manufacturer. Emission of integrating sphere



Spectral intensity of the integrating sphere when the current is 5540 mA.



1.1 27 DEC 2001

All the acquisitions have been done into the laboratory at room temperature, using the Black Body calibration source (M345).

See pictures:

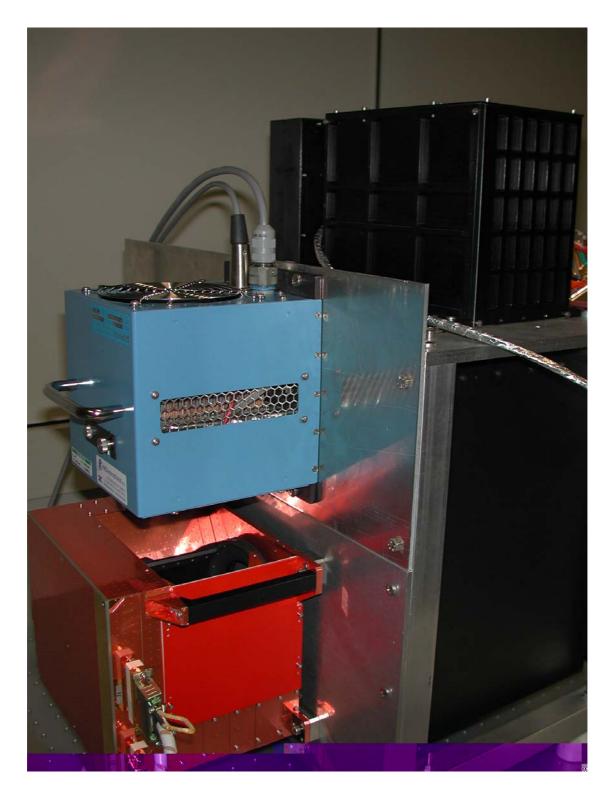


PICTURE 1 PFS and IFSI BB assembly: frontal view.









PICTURE 2 PFS and IFSI BB : lateral view.





PICTURE 3 IFSI BB frontal panel.

File Description:

BB30 011227 100440.ifg

Measurements: 200. Scanner Position: NADIR. Black Body Temperature: 30 °C. Motor: Secondary

BB40_011227_105014.ifg

Measurements: 200.Scanner Position: NADIR. Black Body Temperature: 40 °C. Motor: Secondary

BB50_011227_113148.ifg

Measurements: 200.Scanner Position: NADIR. Black Body Temperature: 50 °C. Motor: Secondary

BBINT_011227_120917.ifg

Measurements: 50.Scanner Position: INTERNAL BLACK BODY. Motor: Secondary

BB50A_011227_124602.ifg

Measurements: 50.Scanner Position: NADIR. Black Body Temperature: 50.5 °C. Motor: Secondary

BB50B 011227 125653.ifg

Measurements: 50.Scanner Position: NADIR. Black Body Temperature: 50.8 °C. Motor: Secondary

BB51_011227_131522.ifg

Measurements: 50.Scanner Position: NADIR. Black Body Temperature: 51 °C. Motor: Secondary

BB51B_011227_133557.ifg

Measurements: 10.Scanner Position: NADIR. Black Body Temperature: 51 °C. Motor: Primary

BB50P 011227 134426.ifg

Measurements: 200.Scanner Position: NADIR. Black Body Temperature: 50 °C. Motor: Primary

BB100_011227_142714.ifg

DD100_011227_142714.llg

Measurements: 200.Scanner Position: NADIR. Black Body Temperature: 100 °C. Motor: Primary

BB149_011227_153737.ifg

Measurements: 200.Scanner Position: NADIR. Black Body Temperature: 148.9 °C. Motor: Primary. SW & LW Gain sensors: 0

BB149B 011227 161518.ifg



Measurements: 50.Scanner Position: NADIR. Black Body Temperature: 148.9 °C. Motor: Primary. SW & LW Gain sensors: 1

BBINTB_011227_162926.ifg

Measurements: 50.Scanner Position: INTERNAL BLACK BODY. Motor: primary. SW & LW Gain sensors: 1

BB30B_011227_164222.ifg

Measurements: 50.Scanner Position: NADIR. Black Body Temperature: 30 °C. Motor: Primary. SW & LW Gain sensors: 1



1.2 28 DEC 2001

All the acquisitions have been done into the laboratory, using the Black Body calibration source (M345) and the russian black body. PES is in the russian green house

PFS is in the russian green house.

See picture:



PICTURE 4 PFS and two BB: the IFSI and the Russian BB assembled inside the "greenhouse" in the laboratory.



File Description:

BBITA150_011228_112317.ifg

Measurements: 200.Scanner Position: NADIR. Black Body Temperature: 150.6 °C. Motor: Primary. SW & LW Gain sensors: 0

BBRUSS-30_011228_120021.ifg

Measurements: 200.Scanner Position: DEEP SPACE (RUSSIAN BB). Black Body Temperature: -30 °C. Motor: Primary. SW & LW Gain sensors: 0

BBINT_011228_123630.ifg

Measurements: 100.Scanner Position: INTERNAL BLACK BODY Motor: Primary. SW & LW Gain sensors: 0

BBRUSS-60_011228_125342.ifg

Measurements: 200.Scanner Position: DEEP SPACE (RUSSIAN BB). Black Body Temperature: -60 °C. Motor: Primary. SW & LW Gain sensors: 0

BBITA150_011228_133306.ifg

Measurements: 109.Scanner Position: NADIR .Black Body Temperature: 150.6 °C. Motor: Primary. SW & LW Gain sensors: 0

BBRUSSRAMP-A_011228_135457.ifg

Measurements: 200.Scanner Position: DEEP SPACE (RUSSIAN BB). Black Body Temperature: from -53 °C to -18 °C. Motor: Primary. SW & LW Gain sensors: 0

BBRUSSRAMP-B_011228_142906.ifg

Measurements: 200.Scanner Position: DEEP SPACE (RUSSIAN BB). Black Body Temperature: from -18 °C to 0 °C. Motor: Primary. SW & LW Gain sensors: 0

BBRUSSRAMP-C_011228_151035.ifg

Measurements: 11.Scanner Position: DEEP SPACE (RUSSIAN BB). Black Body Temperature: from 0 °C to -15 °C. Motor: Primary. SW & LW Gain sensors: 0

BBRUSSRAMP-D_011228_151926.ifg

Measurements: 199.Scanner Position: DEEP SPACE (RUSSIAN BB). Black Body Temperature: from -18 °C to -100 °C. Motor: Primary. SW & LW Gain sensors: 0



BBRUSS-100_011228_155257.ifg

Measurements: 201.Scanner Position: DEEP SPACE (RUSSIAN BB). Black Body Temperature: -100 °C. Motor: Primary. SW & LW Gain sensors: 0

BBRUSSRAMP-E_011228_162757.ifg

Measurements: 32.Scanner Position: DEEP SPACE (RUSSIAN BB). Black Body Temperature: from -100 °C to -85 °C. Motor: Primary. SW & LW Gain sensors: 0

BBRUSSRAMP-F_011228_163729.ifg

Measurements: 59.Scanner Position: DEEP SPACE (RUSSIAN BB). Black Body Temperature: from -80 °C to -55 °C. Motor: Primary. SW & LW Gain sensors: 0

BBRUSSRAMP-G_011228_165402.ifg

Measurements: 200.Scanner Position: DEEP SPACE (RUSSIAN BB). Black Body Temperature: from -80 °C to -55 °C. Motor: Primary. SW & LW Gain sensors: 0

BBINT_011228_174851.ifg

Measurements: 200.Scanner Position: INTERNAL BLACK BODY. Motor: Primary. SW & LW Gain sensors: 0.



1.3 19 JAN 2002

It should be noted that in the above activity the IFSI BB was destroyed because the temperature was set to 150.6 deg C. which a priory should not have been accepted. In January, when trying again the IFSI BB, we discovered that indeed the BB was not usable any more, see the following files:

19 JAN 2002 : All the acquisitions have been done into the laboratory, at room temperature. File Description:

BBEXT51_020119_204408.ifg

Measurements: 50. Scanner Position: NADIR. Black Body Temperature: 51 °C. SW & LW Gain sensors: 0

BBEXT50-80_020119_202836.ifg

Measurements: 50.Scanner Position: NADIR. Black Body Temperature: 50.8 °C.SW & LW Gain sensors: 0

BBEXT50-50_020119_200849.ifg

Measurements: 50.Scanner Position: NADIR. Black Body Temperature: 50.5 °C.SW & LW Gain sensors: 0

BBEXT50_020119_193915.ifg

Measurements: 50. Scanner Position: NADIR. Black Body Temperature: 50 °C. SW & LW Gain sensors: 0

BBEXT33_020119_185818.ifg

Measurements: 50. Scanner Position: NADIR. Black Body Temperature: 33 °C. SW & LW Gain sensors: 0

BBEXT28_020119_183041.ifg

Measurements: 50. Scanner Position: NADIR. Black Body Temperature: 33 °C. SW & LW Gain sensors: 0

BBEXT1_020119_1726004.ifg

Measurements: 100. Scanner Position: NADIR. Black Body Temperature: 28 °C. SW & LW Gain sensors: 0

BBINT_020119_170733.ifg

Measurements: 100. Scanner Position: Internal Black Body. SW & LW Gain sensors: 0

BBINT_020119_175647.ifg

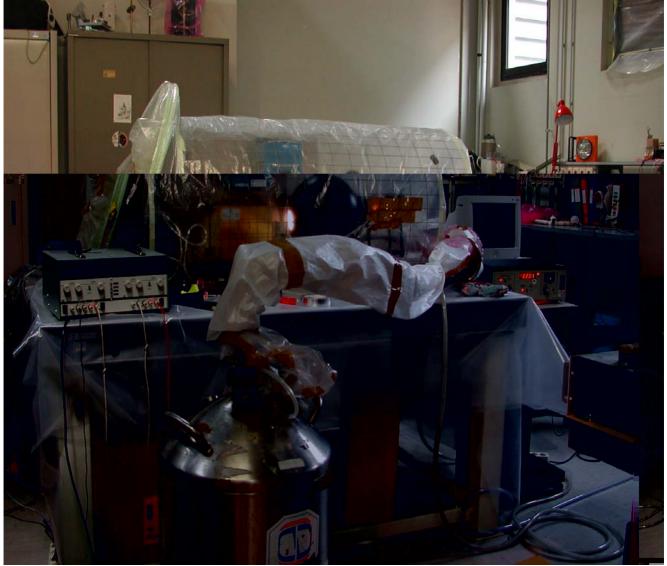
Measurements: 100. Scanner Position: Internal Black Body. SW & LW Gain sensors: 0

BBINT_TEST5_020119_184219.ifg

Measurements: 64. Test 5: n=4. Scanner Position: Internal Black Body. SW & LW Gain sensors: 0



After the work on the LW channel on the 27 and 28 DEC. we tried to work on the SW channel in the greenhouse by cooling the detector.



PICTURE 5 Integrating sphere with the cooled detector in the greenhouse.



1.4 **29 DEC 2001**

All acquisitions have been done into the russian laboratory at room temperature, using the Integrating Sphere, with a constant current set to 5.540 A, while the luminance has been changed about every 100 acquisitions.

File descriptions:

SPHERE500_011229_124020.ifg

50 Acquisitions. Scanner Position: Cold Space. Luminance: 500 footlamberts. SW Gain Sensor: 6. LW Gain Sensor: 0. Motor: Secondary

SPHERE500_011229_125006.ifg

_____6

13 Acquisitions. Scanner Position: Cold Space. Luminance: 500 footlamberts. SW Gain Sensor: 6. LW Gain Sensor: 0. Motor: Secondary

SPHERE500_011229_125356.ifg

100 Acquisitions. Scanner Position: Cold Space. Luminance: 500 footlamberts.SW Gain Sensor:5.LW Gain Sensor: 0. Motor: Secondary

SPHERE500 011229 131202.ifg

26 Acquisitions. Scanner Position: Cold Space. Luminance: 500 footlamberts. SW Gain Sensor: 5. LW Gain Sensor: 0. Motor: Primary

SPHERE400_011229_132537.ifg

100 Acquisitions. Scanner Position: Cold Space. Luminance: 400 footlamberts. SW Gain Sensor: 5. LW Gain Sensor: 0. Motor: Primary

SPHERE300_011229_134318.ifg

100 Acquisitions. Scanner Position: Cold Space. Luminance: 300 footlamberts. SW Gain Sensor: 5. LW Gain Sensor: 0. Motor: Primary

SPHERE200_011229_140323.ifg

100 Acquisitions. Scanner Position: Cold Space. Luminance: 200 footlamberts. SW Gain Sensor: 5. LW Gain Sensor: 0. Motor: Primary

SPHERE100_011229_142042.ifg



66 Acquisitions. Scanner Position: Cold Space. Luminance: 100 footlamberts. SW Gain Sensor: 5. LW Gain Sensor: 0. Motor: Primary

In this activity we had several problems and the motion was not good many times. We decided to take the experiment out.

NOTE ALL THE MEASUREMENTS ON THE INTERNAL BB ARE STILL GOOD

At this point in the calibration activity we had several problems with the double pendulum motion : the motion was not OK in one direction, and often not ok also in the opposite direction. The temporary solution to the problem was to put PFS upside down : not pending from the dampers, but lying on the dampers. This solved the motion problem, which was, indeed, caused by the current being now too low for the motor. Previously had been added 2 coils to module O in order to decrease the inrush current at switch on, as measured in module P, because was too large.

29 Dec 2001

All acquisitions have been done into the russian laboratory at room temperature. cooling the coldfinger with nitrogen. See picture: image01.JPG File descriptions:

CLAND 011000 160652 '6

CLAMP_011229_160653.ifg

58 Acquisitions. Scanner Position: CALIBRATION LAMP.SW Gain Sensor: 0.LW Gain Sensor:0

30 Dec 2001

All acquisitions have been done into the laboratory at room temperature. File descriptions:

CLAMP_PFS_ALIVE_011230_180411.ifg

50 Acquisitions. Scanner Position: CALIBRATION LAMP.SW Gain Sensor: 0.LW Gain Sensor:0. "PFS UPSIDEDOWN"

CLAMP_011229_160653.ifg

_ _ _ C

58 Acquisitions. Scanner Position: CALIBRATION LAMP. SW Gain Sensor: 0.LW Gain Sensor:0

02 Jan 2002

All acquisitions have been done into the laboratory at room temperature. The ADC Serial Clock is set to 4 MHz, instead of 2 MHz.

File descriptions:

CLAMP_020102_160326.ifg

20 Acquisitions. Scanner Position: CALIBRATION LAMP. SW Gain Sensor: 0. LW Gain Sensor: 0



CLAMP_020102_161532.ifg

50 Acquisitions. Scanner Position: CALIBRATION LAMP. SW Gain Sensor: 1. LW Gain Sensor:0

05 Jan 2002

All acquisitions have been done into the laboratory at room temperature, without the scanner. The distance between the Lamp and the Module O is 95 cm.

See picture: image01.jpg

File descriptions:

CERAMIC_020105_122310.ifg

100 Acquisitions. SW Gain Sensor: 3. LW Gain Sensor: 0. Lamp Current: 7 Amps

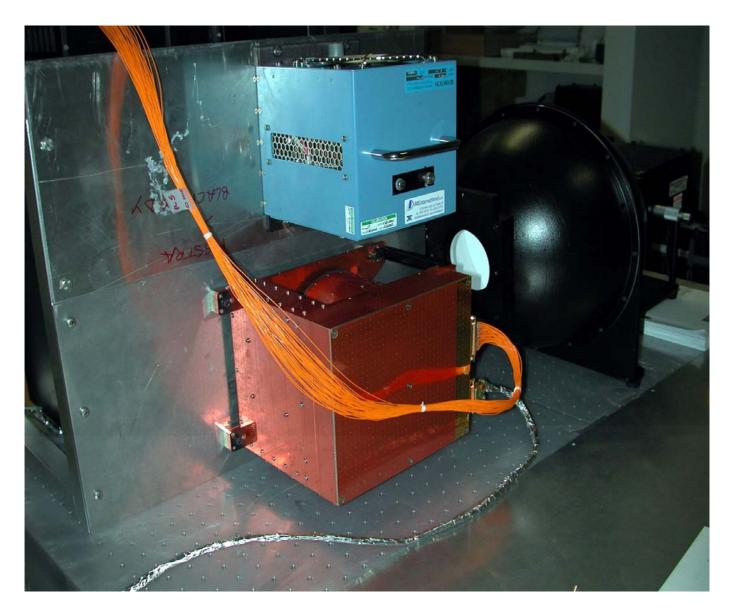


PICTURE 6 PFS in the lab with the Oriel lamp without the scanner.









PICTURE 7 PFS in the lab with the scanner and the integrating sphere.



1.5 02 JAN 2002

All acquisitions have been done into the laboratory at room temperature, without the scanner, using the Integrating Sphere, with a constant current set to 5.540 A. The ADC Serial Clock is set to 4 MHz, instead of 2 MHz.

File descriptions:

_____ SPHERE100 020102 121928.ifg

50 Acquisitions. Luminance: 100 footlamberts. SW Gain Sensor: 0. LW Gain Sensor: 0. Motor: Primarv

------_____

SPHERE200 020102 125326.ifg _____

50 Acquisitions. Luminance: 200 footlamberts. SW Gain Sensor: 0. LW Gain Sensor: 0. Motor: Primary.

_____ SPHERE300 020102 131251.ifg

50 Acquisitions. Luminance: 300 footlamberts. SW Gain Sensor: 0. LW Gain Sensor: 0. Motor: Primary

_____ _____

SPHERE400 020102 133345.ifg _____

50 Acquisitions. Luminance: 400 footlamberts. SW Gain Sensor: 0. LW Gain Sensor: 0. Motor: Primarv

SPHERE500_020102_135121.ifg _____

200 Acquisitions. Luminance: 500 footlamberts. SW Gain Sensor: 0. LW Gain Sensor: 0. Motor: Primary

_____ SPHERE1000 020102 143628.ifg

50 Acquisitions. Luminance: 1000 footlamberts. SW Gain Sensor: 0. LW Gain Sensor: 0. Motor: Primary

_____ SPHERE2000 020102 144813.ifg

50 Acquisitions. Luminance: 2000 footlamberts. SW Gain Sensor: 0. LW Gain Sensor: 0. Motor: Primary

SPHERE3000_020102_150212.ifg _____

50 Acquisitions. Luminance: 3000 footlamberts. SW Gain Sensor: 0. LW Gain Sensor: 0. Motor: Primary

_____ SPHERE4000 020102 151345.ifg



50 Acquisitions. Luminance: 4000 footlamberts. SW Gain Sensor: 0. LW Gain Sensor: 0. Motor: Primary

SPHERE5000_020102_152758.ifg

50 Acquisitions. Luminance: 5000 footlamberts. SW Gain Sensor: 0. LW Gain Sensor: 0. Motor: Primary

SPHERE100_020102_154015.ifg

20 Acquisitions. Luminance: 100 footlamberts. SW Gain Sensor: 5. LW Gain Sensor: 0. Motor: Primary

SPHERE300_020102_154015.ifg

20 Acquisitions. Luminance: 300 footlamberts. SW Gain Sensor: 5. LW Gain Sensor: 0. Motor: Primary

SPHERE500_020102_155210.ifg

20 Acquisitions. Luminance: 500 footlamberts. SW Gain Sensor: 5. LW Gain Sensor: 0. Motor: Primary

1.6 03 JAN 2002

All acquisitions have been done into the laboratory at room temperature, without the scanner, using the Integrating Sphere, with a constant current set to 5.540 A. File descriptions:

SPHERE10A_020103_095638.ifg

50 Acquisitions. Luminance: 10 footlamberts. SW Gain Sensor: 0. Motor: Primary

SPHERE10B_020103_100737.ifg

50 Acquisitions. Luminance: 10 footlamberts. SW Gain Sensor: 1. Motor: Primary

SPHERE10C_020103_102135.ifg

20 Acquisitions. Luminance: 10 footlamberts. SW Gain Sensor: 2. Motor: Primary

SPHERE10D_020103_102848.ifg

20 Acquisitions. Luminance: 10 footlamberts. SW Gain Sensor: 3. Motor: Primary

SPHERE10E_020103_103643.ifg

20 Acquisitions. Luminance: 10 footlamberts. SW Gain Sensor: 4. Motor: Primary

SPHERE10F_020103_104104.ifg

20 Acquisitions. Luminance: 10 footlamberts. SW Gain Sensor: 5. Motor: Primary

SPHERE10G_020103_104539.ifg

20 Acquisitions. Luminance: 10 footlamberts. SW Gain Sensor: 6. Motor: Primary

SPHERE10H_020103_105016.ifg

20 Acquisitions. Luminance: 10 footlamberts. SW Gain Sensor: 7. Motor: Primary

SPHERE30A_020103_105540.ifg

20 Acquisitions. Luminance: 30 footlamberts. SW Gain Sensor: 2. Motor: Primary

SPHERE30B_020103_110205.ifg

20 Acquisitions. Luminance: 30 footlamberts. SW Gain Sensor: 3. Motor: Primary

SPHERE30C_020103_110641.ifg

20 Acquisitions. Luminance: 30 footlamberts. SW Gain Sensor: 5. Motor: Primary

SPHERE30D 020103 111113.ifg

20 Acquisitions. Luminance: 30 footlamberts. SW Gain Sensor: 7. Motor: Primary

SPHERE50A_020103_111528.ifg

20 Acquisitions. Luminance: 50 footlamberts. SW Gain Sensor: 0. Motor: Primary

SPHERE50B_020103_112102.ifg

20 Acquisitions. Luminance: 50 footlamberts. SW Gain Sensor: 1. Motor: Primary

SPHERE50C_020103_112628.ifg

20 Acquisitions. Luminance: 50 footlamberts. SW Gain Sensor: 4. Motor: Primary

SPHERE50D_020103_113101.ifg

20 Acquisitions. Luminance: 50 footlamberts. SW Gain Sensor: 5. Motor: Primary

SPHERE50E_020103_113533.ifg

20 Acquisitions. Luminance: 50 footlamberts. SW Gain Sensor: 6. Motor: Primary

SPHERE50F_020103_114211.ifg



50 Acquisitions. Luminance: 50 footlamberts. SW Gain Sensor: 7. Motor: Primary

SPHERE70A_020103_121126.ifg

_____C

20 Acquisitions. Luminance: 70 footlamberts. SW Gain Sensor: 2. Motor: Primary

SPHERE70B_020103_121527.ifg

20 Acquisitions. Luminance: 70 footlamberts.SW Gain Sensor: 3. Motor: Primary

SPHERE70C_020103_122238.ifg

50 Acquisitions. Luminance: 70 footlamberts. SW Gain Sensor: 5. Motor: Primary

SPHERE70D_020103_123455.ifg

20 Acquisitions. Luminance: 70 footlamberts. SW Gain Sensor: 7. Motor: Primary

SPHERE100A_020103_124248.ifg

50 Acquisitions. Luminance: 100 footlamberts. SW Gain Sensor: 0. Motor: Primary

SPHERE100B_020103_125430.ifg

50 Acquisitions. Luminance: 100 footlamberts. SW Gain Sensor: 1. Motor: Primary

SPHERE100C_020103_130500.ifg

20 Acquisitions. Luminance: 100 footlamberts. SW Gain Sensor: 2. Motor: Primary

SPHERE100D_020103_131128.ifg

50 Acquisitions. Luminance: 100 footlamberts. SW Gain Sensor: 4. Motor: Primary

SPHERE100E 020103 132354.ifg

20 Acquisitions. Luminance: 100 footlamberts. SW Gain Sensor: 5. Motor: Primary

SPHERE100F_020103_133059.ifg

50 Acquisitions. Luminance: 100 footlamberts. SW Gain Sensor: 6. Motor: Primary

SPHERE100G_020103_134126.ifg

20 Acquisitions. Luminance: 100 footlamberts. SW Gain Sensor: 7. Motor: Primary

SPHERE200A_020103_134910.ifg



20 Acquisitions. Luminance: 200 footlamberts. SW Gain Sensor: 0. Motor: Primary

SPHERE200B_020103_135829.ifg

50 Acquisitions. Luminance: 200 footlamberts. SW Gain Sensor: 3. Motor: Primary

SPHERE200C_020103_142646.ifg

20 Acquisitions. Luminance: 200 footlamberts. SW Gain Sensor: 5. Motor: Primary

SPHERE200D_020103_143816.ifg

20 Acquisitions. Luminance: 200 footlamberts. SW Gain Sensor: 7. Motor: Primary

SPHERE300A_020103_144426.ifg

20 Acquisitions. Luminance: 300 footlamberts. SW Gain Sensor: 0. Motor: Primary

SPHERE300B_020103_144841.ifg

20 Acquisitions. Luminance: 300 footlamberts. SW Gain Sensor: 1. Motor: Primary

1.7 04 JAN 2002

All acquisitions have been done into the laboratory at room temperature, without the scanner, using the Integrating Sphere, with a constant current set to 5.540 A. File descriptions:

SPEHERE300C_020104_174823.ifg

50 Acquisitions. Luminance: 300 footlamberts. SW Gain Sensor: 2. Motor: Primary

SPEHERE300D_020104_181108.ifg

20 Acquisitions. Luminance: 300 footlamberts. SW Gain Sensor: 5. Motor: Primary

SPEHERE300E_020104_182507.ifg

20 Acquisitions. Luminance: 300 footlamberts. SW Gain Sensor: 4. Motor: Primary

SPEHERE300F_020104_183526.ifg

20 Acquisitions. Luminance: 300 footlamberts. SW Gain Sensor: 6. Motor: Primary

SPEHERE300G_020104_184150.ifg

20 Acquisitions. Luminance: 300 footlamberts. SW Gain Sensor: 7. Motor: Primary

SPEHERE400A_020104_184150.ifg

20 Acquisitions. Luminance: 400 footlamberts. SW Gain Sensor: 0. Motor: Primary

SPEHERE400B 020104 185355.ifg

20 Acquisitions. Luminance: 400 footlamberts. SW Gain Sensor: 3. Motor: Primary

SPEHERE400C_020104_190014.ifg

20 Acquisitions. Luminance: 400 footlamberts. SW Gain Sensor: 5. Motor: Primary

SPEHERE400D_020104_190816.ifg

20 Acquisitions. Luminance: 400 footlamberts. SW Gain Sensor: 7. Motor: Primary

SPEHERE500A_020104_191341.ifg

20 Acquisitions. Luminance: 500 footlamberts. SW Gain Sensor: 0. Motor: Primary

SPEHERE500B_020104_192136.ifg

20 Acquisitions. Luminance: 500 footlamberts. SW Gain Sensor: 1. Motor: Primary

SPELIERE500C 020104 102045 :f~

SPEHERE500C_020104_192945.ifg

20 Acquisitions. Luminance: 500 footlamberts. SW Gain Sensor: 2. Motor: Primary

1.8 05 JAN 2002

All acquisitions have been done into the laboratory at room temperature, without the scanner, using the Integrating Sphere, with a constant current set to 5.540 A. File descriptions:

SPEHERE500D_020105_085459.ifg

20 Acquisitions. Luminance: 500 footlamberts. SW Gain Sensor: 4. Motor: Primary.

SPEHERE500E_020105_090300.ifg

STERERE300E_020105_070500.llg

20 Acquisitions. Luminance: 500 footlamberts. SW Gain Sensor: 5. Motor: Primary

SPELIERE500E 020105 000012 :6-

SPEHERE500F_020105_090912.ifg

20 Acquisitions. Luminance: 500 footlamberts. SW Gain Sensor: 6. Motor: Primary

SPEHERE500G_020105_091353.ifg



20 Acquisitions. Luminance: 500 footlamberts. SW Gain Sensor: 7. Motor: Primary

SPEHERE1000A_020105_092450.ifg

20 Acquisitions. Luminance: 1000 footlamberts. SW Gain Sensor: 0. Motor: Primary

SPEHERE1000B_020105_092923.ifg

20 Acquisitions. Luminance: 1000 footlamberts. SW Gain Sensor: 3. Motor: Primary

SPEHERE1000C_020105_093453.ifg

20 Acquisitions. Luminance: 1000 footlamberts. SW Gain Sensor: 4. Motor: Primary

SPEHERE2000A_020105_094206.ifg

20 Acquisitions. Luminance: 2000 footlamberts. SW Gain Sensor: 0. Motor: Primary

SPEHERE2000B_020105_094848.ifg

20 Acquisitions. Luminance: 2000 footlamberts. SW Gain Sensor: 1. Motor: Primary

SPEHERE2000C_020105_095331.ifg

20 Acquisitions. Luminance: 2000 footlamberts. SW Gain Sensor: 2. Motor: Primary

SPEHERE2000D_020105_095809.ifg

20 Acquisitions. Luminance: 2000 footlamberts. SW Gain Sensor: 5. Motor: Primary

SPEHERE3000A_020105_100303.ifg

20 Acquisitions. Luminance: 3000 footlamberts. SW Gain Sensor: 0. Motor: Primary

SPEHERE3000B_020105_100919.ifg

20 Acquisitions. Luminance: 3000 footlamberts. SW Gain Sensor: 3. Motor: Primary

SPEHERE3000C_020105_101344.ifg

20 Acquisitions. Luminance: 3000 footlamberts. SW Gain Sensor: 4. Motor: Primary

SPEHERE4000A_020105_101838.ifg

20 Acquisitions. Luminance: 4000 footlamberts. SW Gain Sensor: 0. Motor: Primary

SPEHERE4000B 020105 102713.ifg



20 Acquisitions. Luminance: 4000 footlamberts. SW Gain Sensor: 1. Motor: Primary

SPEHERE4000C_020105_103444.ifg

20 Acquisitions. Luminance: 4000 footlamberts. SW Gain Sensor: 2. Motor: Primary

SPEHERE5000A_020105_103913.ifg

20 Acquisitions. Luminance: 5000 footlamberts. SW Gain Sensor: 0. Motor: Primary

SPEHERE5000B_020105_104626.ifg

20 Acquisitions. Luminance: 5000 footlamberts. SW Gain Sensor: 1. Motor: Primary

SPEHERE5000C_020105_105101.ifg

20 Acquisitions. Luminance: 5000 footlamberts. SW Gain Sensor: 2. Motor: Primary

SPEHERE5000D_020105_105742.ifg

20 Acquisitions. Luminance: 5000 footlamberts. SW Gain Sensor: 3. Motor: Primary

This set completes the SW channel calibration at room temperature. It also contains the non linearity studies with the gain and source intensities.

1.9 FOV MEASUREMENTS

05 Jan 2002

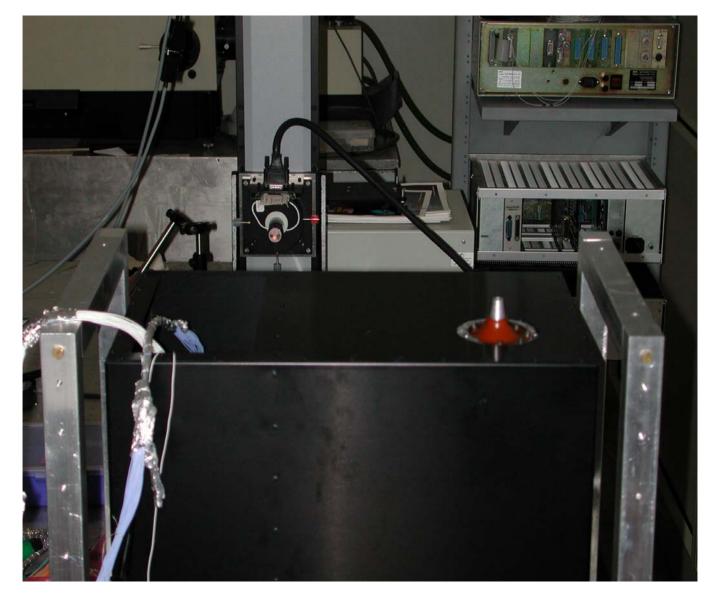
All acquisitions have been done into the laboratory at room temperature, without the scanner. The distance between the Ceramic Element and the Module O is 93 cm, 6 Acquisitions for each step. File descriptions:

SW Gain Sensor: 3. LW Gain Sensor: 3. Lamp Current: 7.5 Amps

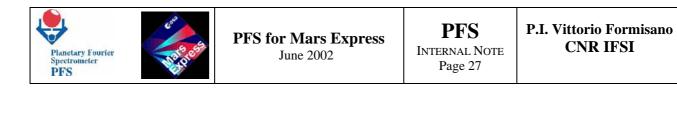
1-FOV001_020105_135001.ifg - 48 Acquisitions 2-FOV002_020105_140141.ifg - 42 Acquisitions 3-FOV003_020105_141228.ifg - 90 Acquisitions 4-FOV004_020105_143245.ifg - 46 Acquisitions (miss last 2) 5-FOV005_020105_144640.ifg - 48 Acquisitions 6-FOV006_020105_154103.ifg - 66 Acquisitions 7-FOV007_020105_155620.ifg - 66 Acquisitions 8-FOV008_020105_161135.ifg - 66 Acquisitions 9-FOV009_020105_162953.ifg - 66 Acquisitions 10-FOV010_020105_164812.ifg - 78 Acquisitions 11-FOV011_020105_183811.ifg - 12 Acquisitions 12-FOV012_020105_184200.ifg - 60 Acquisitions 13-FOV013_020105_185639.ifg - 90 Acquisitions

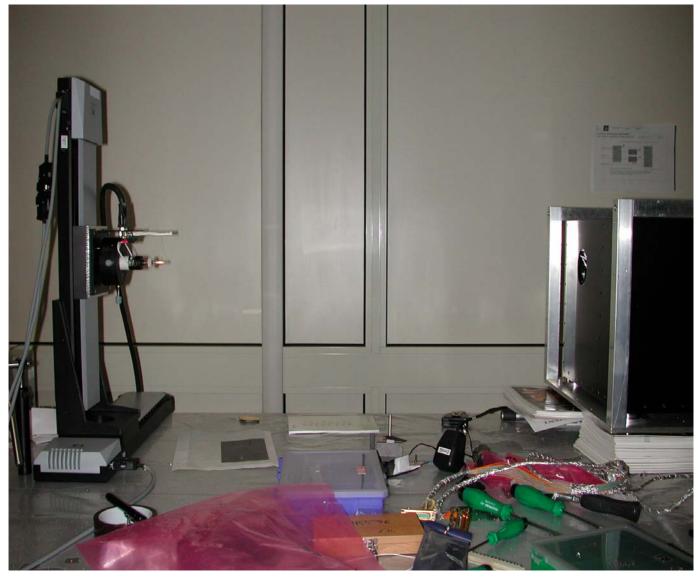


14-FOV014_020105_191701.ifg - 90 Acquisitions (1 more) 15-FOV015_020105_192738.ifg - 24 Acquisitions 16-FOV016_020105_193627.ifg - 114 Acquisitions 17-FOV018_020105_201200.ifg - 6 Acquisitions 18-FOV019_020105_201353.ifg - 84 Acquisitions 19-FOV020_020106_145039.ifg - 66 Acquisitions 20-FOV021_020106_150618.ifg - 66 Acquisitions 21-FOV022_020106_152151.ifg - 66 Acquisitions 22-FOV023_020106_153818.ifg - 66 Acquisitions 23-FOV024_020106_155309.ifg - 48 Acquisitions 24-FOV025_020106_160420.ifg - 24 Acquisitions 25-FOV026_020106_161111.ifg - 36 Acquisitions



PICTURE 8 FOV MEASUREMENTS FRONT VIEW OF THE ASSEMBLY





picture 9

FOV MEASUREMENTS SIDE VIEW OF THE ASSEMBLY



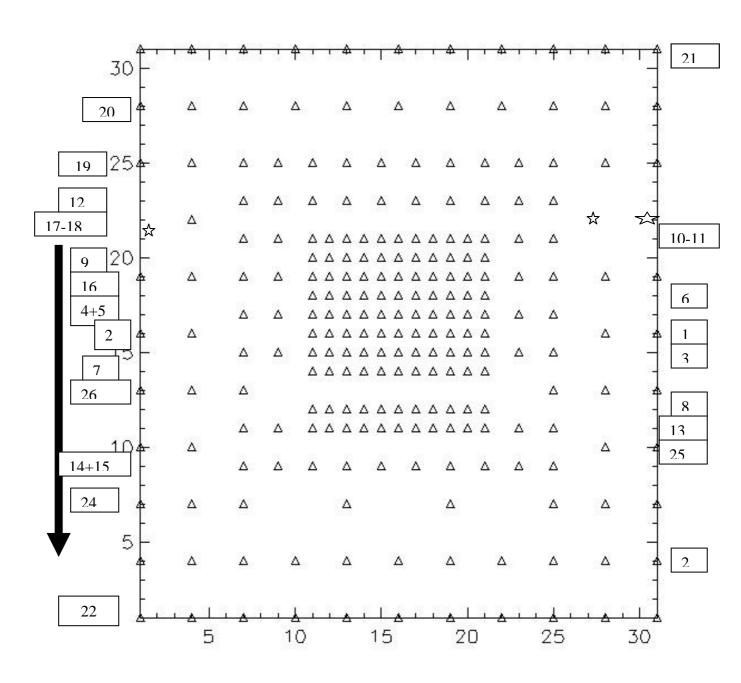


FIGURE 1

LOCATION IN THE X-Y PLANE WHERE THE 6 MEASUREMENTS WERE TAKEN. LOCATION IS IN CM. NOTE UP IS NOW DOWN (SEE ARROW).

Note the number is the file number. Its position gives the direction in which the source was moved along the line to which the number refers. When 2 numbers are present 2 files were needed to complete the line. Only 6 measurements per position were usually taken.



The test was made by putting the Module O on the optical bench, lying on the dampers, as it is shown in the image. The measurement of the FOV and optical axis direction was made in two steps: first we measured the transfer function of the scanner, then we measured the FOV of Module O without the scanner. For the FOV, in each position indicated in the figure, we took10 measurements. The figure identifies a plane at 93 cm from the Module O outer face. Dimensions are in cm. The source (ceramic element) was moved in a plane parallel to the Module O face. The origin (center in the figure) was not in front of the O module optical aperture, but displaced. The normal to O located in the center of the entrance hole, was in the location of the rectangular shape at (3,2) cm away from the center. Here is where we expect the optical axis, and therefore the maximum of the detected signal. The FWHM of the signal detected is shown in the next figure. PFS Field of view

Curves represent surface curves of PFS signal at a level half of the maximum. Scale is 1 cm in the focal actuator plane.

As we can see the optical axis of the two channels are not coincident and are not normal to the Module O "normal" direction. Note that the SW FOV is 38 mm diameter and the LW is 60 mm diameter. To transform these measurements in angle we have to take into account the internal path of module O plus the 930 mm external : for the SW we have a distance of 930 + 489.9 = 1419 This correspond to 1.57 degrees FOV. For the LW we have 930+345,3 = 1275 this corresponds to a FOV of 2,77 degrees. This are FWHM FOV measurements.

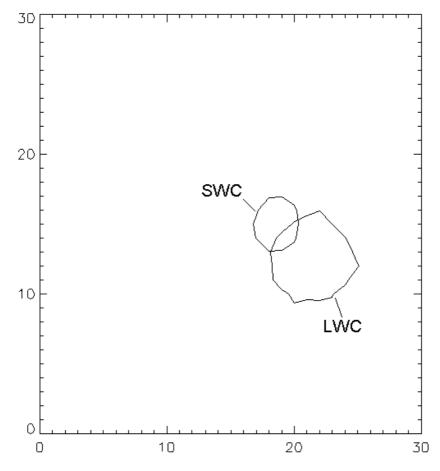
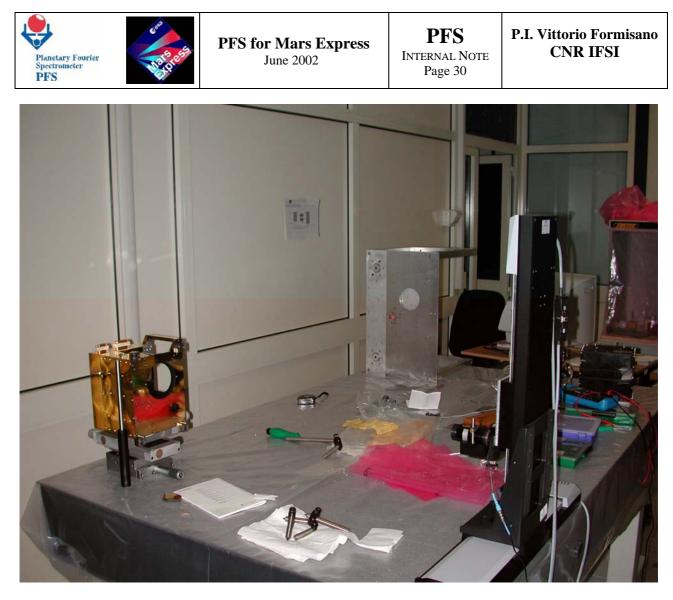


FIGURE 2 FOV of the 2 channels



PICTURE 10 scanner transfer function measurement arrangement.

1.10 SCANNER TRANSFER FUNCTION

First we put the laser beam perpendicular to the 4 feet of the scanner.

We measure the distance from the x-y to the 4 feet : 860 mm.

We measure the distance between the output of the scanner and the plate on which we project the laser "image": it is 1290 mm.

We measure then the height of the laser beam in input : 260 mm at the laser, 261 mm at the scanner: we consider to be OK (parallel to the optical bench).

We measure height of the laser beam at the exit of the scanner 261 mm, and at the plate : 259 mm. It went down only 2 mm.

On the following picture we can see 5 circles-spot, location of the laser in the center and at different location on the scanner mirror. With respect to the continuous line, where the red spot should be, we see a displacement of 12 mm. From this point we move 20 mm up or down the input laser, and the spot moves of the same amount. If we move horizontally by 40 mm, the spot moves by 32, if we move by 20 in the opposite direction the spot moves by almost 20. In conclusion, in one direction is OK, in the horizontal direction the scanner mirror reduces the image (ovalization of the footprint). The deviation is in the Y direction (motion direction) of the spacecraft by 12/1290 mm = 0.55 degrees.







PICTURE 11 Motion of the laser spots for the scanner transfer function measurements.

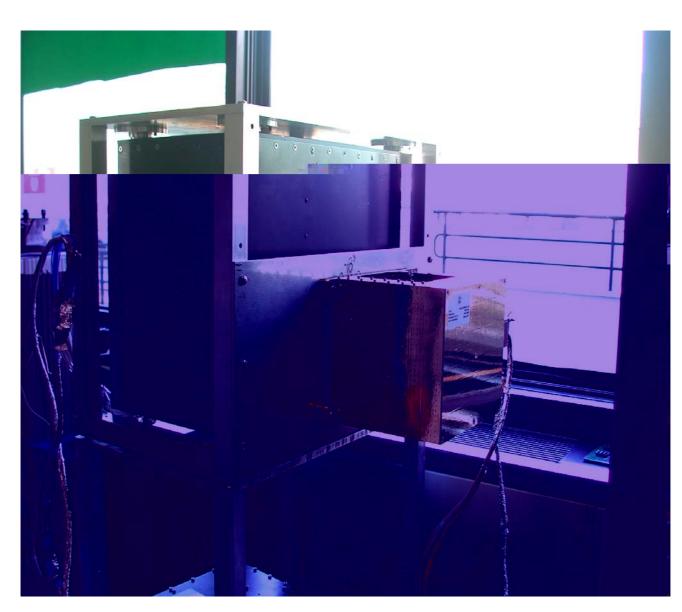
OTHER MEASUREMENTS HAVE BEEN MADE : CUBE MIRROR ON SCANNER AND LASER LIGHT REFLECTED BY SILICON WINDOW.



PICTURE 12

Measurements out of the window: Earth's atmosphere study.





PICTURE 13 Measurements toward Rome.



1.11 OUT OF THE WINDOW MEASUREMENTS.

10 Jan 2002 File descriptions:

FUORI_G00_020110_151119.ifg

Looking at the other building. 3 Acquisitions. SW Gain Sensor: 0. LW Gain Sensor: 0

FUORI G33 020110 152447.ifg

Looking at the other building. 3 Acquisitions. SW Gain Sensor: 3. LW Gain Sensor: 3

FUORI_G53_020110_152730.ifg

Looking at the other building. 3 Acquisitions. SW Gain Sensor: 5. LW Gain Sensor: 3

FUORI_G53_020110_153127.ifg

Looking at the other building. 100 Acquisitions. SW Gain Sensor: 5. LW Gain Sensor: 3

FUORI_G53_020110_154952.ifg

Looking at the atmosphere. 3 Acquisitions. SW Gain Sensor: 5. LW Gain Sensor: 3

FUORI_G53_020110_155221.ifg

Looking at the atmosphere. 10 Acquisitions. SW Gain Sensor: 5. LW Gain Sensor: 3

FUORI_G53_020110_155630.ifg

Looking at the atmosphere. 10 Acquisitions. SW Gain Sensor: 5. LW Gain Sensor: 3

FUORI_G53_020110_155920.ifg

Not looking at the atmosphere (the ceiling). 3 Acquisitions. SW Gain Sensor:5. LW Gain Sensor:3

The measurements were taken first looking straight to IAS building, then looking toward IFA, and far away, finally we tried to look upward toward the sky, but perhaps we see the roof.



PFS for Mars Express June 2002

WE PASS NOW TO THE MEASUREMENTS MADE IN THERMOVACUUM CHAMBER.



PICTURE 14 Measurements in the Thermovac. Mercury lamp in the picture.

1.12 THERMOVACUUM, 08 JAN 2002

All the acquisitions have been done into the thermovacuum chamber. All acquisitions have been done into the laboratory at room temperature, without the scanner, using the Integrating Sphere, the Mercury Lamp, or the Internal Lamp. For the Integrating Sphere the current is set to 5.540 A.

See picture: PICTURE 14

File descriptions:

SPHERE10_020108_113622.ifg



10 Acquisitions. Luminance: 10 footlamberts. SW Gain Sensor: 0. SW Detector Temp.: 286 °K. Motor: Primary

SPHERE10A_020108_113833.ifg

10 Acquisitions. Luminance: 10 footlamberts. SW Gain Sensor: 2. SW Detector Temp.: 286 °K. Motor: Primary

SPHERE10B_020108_154216.ifg

10 Acquisitions. Luminance: 10 footlamberts. SW Gain Sensor: 2. SW Detector Temp.: 240 °K. Motor: Primary

SPHERE10C_020108_154416.ifg

STILLE10C_020100_134410.11g

10 Acquisitions. Luminance: 10 footlamberts. SW Gain Sensor: 0. SW Detector Temp.: 240 °K. Motor: Primary

SPHERE10D_020108_163102.ifg

10 Acquisitions. Luminance: 10 footlamberts. SW Gain Sensor: 4. SW Detector Temp.: 236 °K. Motor: Primary

SPHERE20A 020108 114231.ifg

10 Acquisitions. Luminance: 20 footlamberts. SW Gain Sensor: 2. SW Detector Temp.: 286 °K. Motor: Primary

SPHERE30_020108_113404.ifg

10 Acquisitions. Luminance: 30 footlamberts. SW Gain Sensor: 0. SW Detector Temp.: 286 °K. Motor: Primary

SPHERE30A_020108_114442.ifg

10 Acquisitions. Luminance: 30 footlamberts. SW Gain Sensor: 2. SW Detector Temp.: 286 °K. Motor: Primary

SPHERE30B_020108_154015.ifg

10 Acquisitions. Luminance: 30 footlamberts. SW Gain Sensor: 2. SW Detector Temp.: 240 °K. Motor: Primary

SPHERE30C_020108_154634.ifg

10 Acquisitions. Luminance: 30 footlamberts. SW Gain Sensor: 0. SW Detector Temp.: 240 °K. Motor: Primary



SPHERE30D_020108_163331.ifg

10 Acquisitions. Luminance: 30 footlamberts. SW Gain Sensor: 4. SW Detector Temp.: 236 °K. Motor: Primary

SPHERE50_020108_113200.ifg

10 Acquisitions. Luminance: 50 footlamberts. SW Gain Sensor: 0. SW Detector Temp.: 286 °K. Motor: Primary

SPHERE50A_020108_114653.ifg

10 Acquisitions. Luminance: 50 footlamberts. SW Gain Sensor: 2. SW Detector Temp.: 286 °K. Motor: Primary

SPHERE50B_020108_153815.ifg

10 Acquisitions. Luminance: 50 footlamberts. SW Gain Sensor: 2. SW Detector Temp.: 240 °K. Motor: Primary

SPHERE50C_020108_154836.ifg

10 Acquisitions. Luminance: 50 footlamberts. SW Gain Sensor: 0. SW Detector Temp.: 240 °K. Motor: Primary

SPHERE50D_020108_163542.ifg

10 Acquisitions. Luminance: 50 footlamberts. SW Gain Sensor: 4. SW Detector Temp.: 236 °K. Motor: Primary

SPHERE50D_020108_163542.ifg

10 Acquisitions. Luminance: 50 footlamberts. SW Gain Sensor: 4. SW Detector Temp.: 236 °K. Motor: Primary

SDUEDE70 020100 112042 :6

SPHERE70_020108_112942.ifg

10 Acquisitions. Luminance: 70 footlamberts. SW Gain Sensor: 0. SW Detector Temp.: 286 °K. Motor: Primary

SPHERE70A_020108_114914.ifg

STILLRE/0A_020100_114/14.11g

10 Acquisitions. Luminance: 70 footlamberts. SW Gain Sensor: 2. SW Detector Temp.: 286 °K. Motor: Primary

SPHERE70B_020108_153601.ifg



10 Acquisitions. Luminance: 70 footlamberts. SW Gain Sensor: 2. SW Detector Temp.: 240 °K. Motor: Primary

SPHERE70C_020108_155039.ifg

10 Acquisitions. Luminance: 70 footlamberts. SW Gain Sensor: 0. SW Detector Temp.: 240 °K. Motor: Primary

SPHERE70D_020108_163805.ifg

100 Acquisitions. Luminance: 70 footlamberts. SW Gain Sensor: 4. SW Detector Temp.: 236 °K. Motor: Primary

SPHERE100_020108_112721.ifg

STHERE100_020100_112721.itg

10 Acquisitions. Luminance: 100 footlamberts. SW Gain Sensor: 0. SW Detector Temp.: 286 °K. Motor: Primary

SPHERE100A_020108_115137.ifg

10 Acquisitions. Luminance: 100 footlamberts. SW Gain Sensor: 2. SW Detector Temp.: 286 °K. Motor: Primary

SPHERE100B_020108_153240.ifg

10 Acquisitions. Luminance: 100 footlamberts. SW Gain Sensor: 2. SW Detector Temp.: 240 °K. Motor: Primary

SPHERE100C_020108_155249.ifg

10 Acquisitions. Luminance: 100 footlamberts. SW Gain Sensor: 0. SW Detector Temp.: 240 °K. Motor: Primary

SPHERE100D_020108_164945.ifg

10 Acquisitions. Luminance: 100 footlamberts. SW Gain Sensor: 4. SW Detector Temp.: 236 °K. Motor: Primary

SPHERE200_020108_112431.ifg

10 Acquisitions. Luminance: 200 footlamberts. SW Gain Sensor: 0. SW Detector Temp.: 286 °K. Motor: Primary

SPHERE200A_020108_115406.ifg

10 Acquisitions. Luminance: 200 footlamberts. SW Gain Sensor: 2. SW Detector Temp.: 286 °K. Motor: Primary

SDUEDE200D 020100 152021 :6

SPHERE200B_020108_153031.ifg



10 Acquisitions. Luminance: 200 footlamberts. SW Gain Sensor: 2.SW Detector Temp.: 240 °K. Motor: Primary

SPHERE200C_020108_155457.ifg

10 Acquisitions. Luminance: 200 footlamberts. SW Gain Sensor: 0. SW Detector Temp.: 240 °K. Motor: Primary

SPHERE300_020108_112145.ifg

10 Acquisitions. Luminance: 300 footlamberts. SW Gain Sensor: 0. SW Detector Temp.: 286 °K. Motor: Primary

SPHERE300A_020108_115613.ifg

10 Acquisitions. Luminance: 300 footlamberts. SW Gain Sensor: 2. SW Detector Temp.: 286 °K. Motor: Primary

SPHERE300B_020108_152826.ifg

10 Acquisitions. Luminance: 300 footlamberts. SW Gain Sensor: 2. SW Detector Temp.: 240 °K. Motor: Primary

SPHERE300C_020108_160400.ifg

10 Acquisitions. Luminance: 300 footlamberts. SW Gain Sensor: 0. SW Detector Temp.: 240 °K. Motor: Primary

SPHERE400_020108_111901.ifg

10 Acquisitions. Luminance: 400 footlamberts. SW Gain Sensor: 0. SW Detector Temp.: 286 °K. Motor: Primary

SPHERE400A_020108_115824.ifg

10 Acquisitions. Luminance: 400 footlamberts. SW Gain Sensor: 2. SW Detector Temp.: 286 °K. Motor: Primary

SPHERE400B_020108_152622.ifg

10 Acquisitions. Luminance: 400 footlamberts. SW Gain Sensor: 2. SW Detector Temp.: 240 °K. Motor: Primary

SPHERE400C_020108_160621.ifg

10 Acquisitions. Luminance: 400 footlamberts. SW Gain Sensor: 0. SW Detector Temp.: 240 $^{\circ}\text{K}.$ Motor: Primary



SPHERE500_020108_111555.ifg

10 Acquisitions. Luminance: 500 footlamberts. SW Gain Sensor: 0. SW Detector Temp.: 286 °K. Motor: Primary

SPHERE500A_020108_120033.ifg

10 Acquisitions. Luminance: 500 footlamberts. SW Gain Sensor: 2. SW Detector Temp.: 286 °K.Motor: Primary

SPHERE500B_020108_152329.ifg

10 Acquisitions . Luminance: 500 footlamberts. SW Gain Sensor: 2. SW Detector Temp.: 240 °K. Motor: Primary

SPHERE500C_020108_160847.ifg

10 Acquisitions. Luminance: 500 footlamberts. SW Gain Sensor: 0. SW Detector Temp.: 240 °K. Motor: Primary

SPHERE1000_020108_111323.ifg

10 Acquisitions. Luminance: 1000 footlamberts. SW Gain Sensor: 0. SW Detector Temp.: 286 °K. Motor: Primary

SPHERE1000C_020108_161106.ifg

10 Acquisitions. Luminance: 1000 footlamberts. SW Gain Sensor: 0. SW Detector Temp.: 240 °K. Motor: Primary

CDUEDE2000 020100 111010 :f.

SPHERE3000_020108_111019.ifg

10 Acquisitions. Luminance: 3000 footlamberts. SW Gain Sensor: 0. SW Detector Temp.: 286 °K. Motor: Primary

SPHERE3000C_020108_161307.ifg

10 Acquisitions. Luminance: 3000 footlamberts. SW Gain Sensor: 0. SW Detector Temp.: 240 °K. Motor: Primary

SPHERE5000_020108_110554.ifg

51112122000_020100_11000 hing

10 Acquisitions. Luminance: 5000 footlamberts. SW Gain Sensor: 0. SW Detector Temp.: 286 °K. Motor: Primary

MERCURY_020108_094128.ifg

3 Acquisitions. Motor: Primary



MERCURY_020108_103149.ifg

5 Acquisitions. Motor: Primary

MERCURY_TEST6_020108_094425.ifg

- - -

210 Acquisitions - TEST 6. Motor: Primary

1.13 10 JAN 2002

All the acquisitions have been done into the thermovacuum chamber. All acquisitions have been done into the laboratory at room temperature, without the scanner, using the Mercury Lamp, or the Internal Lamp.

See picture: PICTURE 14

File descriptions:

CLAMP_020110_094859.ifg

38 Acquisitions. SW Gain Sensor: 0. SW Detector Temp.: 295 °K. Motor: Primary

CLAMP 020110 100310.ifg

CLAMP_020110_100310.1fg

60 Acquisitions. SW Gain Sensor: 0. SW Detector Temp.: 295 °K. Motor: Primary

CLAMP_SWDETRAMP_020110_101627.ifg

200 Acquisitions. SW Gain Sensor: 0. SW Detector Temp.: RAMP TO 250 °K. Motor: Primary

CLAMP_SWDETRAMP_020110_105332.ifg

92 Acquisitions. SW Gain Sensor: 0. SW Detector Temp.: RAMP TO 250 °K. Motor: Primary

CLAMP_SWDETRAMP_020110_112402.ifg

100 Acquisitions. SW Gain Sensor: 0. SW Detector Temp.: RAMP TO 250 °K. Motor: Primary

CLAMP_SWDET-23_020110_120131.ifg

100 Acquisitions. SW Gain Sensor: 0. SW Detector Temp.: 250 °K. Motor: Primary

CLAMP_TEST5_020110_114348.ifg

96 Acquisitions. SW Gain Sensor: 0. SW Detector Temp.: 250 °K. Motor: Primary

MERCURY_020110_122007.ifg

3 Acquisitions. SW Gain Sensor: 0. SW Detector Temp.: 250 °K. Motor: Primary

MERCURY_020110_122336.ifg

_ _ _ C

3 Acquisitions. SW Gain Sensor: 0. SW Detector Temp.: 250 °K. Motor: Primary

MERCURY_TEST6_020110_122521.ifg

84 Acquisitions. SW Gain Sensor: 0. SW Detector Temp.: 250 °K. Motor: Primary

1.14 21 JAN 2002

All the acquisitions have been done into the thermovacuum chamber, using the Integrating Sphere with the current set to 5.540 A.

File descriptions:

FONDO_020121_113645.ifg

20 Acquisitions. Luminance: 0 footlamberts. SW Gain Sensor: 0. SW Detector Temp.: 280°K. Motor: Primary

FONDO_020121_122532.ifg

20 Acquisitions. Luminance: 0 footlamberts. SW Gain Sensor: 2. SW Detector Temp.: 280 °K. Motor: Secondary

FONDO_020121_121919.ifg

20 Acquisitions. Luminance: 0 footlamberts. SW Gain Sensor: 4. SW Detector Temp.: 280 °K. Motor: Secondary

SPHERE_020121_123331.ifg

17 Acquisitions. Luminance: 10 footlamberts. SW Gain Sensor: 0. SW Detector Temp.:280°K. Motor: Secondary

SNUEDE 020121 122/21 16

SPHERE_020121_123621.ifg

20 Acquisitions. Luminance: 10 footlamberts. SW Gain Sensor: 2. SW Detector Temp.: 280 °K. Motor: Secondary

SPHERE_020121_124014.ifg



20 Acquisitions. Luminance: 10 footlamberts. SW Gain Sensor: 4. SW Detector Temp.: 280 °K. Motor: Secondary

SPHERE_020121_124403.ifg

20 Acquisitions. Luminance: 50 footlamberts. SW Gain Sensor: 0. SW Detector Temp.: 280 °K. Motor: Secondary

SPHERE_020121_124758.ifg

20 Acquisitions. Luminance: 50 footlamberts. SW Gain Sensor: 2. SW Detector Temp.: 280 °K. Motor: Secondary

SPHERE_020121_125145.ifg

20 Acquisitions. Luminance: 50 footlamberts. SW Gain Sensor: 4. SW Detector Temp.: 280 °K. Motor: Secondary

SPHERE_020121_125549.ifg

20 Acquisitions. Luminance: 100 footlamberts. SW Gain Sensor: 0. SW Detector Temp.: 280 °K. Motor: Secondary

SPHERE_020121_125957.ifg

20 Acquisitions. Luminance: 100 footlamberts. SW Gain Sensor: 2. SW Detector Temp.: 280 °K. Motor: Secondary

SPHERE_020121_130345.ifg

20 Acquisitions. Luminance: 100 footlamberts. SW Gain Sensor: 4. SW Detector Temp.: 280 °K Motor: Secondary

SPHERE_020121_130734.ifg

20 Acquisitions. Luminance: 500 footlamberts. SW Gain Sensor: 0. SW Detector Temp.: 280 °K Motor: Secondary

SPHERE_020121_131155.ifg

20 Acquisitions. Luminance: 500 footlamberts. SW Gain Sensor: 2. SW Detector Temp.: 280 °K Motor: Secondary

SPHERE_020121_131548.ifg

20 Acquisitions. Luminance: 500 footlamberts. SW Gain Sensor: 4. SW Detector Temp.: 280 $^{\circ}\text{K}$ Motor: Secondary



SPHERE_020121_131935.ifg

20 Acquisitions. Luminance: 1000 footlamberts. SW Gain Sensor: 0. SW Detector Temp.: 280 °K Motor: Secondary

SPHERE_020121_132352.ifg

20 Acquisitions. Luminance: 1000 footlamberts. SW Gain Sensor: 2. SW Detector Temp.: 280 °K Motor: Secondary

SPHERE 020121 132743.ifg

20 Acquisitions. Luminance: 1000 footlamberts. SW Gain Sensor: 4. SW Detector Temp.: 280 °K Motor: Secondary

SPHERE_020121_133139.ifg

20 Acquisitions. Luminance: 5000 footlamberts. SW Gain Sensor: 0. SW Detector Temp.: 280 °K Motor: Secondary

SPHERE_020121_133546.ifg

20 Acquisitions. Luminance: 5000 footlamberts. SW Gain Sensor: 2. SW Detector Temp.: 280 °K Motor: Secondary

SPHERE_020121_133936.ifg

20 Acquisitions. Luminance: 5000 footlamberts. SW Gain Sensor: 4. SW Detector Temp.: 280 °K Motor: Secondary

BBINT_020121_134322.ifg

20 Acquisitions. SW Gain Sensor: 4. SW Detector Temp.: 280 °K. Motor: Secondary

BBINT_020121_134840.ifg

20 Acquisitions. SW Gain Sensor: 2. SW Detector Temp.: 280 °K. Motor: Secondary

BBINT_020121_135233.ifg

20 Acquisitions. SW Gain Sensor: 0. SW Detector Temp.: 280 °K. Motor: Secondary

CLAMP_020121_135622.ifg

20 Acquisitions. SW Gain Sensor: 0. SW Detector Temp.: 280 °K. Motor: Secondary

BBINT-RAMP_020121_140147.ifg

300 Acquisitions. SW Gain Sensor: 0. SW Detector Temp.: from 280 °K to ? °K. Motor: Secondary

BBINT-RAMP_020121_152718.ifg

100 Acquisitions . SW Gain Sensor: 0. SW Detector Temp.: from ? °K to ? °K. Motor: Secondary

BBINT-RAMP_020121_154539.ifg

20 Acquisitions. SW Gain Sensor: 0. SW Detector Temp.: from ? °K to ? °K. Motor: Secondary

BBINT-RAMP_020121_155057.ifg

20 Acquisitions. SW Gain Sensor: 0. SW Detector Temp.: from ? °K to ? °K. Motor: Secondary

FONDO_020121_183028.ifg

20 Acquisitions. Luminance: 0 footlamberts. SW Gain Sensor: 0. SW Detector Temp.: 260 °K Motor: Secondary

SPHERE_020121_160224.ifg

20 Acquisitions. Luminance: 10 footlamberts. SW Gain Sensor: 0 SW Detector Temp.: 260 °K Motor: Secondary

SPHERE_020121_160559.ifg

20 Acquisitions. Luminance: 50 footlamberts. SW Gain Sensor: 0. SW Detector Temp.: 260 °K Motor: Secondary

SPHERE_020121_161004.ifg

20 Acquisitions. Luminance: 100 footlamberts. SW Gain Sensor: 0. SW Detector Temp.: 260 °K Motor: Secondary

SPHERE_020121_161417.ifg

20 Acquisitions. Luminance: 500 footlamberts. SW Gain Sensor: 0. SW Detector Temp.: 260 °K Motor: Secondary

SPHERE_020121_161852.ifg

20 Acquisitions. Luminance: 1000 footlamberts. SW Gain Sensor: 0. SW Detector Temp.: 260 °K Motor: Secondary

SPHERE_020121_162247.ifg

20 Acquisitions. Luminance: 5000 footlamberts. SW Gain Sensor: 0. SW Detector Temp.: 260 °K Motor: Secondary

SPHERE_020121_162707.ifg



20 Acquisitions. Luminance: 5000 footlamberts. SW Gain Sensor: 2. SW Detector Temp.: 264.3 °K Motor: Secondary

SPHERE_020121_163146.ifg

20 Acquisitions. Luminance: 1000 footlamberts. SW Gain Sensor: 2. SW Detector Temp.: 264.3 °K Motor: Secondary

SPHERE_020121_163538.ifg

20 Acquisitions. Luminance: 500 footlamberts. SW Gain Sensor: 2. SW Detector Temp.: 265.3 °K Motor: Secondary

CLAMP_020121_163938.ifg

20 Acquisitions. SW Gain Sensor: 0. SW Detector Temp.: 265.3 °K. Motor: Secondary

SPHERE_020121_173737.ifg

22 Acquisitions. Luminance: 100 footlamberts. SW Gain Sensor: 2. SW Detector Temp.: 266 °K Motor: Secondary

SPHERE_020121_174130.ifg

20 Acquisitions. Luminance: 50 footlamberts. SW Gain Sensor: 2. SW Detector Temp.: 266 °K Motor: Secondary

SPHERE_020121_174521.ifg

20 Acquisitions. Luminance: 10 footlamberts. SW Gain Sensor: 2. SW Detector Temp.: 266 °K Motor: Secondary

SPHERE_020121_174919.ifg

20 Acquisitions. Luminance: 10 footlamberts. SW Gain Sensor: 4. SW Detector Temp.: 266 °K Motor: Secondary

SPHERE_020121_175324.ifg

20 Acquisitions. Luminance: 50 footlamberts. SW Gain Sensor: 4. SW Detector Temp.: 266 °K Motor: Secondary

SPHERE_020121_175720.ifg

20 Acquisitions. Luminance: 100 footlamberts. SW Gain Sensor: 4. SW Detector Temp.: 266 °K Motor: Secondary

SPHERE_020121_180110.ifg



20 Acquisitions. Luminance: 500 footlamberts. SW Gain Sensor: 4. SW Detector Temp.: 266 °K Motor: Secondary _____

SPHERE_020121_180453.ifg

20 Acquisitions. Luminance: 1000 footlamberts. SW Gain Sensor: 4. SW Detector Temp.: 266 °K Motor: Secondary

SPHERE_020121_180840.ifg _____

20 Acquisitions. Luminance: 5000 footlamberts. SW Gain Sensor: 4. SW Detector Temp.: 266 °K Motor: Secondary

FONDO_020121_182153.ifg _____

20 Acquisitions. Luminance: 0 footlamberts. SW Gain Sensor: 4. SW Detector Temp.: 266 °K Motor: Secondary

_____ _____

FONDO_020121_182631.ifg _____

1

20 Acquisitions. Luminance: 0 footlamberts. SW Gain Sensor: 2. SW Detector Temp.: 266 °K Motor: Secondary

2	Parent Directory	10-Jan-2002	12:11	-
	TVBalance_011215_165>	03-Jan-2002	16:03	161k
	TVBalance_011215_182>	03-Jan-2002	16:03	926k
	TVBalance_011215_190>	03-Jan-2002	16:03	2.9M
	TVBalance_011215_202>	03-Jan-2002	16:03	362k
ľ	TVBalance_011215_202>	03-Jan-2002	16:03	1.8M

2	Parent Dir	rectory		10-Jan-2002	12:11	-
	TVBalance	011216	151>	03-Jan-2002	16:03	2.3M
ľ	TVBalance	011216	160>	03-Jan-2002	16:03	2.3M
ľ	TVBalance	011216	<u>171></u>	03-Jan-2002	16:03	1.2M
ľ	TVBalance	011216	<u>185></u>	03-Jan-2002	16:03	886k



1.15

05 Jan 2002

All acquisitions have been done into the laboratory at room temperature, without the scanner. The distance between the Lamp and the Module O is 103 cm.

File descriptions:

ORIEL LAMP

ORIEL_020105_114934.ifg

100 Acquisitions. SW Gain Sensor: 3. LW Gain Sensor: 0. Lamp Current: 5 Amps

1.1	6 - MINERALS			
2	Parent Directory	20-Jan-2002	19:17	-
Ĩ	ALBITE.ifg	17-Jan-2002	16:14	81k
	GESSO.ifg	17-Jan-2002	16:14	81k
ľ	KBR.ifg	17-Jan-2002	16:14	309k

18 Jan 2002

All the acquisitions have been done into the laboratory, at room temperature.

See picture: minerals.jpg

File Description:

ALBITE_020118_142603.ifg

Measurements: 50. ALBITE (Ontario) very fine grain, 0.50 mg dispersi in 253.1 mg of KBr.

KBr1_020118_145325.ifg

Measurements: 50. KBr n°1 from 252.8 mg

------HEMATITE1_020118_150316.ifg

Measurements: 50. Hematite South Dakota (raw). 0.67 mg in 263.3 mg of KBr

DOL1_020118_153351.ifg

Measurements: 50. Dolomite Bayern (raw). 0.69 mg in 251.2 mg of KBr

KBr2_020118_162223.ifg

Measurements: 50. KBr n°2 from 255.7 mg

QUARZO1_020118_163533.ifg

Measurements: 50. quartz dust Dr. Krantz. 0.67 mg in 248.8 mg of KBr

KAOLINITE1_020118_170007.ifg

Measurements: 50. Kaolinite New Mexico. Size of grain 20-50 um 0.66 mg in 248.8 mg of KBr

KBr1A_020118_171538.ifg

Measurements: 50. KBr $n^{\circ}1$ from 252.8 mg

GYPSUM_020118_174539.ifg

Measurements: 50. GYPSUM 1 Colorado. Size of grain 20-50 um 0.75 mg in 251.8 mg of KBr

OLIVINA2_020118_182438.ifg

Measurements: 50. OLIVINA 2 Eifel. Size of grain < 20 um, 0.72 mg in 249.5 mg of KBr

KBr2A_020118_184009.ifg

Measurements: 50. KBr n°2 from 255.7 mg

CALCITE_020118_190011.ifg

Measurements: 50. CALCITE 1 Lecce (raw). 0.64 mg in 265.5 mg of KBr

HEMATITE2_020118_191521.ifg

Measurements: 50. Hematite 2 Michigan (raw).0.72 mg in 254.2 mg of KBr

KBr1B_020118_193042.ifg

Measurements: 50. KBr n°1 from 252.8 mg

OLIVINA 1 020119 104540 ;fa

OLIVINA1_020118_194549.ifg

Measurements: 50. OLIVINA 1 Eifel. Size of grain 100-200 um, 0.68 mg in 249.9 mg of KBr



19 JAN 2002

All the acquisitions have been done into the laboratory, at room temperature.

File Description:

FONDO_020119_110418.ifg

50 Acquisitions. SW Gain Sensor: 2.LW Gain Sensor: 2. Lamp OFF

LAMP_020119_115249.ifg

50 Acquisitions. SW Gain Sensor: 1. LW Gain Sensor: 3. Lamp ON

GYPSUM1_020119_122454.ifg

Measurements: 50. gYPSUM 1 Colorado. Size of grain 20-50 um 0.75 mg in 251.8 mg of KBr

KBr2_020119_123635.ifg

Measurements: 50. KBr n°2 from 255.7 mg

CALCITE1_020119_125027.ifg

Measurements: 50. CALCITE 1 Lecce (raw). 0.64 mg in 265.5 mg of KBr

KBr1_020119_130303.ifg

Measurements: 50. KBr n°1 from 252.8 mg

FONDOL2_020119_131639.ifg

Measurements: 50. Buio lampada 2 (ceramic element)

LAMP2_020119_135825.ifg

Measurements: 50. Lamp ON (ceramic element)

ALBITE_020119_211450.ifg

Measurements: 50. ALBITE (Ontario) very fine grain, 0.50 mg dispersi in 253.1 mg of KBr.

KBr1_020119_213045.ifg

Measurements: 50. KBr n°1 from 252.8 mg

5	Planetary Fourier spectrometer PFS	PFS for Mars Ex June 2002	press	PFS Internal Page 5	Note	P.I. Vittorio Formisano CNR IFSI
	Parent Directory	20-Jan-2002	19:17	-		
ľ	ALBITE_BB_020120_113.	<u>.></u> 20-Jan-2002	15:19	2.OM		
ľ	DOLOMITE_BB_020120_1.	<u>.></u> 20-Jan-2002	15:19	2.OM		
ľ	EMPTY-CSI1_LAMP_0201.	<u>.></u> 20-Jan-2002	15:59	2.OM		
	EMPTY-CSI2_LAMP_0201.	<u>.></u> 20-Jan-2002	15:59	2.OM		
	EMPTY-CSI_BB_020120	<u>.></u> 20-Jan-2002	15:19	2.OM		
ľ	GYPSUM_BB_020120_115.	<u>.></u> 20-Jan-2002	15:19	2.OM		
Į	OLIVINE_LAMP_020120	<u>.></u> 20-Jan-2002	15:59	2.OM		
	ORPIMENT_LAMP_020120.	<u>.></u> 20-Jan-2002	15:59	2.OM		
Ĩ	ORPI_BB_020120_12332.	<u>.></u> 20-Jan-2002	15:19	2.OM		
Î	SPODUMENE_LAMP_02012.	<u>.></u> 20-Jan-2002	15:59	1.9M		
	TITANITE_LAMP_020120.	<u>.></u> 20-Jan-2002	15:59	2.OM		
S	pasticche.jpg	20-Jan-2002	19:15	852k		

16 JAN 2002: IFSI Clean Room 18 JAN 2002: IFSI Laboratory - Ceramic Element 19 JAN 2002: IFSI Laboratory - Oriel Lamp

1.17 DATA TAKEN IN ALENIA IN TURIN

Stand Alone Procedure

The procedure done for the stand alone test before the integration to the spacecraft is the follow:

- 1. Test 1 as GONOGO Test
- 2. Unblock the Double Pendulum
- 3. Test 5 with the scanner pointing a source (a mercury lamp)
- 4. Test 6 with the scanner pointing a source (a mercury lamp)
- 5. Block the Double Pendulum
- 6. Run automatically the FPT test sequence (in nominal and in the redundand power)

Test 1 as GONOGO Test

During the test we checked the current and the voltage values and are in the right range value. Also the Housekeeping was regular.

GONOGO TEST IS OK

UNBLOCK DOUBLE PENDULUM

Unblocking the pendulum the current values are correct and after about 12 min. the pendulum was correctly unblocked. We received the event packet meaning that the double pendulum was unblocked, and also the remarks in the housekeeping meant that the pendulum was really unblocked.



Test 5 and Test 6

Done with the MERCURY LAMP : DATA ARE IN THE FILE S: MERCURY_020205_12213..> 25-Feb-2002 19:32 163k MERCURY 020205 12285..> 25-Feb-2002 19:32 163k

MERCURY_020205_12341..> 25-Feb-2002 19:32 33k

MERCURY 020205 12370..> 25-Feb-2002 19:32 49k

E. MERCURY 020205 12402...> 25-Feb-2002 19:32 163k

? Stand-Alone-Report.doc 25-Feb-2002 18:01 103k

TEST5 020205 115841.ifg 25-Feb-2002 19:35 1.3M

TEST6_020205_121131.ifg 25-Feb-2002 19:32 1.7M

Both tests are OK

BLOCK DOUBLE PENDULUM

The block is OK **Full Performance Test** Finally we performed the FPT test sequence. FPT Input File: /Mex/CmdFiles/PFS FPT.txt **Test Result for Nominal Power** Telecommands sent file name: /Mex/Archive/TELECOMMAND/C020205_133159.cmd Log file name: /Mex/Archive/TELECOMMAND/L020205_133243.log For Redundant Power the test result is ok (similar to the nominal). **CONCLUSION** The Stand Alone Test is OK

The file with the stand alone interferograms is : no file : science telemetry was not on.

Second we performed the Full performance test after integration on the spacecraft : the file with the data is :

- EFPT 020212_052243.ifg 25-Feb-2002 18:18 443k



2 TYPICAL RAW SPECTRA OBTAINED.

2.1 THERMAL LW CHANNEL.

We shall now present some raw data obtained with PFS. The raw data are interesting to note instrumental effects, when present, To have an idea of the intensity of the signal measured, to observe unexpected features and so on. In general raw data are useful to become familiar with the data themselves.

In the following figures we present data from the LW channel obtained with the IFSI BB. The measurements were taken in air, therefore atmospheric features (water and CO2 lines) are present. The side panels give the PFS measurement conditions. See plot 1,2,3,4,5.

We note that the spectrum covers 2.5 decades, the spectral range is between 250 and 1750 cm-1. The spectra show the typical modulation induced by the optical element. As the IFSI BB was hold at temperatures larger than ambient temperature, but decreasing from 150 to 30, we could see that the signal itself was decreasing: 10 DN at max, and 0.5 at 30 deg.

Note that a spurious peak at 1300 cm-1 appears due to bad motion in the measurements taken at low radiation intensity on December 27 and 28. Later the current of the motors was increased and the peak disappeared. This peak was present with the secondary motor (lowest current). The current was too low after the addition of two big inductance to decrease the inrush current l, as requested by module P. After January 2 the current was increased and the effect disappeared for ever. Plot 6 and 7 summarize and compare the measurements.

When comparing the raw data spectra, we should not expect the curves to be linearly ordered : planckians are not linear with T.

When we compare the IFSI BB at 30 C with internal BB (green and red in the second color figure) we see that, with respect to the detector we have + 7 deg and - 5 deg: the 2 curves almost coincide, meaning that the instrument temperature is also important and it has been neglected in this simple comparison.

The last color figure (plot 8) of the LW channel spectra shows 3 measurements of a 50 C source. The 3 are in relatively agreement, the differences being due to the fact that we had not waited 15 minutes for the measurement, as required by the stability of the source.

It is interesting to note, finally, the behaviour of the atmospheric features in the spectra : CO2 almost completely absent (only the Q branch visible : on the short path of 10-20 cm there is not enough CO2 to be measured, or the temperature is uniform). The water lines are more evident in the spectra of the hot source. (The termal difference increase the contrast).

2.2 THERMAL SW CHANNEL.

In the next set of figures (plot 9,10,11,12,13) we plot the contribution in the SW channel of the same measurements presented above. We see that we have now a 2 orders of magnitude spectra, ranging from 2000 to 4000 cm-1. The CO2 main band is clearly visible now, and we shall come back to comment it. We see that the SW channel is not able to see thermal differences less than 10 degrees (in a single spectrum). The important point is that we can use the same procedure for SW and LW to calibrate the LW channel and this portion of the SW channel. (see plot 14, 15, 16).

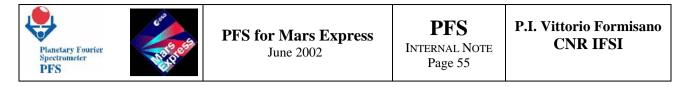
The blow up of the CO2 band shows several things : first PFS is able to resolve rotovibrational features (see plot 17). Second we discover that the central peak of the CO2 band is in some cases shifted by 5 cm-1. Details of this effect can be seen in the next 3 figures. See plot 18, 19, 20. xWe start with some water lines in the LW channel : it is clear that the water line does not move, but the laser diode reference wavelength is changing. The second figure shows similarly the CO2 Q branch: apparently tends to move and its depth is reduced : this is due to changing of the laser diode

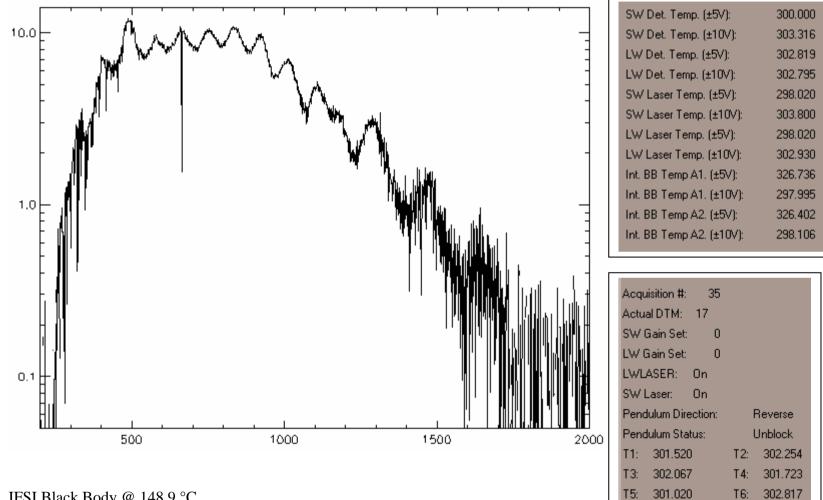


wavelength and to the response of the instrument to monochromatic sources. We shall come back later to this point, when studying the mercury lamp measurements.

In the cases of figures the temperature of the laser diodes was changing by only 0.5 - 0.6 degrees, but this was enough to change the wavelength DISCONTINUOUSLY. It is clear that if we mix all the measurements, with different motion (forward-reverse) and different laser diode wavelength, we will get wrong responsivity and NER.

It is obvious, therefore that before averaging data and computing whatever quantity, the data must be screened one by one and corrected for the laser diode wavelength. Also, the behaviour of the laser diode must be well studied, and the response of PFS to monochromatic features must be well understood.





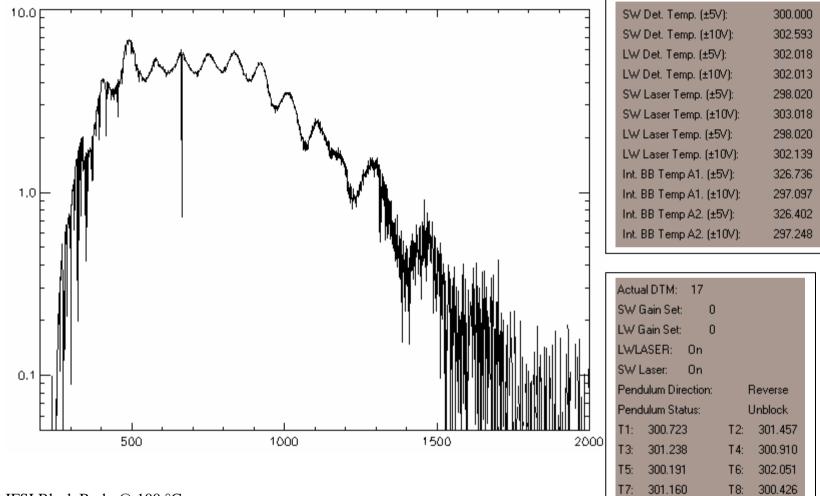
T7: 301.957

T8: 301.223

IFSI Black Body @ 148,9 °C LWC FFT spectrum

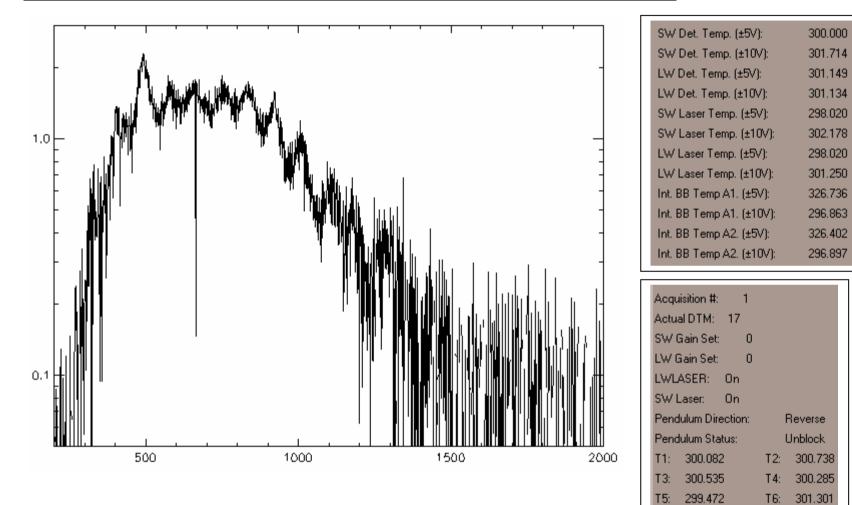
Plot - 1

Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 56	P.I. Vittorio Formisano CNR IFSI
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IFSI Black Body @ 100 °C LWC FFT spectrum

Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 57	P.I. Vittorio Formisano CNR IFSI
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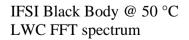


T7:

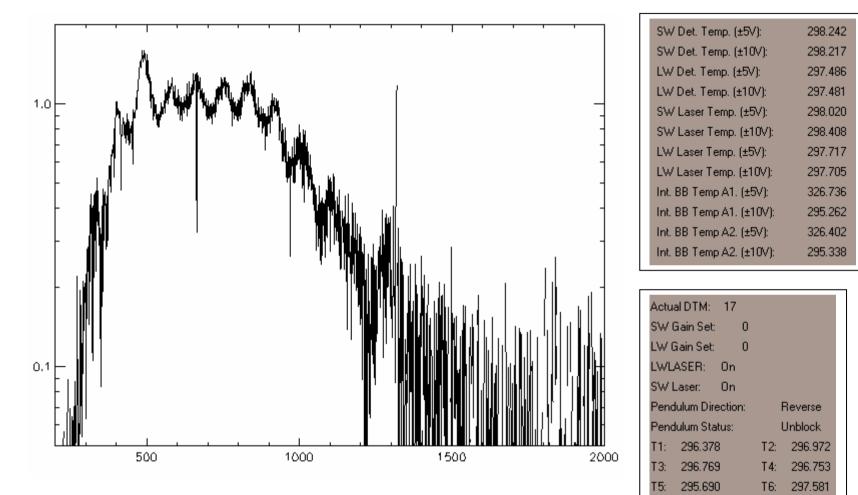
300.441

T8:

299.691



Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS INTERNAL NOTE Page 58	P.I. Vittorio Formisano CNR IFSI
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T7:

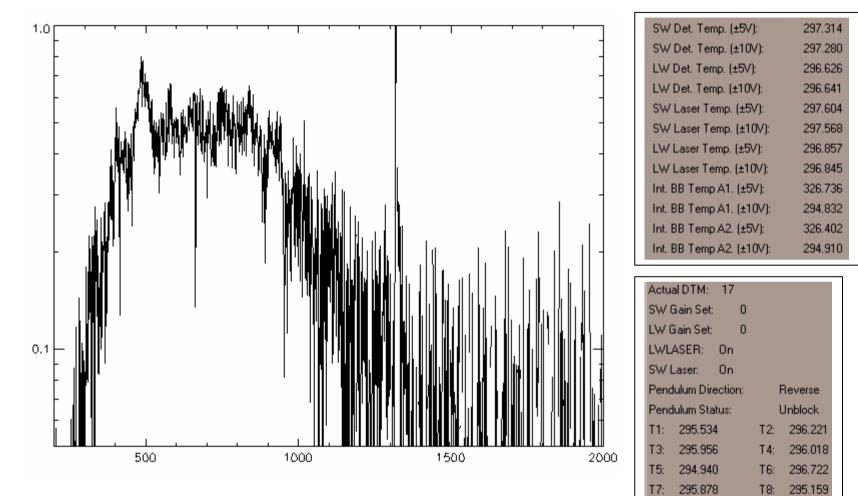
296.690

T8: 295.940

IFSI Black Body @ 40 °C LWC FFT spectrum

NOTE: Secondary motor was used. Plot – 4

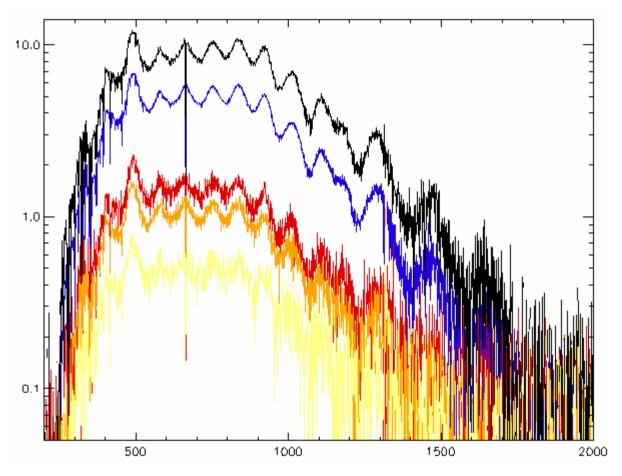
Planetary Fourier Spectrometer PFS PFS PFS PFS PFS PFS PFS PFS PFS PFS	PFS Internal Note Page 59	P.I. Vittorio Formisano CNR IFSI
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IFSI Black Body @ 30 °C LWC FFT spectrum

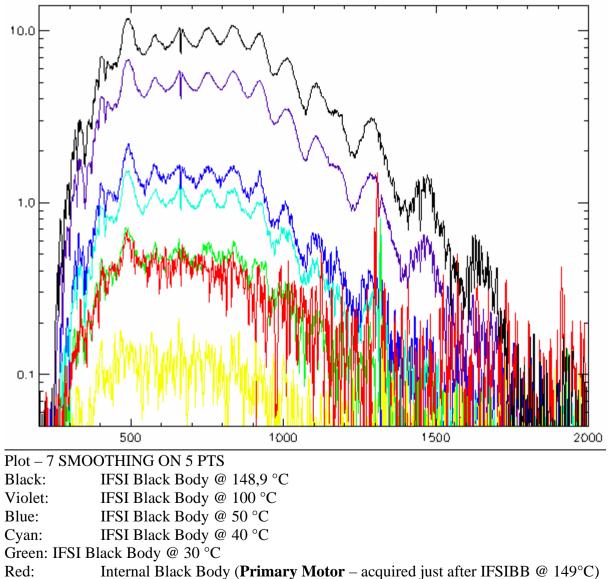
NOTE: Secondary motor was used. Plot – 5





Black:	IFSI Black Body @ 148,9 °C
Blue:	IFSI Black Body @ 100 °C
Red:	IFSI Black Body @ 50 °C
Orange:	IFSI Black Body @ 40 °C
Yellow:	IFSI Black Body @ 30 °C.LWC FFT spectra
Plot – 6	

Planetary Fourier Spectrometer PFS Provide PFS for Mars Express June 2002	PFS Internal Note Page 61	P.I. Vittorio Formisano CNR IFSI
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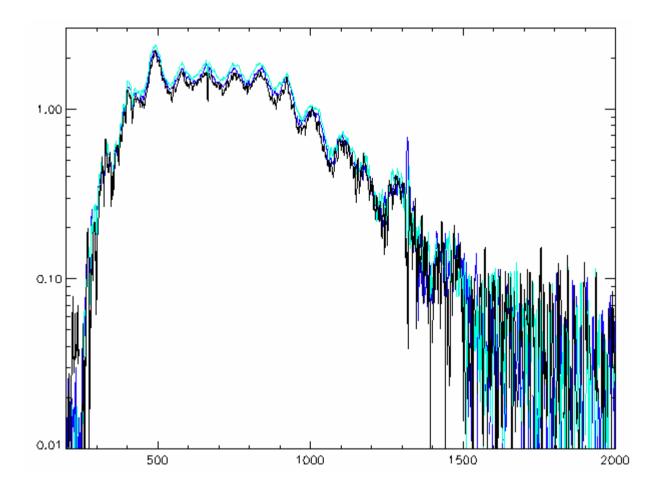


Yellow: Internal Black Body (Secondary Motor - acquired just after IFSIBB @ 50°C)

Int.BB after IFSIBB @	148,9°C
SW Det. Temp. (±5V):	300.000
SW Det. Temp. (±10V):	303.941
LW Det. Temp. (±5V):	303.386
LW Det. Temp. (±10V):	303.381
SW Laser Temp. (±5V):	298.020
SW Laser Temp. (±10V):	304.425
LW Laser Temp. (±5V):	298.020
LW Laser Temp. (±10V):	303.546
Int. BB Temp A1. (±5V):	326.736
Int. BB Temp A1. (±10V):	298.932
Int. BB Temp A2. (±5V):	326,402
Int. BB Temp A2. (±10V):	298.963

Int.BB after IFSIBB @ 50°C					
SW Det. Temp. (±5V):	300.000				
SW Det. Temp. (±10V):	299.995				
LW Det. Temp. (±5V):	299.352				
LW Det. Temp. (±10V):	299.337				
SW Laser Temp. (±5V):	298.020				
SW Laser Temp. (±10V):	300.381				
LW Laser Temp. (±5V):	298.020				
LW Laser Temp. (±10V):	299.541				
Int, BB Temp A1, (±5V):	326.736				
Int, BB Temp A1, (±10V):	296.199				
Int. BB Temp A2. (±5V):	326.402				
Int. BB Temp A2. (±10V):	296.352				

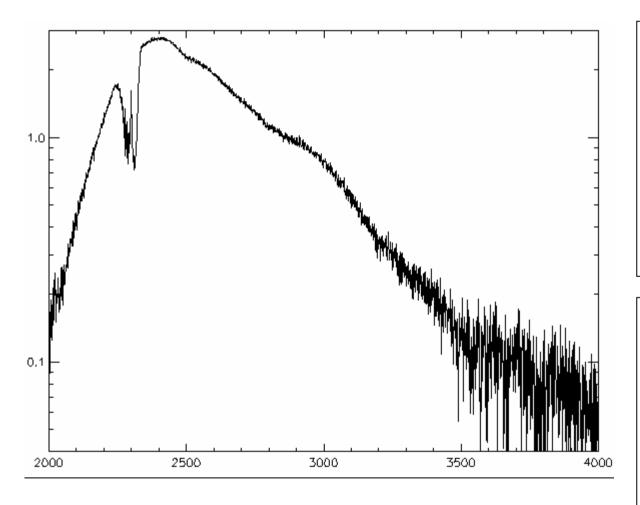
Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 62	P.I. Vittorio Formisano CNR IFSI
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Cyan:IFSI Black Body @ 50°C (secondary motor)Blue:IFSI Black Body @ 50,5°CBlack:IFSI Black Body @ 50°C

Plot-8

Planetary Fourier Spectrometer PFS Provider June 2002	PFS INTERNAL NOTE Page 63	P.I. Vittorio Formisano CNR IFSI
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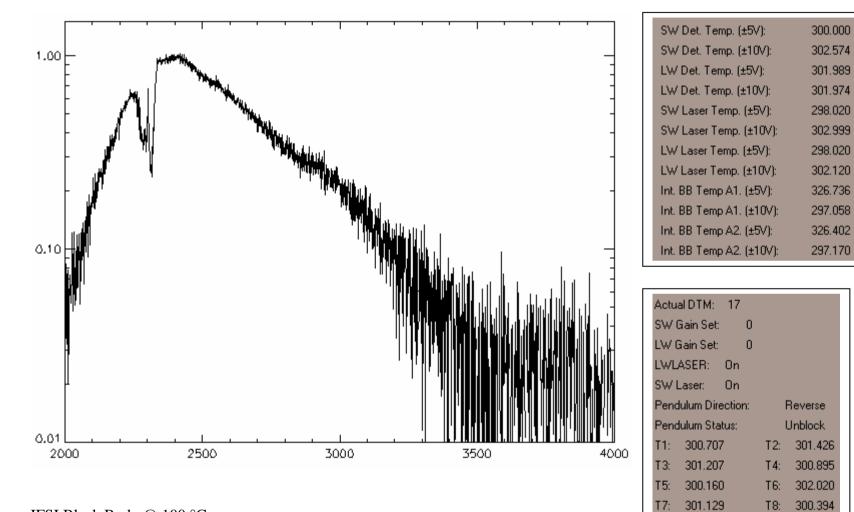


IFSI Black Body @ 148,9 °C SWC FFT spectrum Plot – 9

SW Det. Temp. (±5V):	300.000	
SW Det. Temp. (±10V):	303.199	
LW Det. Temp. (±5V):	302.722	
LW Det. Temp. (±10V):	302.717	
SW Laser Temp. (±5V):	298.020	
SW Laser Temp. (±10V):	303.692	
LW Laser Temp. (±5V):	298.020	
LW Laser Temp. (±10V):	302.794	
Int. BB Temp A1. (±5V):	326.736	
Int. BB Temp A1. (±10V):	297.917	
Int. BB Temp A2. (±5V):	326.402	
Int. BB Temp A2. (±10V):	297.989	
	SW Det. Temp. (±10V): LW Det. Temp. (±5V): LW Det. Temp. (±10V): SW Laser Temp. (±5V): SW Laser Temp. (±10V): LW Laser Temp. (±5V): LW Laser Temp. (±10V): Int. BB Temp A1. (±5V): Int. BB Temp A1. (±10V): Int. BB Temp A2. (±5V):	SW Det. Temp. (±10V): 303.199 LW Det. Temp. (±5V): 302.722 LW Det. Temp. (±10V): 302.717 SW Laser Temp. (±10V): 298.020 SW Laser Temp. (±10V): 303.692 LW Laser Temp. (±10V): 298.020 LW Laser Temp. (±10V): 302.794 Int. BB Temp A1. (±5V): 326.736 Int. BB Temp A2. (±5V): 326.402

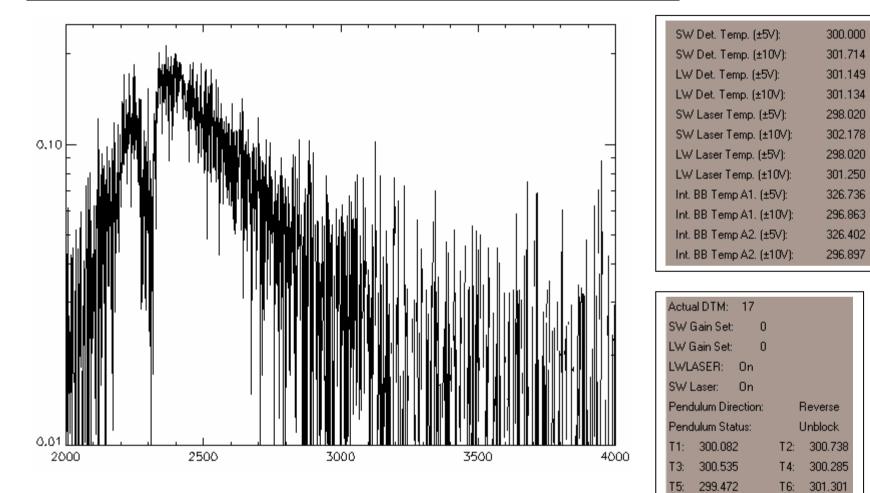
Actua	al DTM:	17				
SW 0	ain Set:		0			
LW G	iain Set:		0			
LWD	ASER:	On				
SW L	.aser:	On				
Pendulum Direction:			Reverse			
Pend	ulum Sta	tus:		ι	Jnblock	
T1:	301.473	3		T2:	302.223	
T3:	302.035	5		T 4:	301.676	
T5:	300.973	3		T6:	302.786	
T7:	301.926	5		T8:	301.176	

Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS INTERNAL NOTE Page 64	P.I. Vittorio Formisano CNR IFSI
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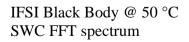
IFSI Black Body @ 100 °C SWC FFT spectrum

Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS INTERNAL NOTE Page 65	P.I. Vittorio Formisano CNR IFSI
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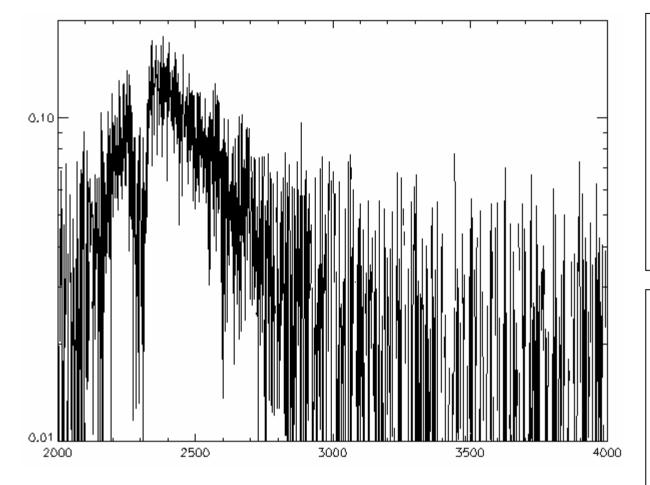


T7: 300.441

T8: 299.691



Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS INTERNAL NOTE Page 66	P.I. Vittorio Formisano CNR IFSI
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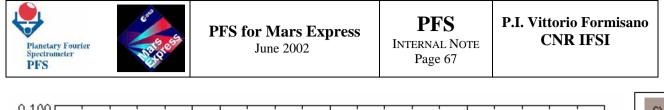


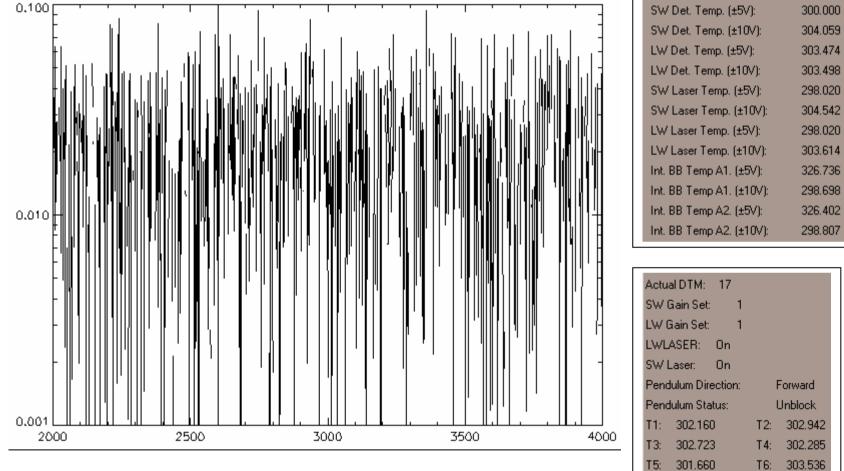
IFSI Black Body @ 40 °C SWC FFT spectrum

NOTE: Secondary motor was used. Plot -12

SW Det. Temp. (±5V):	298.212
SW Det. Temp. (±10V):	298.178
LW Det. Temp. (±5V):	297,466
LW Det. Temp. (±10V):	297.481
SW Laser Temp. (±5V):	298.020
SW Laser Temp. (±10V):	298.388
LW Laser Temp. (±5V):	297.697
LW Laser Temp. (±10V):	297.675
Int. BB Temp A1. (±5V):	326,736
Int. BB Temp A1. (±10V):	295,262
Int. BB Temp A2. (±5V):	326,402
Int. BB Temp A2. (±10V):	295.338

Actual DTM:	17				
SW Gain Set:		0			
LW Gain Set:		0			
LWLASER: 0	Dn				
SW Laser: 0	Dn				
Pendulum Direc	tior	n:	F	leverse	
Pendulum Statu	is:		L	Jnblock	
T1: 296.347			T2:	296.956	
T3: 296.737			T4:	296.722	
T5: 295.674			T6:	297.550	
T7: 296.659			T8:	295.909	





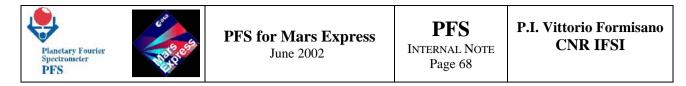
302.629

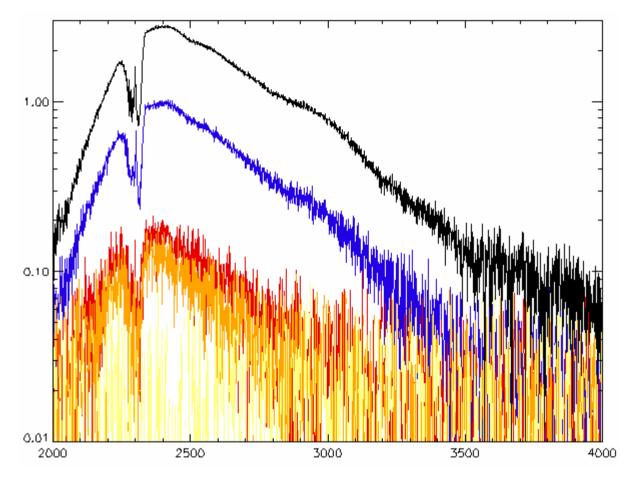
T7:

T8: 301.879

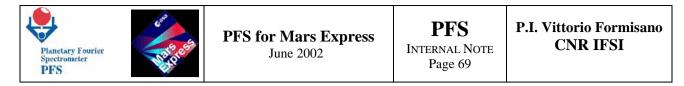
IFSI Black Body @ 30 °C SWC FFT spectrum

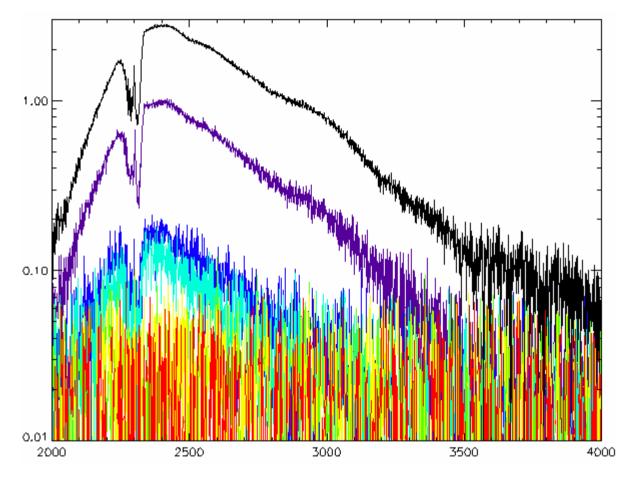
Plot-13





Black:IFSI Black Body @ 148,9 °CBlue:IFSI Black Body @ 100 °CRed:IFSI Black Body @ 50 °COrange:IFSI Black Body @ 40 °CYellow:IFSI Black Body @ 30 °CSC FFT spectraPlot – 14





Plot-15

Black: IFSI Black Body @ 148,9 °C

Violet: IFSI Black Body @ 100 °C

Blue: IFSI Black Body @ 50 °C

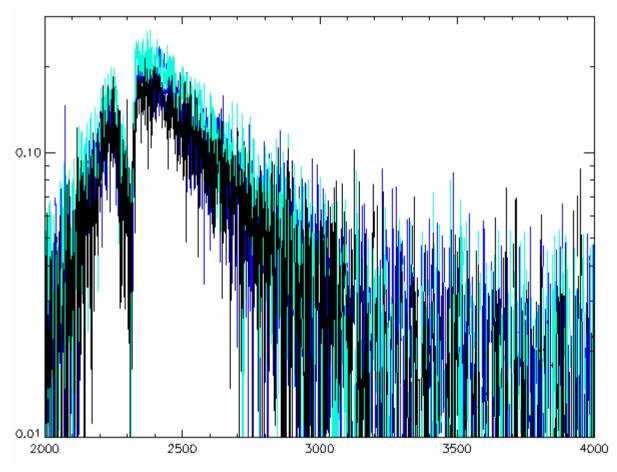
Cyan: IFSI Black Body @ 40 °C

Green: IFSI Black Body @ 30 °C

Red: Internal Black Body (**Primary Motor** – acquired just after IFSIBB @ 50 °C)

Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 70	P.I. Vittorio Formisano CNR IFSI
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Yellow: Internal Black Body (Secondary Motor - acquired just after IFSIBB @ 148,9 °C)

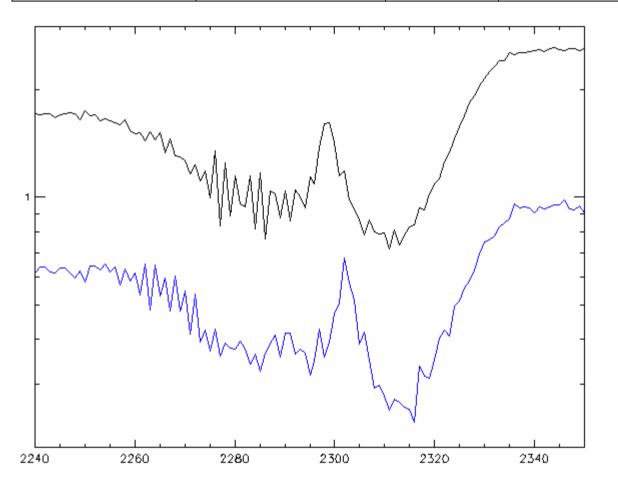




IFSI Black Body @ 50°C (secondary motor) IFSI Black Body @ 50,5°C IFSI Black Body @ 50°C Cyan:

- Blue:
- Black:

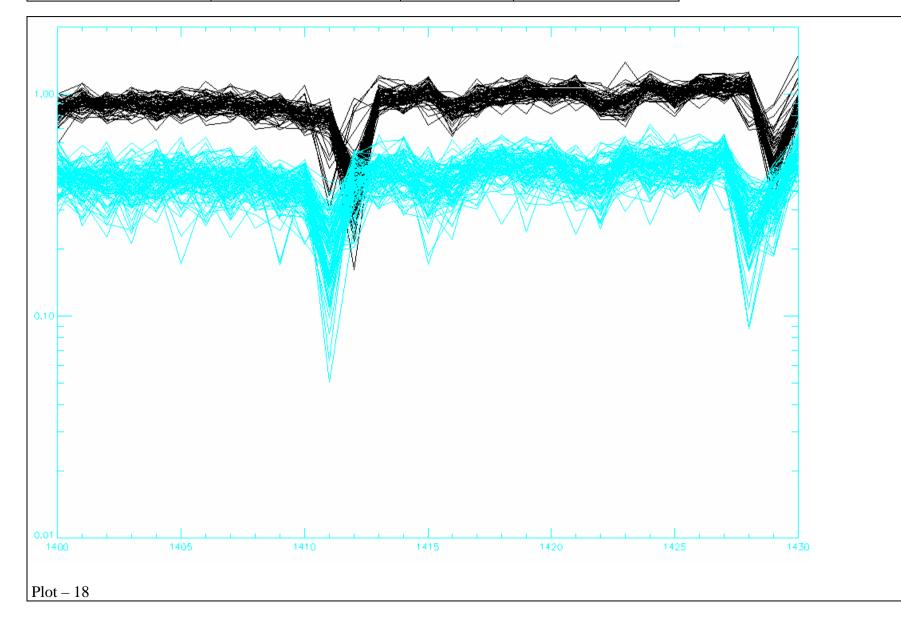
Planetary Fourier Spectrometer PFS PFS PFS PFS PFS PFS PFS PFS PFS PFS	PFS INTERNAL NOTE Page 71	P.I. Vittorio Formisano CNR IFSI
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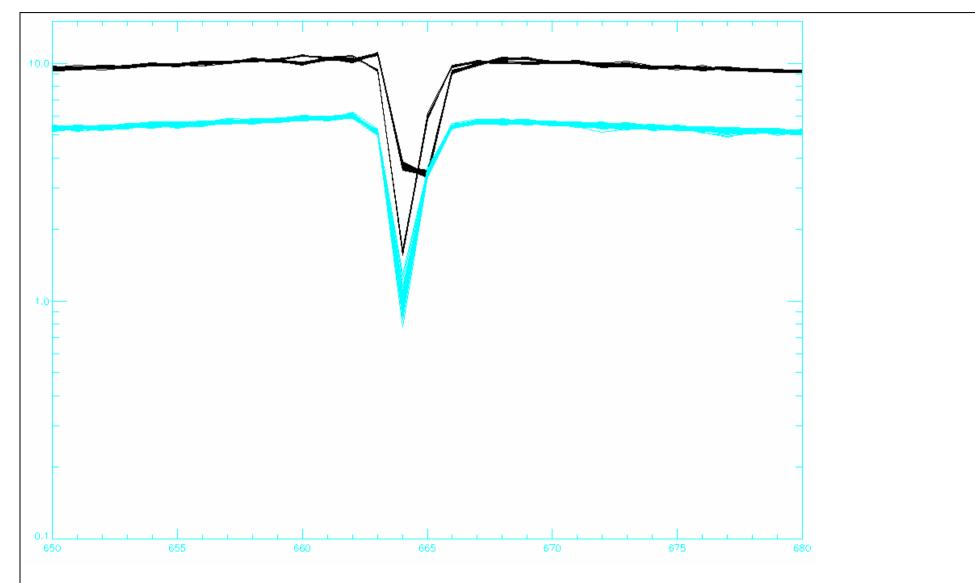
 $2240 \div 2350 \text{ cm}^{-1}$ Short Wavelength Channel Blow-up

Black:IFSI Black Body @ 148,9 °CBlue:IFSI Black Body @ 100 °CPlot - 17

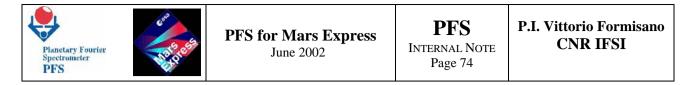
Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 72	P.I. Vittorio Formisano CNR IFSI
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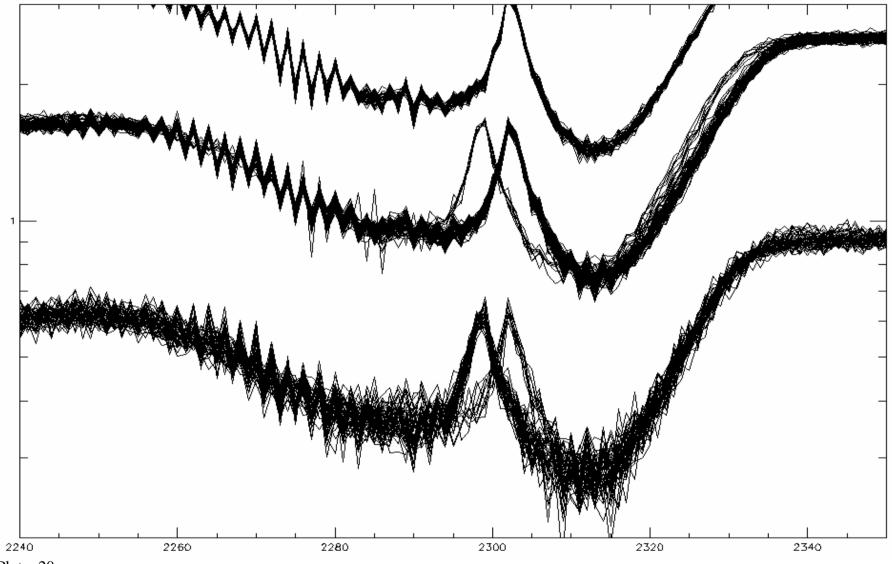


Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 73	P.I. Vittorio Formisano CNR IFSI
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Plot – 19









2.3 RUSSIAN BLACK BODY

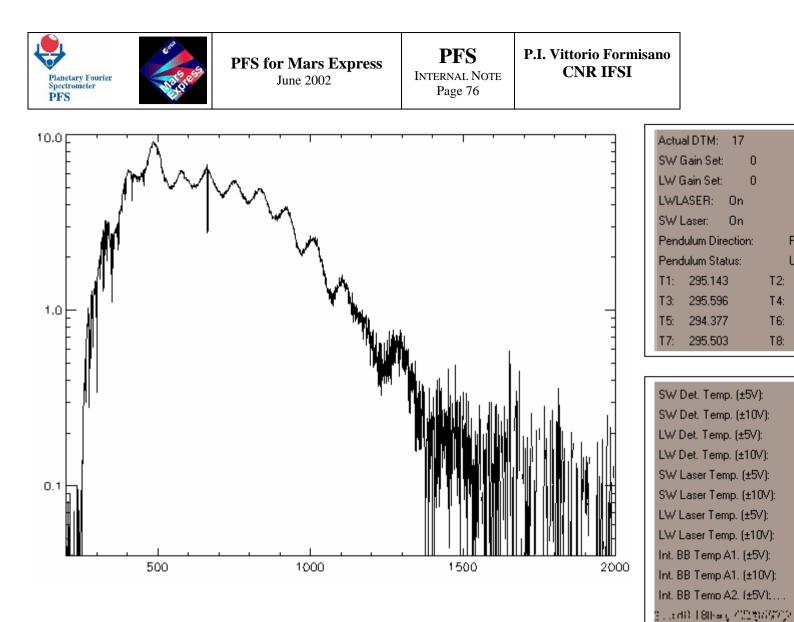
We now turn to some raw spectra of Russian BB. Temperatures of -100, -60, -30 deg were measured. Note that the spectra with -30 deg (240 Kelvin) should be similar to what we shall see at Mars in terms of radiance and DT. The spectra do not show any important difference with the similar spectra taken with the IFSI BB. See plots 21,22,23, 24, 25, 26, 27.

2.4 GOSTS FEATURES

As it can be seen in the next figure, (see plot 28,29,30,31, 32, 33, 34, 35) when we look at a 149 deg IFSI BB, the spectrum does not stop at 4000 cm -1, but goes up to 6000 with the shape of a strange square wave. This radiation depends on the source intensity, as it disappears when the IFSI BB temperature goes to 50 celsius. The next figure shows that there are some spectral features which are strange and need to be understood. The two following figures show that the first half of the platueaughost can be explained with a 4 % ghost – repetition of the main spectrum at x2 resolution and x2 wavenumber. As in the case of IRIS Mariner 9 studied in the past we think that his ghost is due to internal reflections: we have tilted the entrance silicon window, but we have not considered that the detector itself has a saffire window which can reflect. The present ghost, however, is not similar to the one of IRIS, and the rest of the plateau is not explained. It has been suggested that the introduction of a third spectrum with opposite wavenumber scale can explain the observations. We note that the absorption band that could be another ghost of the CO2 band, and that is observed before 6000 cm-1, does not have the central peak, therefore it is still to be demonstrated and understood the cause of the plateau observed in the SW channel when the IFSI BB was observed at 150 celsius. A very important point is to keep in mind that the shape of the CO2 ghost band is not identical to the main one, as a peak appears at 2330 cm-1 where the second minimum of the band should be. The summation of the absorption due to the CO2 ghost with some other spectrum should not give a peak, unless there is emission : at the moment, therefore the entire SW spectrum between 4000 and 8000 cm-1 is not understood. Finally it is worth noting that the absorption band seen at the second edge of the plateau, does not look like a CO2 absorption band, so is not related to the ghost. We have made the exercise to subtract the ghost, computed assuming 3% of the original spectrum, at 2x resolution and 2x wavenumber. The result, shown in the figure, illustrates that other bands may be present, but for the moment it is not clear that what is left can be considered another spectrum with reversed wavenumber scale.

Note also that the curve subtracted and computed as 3 % of the spectrum, because decreases the noise level at 6300 cm-1, should be considered an upper limit to the ghost intensity, a value 2 or 2.5 being a better estimate.

We have searched for the ghost in other measurements, to be sure it is present and it is of the intensity (4%) shown above.



Forward

Unblock

294.096

296.190

294.247

294.212 295.943

295.938 297.307

297.275

295,900 295.888

326.736

286.749 326 402

30905

T2: 295.487

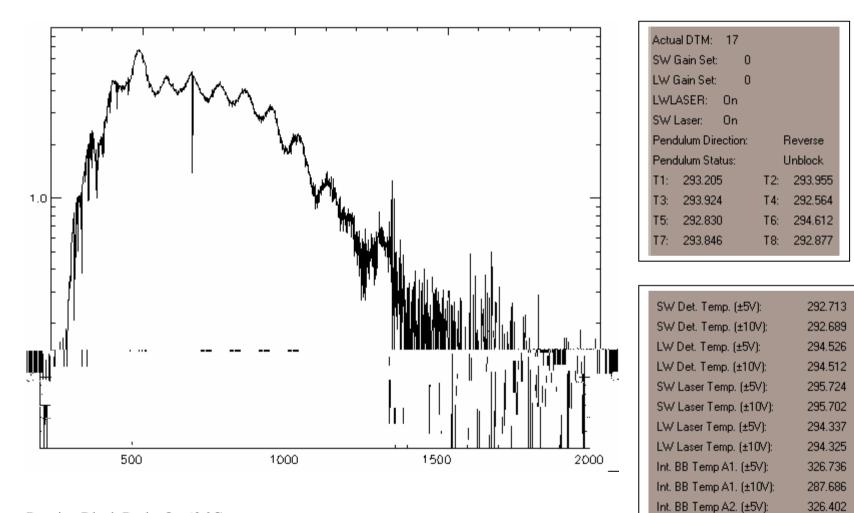
T8: 294.377

T4:

T6:

Russian Black Body @ -100 °C

Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 77	P.I. Vittorio Formisano CNR IFSI
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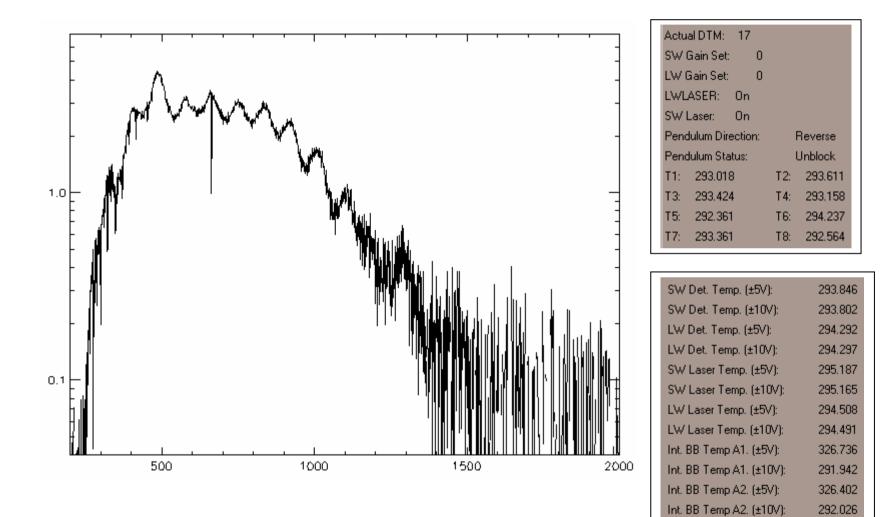
Int. BB Temp A2. (±10V):

287.738





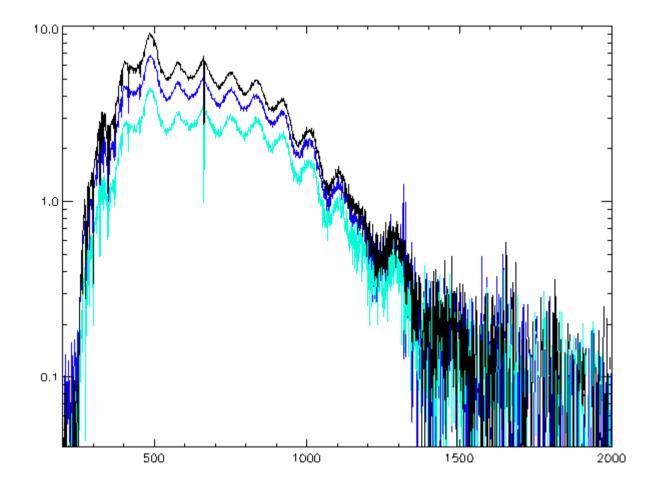
Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 78	P.I. Vittorio Formisano CNR IFSI
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Russian Black Body @ -30 °C

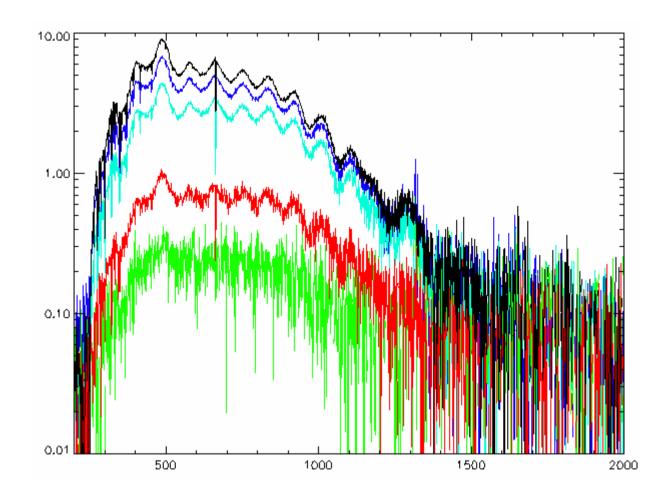
Plot - 23

Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 79	P.I. Vittorio Formisano CNR IFSI
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Black:Russian Black Body @ -100°CBlue:Russian Black Body @ -60°CCyan:Russian Black Body @ -30°CPlot - 24Plot - 24

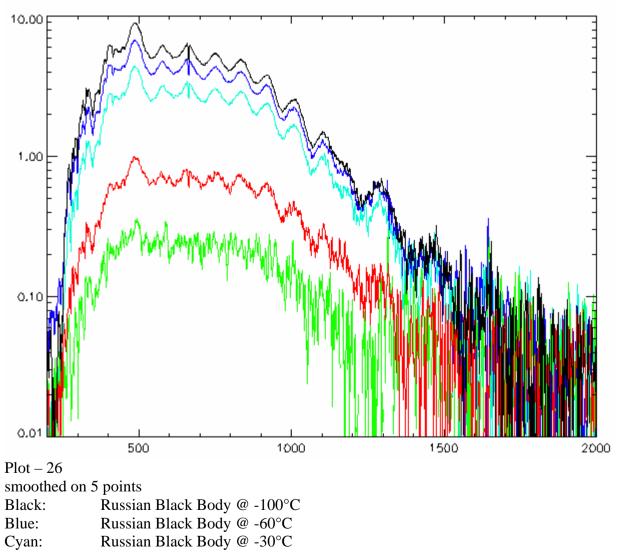
Planetary Fourier Spectrometer PFS Planetary Fourier Spectrometer PFS	oress PFS INTERNAL NOTE Page 80	P.I. Vittorio Formisano CNR IFSI
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Plot-25

- Black: Russian Black Body @ -100°C
- Blue: Russian Black Body @ -60°C
- Cyan: Russian Black Body @ -30°C
- Red: Russian Black Body (acquired just after Russian BB RAMP from -80 °C to -55 °C)
- Green: Russian Black Body (acquired just after Russian BB @ -30 °C)

Planetary Fourier Spectrometer PFS	for Mars Express June 2002	PFS INTERNAL NOTE Page 81	P.I. Vittorio Formisano CNR IFSI
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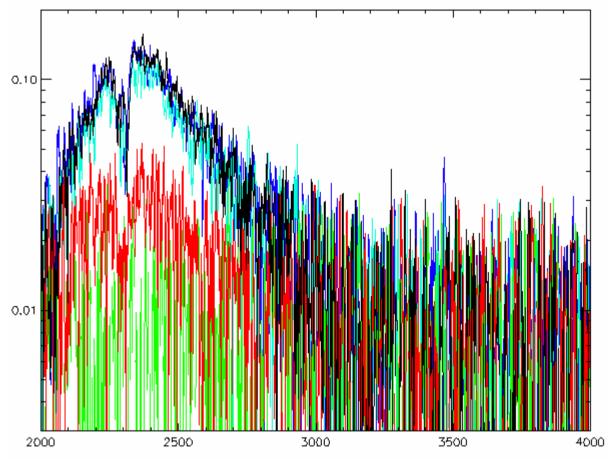
Red: Internal Black Body (acquired just after Russian BB RAMP from -80	$^{\circ}$ C to $-55 ^{\circ}$ C)
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Green: Internal Black Body (acquired just after Russian BB @ -30 °C)

Int. BB after Ramp				
SW Det. Temp. (±5V):	291.219			
SW Det. Temp. (±10V):	291.184			
LW Det. Temp. (±5V):	292.563			
LW Det. Temp. (±10V):	292.538			
SW Laser Temp. (±5V):	293.575			
SW Laser Temp. (±10V):	293,553			
LW Laser Temp. (±5V):	292.286			
LW Laser Temp. (±10V):	292.283			
Int. BB Temp A1. (±5V):	316.075			
Int. BB Temp A1. (±10V):	281.282			
Int. BB Temp A2. (±5V):	316,151			
Int. BB Temp A2. (±10V):	281.346			

Int. BB after –30) °C	
SW Det. Temp. (±5V):	293.368	
SW Det. Temp. (±10V):	293.333	
LW Det. Temp. (±5V):	294.468	
LW Det. Temp. (±10V):	294.453	
SW Laser Temp. (±5V):	295.607	
SW Laser Temp. (±10V):	295.575	
LW Laser Temp. (±5V):	294.425	
LW Laser Temp. (±10V):	294.413	
Int. BB Temp A1. (±5V):	326.736	
Int. BB Temp A1. (±10V):	289.951	
Int. BB Temp A2. (±5V):	326,402	
Int. BB Temp A2. (±10V):	289.999	

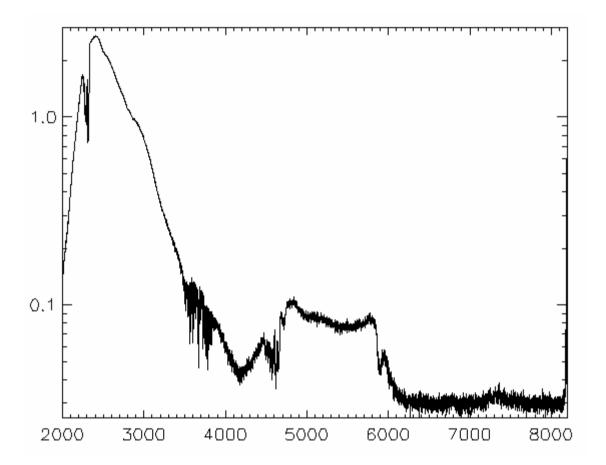
Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 82	P.I. Vittorio Formisano CNR IFSI
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Plot - 27. SMOOTHED ON 5 POINTS

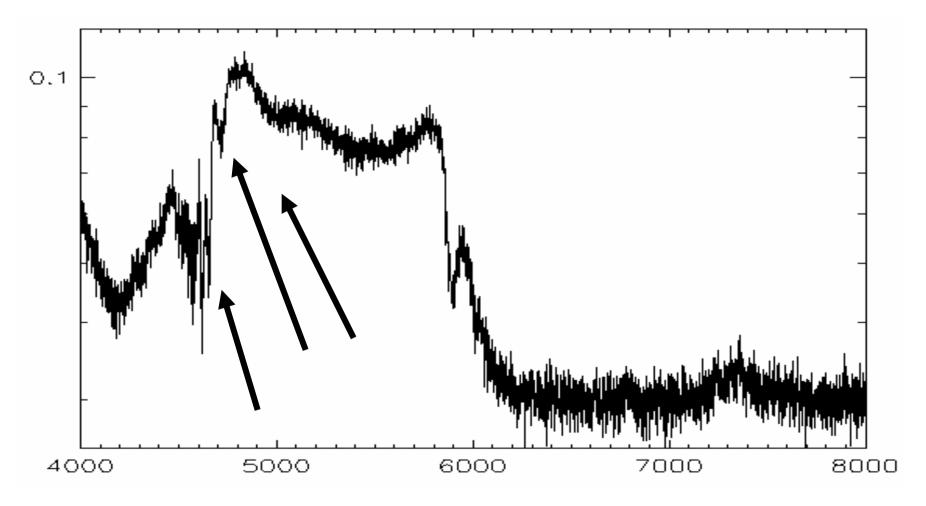
- Black: Russian Black Body @ -100 °C
- Blue: Russian Black Body @ -60 °C
- Cyan: Russian Black Body @ -30 °C
- Red: Internal Black Body (acquired just after Russian BB RAMP from -80 °C to -55 °C)
- Green: Internal Black Body (acquired just after Russian BB @ -30 °C)

Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 83	P.I. Vittorio Formisano CNR IFSI
--	--	--	-------------------------------------



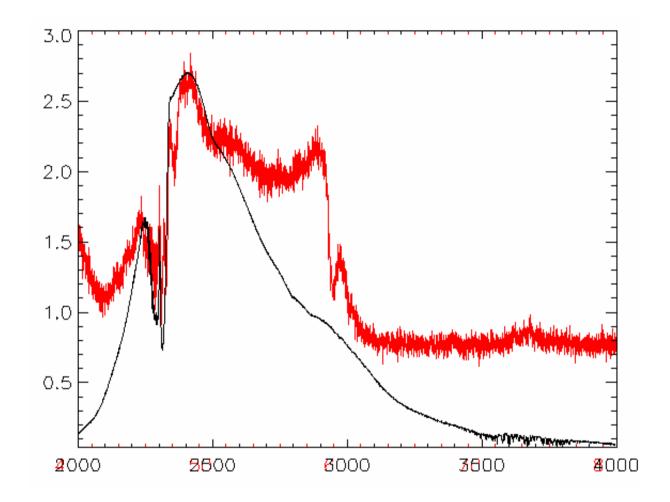
IFSI BB @ 148.9 °C - Average on 95 spectra SW spectrum in the entire wavenumber range : some extra radiation is present. Plot – 28

Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 84	P.I. Vittorio Formisano CNR IFSI
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Plot – 29 IFSI BB @ 148.9 °C - Average on 95 spectra Blow-up within 4000÷8000 cm⁻¹ Some real radiation enable to see not only the ghost of the CO2 4.3 micron band, but also the 3 two micron bands indicated by arrows.

Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 85	P.I. Vittorio Formisano CNR IFSI
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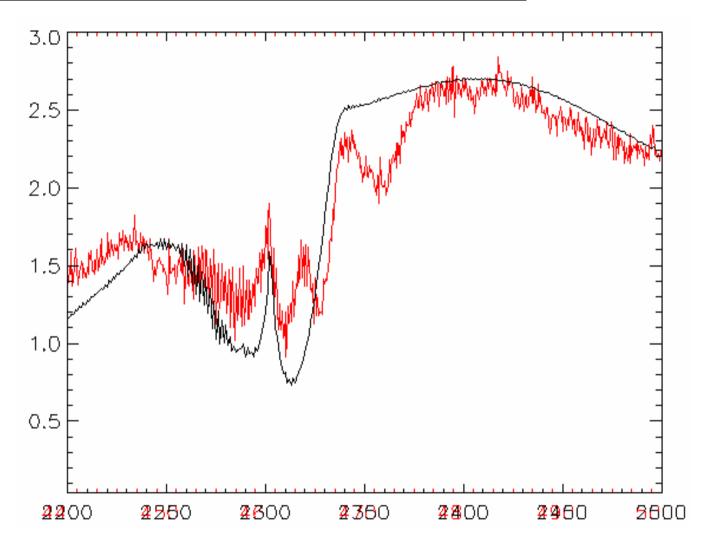




Black:IFSI BB @ 148.9 °C within the range $2000 \div 4000 \text{ cm}^{-1}$ Red:IFSI BB @ 148.9 °C within the range $4000 \div 8000 \text{ cm}^{-1}$

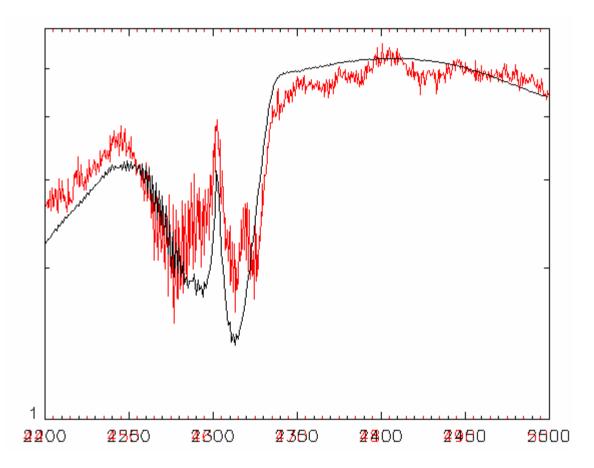
The red line was multiplied by 25.7277 (see previous page for details).

Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 86	P.I. Vittorio Formisano CNR IFSI
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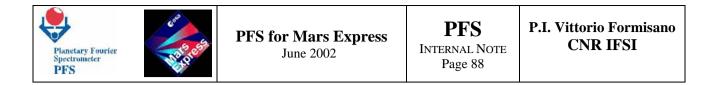


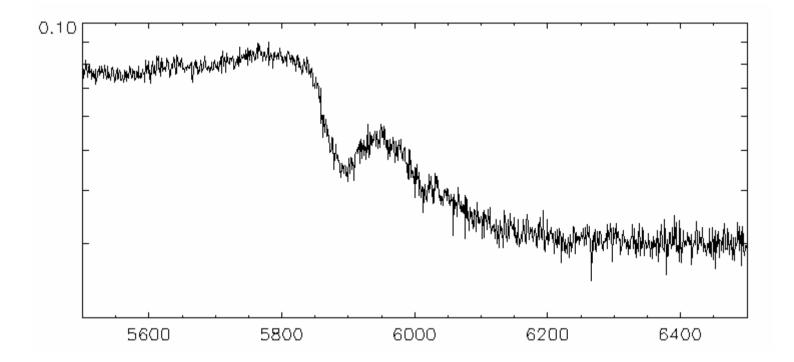
Same as fig. 6; Blow-up within the range $2200 \div 2500 \text{ cm}^{-1}$ (black curve) and $2000 \div 4000 \text{ cm}^{-1}$ (red curve). The intensity of the ghost is then about the 3.9% of the intensity of the main spectrum. Plot – 31

S 65	Mars Express ne 2002 PFS INTERNAL NOTE Page 87	P.I. Vittorio Formisano CNR IFSI
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Same as fig. 6, but for the IFSI BB @ 148.9 °C – Gain 1 The red line was multiplied by 27.1418 Plot – 32

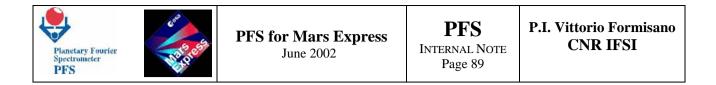


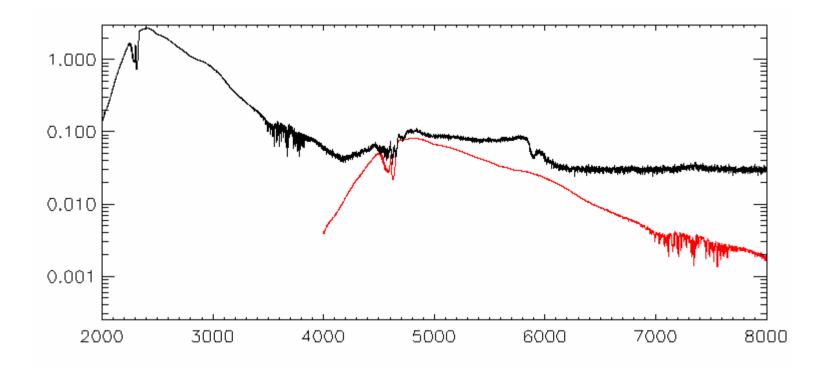


Zoom of the "reversed" ghost

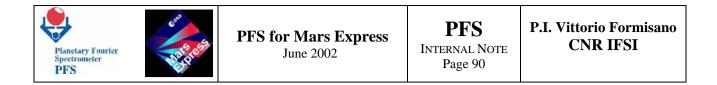
The central line is absent.

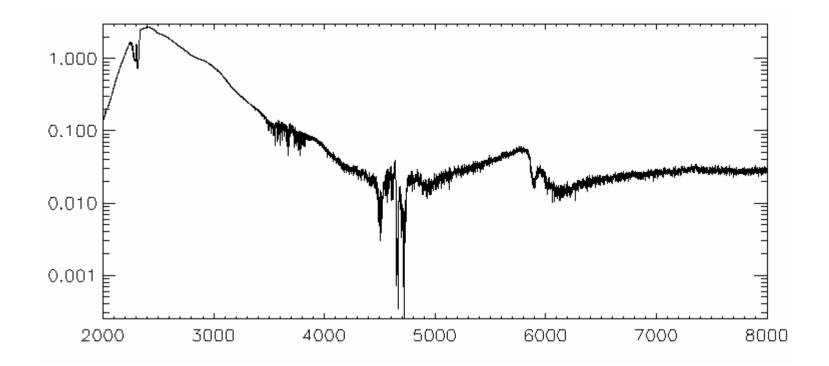
Plot - 33











Plot - 35



2.5 ORIEL AND CERAMIC ELEMENT LAMP.

We report now on SW spectra taken with the ceramic element lamp and with the Oriel lamp(See Plot 36, 37, 38, 39, 40, 41, 42. The last one should have a spectrum more intense in the short wavelenght range. We start first showing single spectrum, and its blow up around the CO2 absorption band, than we make averages. On the average over 25 spectra we see that the CO2 band may have a ghost 2x, therefore we again blow up and overplot the ghost to see how it matchs the main band. The ghost is not very clear, being noisy, but it is certainly there.

When looking at the Oriel lamp we discover a new effect : the interferogram is NOT SATURATED, but we see a very strong signal below 2000 cm-1. This effect has been interpreted occasionally as saturation of the detector without saturation of the interferogram. Alternatively can be interpreted as strong non linearity of the detector. We shall study in more details later this effect.

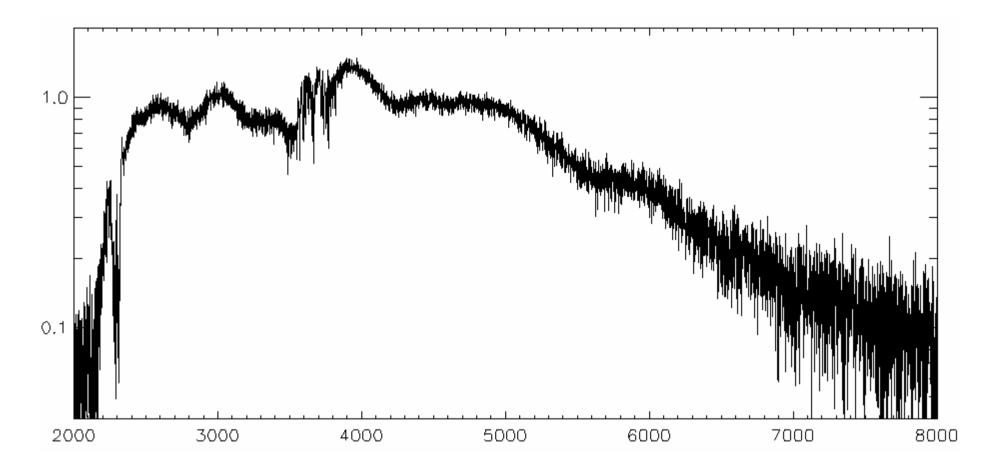
2.6 FORWARD AND REVERSE MOTION.

Plot 43, 44, 45, 46, 47 are shown to compare the raw data from forward and reverse motion. In the cases considered here seems to be essentially identical spectrum for the measurements taken from forward and for reverse motion. This indeed should be the case, but we have seen occasionally important differences, specially when the movement was not OK. All the average spectra shown for both SW and LW channel demonstrated that the direction of motion is not important for the measurement of the spectrum, however we know that the interferogram shows that the central part is displaced on one side or another depending on the motion direction.

As the SW spectra shown are from BB observations, it is important to look at some other source to confirm that the forward-reverse motion gives the same spectrum. We have considered the integrating sphere data. The following figures show the SW spectra for different source intensity. We see that the reverse motion shows a spectrum that does not change much the intensity close to 8000 cm -1, while the forward spectrum shows an intensity change proportional to the integrating sphere input (up to 5). The conclusion is that the reverse motion may have a different speed, therefore the electronic filter acts more strongly to decrease the signal. We have essentially two different instruments, one forward, one reverse motion : responsivity and ner should be computed separately for the 2. The effect in theory should be visible also in the LW channel, but there the signal in the range 1800-2000 cm-1 is absent, and the filter shape may also be different : a fact is that we are not able to see a substantial difference in the LW forward or reverse motion direction, while in the SW channel a strong effect appears in the range 7500- 8200 cm-1. On the cause of this effect we should say the following :

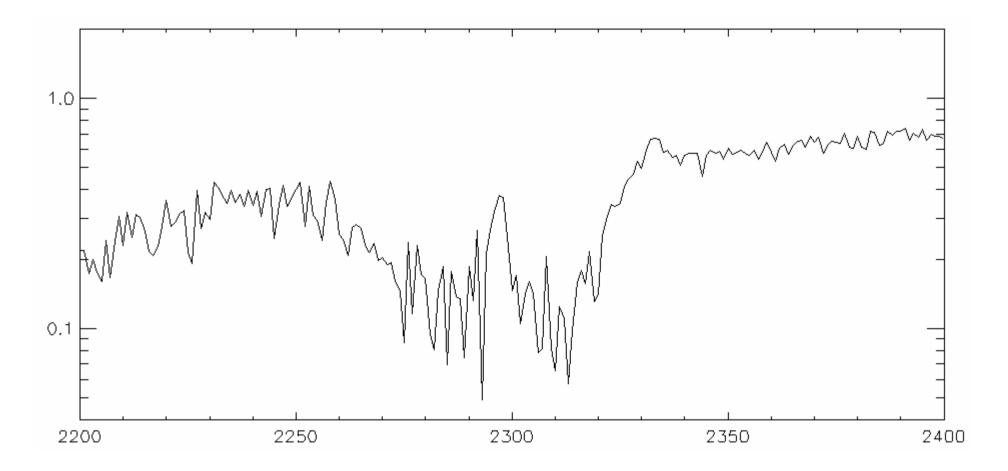
The double pendulum has a motor (with an identical spare for redundancy) which is used, no to the full possible power, but to the almost minimum power. On January 2, 2002 the power (current) was increased because motion was not OK after the inclusion of two large inductances in Module O. Also with the new current level the motion in all the laboratory measurements can be and it is affected by the presence of the gravity. As the Module O is never perfectly horizonthal, the motion forward (or reverse) can be against gravity, while the opposite is with gravity in favour. The motion with gravity in favour has obviously higher speed and therefore will show the effect of the electronic filter cutting the signal. The opposite motion will not show the effect. As in the Lab we have worked occasionally with PFS upside down (specially in early January - 2/5 jan) we may have data with reverse motion showing the filter cutting, other data may have the forward motion with the filter cutting, if PFS was upside down. It is important to note that in space this difference forward reverse should not exists.

Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 92	P.I. Vittorio Formisano CNR IFSI
--	--	--	-------------------------------------



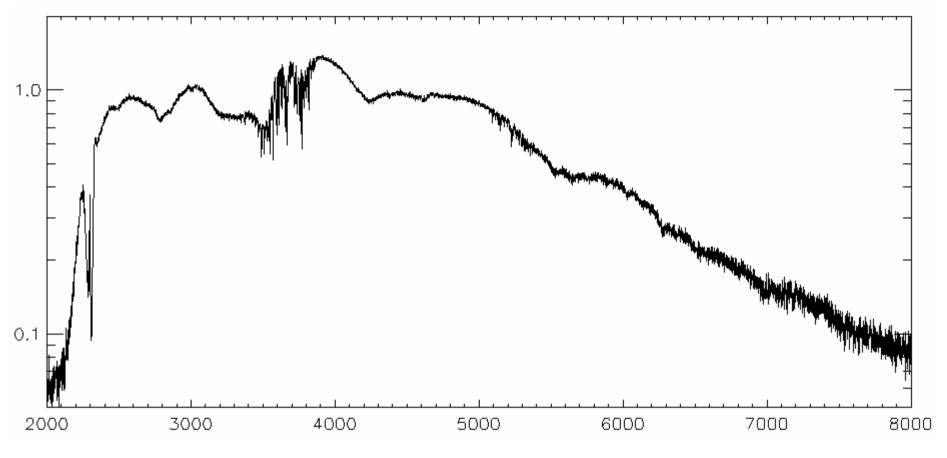
Ceramic Element – Single FFT Spectra – SWC Plot – 36

Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 93	P.I. Vittorio Formisano CNR IFSI
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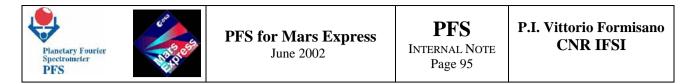


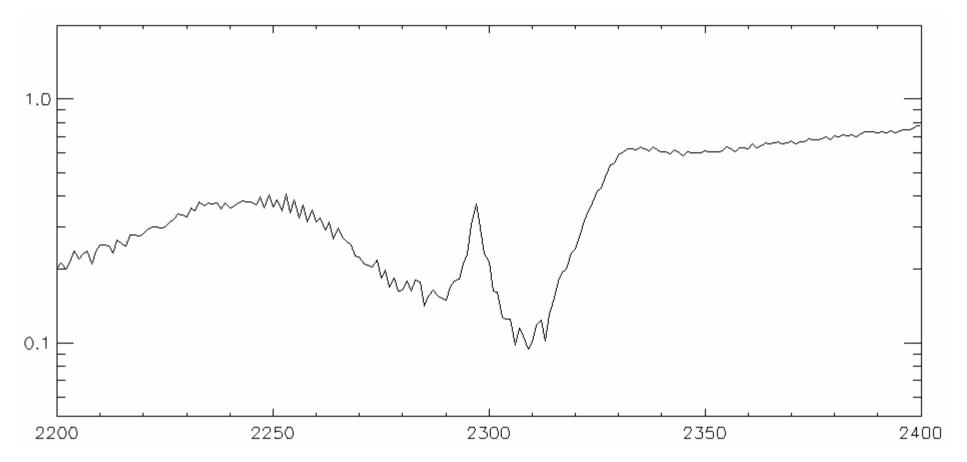
 $Ceramic \ Element - Single \ FFT \ Spectra - SWC - Blow-Up \ of \ the \ CO_2 \ Absorption \ Bands \ Plot - 37$

Planetary Fourier Spectrometer PFS PFS	for Mars Express June 2002	PFS Internal Note Page 94	P.I. Vittorio Formisano CNR IFSI
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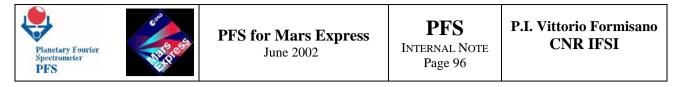
Plot – 38 Ceramic Element – Average on 25 FFT Spectra – SWC

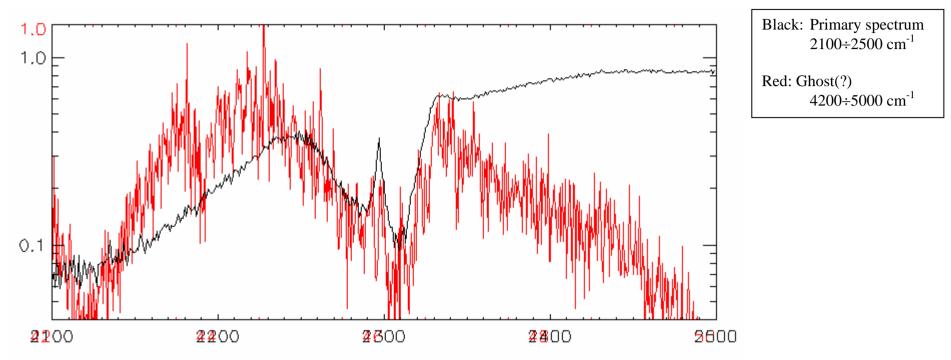




Ceramic Element – Average on 25 FFT Spectra – SWC - Blow-Up of the CO₂ Absorption Bands

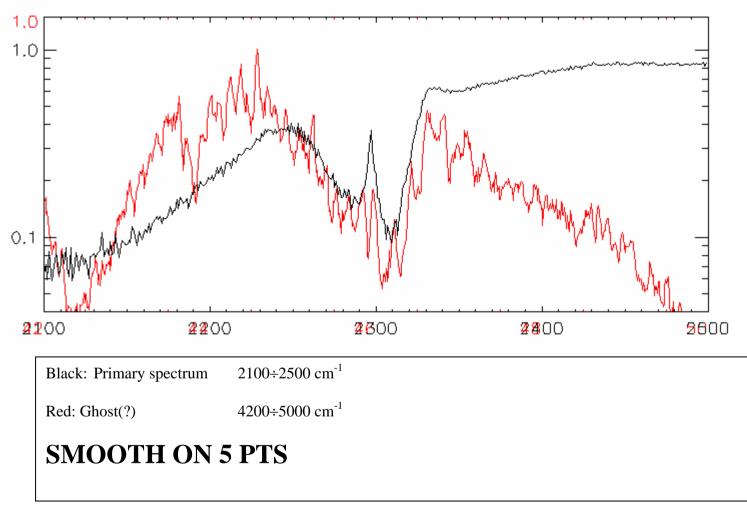
Plot – 39



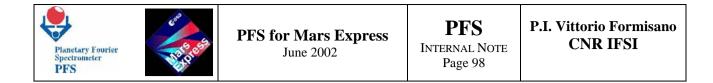


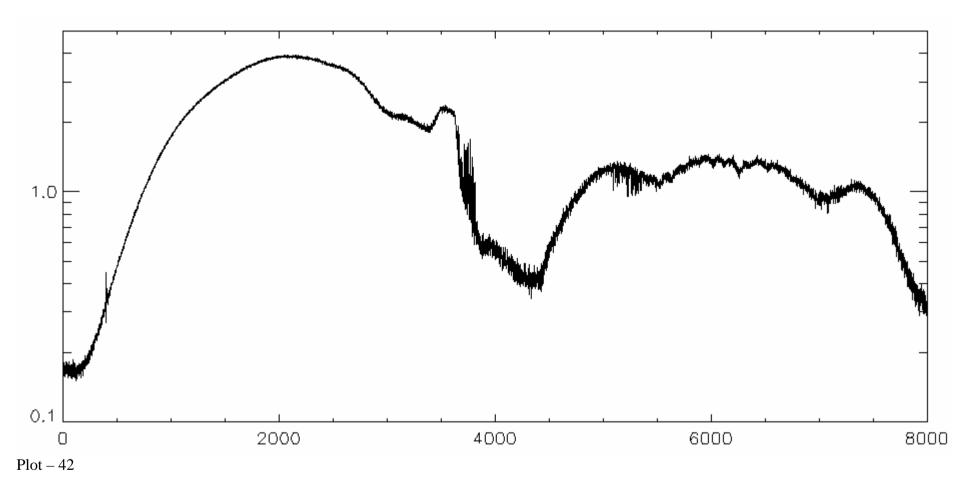


Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 97	P.I. Vittorio Formisano CNR IFSI
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Plot-41

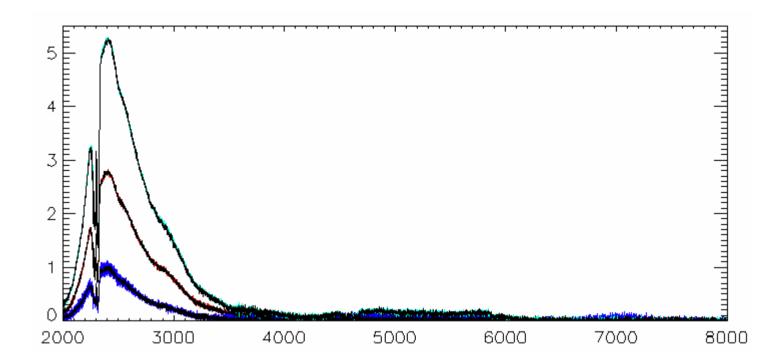




Oriel Lamp – Average on 25 FFT spectra – SWC

NOTE: Something went obviously wrong! We have high signals even under 1000cm⁻¹

Planetary Fourier Spectrometer PFS	FS for Mars Express June 2002	PFS Internal Note Page 99	P.I. Vittorio Formisano CNR IFSI
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Blue: Forward direction IFSI BB @ 100 °C

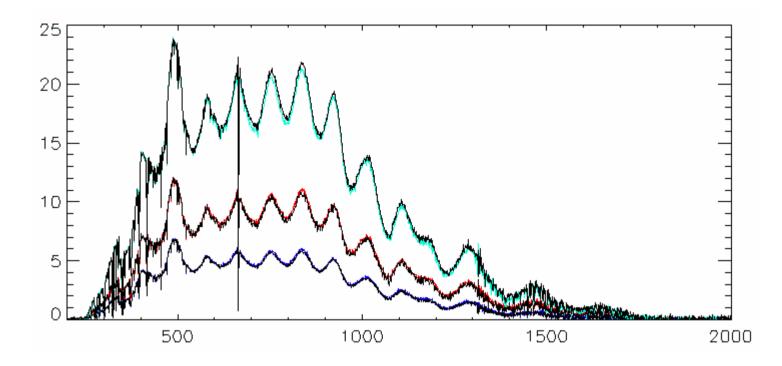
Red: Forward direction IFSI BB @ 148.9 °C

Cyan: Forward direction IFSI BB @ 148.9 °C – Gain 1

SWC

Plot- 43

Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 100	P.I. Vittorio Formisano CNR IFSI
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Black: Reve	erse directions
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Blue:	Forward direction	IFSI BB @ 100 °C	
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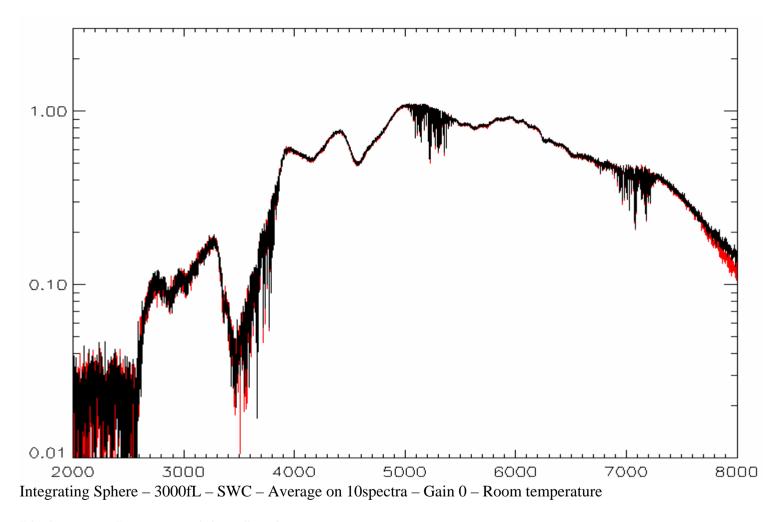
Red: Forward direction IFSI BB @ 148.9 °C

Cyan: Forward direction IFSI BB @ 148.9 °C – Gain 1

LWC

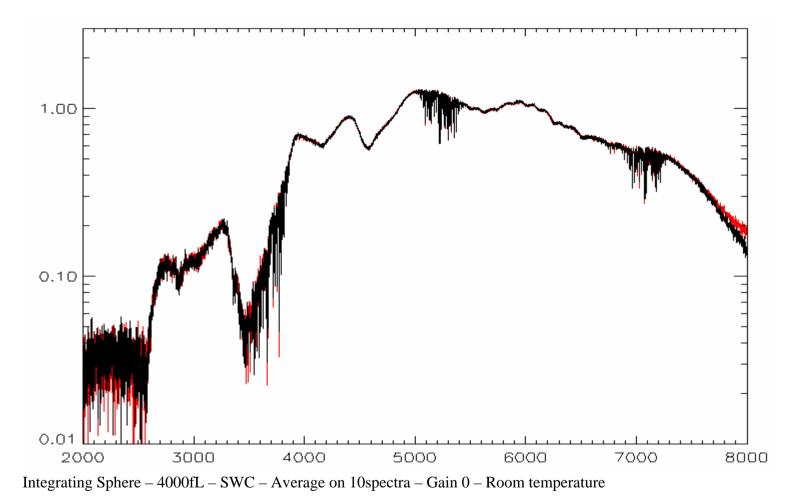
Plot - 44

Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 101	P.I. Vittorio Formisano CNR IFSI
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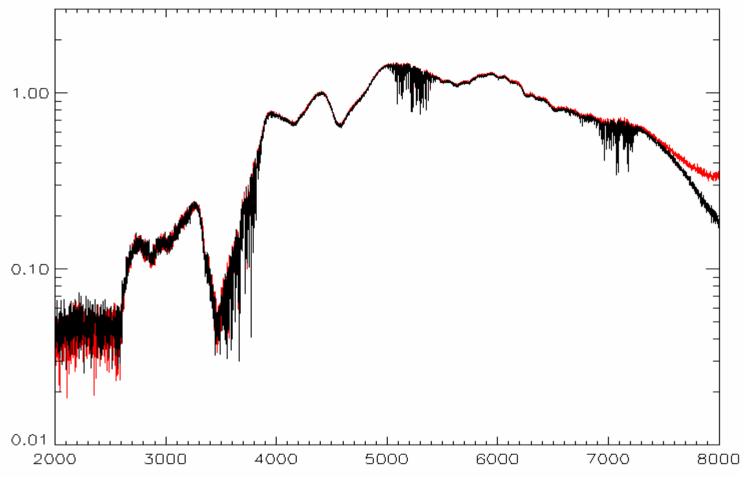
Black:Reverse pendulum directionRed:Forward pendulum directionPlot - 45Forward pendulum direction

Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 102	P.I. Vittorio Formisano CNR IFSI
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Black:Reverse pendulum directionRed:Forward pendulum directionPlot - 46Forward pendulum direction

Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 103	P.I. Vittorio Formisano CNR IFSI
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Integrating Sphere – 5000fL – SWC – Average on 10spectra – Gain 0 – Room temperature

Black:Reverse pendulum directionRed:Forward pendulum directionPlot - 47



2.7 INTEGRATING SPHERE

In the following we show measured spectra taken using the integrating sphere (See plot 48, 49, 50, 51). We show 2 figures in each case : spectra average over 10 measurements (all reverse motion) for different gain factors first. The same figure is then shown with the spectra smoothed. The luminosity was fixed at 100 fotlam. Note that in the last measurements (gain 7) probably we have saturation, or we are very close to saturation, and the spectrum is deformed. The integrating sphere should not give signal below 4000 cm-1. We do see something but we do not understand the cause of that signal. The spectra show that indeed there is a factor 2 increase in gain going from gain 0 to 1,2,3,4,5,6,7 (each step a factor 2). It is interesting to note that 100 Fl are equivalent to 1-2 ergs/s sr.s cm-1, like at Mars. At Mars, therefore we may have to use gain 4 to get good signal to noise ratio without saturation.

After that we show the gain 0 measurements for different luminosity, from 10 to 5000ftl. We note here that in the region 7500 - 8000 cm-1 the spectral shape may depend on the fact that the motion of the double pendulum was against gravity or not.

In both cases we not that the spectra can span over 2-3 decades.

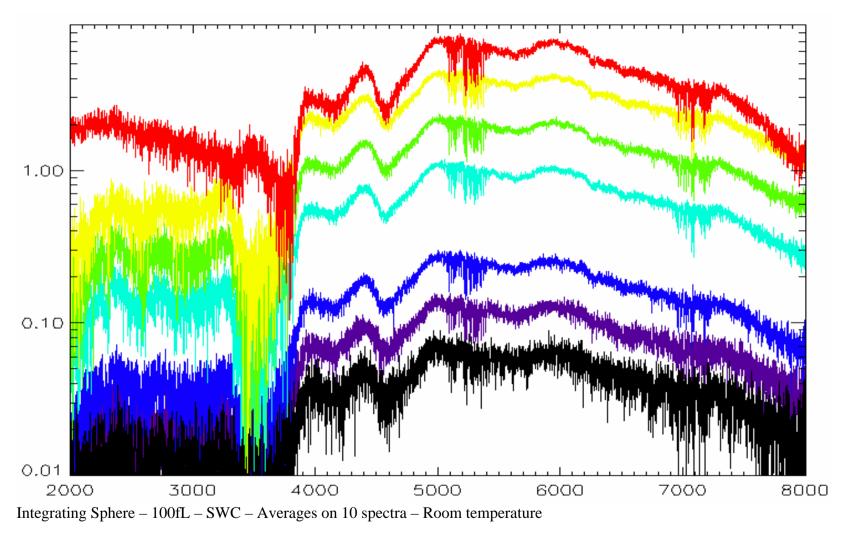
We show now some strange spectra which indicate the presence of other instrumental effects. See plot 52, 53, 54, 55, 56, 57, 58. The next figure shows a spectrum average over 5, and smoothed over 10 points. The spectrum was taken in the Thermovac, with the Sw detector cooled at -37 celsius. PFS was looking at the integrating sphere out of the window. The signal from the sphere was very small and the radiance from the window was dominating. Gain 4 was used and the interferogram was not saturated, being 23700 at max intensity. A strong signal between 1000 and 2000 is observed. A priory no signal is detectable between 0 and 2000 cm-1. There is only one way to generate a false signal below 2000 cm-1, and this is by means of strong distortion of the central part of the interferogram, which is the part that controls the low frequencies of the spectrum. This effect can be explained as a non linearity of the detector for large signals. That this is the case is demonstrated also by the following figure : we are now looking at the internal calibration lamp at gain zero and with the detector at decreasing temperatures. We see that the spectral intensity increases as the responsivity increases, as expected. What is not expected is the increase of the signal below 2000 cm-1, which becomes non negligible at – 46 Celsius.

We shall in the following study the linearity of the detector and the increase of the responsivity at decreasing detector temperature. Here we simply conclude that the detector may be non linear for intense signals, and this effect may deform the shape of the spectrum.

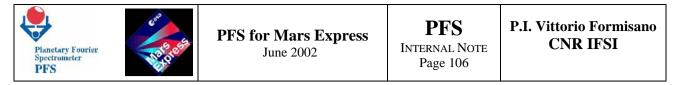
2.8 INTERNAL LAMP

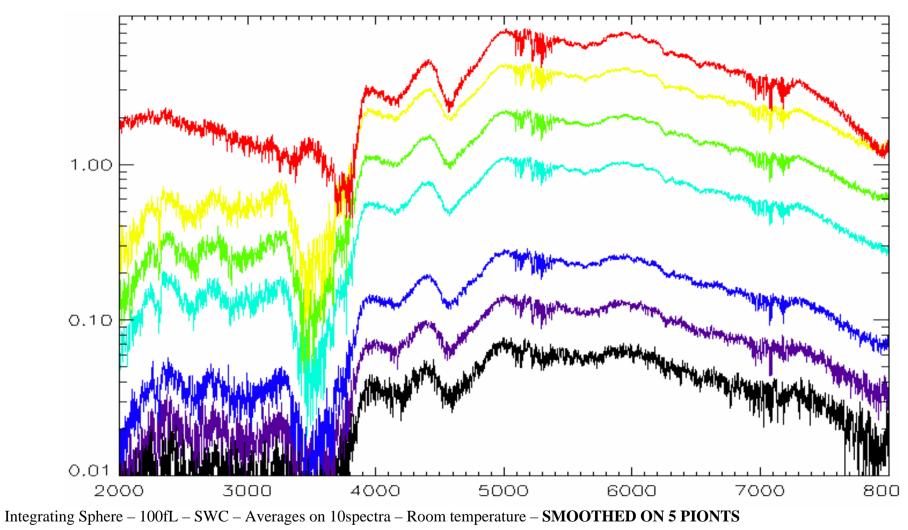
In the following figures (see plot 59, 60, 61, 62, 63) we show other spectra of the internal lamp. The main point to be stressed here is that the internal lamp can be observed at room temperature only with gain 0 and 1. at all the other gains the interferogram is saturated in spite of the current reduction from 16 to 10 ma. It is obvious that when we go to cooled SW detector, the responsivity increases and the interferogram for the internal calibration lamp can be saturated for all the gain factors available. This probably means that in space we will not have any calibration source for the SW channel.

Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 105	P.I. Vittorio Formisano CNR IFSI
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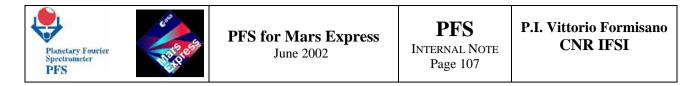


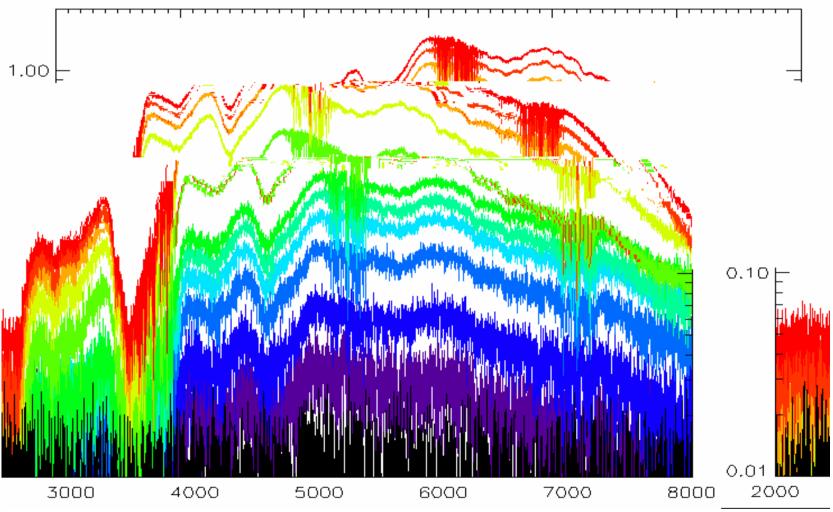
Black:	Gain 0	Cyan:	Gain 4	Red:	Gain 7
Violet:	Gain 1	Green:	Gain 5		
Blue:	Gain 2	Yellow:	Gain 6		
Plot - 48					





Black:	Gain 0	Cyan:	Gain 4	Red:	Gain 7
Violet:	Gain 1	Green:	Gain 5		
Blue:	Gain 2	Yellow:	Gain 6		
Plot - 49					



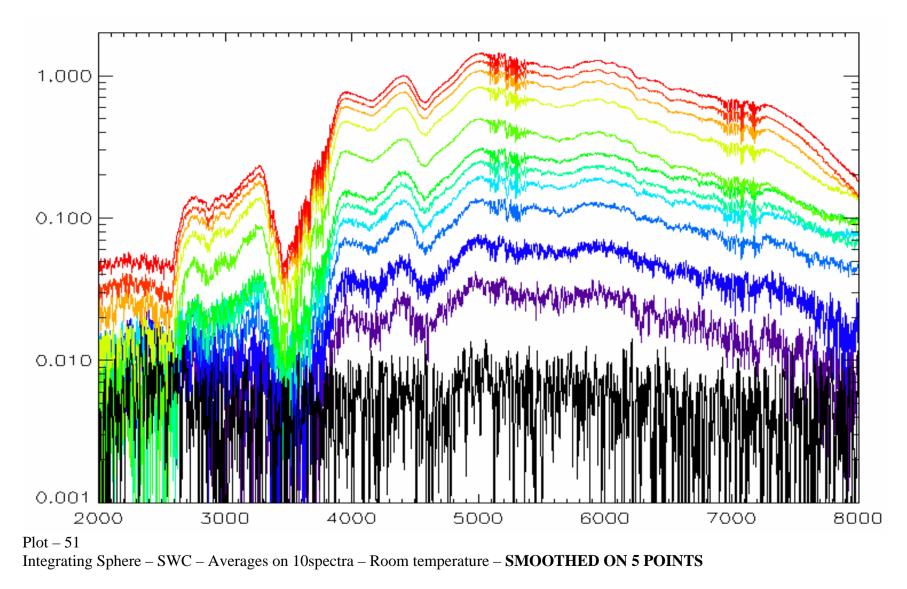


Plot - 50

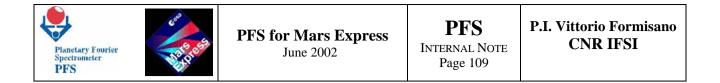
Integrating Sphere – SWC – Averages on 10spectra – Room temperature

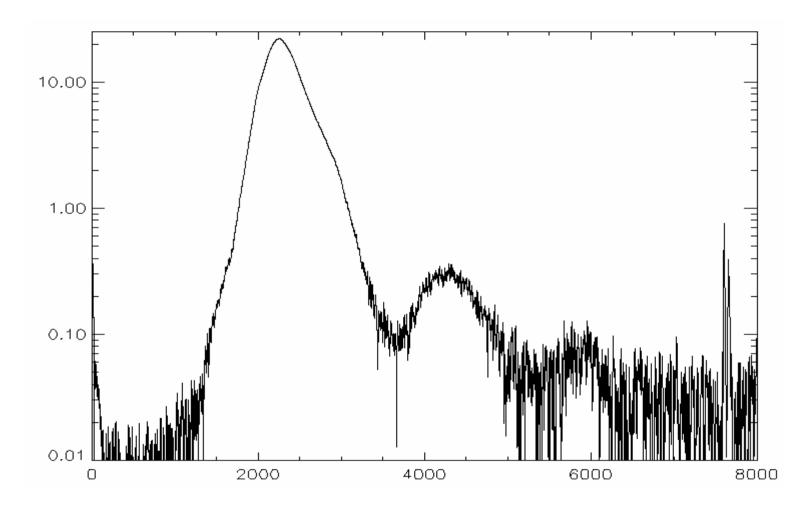
From bottom to top: 10fL, 50fL, 100fL, 200fL, 300fL, 400fL, 500fL, 1000fL, 2000fL, 3000fL, 4000fL, 5000f

Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 108	P.I. Vittorio Formisano CNR IFSI
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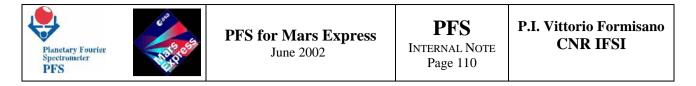


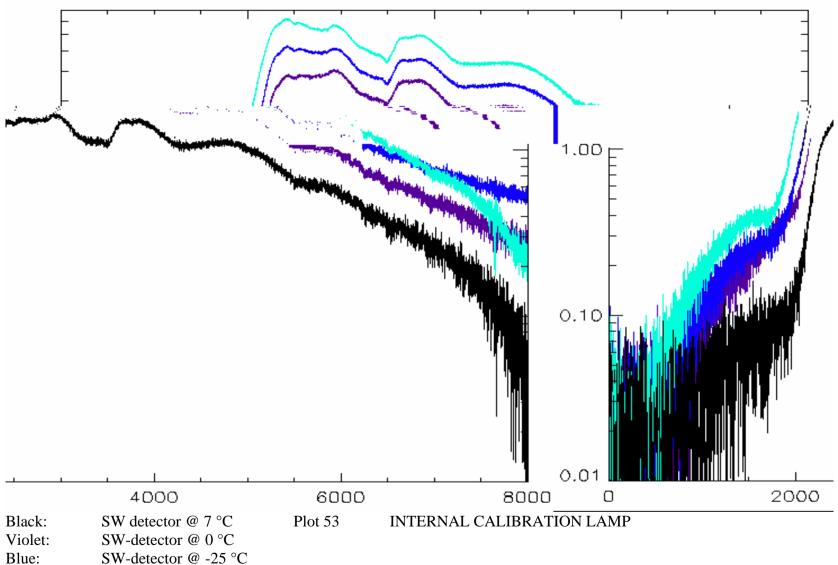
From bottom to top: 10fL, 50fL, 100fL, 200fL, 300fL, 400fL, 500fL, 1000fL, 2000fL, 3000fL, 4000fL, 5000fL





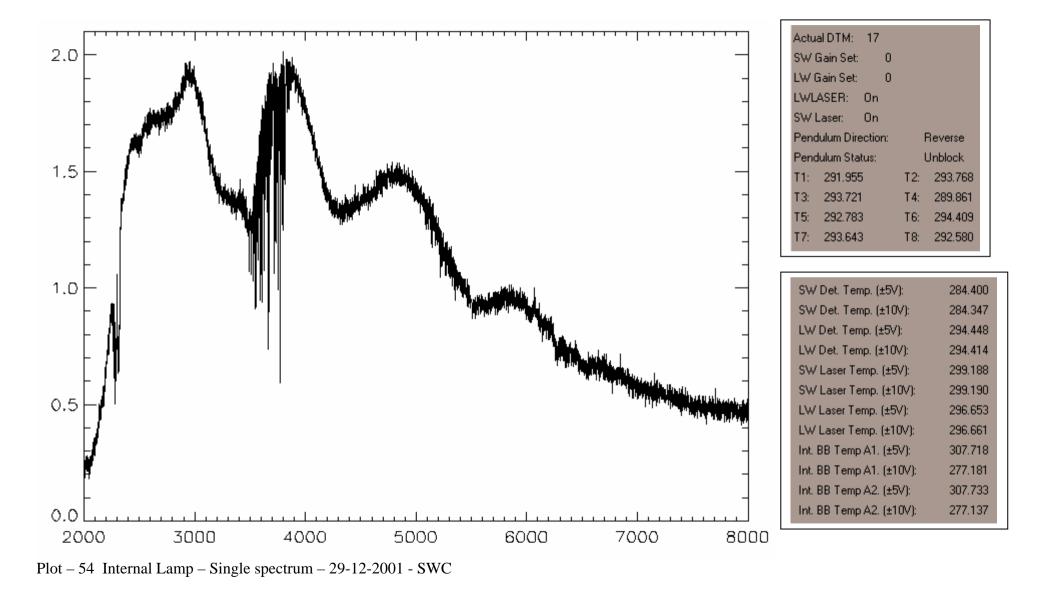
Integrating Sphere – 100fL - SWC - Gain 4 - AVERAGE OVER 5 SPECTRA – Smoothed on 10 points. Very low signal from the sphere. Thermal signal from chamber window. Detector SW cooled to –40 celsius. Plot – 52

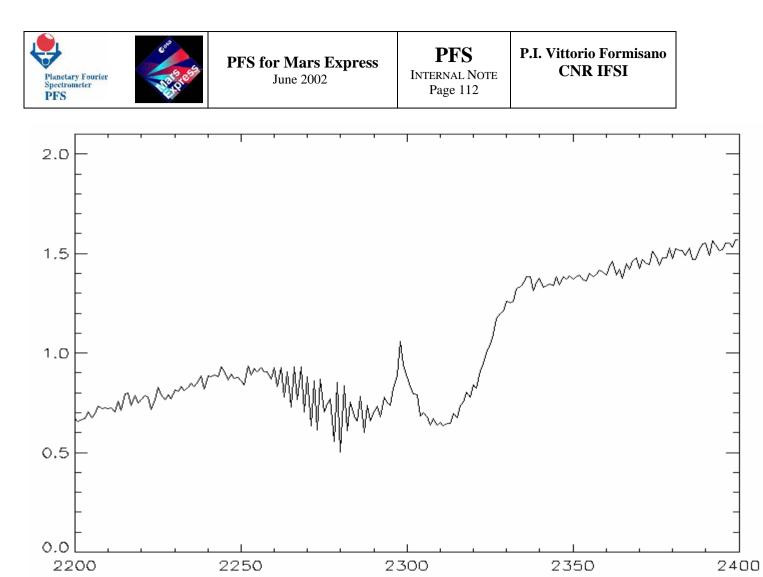




Cyan: SW-detector @ -46 °C

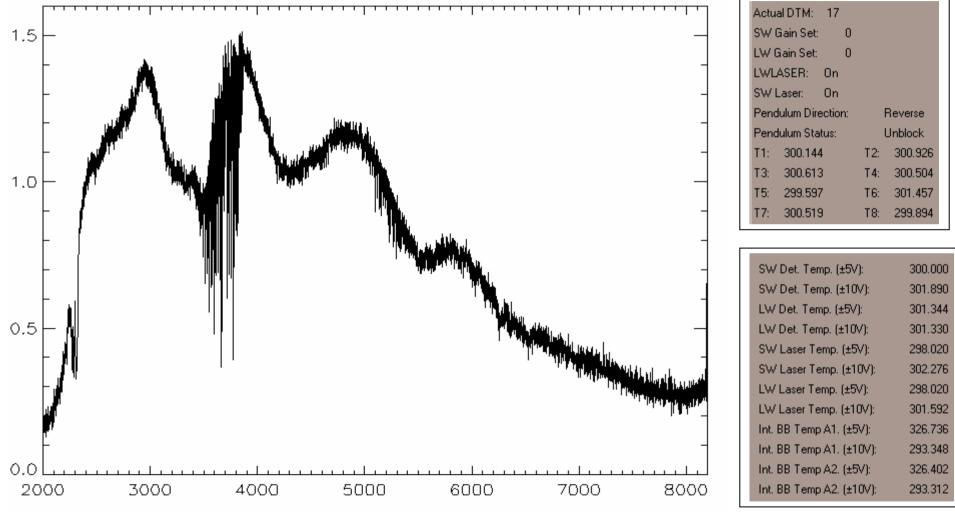
Planetary Fourier Spectrometer PFS PFS PFS	for Mars Express June 2002	PFS Internal Note Page 111	P.I. Vittorio Formisano CNR IFSI
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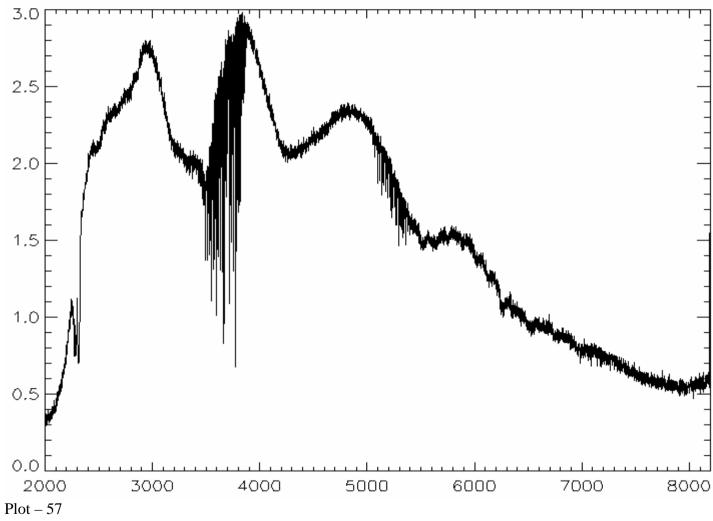
Internal Lamp – Single spectrum – 29-12-2001 – SWC – CO_2 band No line-shifting is present

S 55	Mars Express ine 2002 PFS INTERNAL NOTH Page 113	P.I. Vittorio Formisano CNR IFSI
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Plot - 56 Internal Lamp – Single spectrum – 02-01-2002 – SWC No line-shifting is present.

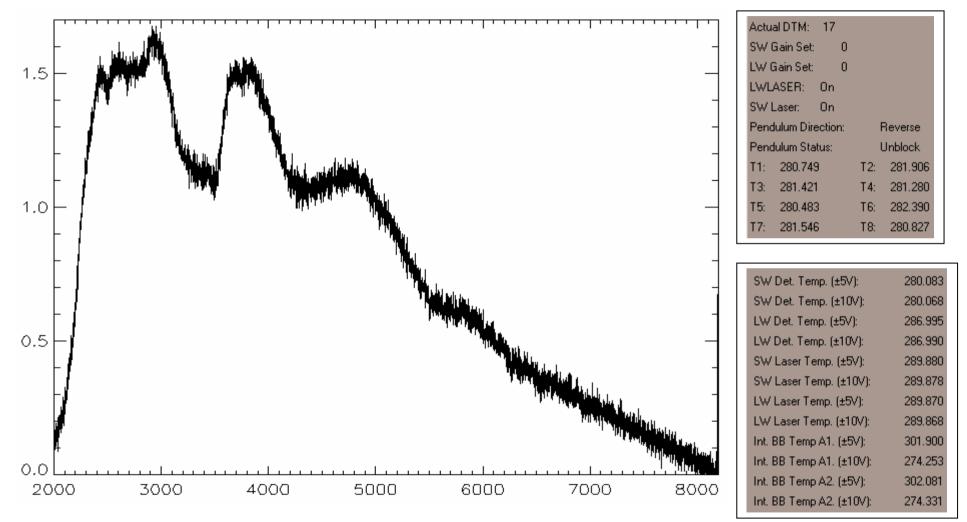
Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 114	P.I. Vittorio Formisano CNR IFSI
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 $Internal\ Lamp-Single\ spectrum-02-01-2002-SWC-Gain\ 1$

No line-shifting is present.

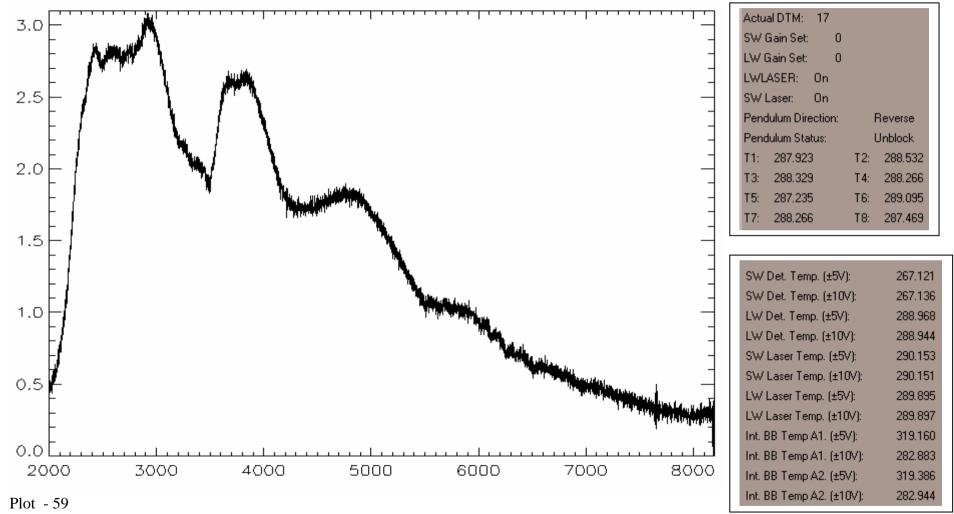
Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 115	P.I. Vittorio Formisano CNR IFSI
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Internal Lamp – Single spectrum – 08-01-2002 – SWC – In the thermo-vacuum Chamber

SW-detector @ room temperature Plot - 58

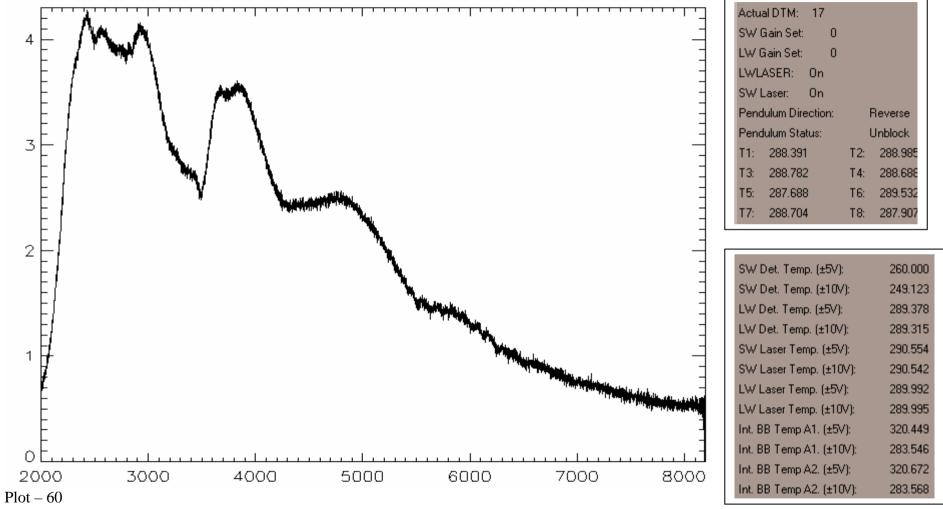
Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 116	P.I. Vittorio Formisano CNR IFSI
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Internal Lamp – Single spectrum – 08-01-2002 – SWC – In the thermo-vacuum Chamber

SW-detector @ 0 °C

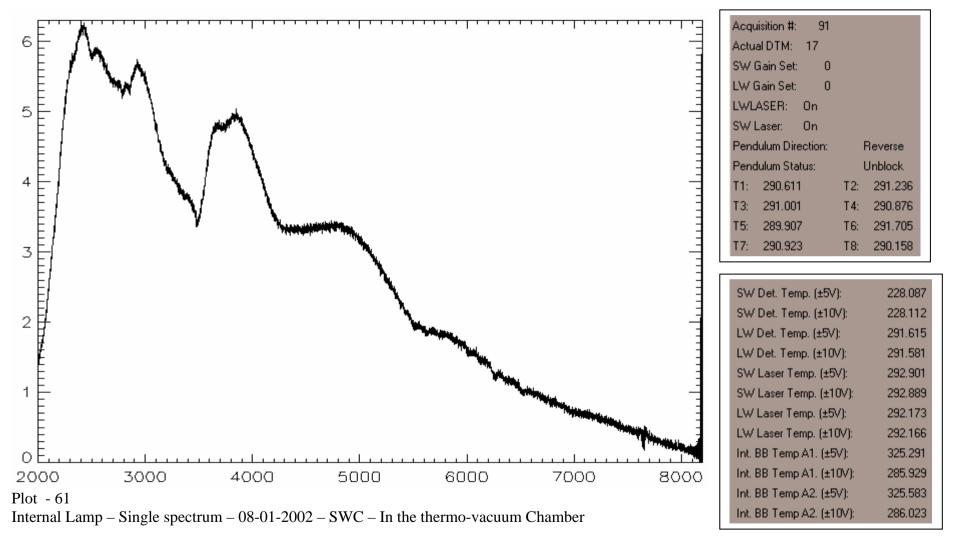
Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 117	P.I. Vittorio Formisano CNR IFSI
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Internal Lamp – Single spectrum – 08-01-2002 – SWC – In the thermo-vacuum Chamber

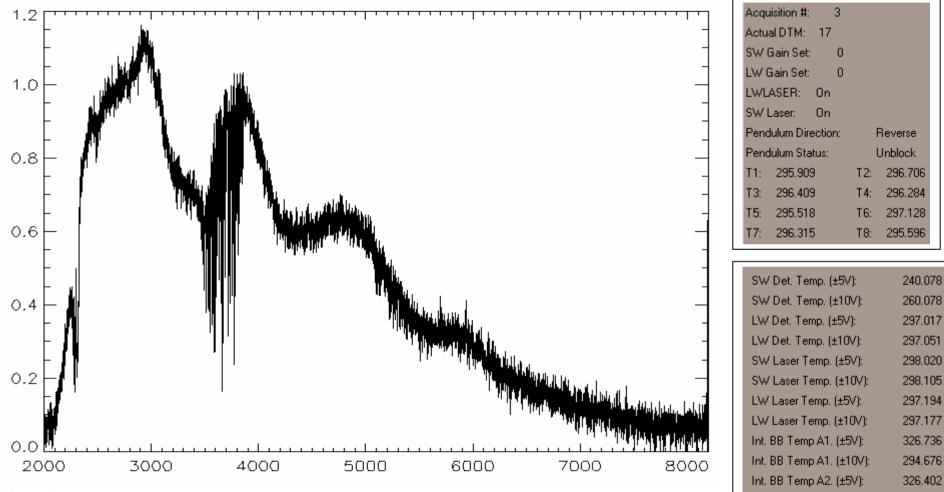
SW-detector @ -25 °C

Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 118	P.I. Vittorio Formisano CNR IFSI
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SW-detector @ -46 °C

Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 119	P.I. Vittorio Formisano CNR IFSI
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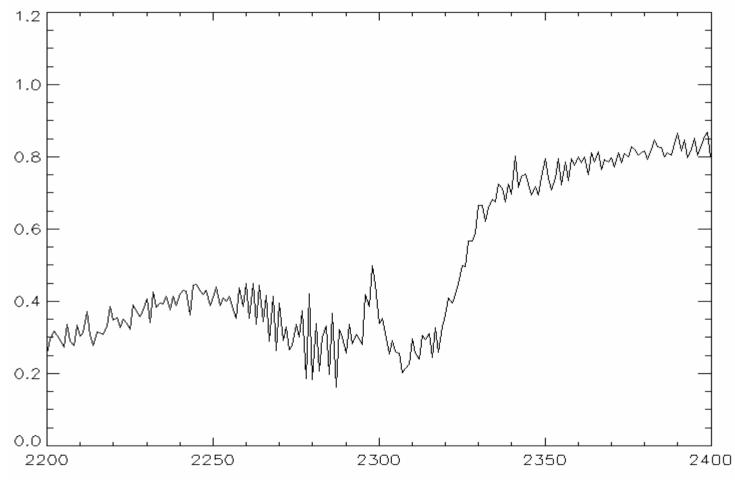


Int. BB Temp A2. (±10V):

294.793

Plot – 62 Internal Lamp – Single spectrum Alenia Spazio Torino 11-12/02/2002

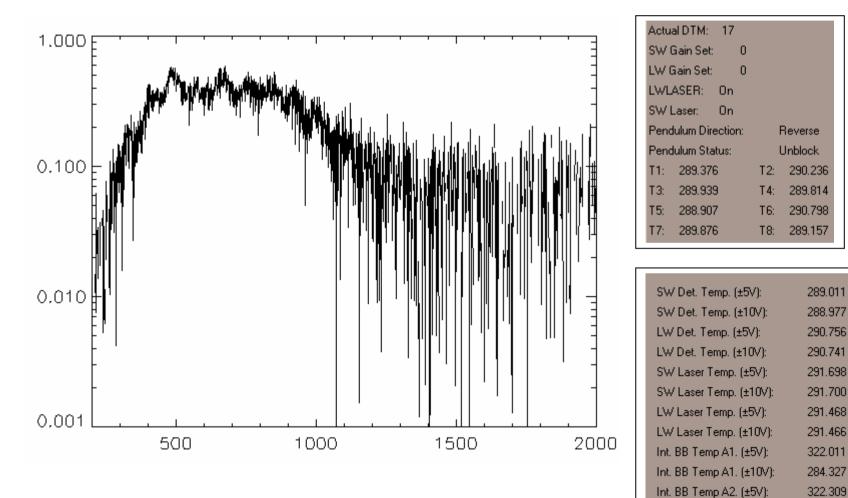
Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 120	P.I. Vittorio Formisano CNR IFSI
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Plot – 63 Internal Lamp – Single spectrum Alenia Spazio Torino 11-12/02/2002

Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 121	P.I. Vittorio Formisano CNR IFSI
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2.9 INTERNAL BLACK BODY



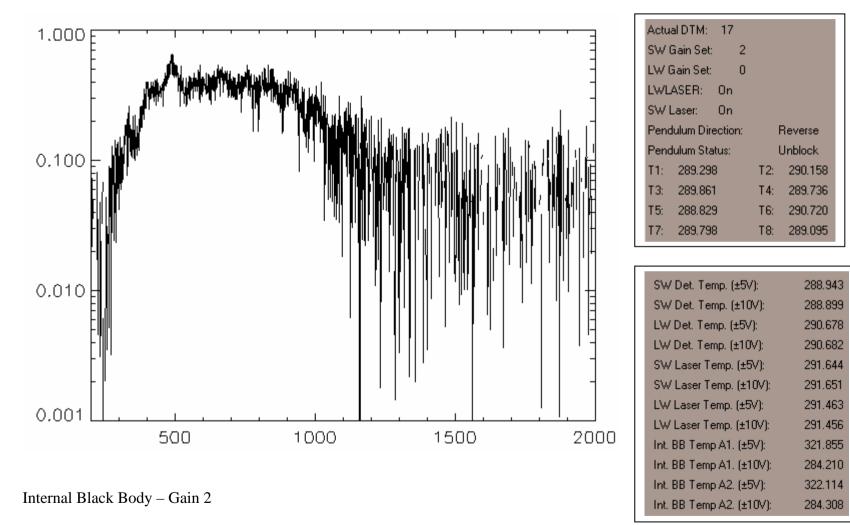
Int. BB Temp A2. (±10V):

284.386

Internal Black Body – Gain 0

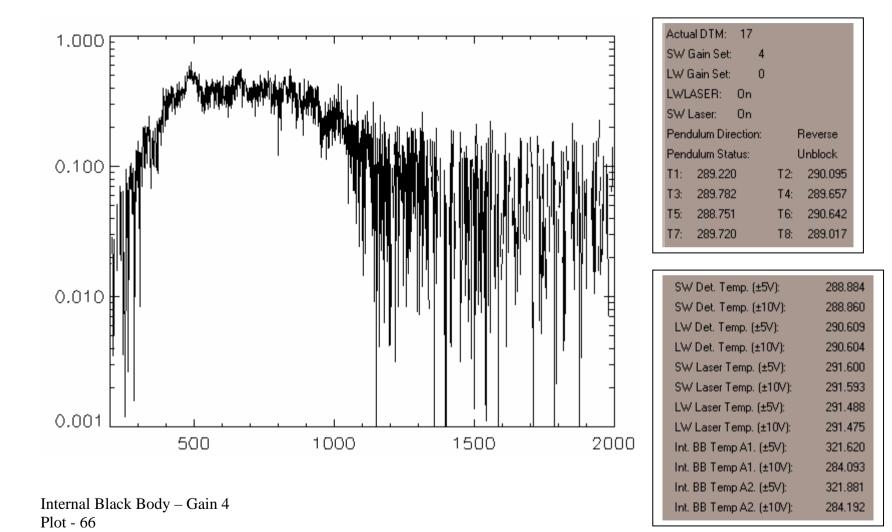
Plot - 64

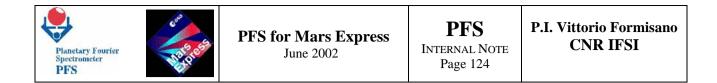
Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 122	P.I. Vittorio Formisano CNR IFSI
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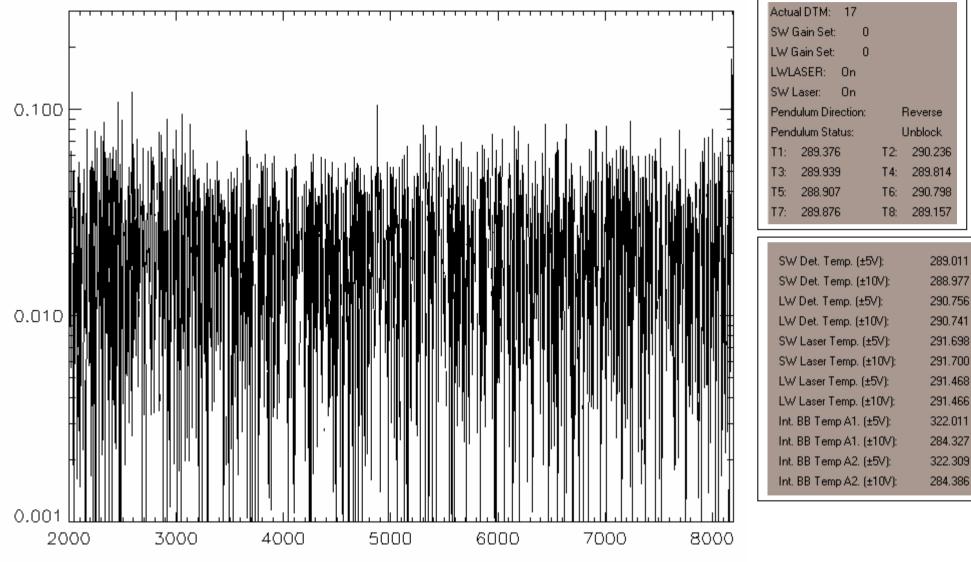




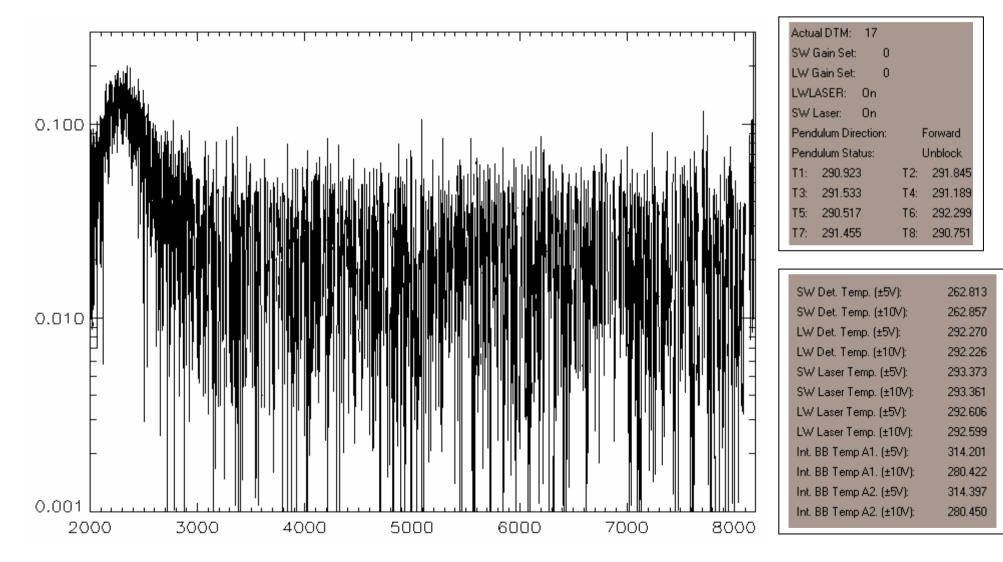
Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 123	P.I. Vittorio Formisano CNR IFSI
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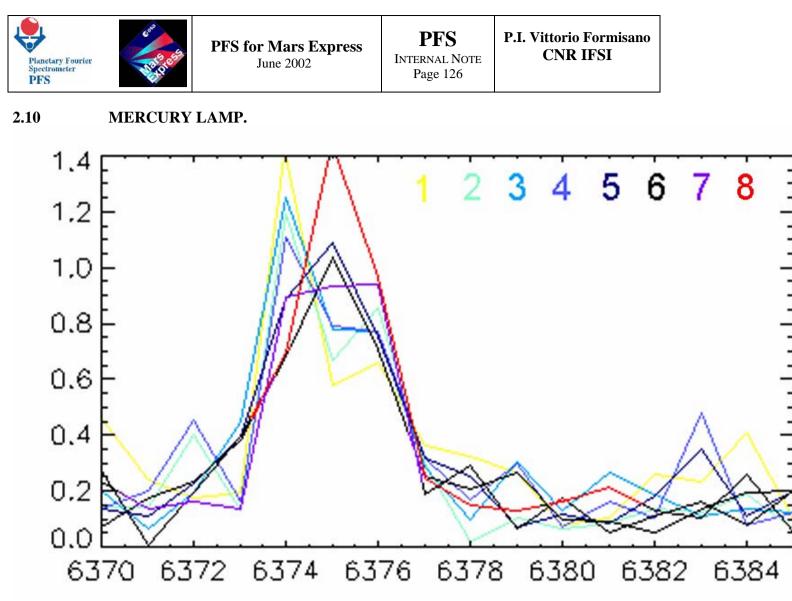






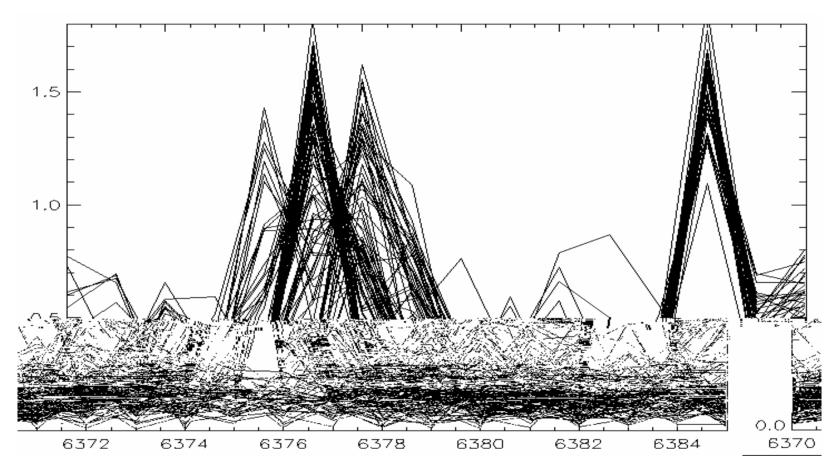
Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 125	P.I. Vittorio Formisano CNR IFSI
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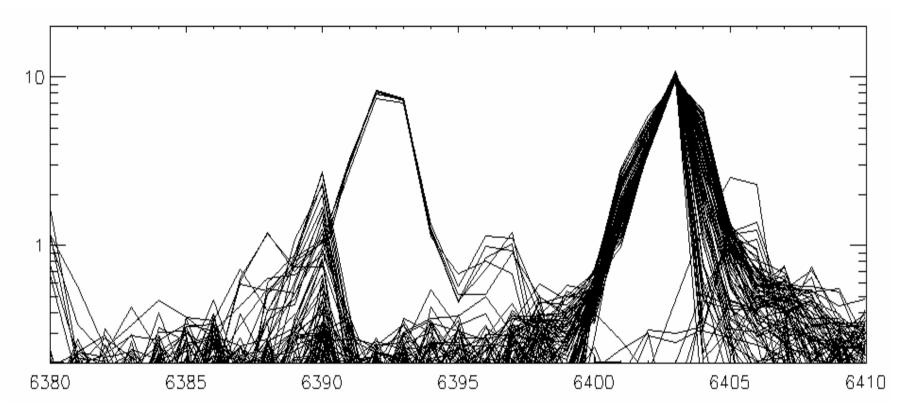
Mercury Lamp single channel line shifting – Gain 4 – SW Detector @ 280K

Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 127	P.I. Vittorio Formisano CNR IFSI
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Mercury Lamp line shifting from "MERCURY_TEST6_020108_094425-210.ifg" – Gain 4 – SW Detector @ 280K

Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 128	P.I. Vittorio Formisano CNR IFSI
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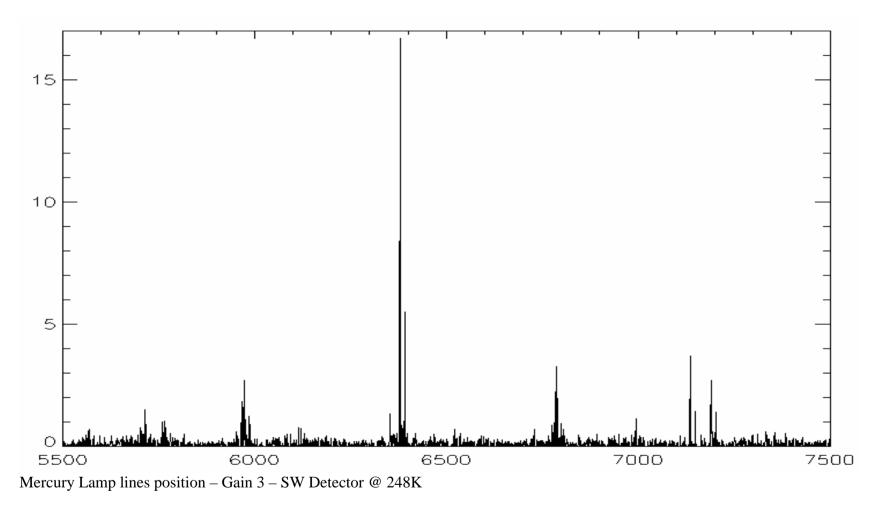


Mercury Lamp line shifting from "MERCURY_TEST6_020110_122521-84.ifg" – Gain 3 – SW Detector @ 248K Plot - 71

PFS for Mars Express June 2002 PFS

INTERNAL NOTE PFS fo6()6(Page 1)4.6(29 f 0 16.02 -16.02 001.1 4001.1 5055.5 439.31Tc -0.0003 S)Tj 003

Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 130	P.I. Vittorio Formisano CNR IFSI
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2.9 INTERNAL BLACK BODY

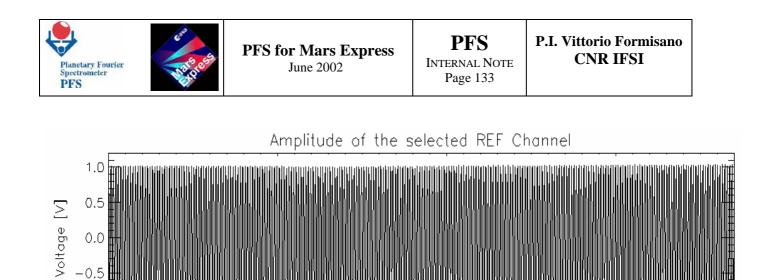
All the calibration activity is aiming to get knowledge of the calibration sources that we carry in the experiment. It is natural, therefore that we look also at the Internal Black body to see how it appears in the data.(See plot 64, 65, 66, 67, 68). The internal BB, being almost at the same temperature of the instrument gives a very small signal in the LW channel. In the SW channel, on the contrary, the signal depends strongly on the SW detector temperature and this may be very important in space, where theSW detector is going to be cooled at 200 K. The SW channel may complement very well to determine the BB temperature.

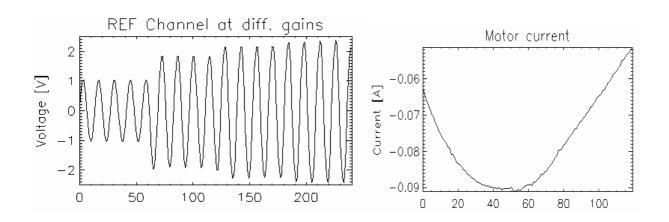
2.10 MERCURY LAMP

We have taken several measurements in Test 6 with a mercury lamp.See plot 69,70,71,72,73. Test 6 is supposed to heat the laser diode by 10 degrees in steps of 0.5 deg, and at each step take a number of measurements. We have not been able to really get the 10 deg temperature increase, because either not enough time was given to the laser for temperature adjustement, or the laser was not really controlled because ambient temperature was larger than the set temperature. In any case this test was very important because by changing the laser wavelength we have also the instrument response to monochromatic features. In the color figure, as well as in the test made in Alenia, the shift of the line of the mercury lamp is by only 1 cm-1. We see that before the peak moves to the next wavenumber channel it decreases to half intensity (energy conservation ?). We see also that the laser diode wavelength does not move continuously, but jumps discontinuously. This is a very important fact that we must further investigate in the future.

2.11 AUTOTEST DATA

PFS has an automatic test called autotest in which instrumental information is sampled and transmitted. We show here two autotest performed one in the IFSI lab(see plot 74, 75), and one when integrated on the spacecraft. From the autotest, performed in 2 complete motions of the double pendulum, we get 4 measurements. The first is the speed of the double pendulum (first panel). The second is the shape and stability of the sine wave, laser diode interferogram. The third is the current used for the motion, the last is the sine wave at different amplification factors. In the following figures we see that the motion on the spacecraft was not perfect (we cannot put the spacecraft exactly perpendicular, to minimise gravity effects), although the data taken were all ok. The fact is that out of 18000 sine periods, we are sampling the first 1000 only, while the interferogram is sampled when the motion is stabilised.

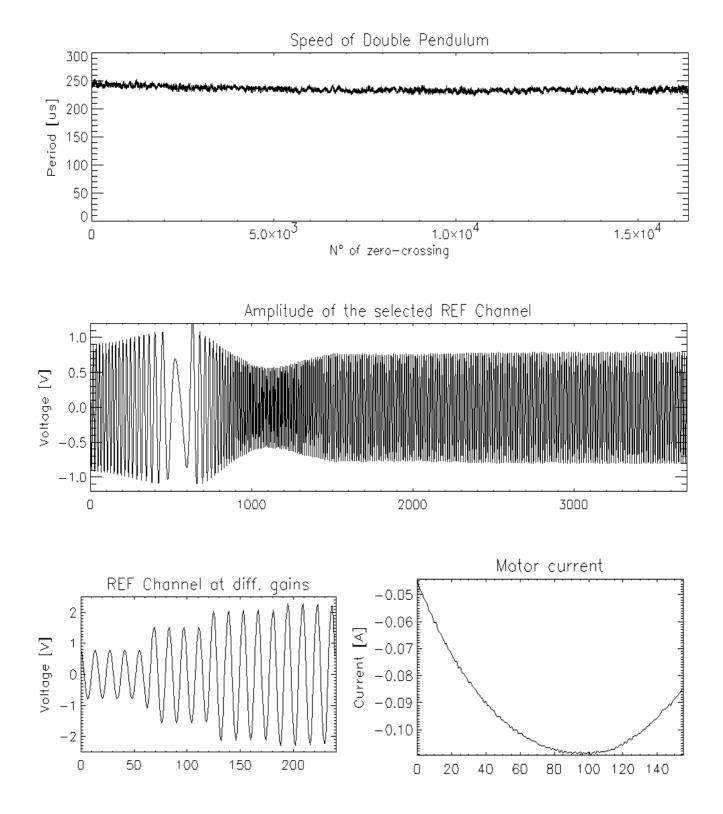




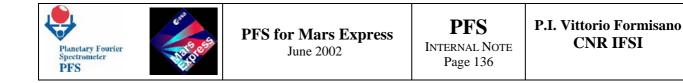
Plot – 74 AUTOTEST IN THE LAB IFSI

-1.0



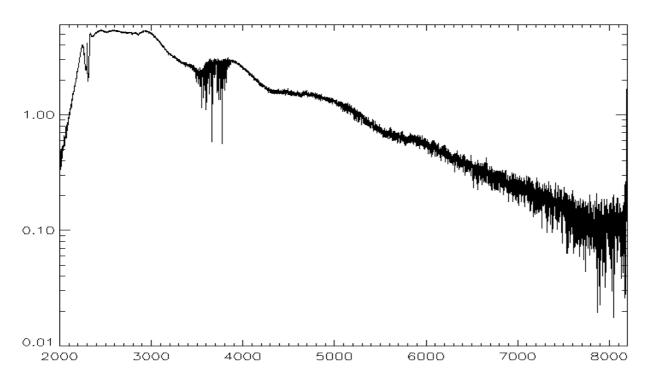


Plot - 75 AUTOTEST AFTER INTEGRATION ON THE SPACECRAFT.

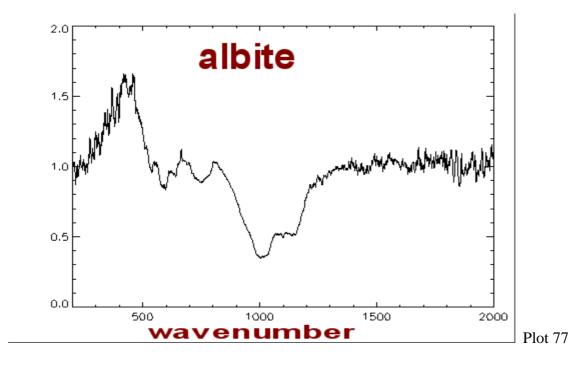


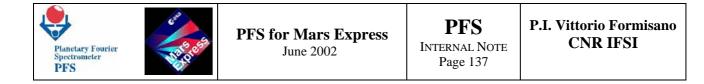
2.12 MINERALS

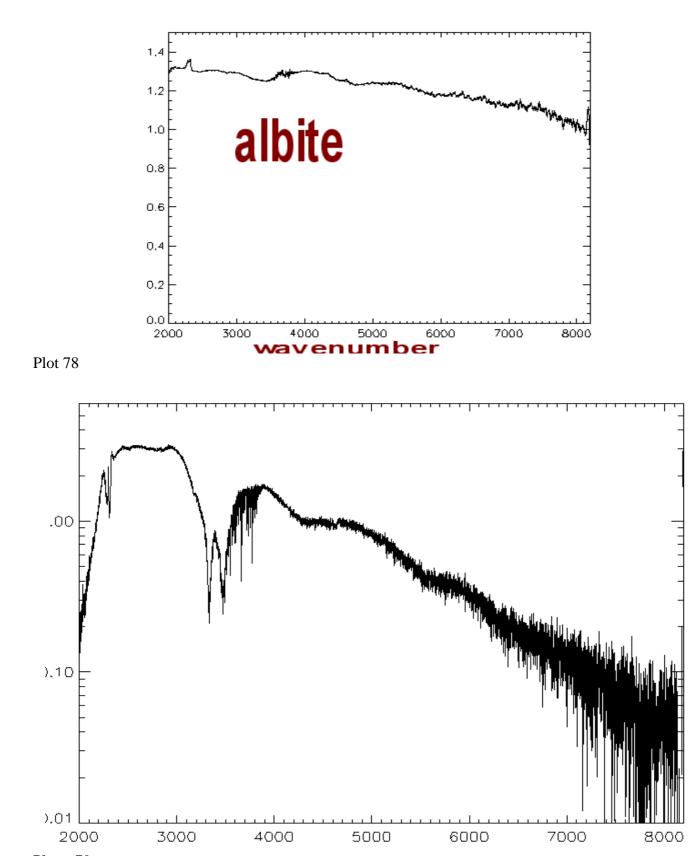
A number of trasmissivity spectra for minerals were taken(see plot 76, 77, 78). The minerals were either prepared as dust elements included in KBR diskette or as dust element distributed over a CaF2 disk. Here we present only some spectra taken as diskette. A full report will be prepared for publication.



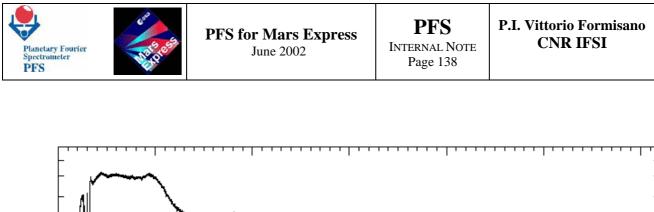
Plot- 76 Albite Illuminated by a ceramic element.

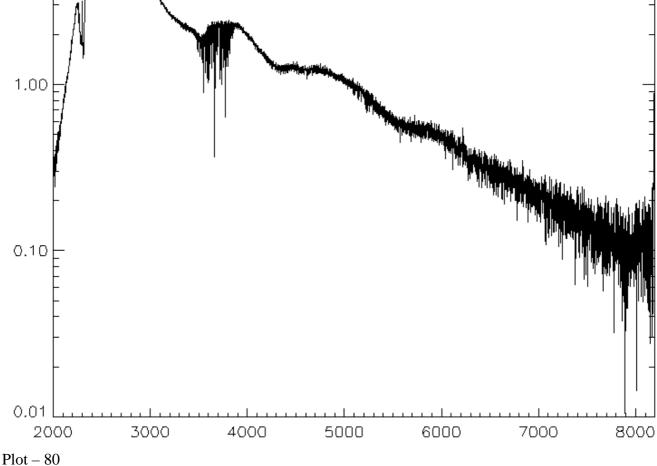






Plot – 79 Gypsum illuminated by a ceramic elemen





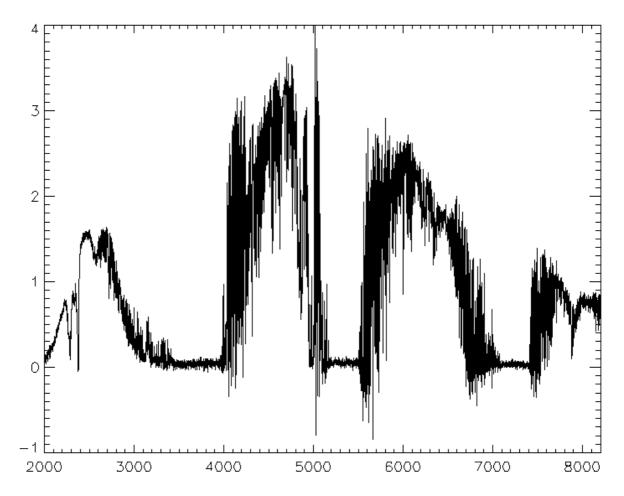
KBR illuminated by a ceramic element.

Plot 76,79, 80 show the obtained spectrum, while plot 78,79 show the ratio with Kbe spectrum, and therefore gives the trasmissivity of albite mineral.



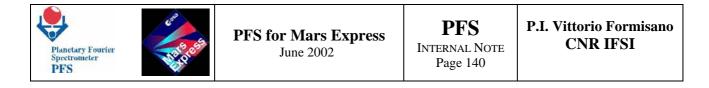
2.13 EARTH'S ATMOSPHERE

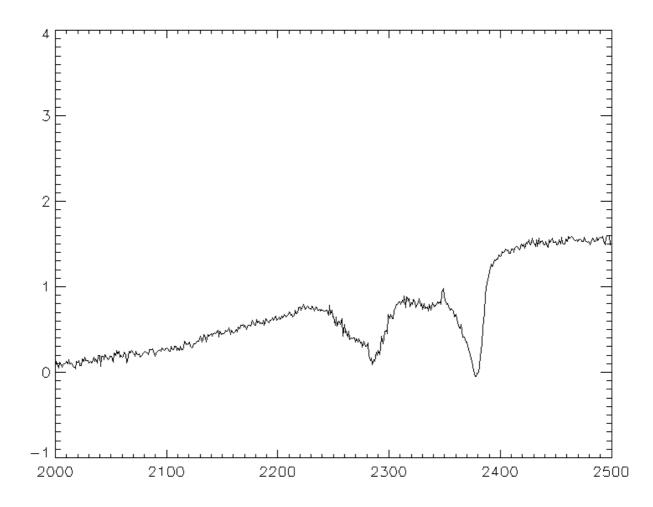
We have used PFS to study the Earth's atmosphere by looking out of the window in daytime. The average over 50 spectra in the SW channel is shown below. The 3 saturated water band are evident, however there are, in the 4 trasmitting windows of the atmosphere many important bands to study : CO2, H2O, HDO, O2, CH4 etc.



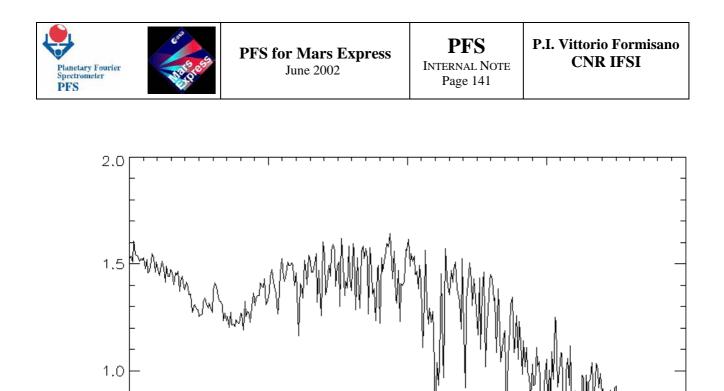
SW channel spectrum, average over 50 measurements. Gain 5. Plot-81

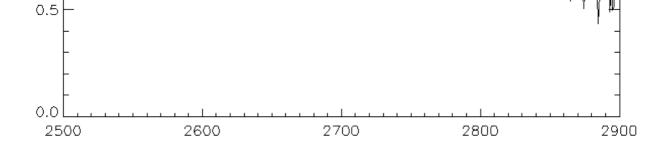
81



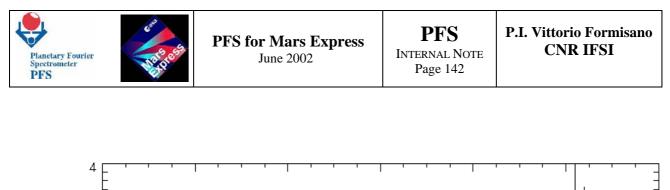


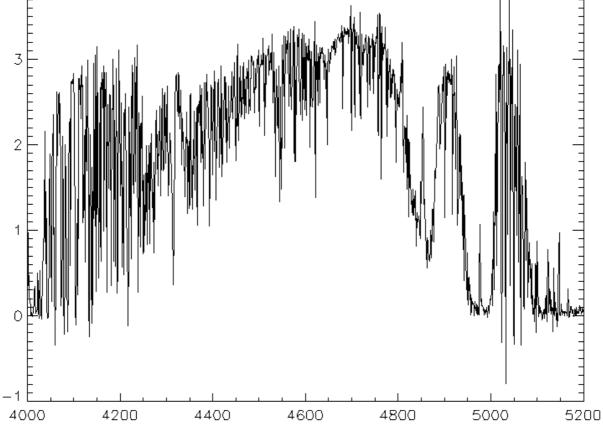
SW channel spectrum, average over 50 measurements. Gain 5. Blow up around the CO2 absorption band. Note that the band is almost completely masked by apparent emission : Between the source and the detector the thermal situation is very complex : Note the absence of the usual thermal emission rising toward the 2000 cm-1.



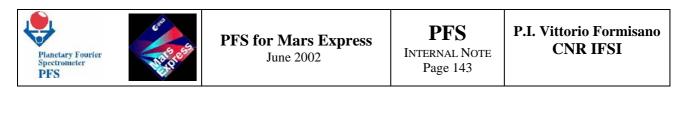


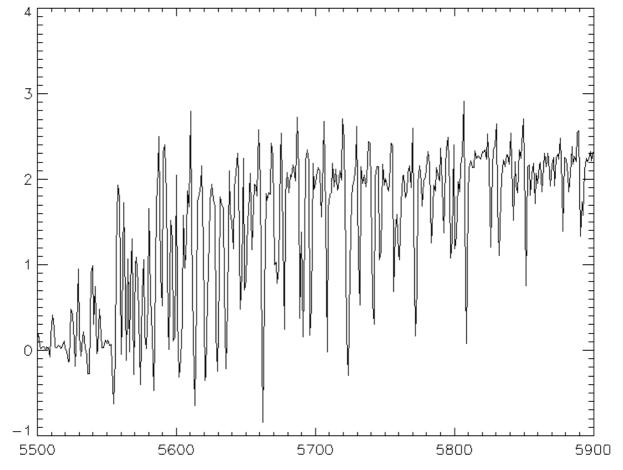
Blow up of the average spectrum in the $2500 - 2900 \text{ cm}^{-1}$ region. 2 major features should be noted here: the NO₂ band at 2540-2590 cm⁻¹ (another similar band at 3500 cm⁻¹ disappears in the H₂O band), and the deuterium band at 2720 cm⁻¹. Other water bands are also evident. Plot - 83



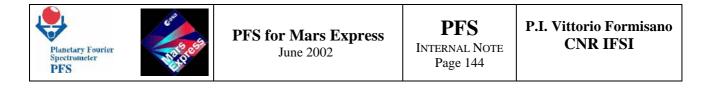


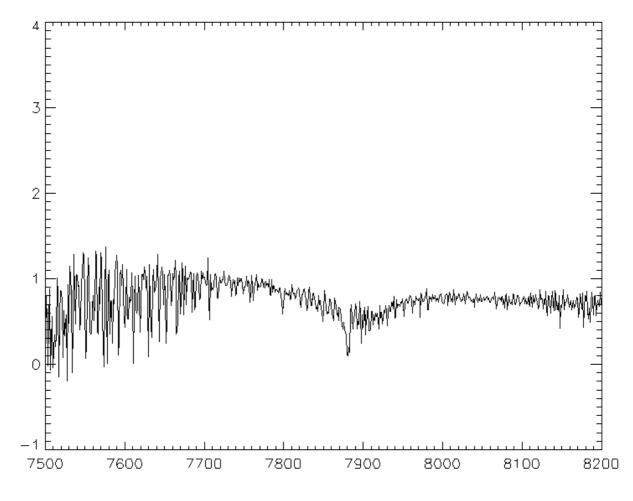
Blow up of the atmospheric spectrum in the 4000-5200 cm⁻¹ region : In the region 4300- 4500 cm⁻¹ water has no major features : the bands observed are due to methane. In the region 4800- 5200 cm⁻¹ are the 3 major CO_2 bands, the third one being absorbed in the big water band absorption. Plot - 84





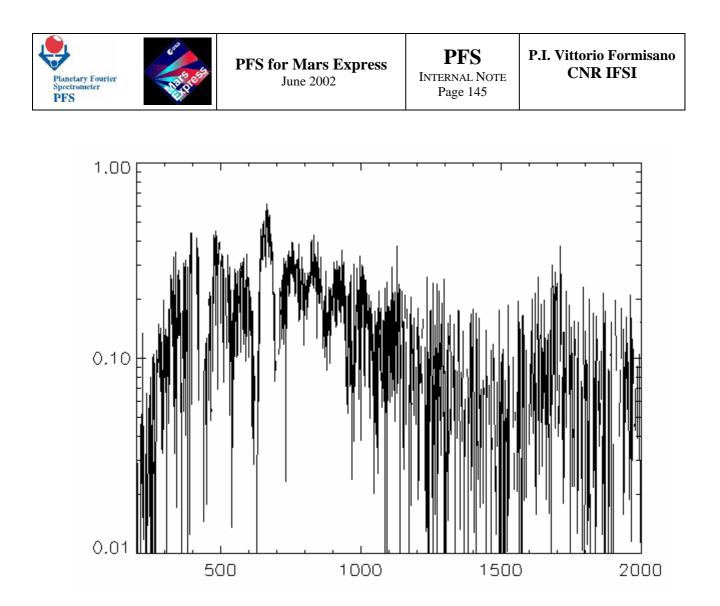
Blow up in the 5500- 5900 cm-1 zone. Many water lines are clearly visible.





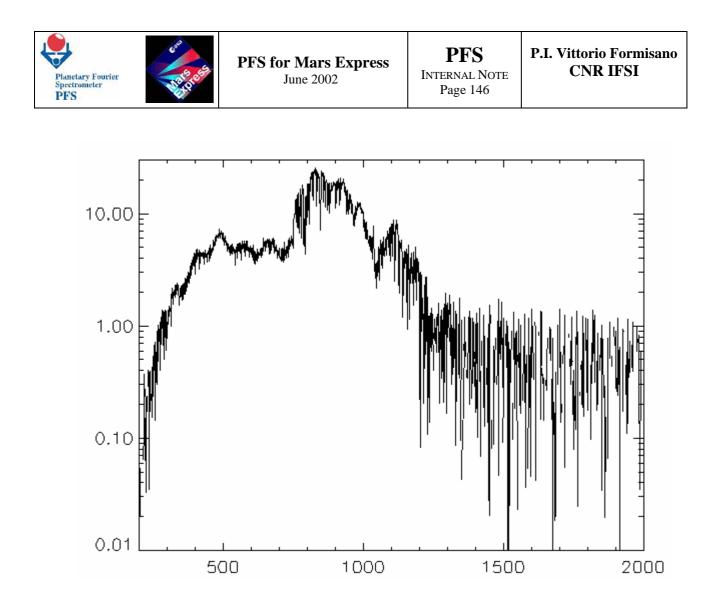


Blowup of the spectrum in the wavenumber region 7500- 8200 cm-1. Note the O2 line and band clearly visible at 7880 cm-1. This is the FM at gain 5 SW channel.



Plot 87

The LW channel looking out of the window in the direction of IAS. Note the CO2 band in emission ,as in the case of the SW channel.



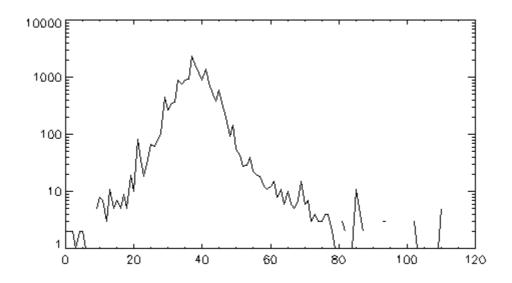
Plot 88

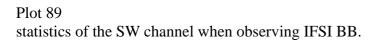
The LW channel when looking in the direction of Rome. The soil is out of the field of view. Over Rome there was a thermal inversion stratification: aerosols emission at higher temperature is evident.

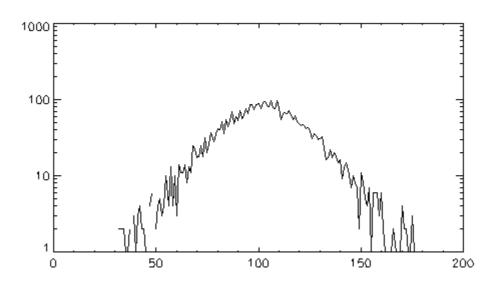


2.14 INTERFEROGRAMS STATISTICS

In order to study the behaviour of the ADC, we have performed statistical study of the values obtained by the interferograms, specially the values of the long wings. From our study it is evident that there is no bias and that the ADC are working correctly.

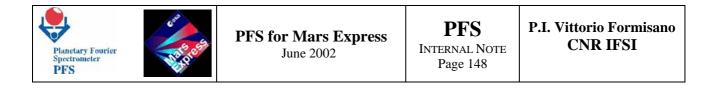


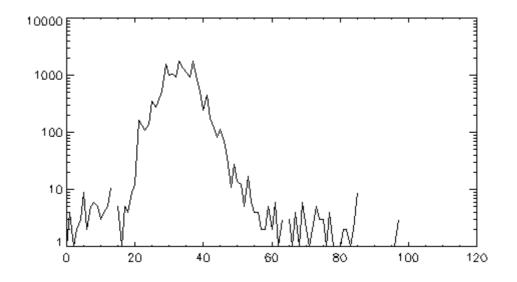


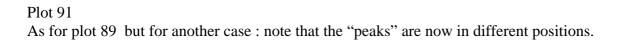


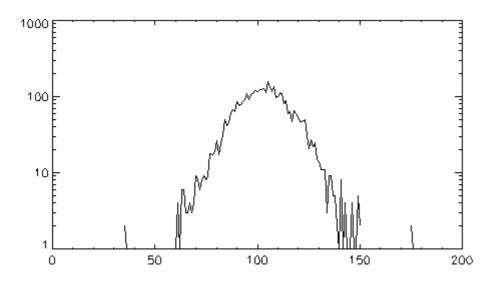
Plot 90

Statistic of the LW channel when observing the IFSI BB. Both histograms 89 and 90 are bin 1 and centred roughly around the zeros of the interferograms.







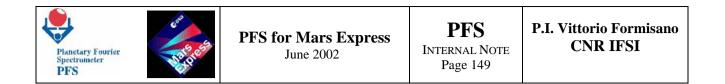


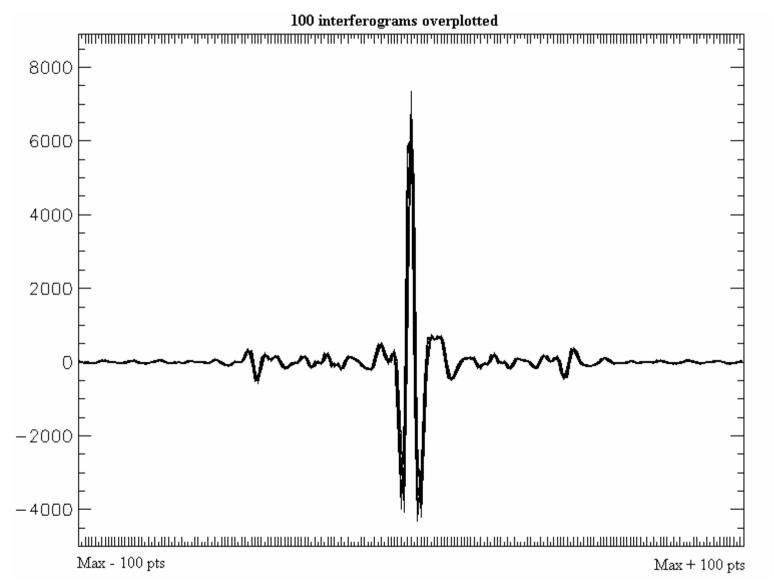
Plot 92 As for Plot 90 LW but for another case. Note the Gaussian distribution.

2.15 STABILITY OF THE EXPERIMENT

LW Channel Stability

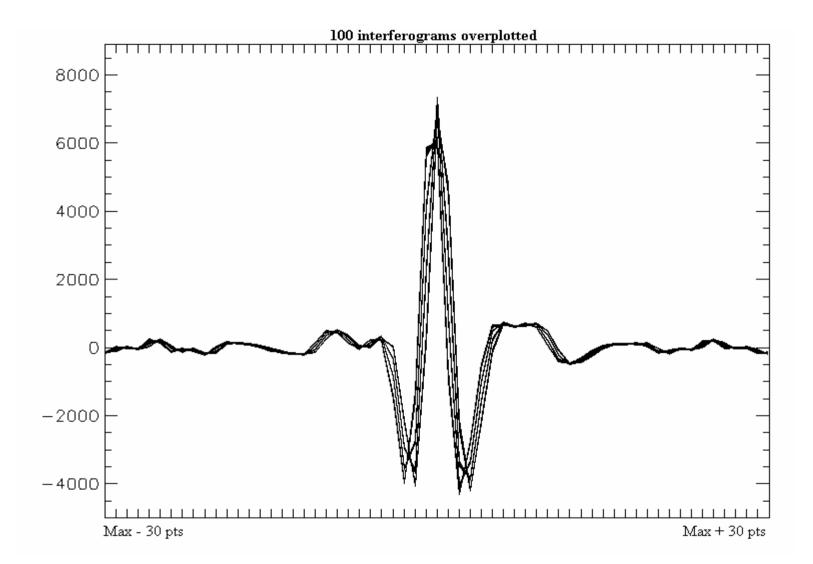
In order to study the stability of the experiment during one set of measurements, we over plot the center part of the interferograms of that session, selecting forward or reverse motion, and check that the interferograms are stable.





Plot 93 : 100 LW interferograms overplotted. +- 100 points are shown around the zero optical path difference.

Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 150	P.I. Vittorio Formisano CNR IFSI
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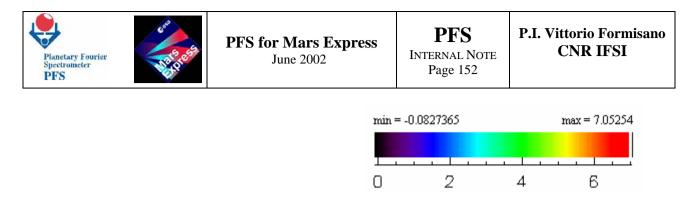


Plot 94 – The same data as plot 93 but a blow up around the ZOPD : Note theshift caused by the laser diode change.

Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 151	P.I. Vittorio Formisano CNR IFSI
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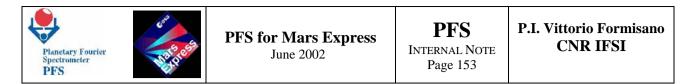
1100cm⁻¹



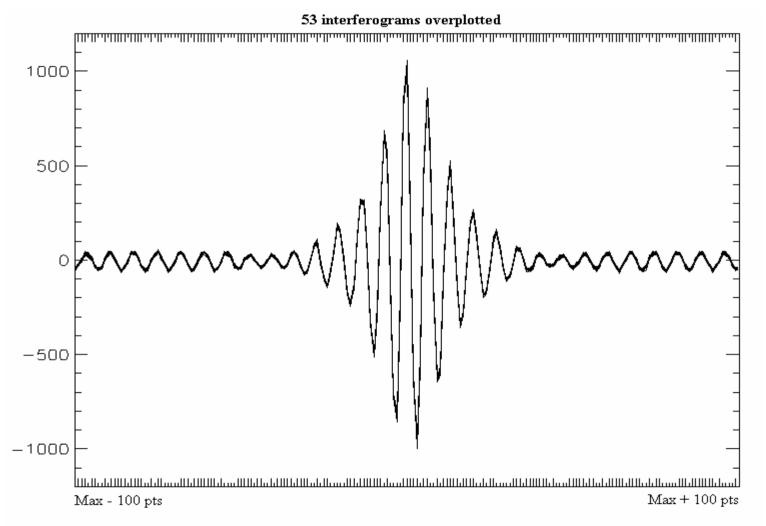
Plot 95.

Spectral shape stability over 100 measurements : Each line in the color figure above is one spectrum. The color code is given above. Note the wavy features of the spectra in the range of wavenumbers shown. The CO2 Qbranch is evident as thin vertical line.

The conclusion is that the LW channel is very stable and therefore very reliable instrument.

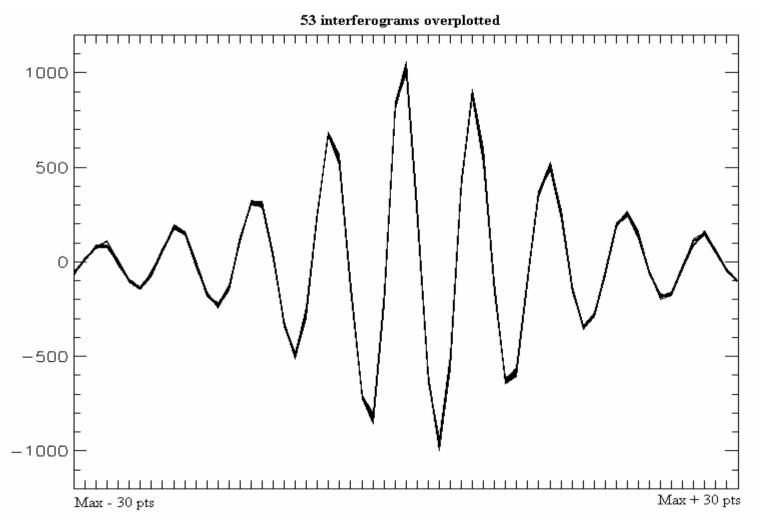


SW Channel Stability



Plot 96 : 53 interferograms SW overplotted.

Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 154	P.I. Vittorio Formisano CNR IFSI
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Plot 97 : The same interferograms of plot 96: blow up of the central part.

Planetary Fourier Spectrometer PFS	PFS for Mars Express June 2002	PFS Internal Note Page 155	P.I. Vittorio Formisano CNR IFSI
2000cm^{-1}			
	min = -0.0235690		max = 1,064929
	0.0 0.2 0.4 0.6 0.8 1.0		

Plot 98 :

Color plotting of the 53 spectra (only the thermal part, as the source was the IFSI BB). Note the presence of the CO2 band.

2800cm⁻¹

It is clear that the conclusion is that also the SW channel is very stable.

PFS for Mars Express June 2002

PFS

INTERNAL NOTE Page 156 P.I. Vittorio Formisano CNR IFSI



-LASER WAVELENGTH JUMPING

we have shown a few spectra in which the laser diode was jumping from one wavelength to another. this effects can be identified because the natural wavenumber of co2 or h2o features are fixed and do not move. we have built in pfs a test called test 6 in which the temperature of the laser diodes is increased by 10 degrees in step of 0.5 degrees. this test has not always worked well because depends on the ambient temperature and on the temperature set for the laser diodes at the beginning. we should, however have enough data to study the laser diode wavelength as function of the temperature. it is obvious that no averaging of the data should be performed a priori, but only after selection of measurements with the laser in the same conditions. the result could be, otherwise, an increase of the apparent noise.

-SATURATION OF THE INTERNAL CALIBRATION LAMP

This is a different instrumental effect, in the sense that it is not observed always on ground but it is expected to be so in space also for gain 0, simply because the SW detector responsivity shall increase by a large factor. Note that with respect to the original 16 mA current we have tested several values (16, 8, 12, 10 mA) 8 mA was not acceptable because the lamp was notable to switch on, and acted as a thermal source only. We shall report on measurements at 10 and 12 mA. But is a fact that when the SW detector was cooled gain 1 was always saturated, and the temperatures reached were far from the temperatures in space. Further cooling for the detector is expected to give a better signal to noise ratio for the observations of Mars.

-STABILITY OF THE EXPERIMENT

The experiment is very stable when a sequence of measurements in the same conditions are made.