ROSETTA-RPC-IES
PLANETARY SCIENCE ARCHIVE
INTERFACE CONTROL DOCUMENT

MAY 2016

SwRI® Project 10991

Document No. 10991-IES-EAICD-02

Contract JPL 1200670

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REVISION NOTICE

Initial Issue: September 2005.

Revision 1: Updated for version 2 archive products. August 2007

Revision 2: Updated in response to comet phase archive review. May 2016.

1. INTRODUCTION

1.1 Purpose and Scope

The purpose of this EAICD (Experimenter to (Science) Archive Interface Control Document) is to provide users of the RPC-IES instrument data with detailed description of the product and a description of how it was generated, including data sources and destinations. It is the official interface between the instrument team and the archiving authority.

1.2 Archiving Authorities

The Planetary Data System Standard is used as archiving standard by

- NASA for U.S. planetary missions, implemented by PDS
- ESA for European planetary missions, implemented by the Research and Scientific Support Department (RSSD) of ESA

1.2.1 ESA's Planetary Science Archive (PSA)

ESA implements an online science archive, the PSA,

- to support and ease data ingestion
- to offer additional services to the scientific user community and science operations teams as e.g.
    - search queries that allow searches across instruments, missions and scientific disciplines
    - several data delivery options as
        - direct download of data products, linked files and data sets
        - ftp download of data products, linked files and data sets

The PSA aims for online ingestion of logical archive volumes and will offer the creation of physical archive volumes on request.

1.3 Contents

This document describes the data flow of the IES instrument on the Rosetta mission from the spacecraft until the insertion into the PSA for ESA. It includes information on how data were processed, formatted, labeled and uniquely identified. The document discusses general naming schemes for data volumes, data sets, data and label files. Standards used to generate the product are explained.

The design of the data set structure and the data product is given. An example data product is given in section 4.3 Data Product Design.

1.4 Intended Readership

This document's intended readership includes the staff of the archiving authority (Planetary Science Archive, ESA, RSSD, design team) and any potential user of the RPC-IES data.

1.5 Applicable Documents

Planetary Data System Data Archive Preparation Guide April 1, 2010, Version 1.4, JPL D31224

Planetary Data System Standards Reference, February 27, 2009, Version 3.8, JPL, D-7669, Part 2

Rosetta Archive Generation, Validation and Transfer Plan, January 2006, RO-EST-PL-5011

Rosetta Plasma Consortium Users' Manual, V 2.18, October 2011, RO-RPC-UM

Ion and Electron Sensor (IES) Flight Software Requirements Document, November 14, 2000, Rev. 0 Change 0, SWRI, Document No. 8182-FSRD-01

1.6 Relationships to Other Interfaces

N/A

1.7 Acronyms and Abbreviations

CCSDS	Consultative Committee for Space Data Systems
DDS	Data Distribution System
ESA	European Space Agency, Electrostatic Analyzer
ESOC	European Space Operations Centre
FOV	Field of View
HGA	High Gain Antenna
HGRTN	Heliocentric Radial-Tangential-Normal
IES	Ion and Electron Sensor
IESGS	IES Ground System
MCP	Microchannel Plate
NAIF	Navigation and Ancillary Information Facility
PDS	Planetary Data System
PSA	Planetary Science Archive
RDDS	Rosetta Data Distribution System
RPC	Rosetta Plasma Consortium
SPICE	NAIF information system

1.8 Contact Names and Addresses

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2. OVERVIEW OF INSTRUMENT DESIGN, DATA HANDLING PROCESS AND PRODUCT GENERATION

2.1 Instrument Design

The IES for Rosetta is an electrostatic analyzer (ESA), featuring electrostatic angular deflection to obtain a field of view of 90º x 360º. The instrument objective is to obtain ion and electron distribution functions over the energy range extending from 1 eV/e up to 17.67 keV/e. The angular resolution for electrons is 5º x 22.5º (16 azimuthal and 16 polar-angle sectors). For ions the angular resolution is 5º x 45º (16 azimuthal and 8 polar-angle sectors) with additional segmentation to 5º x 5º in the 45º polar-angle sector most likely to contain the solar wind (giving a total of 16 polar-angle sectors for ions). The back-to-back top hat geometry of the IES electrostatic analyzer allows it to analyze both electrons and positive ions with a single entrance aperture. The IES top hat analyzers have toroidal geometry with a smaller radius of curvature in the deflection plane than in the orthogonal plane. This toroidal feature results in a flat deflection plate geometry at the poles of the analyzers and has the advantage that the focal point is located outside the analyzers rather than within them, as is the case with spherical top hat analyzers. Particles within a narrow 8% energy pass band will pass through the analyzers and be focused onto the electron and ion microchannel plates (MCPs), which produce charge pulses on 16 discrete anodes, which define the azimuth acceptance angles. In addition, the IES entrance aperture contains electrostatic deflection electrodes, which expand its elevation angle field of view to ±45º. With the typical top hat polar-angle field of view of 360º, the IES acquires a total solid angle of 2.8 pi steradians.

Operation of IES is controlled by its on-board software in conjunction with sets of (selectable) look up tables. A table in one set determines the sequence of voltages applied to the electrostatic analyzer, thereby selecting the energy/charge of electrons and ions entering the sensor. Likewise, a table in another set determines the sequence of voltages applied to the deflector plates, thereby defining the acceptance angle of the particles. In the typical operating mode, for each deflector voltage chosen the ESA is stepped over its range, the deflector voltage is stepped to its next value, and so on. A complete 2 voltage sequence thus determines a complete measurement cycle. Several versions of each table are stored in the

instrument so different operating modes can be easily chosen. In addition, new tables can be uploaded if desired.

During a measurement cycle the instrument obtains a full measurement of ion and electron flux within 16 azimuthal bins, 16 elevation bins and 128 energy bins, for a total of 65536 values (2x16x16x128) per measurement. To fit within the data volume allocated to IES, blocks of adjacent angle/energy bins are summed together. The details of this summation are mode-dependent, but this collapse and the 128 or 1024 second accumulation time are the only differences between IES operations in different modes.

2.2 Scientific Objectives

IES supports the RPC science goals by measurements of three-dimensional ion and electron velocity distributions and the derived quantities such as plasma density and flow velocity.

2.3 Data Handling Process

All RPC data packets are transmitted together during downlinks with Rosetta. RPC data is retrieved from the DDS at ESOC to a central RPC data server at Imperial College in London. Data for IES is copied from the RPC central data server by IESGS at Southwest Research Institute. Please see section 2.5.7 for an overview of IESGS and section 1.8 for contact information.

2.4 Data Products

2.4.1 Pre-Flight Data Products

None. Raw calibration data will be generated in the archive format for internal use, but there are no current plans to submit these data to the PSA.

2.4.2 Instrument Calibrations

IES calibration data will be added during a later release.

2.4.3 Other Files written during Calibration

None

2.4.4 In-Flight Data Products

To ensure that the IES goals can be achieved, data will be archived as:

- Edited raw data (CODMAC level 2) – the science data stream converted to human and PDS readable format.
- Calibrated data (CODMAC level 3) – the contents of the edited raw data with calibration information included.
- Derived higher level data (CODMAC level 4) – quantities calculated from phase space density, such as plasma density, flow speed, ion and electron pressure, or electron pitch angle distributions. (To be included in a future release)

These data may be used for cross-instrument calibrations, and both stand-alone and cross-instrument scientific analysis.

Table 1: Spacecraft Science Data Products in IES Data Sets			
Product	**Data Set Type**	**Maximum (MB / Day)**	**Mission Total, estimated (MB)**
ELC	Un-calibrated	33	26500
	Calibrated	66	
ION	Un-calibrated	25	21500
	Calibrated	50	
DERIVED	Moments	0.25	150
	Electron pitch angle distributions	1.2	650

2.4.5 Software

We do not intend to deliver any software.

2.4.6 Calibration Software

There is no calibration software that is applicable to IES at this time.

2.4.7 Pipeline Processing Software

The pipeline processing software is the IES Ground System (IESGS). IESGS extracts IES CCSDS packets from the RPC collective data files stored on the RPC central data server at Imperial College. These packets are used to build ion and electron data products. The data products are grouped by date and written out to PDS compliant archive data files. One data file is created for each mode used in each day. IESGS also generates the labels for the archive data files. IES science products, archive and label files, and limited spectrograms are available to team scientists on the IESGS website.

2.4.8 Scientific Analysis Software

Spectrograms can be generated from the IES archive data. These spectrograms can illustrate electron and ion counts per energy level, elevation angle, or azimuth bin.

2.4.9 Documentation

The document directory contains documentation that is considered to be either necessary or simply useful for users to understand the archive data set. These documents are not necessarily appropriate for inclusion in the PDS catalog. Documents may be included in multiple forms (ASCII, PDF, MS Word, HTML with image file pointers, etc.). PDS standards require that any documentation deemed required for use of the data be available in some ASCII format. HTML and PostScript are acceptable as ASCII formats in addition to plain text. Images and drawings are included as separate files.

There will be a separate directory for each document that is to be archived. Each of the document directories will include the document in plain text (ASCII) and the document in another format (i.e. .DOC or .PDF). There will also be a single label file that describes all the different formats of the included

documents. When reformatting to plain text affects the information content, this will be noted in the label file.

2.4.10 Derived and other Data Products

The IES higher level (derived) data products may include plasma density, flow velocity, ion and electron pressure, ion and electron temperature, and ion and electron pitch angle distributions. Many of these calculations will require co-operation with other RPC instruments: Calculations of ion moments require some composition data (e.g. the mean mass to charge ratio)..

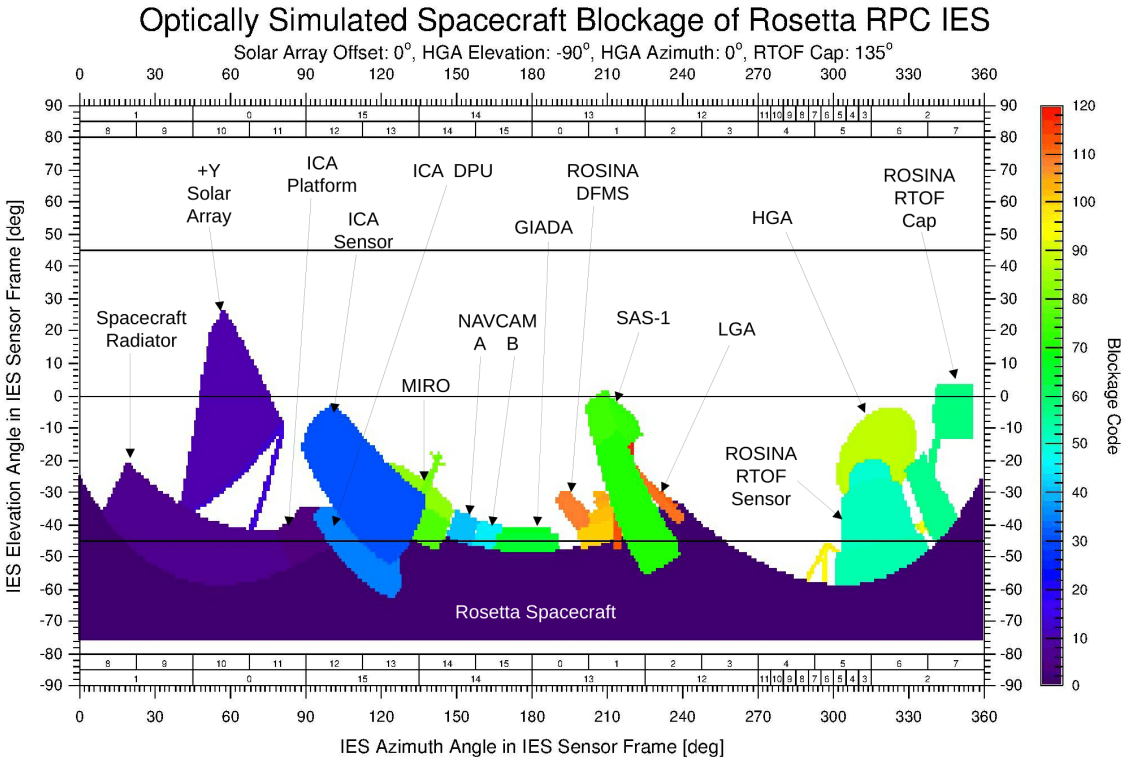
2.4.11 Ancillary Data Usage

Information on additional events may be desirable, if these events affect IES data (e.g. sweeping of the LAP voltages may affect the spacecraft electron sheath and therefore IES electron data.)

2.5 Data Caveats

2.5.1 Blockage of Some Elevation Angle Bins

Several S/C structures and parts of other instruments block the IES FOV in portions of the most negative elevation angles. See the accompanying figure for an example. In particular, note that the positions of the solar arrays and the HGA in the IES FOV vary throughout the mission. Geometric Factor
The documented IES geometric factor was obtained by combining results of ray tracing, flight calibration, and laboratory measurements.



2.5.2 Geometric Factor

The documented IES geometric factor was obtained by combining results of ray tracing, flight calibration, and laboratory measurements.

2.5.3 Individual Anode (Azimuth) Characteristics

Sometimes anode data are combined on board in certain operating modes because of telemetry limitations. In those cases each of the anodes is given identical values (sum/number of anodes).

2.5.3.1 Ions

Anodes 13 and 14 are often noisy so caution is recommended when using data from them.

The so-called "fine anodes" (3 - 11) are sometimes combined on board in certain operating modes because of telemetry limitations. In those cases each of these 9 anodes are given identical values (the sum/9).

The data from individual fine anodes, when provided, suffer from crosstalk between these anodes and thus may not be reliable.

It was discovered after launch, apparently as a result of a light leak, that when the sun is incident between anodes 2 and 3, a high count rate is seen in anode 12. It is also seen in anode 13 when counts in 12 and 13 are combined.

2.5.3.2 Electrons

Occasionally the electron data exhibit a signal at narrow energy ranges between 200-2000 eV, appearing as short dashes in spectrograms, as a result of interference from a neighboring instrument (ICA). The occurrences are indicated in the flag column.

Anode 11 became noisy shortly after launch and since 16 September 2007 data from that anode have not been downloaded. In those operating modes for which data from 2 or more anodes are combined will result in the absence of data from one or more anodes adjacent to number 11.

It was discovered after launch, apparently as a result of a light leak, that when the sun is incident between elevation steps 6 and 7 a high count rate may be seen in anodes 8 - 10.

3. ARCHIVE FORMAT AND CONTENT

3.1 Format and Conventions

3.1.1 Deliveries and Archive Volume Format

The IES team will submit the archive to PSA and PDS electronically. PSA and PDS will be responsible for creating the physical volumes used for deep archiving.

3.1.2 Data Set ID Formation

RO-E/M/A/C/CAL/X/SS/D-RPCIES-x-phase-Vn.m

where:

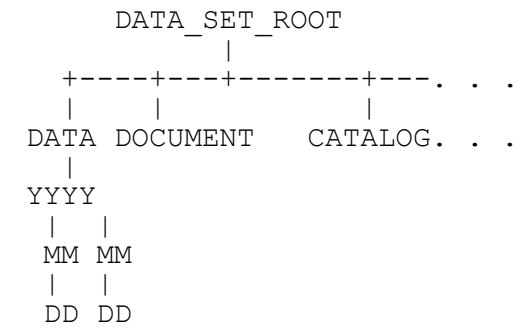
RO	= INSTRUMENT_HOST_ID
E/M/A/C/CAL/X/SS/D	= TARGET_ID (Earth/Mars/Asteroid/Comet/Calibration/Checkout/Solar System/Dust)
RPCIES	= INSTRUMENT_ID
x	= {2,3,5} CODMAC data processing level numbers.
phase	= Mission phase abbreviation (GRND, LEOP, CVP, CR1, EAR1, etc)
n.m	= Version number

Within each data set TARGET_NAME and TARGET_TYPE will then be used to identify the current target.

(Thus they will not stay the same within one data set, but data set id will.)

3.1.3 Data Directory Naming Convention

We intend to use a year/month/day directory hierarchy. The directory structure is covered in more detail in section 3.4.3.



3.1.4 Filenaming Convention

For uncalibrated and calibrated data there will be two IES data files generated per mode per day. There will be one file for electron data and one file for ion data. The file names will follow the following naming convention:

```
POSITION: 0123456789012345678.0123
FILENAME: RPCIESYYYYDDD_nnn_VV.EXT

where:

YYYY = Year
DDD  = Day of Year
```

```
nnn  = ELC (electron) or ION (ion)
VV   = Archive product version
EXT  = LBL or TAB
```

3.2 Standards Used in Data Product Generation

3.2.1 PDS Standards

IES complies to PDS version 3, and we use version 3.6 of the PDS standard reference.

3.2.2 Time Standards

Time(UTC) in LBL files: yyyy-doyThh:mm:ss.sss
Time(UTC) in TAB files: yyyy-doyThh:mm:ss.sss
Spacecraft Clock (OBT) in LBL files: "1/nnnnnnnnnn"
Spacecraft Clock (OBT) in TAB files: nnnnnnnnnn

3.2.3 Reference Systems

IES field of view definitions, anode and elevation sizes, their orientations with respect to the spacecraft and frame definitions are described in the SPICE kernels defined and provided by NAIF. At the time of writing of this document, the filenames for the latest versions of these instrument kernel and frame kernel files were ROS_RPC_V17.TI and ROS_V25.TF respectively.
These kernels can be loaded and used to determine the position and attitude of the spacecraft, orientation of the IES instrument, look directions of anodes and elevations in any coordinate system at any given time. The reference frame internal to the instrument is called ROS_RPC_IES. This can be used as a reference when converting pointing information to and from other reference frames such as the Rosetta spacecraft, comet-centric or helio-centric frames.

Data Validation

Data will be scanned for internal consistency when decommutating to edited raw format. Derived data will be compared to independent measurements by other instruments when possible. Before archiving a data set from some mission phase, this set will have been used internally by RPC scientists. It is planned to base all scientific analysis on the data products formatted. To actually have the data used by scientists before delivery to archive is considered the best way of revealing problems, and this is the approach taken by IES.

After submission a PDS peer review will assess the data set and documentation for compliance and scientific usability. The peer review is typically done once for the initial submission and all subsequent submissions are merely checked for conformance to the standards put forth in this document. There will also be peer reviews from the Rosetta archive team as the data is made ready for ingestion into the PSA.

3.3 Content

3.3.1 Volume Set

The IES archive will be submitted electronically, so there will initially be one volume for the entire dataset. PDS will create physical volumes for deep archiving. PSA requires no physical volumes, as the PSA is a completely online system.

3.3.2 Data Set

Our naming convention for the data set will follow the same principles as the DATA_SET_ID thus.

```
DATA_SET_NAME="ROSETTA-ORBITER E/M/A/C/CAL/X/SS/D RPCIES d PHASE
Vm.n"

where:

ROSETTA-ORBITER          = INSTRUMENT_HOST_NAME

E/M/A/C/CAL/X/SS/D       = TARGET_NAME (EARTH MARS ASTEROID COMET
                         CALIBRATION CHECKOUT SOLAR SYSTEM DUST)


RPCIES                   = INSTRUMENT_ID

d                        = CODMAC data processing level numbers 2,3
 or 5.

PHASE                    = Mission phase abbreviation (GRND, LEOP,
 CVP, CR1, EAR1, etc)

Vm.n                     = Version number
```

One data set will be used for each processing level. Multiple targets will be used for each data set and within each data set TARGET_ID will be used to identify the current target (Thus they will not stay the same within one data set, but data set id will). The data set name fits in the full length thus 60 characters.

3.3.3 Directories

3.3.3.1 Root Directory

Table 2: Root Directory Contents

File Name	File Contents
AAREADME.TXT	This file completely describes the Volume organization and contents
VOLDESC.CAT	A description of the contents of this Volume in a PDS format readable by both humans and computers
CATALOG/	Catalog directory
DOCUMENT/	Document directory
INDEX/	Index directory
DATA/	Data directory
BROWSE/	Browse directory
CALIB/	Calibration data directory

3.3.3.2 Catalog Directory

Table 3: Catalog Directory Contents	
File Name	**File Contents**
CATINFO.TXT	A description of the contents of this directory
DATASET.CAT	PDS Data Set catalog description of all the IES data files
INSTHOST.CAT	PDS instrument host (spacecraft) catalog description of the Rosetta orbiter spacecraft
INST.CAT	PDS instrument catalog description of the IES instrument
MISSION.CAT	PDS mission catalog description of the Rosetta mission
PERSON.CAT	PDS personnel catalog description of IES Team members and other persons involved with generation of IES Data Products
REF.CAT	IES-related references mentioned in other *.CAT files
SOFTWARE.CAT	Software catalog file

3.3.3.3 Index Directory

This directory contains the index files generated by the ESA S/W PVV.

3.3.3.4 Browse Directory and Browse Files

Spectrograms or other browse products may be introduced at a later date.

3.3.3.5 Geometry Directory

The geometry directory will contain any necessary instrument or frame kernels.

3.3.3.6 Document Directory

Table 4: Document Directory Contents	
File Name	**File Contents**
DOCINFO.TXT	A description of the contents of this directory and all subdirectories.
IES_EAICD/	Directory containing the IES EAICD document
IES_EAICD/IES_EAICD.DOC	The IES Experiment-Archive Interface Control Document as a MS Word doc
IES_EAICD/IES_EAICD.TXT	The IES Experiment-Archive Interface Control Document in plain text
IES_EAICD/IES_EAICD.LBL	A PDF detached label that describes IES_EAICD.HTM, IES_EAICD.ASC, and IES_EAICD.PDF

3.3.3.7 Data Directory

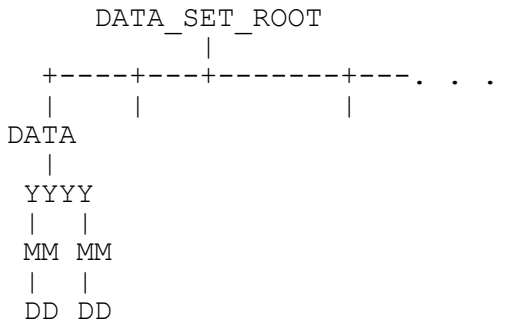
The data directory will contain .TAB files that have the archive data in fixed width, comma separated columns corresponding to PDS table objects. Accompanying each .TAB file will be a label file (.LBL) containing metadata about the archive.

4. DETAILED INTERFACE SPECIFICATIONS

4.1 Structure and Organization Overview

See section 3.1.3 for general overview.

Now as defined in section 3.1.3 we have the following structure for the DATA directory.



4.2 Data Sets, Definition and Content

IES data is archived in PDS table objects. Each line represents a set of electron or ion counts for the azimuth bin groups at a given time, energy, and elevation. The following columns will be first in each archive file:

Spacecraft Event Time (UTC)	UTC time at the beginning of sample integration. UTC time is converted from the spacecraft clock time using the SPICE toolkit.
Mode	Instrument mode, which defines the structure of the energy-elevation-azimuth collapse for the counts.
Energy Start Step	Each electron or ion count occurs within a specified energy range. This is the number of the step that defines the start of the range of energy values.
Energy Stop Step	Each electron or ion count occurs within a specified energy range. This is the number of the step that defines the end of the range of energy values.
Angle Start Step	Each electron or ion count occurs within a specified elevation angle range. This is the number of the step that defines the start of the range of angle values.

Angle Stop Step	Each electron or ion count occurs within a specified elevation angle range. This is the number of the step that defines the end of the range of angle values.

Following these columns is a series of azimuth columns. The value represents the number of electrons or ions observed in the azimuth bin (commonly referred to as "counts") at the given time, energy, and elevation. These values are transmitted in groups of azimuth bins, which we expand by dividing the value by the number of azimuth bins in the group.

4.3 Data Product Design

4.3.1 Data Product Uncalibrated Design

Example of edited raw data detached label file (e.g. RPCIES2014323_ELC_V2.LBL):

```
PDS_VERSION_ID               = PDS3
DATA_SET_ID                  = "RO-C-RPCIES-2-ESC1-V1.0"
DATA_SET_NAME                = "
    ROSETTA-ORBITER 67P RPCIES 2 ESC1 V1.0"
STANDARD_DATA_PRODUCT_ID     = "ELECTRON"
PRODUCT_ID                   = "RPCIES2014323_ELC_V2"
PRODUCT_TYPE                 = "EDR"
PROCESSING_LEVEL_ID          = "2"
PRODUCT_CREATION_TIME        = 2015-10-30T15:48:24.177
PRODUCT_VERSION_ID           = "1.0"
LABEL_REVISION_NOTE          = "RELEASE VERSION 1.0"
INSTRUMENT_MODE_ID           = "N/A"
INSTRUMENT_MODE_DESC         = "N/A"
ROSETTA:PIPELINE_VERSION_ID  = "3.7"

RECORD_TYPE                  = FIXED_LENGTH
RECORD_BYTES                 = 387
FILE_RECORDS                 = 174080
MD5_CHECKSUM                 = "fcea5a47790a89cb12f4bd23a8954800"

START_TIME                   = 2014-11-19T00:00:34.336
STOP_TIME                    = 2014-11-19T23:54:10.365
SPACECRAFT_CLOCK_START_COUNT = "1/374975963"
```

```
SPACECRAFT_CLOCK_STOP_COUNT  = "1/375061979"

MISSION_NAME                 = "INTERNATIONAL ROSETTA MISSION"
MISSION_ID                   = "ROSETTA"
MISSION_PHASE_NAME           = "COMET ESCORT 1"
TARGET_NAME                  = "67P/CHURYUMOV-GERASIMENKO 1 (1969 R1)"
TARGET_TYPE                  = "COMET"
INSTRUMENT_HOST_NAME         = "ROSETTA-ORBITER"
INSTRUMENT_HOST_ID           = "RO"
INSTRUMENT_ID                = "RPCIES"
INSTRUMENT_NAME              = "
    ROSETTA PLASMA CONSORTIUM - ION AND ELECTRON SENSOR"
INSTRUMENT_TYPE              = "PLASMA INSTRUMENT"

COORDINATE_SYSTEM_ID         = "N/A"
COORDINATE_SYSTEM_NAME       = "N/A"
NOTE                         = "The values of the keywords
  SC_SUN_POSITION_VECTOR, SC_TARGET_POSITION_VECTOR,
  SC_TARGET_VELOCITY_VECTOR are related to the equatorial J2000
inertial
  frame. The values of SUB_SPACECRAFT_LATITUDE and
SUB_SPACECRAFT_LONGITUDE
  refer to the Cheops reference frame. All values are computed for the
time
  t=START_TIME. Distances are given in <km>, velocities in <km/s>, and
  angles in <deg>.
  Unit for SC_SUN_POSITION_VECTOR is km
  Unit for SC_TARGET_POSITION_VECTOR is km
  Unit for SC_TARGET_VELOCITY_VECTOR is km/s
  Unit for SPACECRAFT_ALTITUDE is km"

PRODUCER_ID                  = "RPC_IES_TEAM"
PRODUCER_FULL_NAME           = "BRAD TRANTHAM"
```

```
PRODUCER_INSTITUTION_NAME       = "SOUTHWEST RESEARCH INSTITUTE, SAN
ANTONIO"
DATA_QUALITY_ID             = "0"
DATA_QUALITY_DESC           = "Data quality not assessed"

SC_SUN_POSITION_VECTOR      =
    (-2.455E8, 3.119E8, 1.919E8)

SC_TARGET_POSITION_VECTOR   =
    (1.921E1, 2.136E1, -1.026E1)

SC_TARGET_VELOCITY_VECTOR   =
    (1.921E1, 2.136E1, -1.026E1)

SPACECRAFT_ALTITUDE         = 2.855E1

SUB_SPACECRAFT_LATITUDE     = 1.04E-1
SUB_SPACECRAFT_LONGITUDE    = -6.972E1

DESCRIPTION                 = "
    This file contains IES raw electron sensor counts acquired during
    the Comet Escort 1 between 2014-11-19T00:00:34.336 and
    2014-11-19T23:54:10.365."

^HEADER                     = ("RPCIES2014323_ELC_V2.TAB", 1)
^TABLE                      = ("RPCIES2014323_ELC_V2.TAB", 2)
OBJECT                      = HEADER
  HEADER_TYPE               = "TEXT"
  INTERCHANGE_FORMAT        = ASCII
  BYTES                     = 387
  RECORDS                   = 1
```

```
  DESCRIPTION                 = "Row of comma delimited, quoted column
names"
END_OBJECT                  = HEADER

OBJECT                      = TABLE
  INTERCHANGE_FORMAT        = ASCII
  ROWS                      = 174080
  COLUMNS                   = 23
  ROW_BYTES                 = 387

  OBJECT                    = COLUMN
    NAME                    = "SPACECRAFT EVENT TIME (UTC)"
    COLUMN_NUMBER           = 1
    DATA_TYPE               = TIME
    START_BYTE              = 1
    BYTES                   = 21
    FORMAT                  = "A21"
    DESCRIPTION             = "
      This field contains the UTC time at the spacecraft at the
beginning
      of the sample integration. This field has been generated from
      the spacecraft clock counter using the SPICE toolkit and
appropriate
      leap seconds and spacecraft clock kernels. Time is provided in
the
      standard PDS month/day format (i.e. 2005-03-05T00:00:00.215).
      All records from a single integration are assigned the same
time.
      The amount of integration time is governed by the science mode.
      Details can be found in CALIB\ENERGY_STEPS.TAB. A complete
integration
      requires the instrument to sweep through 16 azimuth directions
      azimuth directions for each of the 128 energy steps. Each
azimuth
```

```
    step takes 1/16th of an energy step to complete.."

END_OBJECT                  = COLUMN

OBJECT                      = COLUMN
  NAME                      = "MODE"
  COLUMN_NUMBER             = 2
  START_BYTE                = 23
  BYTES                     = 11
  FORMAT                    = "A11"
  DATA_TYPE                 = CHARACTER
  DESCRIPTION               = "
    Instrument mode, which determines the values used for the energy
    and elevation steps."
END_OBJECT                  = COLUMN

OBJECT                      = COLUMN
  NAME                      = "ENERGY_START_STEP"
  COLUMN_NUMBER             = 3
  START_BYTE                = 35
  BYTES                     = 16
  DATA_TYPE                 = ASCII_INTEGER
  FORMAT                    = "I16"
  DESCRIPTION               = "
    The number of the energy step that starts this range"
END_OBJECT                  = COLUMN

OBJECT                      = COLUMN
  NAME                      = "ENERGY_STOP_STEP"
  COLUMN_NUMBER             = 4
  START_BYTE                = 52
  BYTES                     = 16
```

```
    DATA_TYPE                = ASCII_INTEGER
    FORMAT                   = "I16"
    DESCRIPTION              = "
      The number of the energy step that ends this range"
  END_OBJECT                 = COLUMN

  OBJECT                     = COLUMN
    NAME                     = "ANGLE_START_STEP"
    COLUMN_NUMBER            = 5
    START_BYTE               = 69
    BYTES                    = 16
    DATA_TYPE                = ASCII_INTEGER
    FORMAT                   = "I16"
    DESCRIPTION              = "
      The number of the elevation step that starts this range"
  END_OBJECT                 = COLUMN

  OBJECT                     = COLUMN
    NAME                     = "ANGLE_STOP_STEP"
    COLUMN_NUMBER            = 6
    START_BYTE               = 86
    BYTES                    = 16
    DATA_TYPE                = ASCII_INTEGER
    FORMAT                   = "I16"
    DESCRIPTION              = "
      The number of the elevation step that ends this range"
  END_OBJECT                 = COLUMN

  OBJECT                     = COLUMN
    NAME                     = "AZIMUTH 0 COUNTS"
    COLUMN_NUMBER            = 7
    START_BYTE               = 103
```

```
    BYTES                    = 16
    DATA_TYPE                = ASCII_REAL
    FORMAT                   = "F16.4"
    MISSING_CONSTANT         = "-1.000"
    DESCRIPTION              = "
      This field contains electron counts observed in
      azimuth bin 0 divided by the size of the azimuth
      bin grouping. A fill value of -1 is used when data is
      not available for this bin."
  END_OBJECT                 = COLUMN
  OBJECT                     = COLUMN
    NAME                     = "AZIMUTH 1 COUNTS"
    COLUMN_NUMBER            = 8
    START_BYTE               = 120
    BYTES                    = 16
    DATA_TYPE                = ASCII_REAL
    FORMAT                   = "F16.4"
    MISSING_CONSTANT         = "-1.000"
    DESCRIPTION              = "
      This field contains electron counts observed in
      azimuth bin 1 divided by the size of the azimuth
      bin grouping. A fill value of -1 is used when data is
      not available for this bin."
  END_OBJECT                 = COLUMN
  OBJECT                     = COLUMN
    NAME                     = "AZIMUTH 2 COUNTS"
    COLUMN_NUMBER            = 9
    START_BYTE               = 137
    BYTES                    = 16
    DATA_TYPE                = ASCII_REAL
    FORMAT                   = "F16.4"
    MISSING_CONSTANT         = "-1.000"
```

```
    DESCRIPTION              = "
      This field contains electron counts observed in
      azimuth bin 2 divided by the size of the azimuth
      bin grouping. A fill value of -1 is used when data is
      not available for this bin."
  END_OBJECT                 = COLUMN
  OBJECT                     = COLUMN
    NAME                     = "AZIMUTH 3 COUNTS"
    COLUMN_NUMBER            = 10
    START_BYTE               = 154
    BYTES                    = 16
    DATA_TYPE                = ASCII_REAL
    FORMAT                   = "F16.4"
    MISSING_CONSTANT         = "-1.000"
    DESCRIPTION              = "
      This field contains electron counts observed in
      azimuth bin 3 divided by the size of the azimuth
      bin grouping. A fill value of -1 is used when data is
      not available for this bin."
  END_OBJECT                 = COLUMN
  OBJECT                     = COLUMN
    NAME                     = "AZIMUTH 4 COUNTS"
    COLUMN_NUMBER            = 11
    START_BYTE               = 171
    BYTES                    = 16
    DATA_TYPE                = ASCII_REAL
    FORMAT                   = "F16.4"
    MISSING_CONSTANT         = "-1.000"
    DESCRIPTION              = "
      This field contains electron counts observed in
      azimuth bin 4 divided by the size of the azimuth
      bin grouping. A fill value of -1 is used when data is
```

```
    not available for this bin."
  END_OBJECT                = COLUMN
  OBJECT                    = COLUMN
    NAME                    = "AZIMUTH 5 COUNTS"
    COLUMN_NUMBER           = 12
    START_BYTE              = 188
    BYTES                   = 16
    DATA_TYPE               = ASCII_REAL
    FORMAT                  = "F16.4"
    MISSING_CONSTANT        = "-1.000"
    DESCRIPTION             = "
      This field contains electron counts observed in
      azimuth bin 5 divided by the size of the azimuth
      bin grouping. A fill value of -1 is used when data is
      not available for this bin."
  END_OBJECT                = COLUMN
  OBJECT                    = COLUMN
    NAME                    = "AZIMUTH 6 COUNTS"
    COLUMN_NUMBER           = 13
    START_BYTE              = 205
    BYTES                   = 16
    DATA_TYPE               = ASCII_REAL
    FORMAT                  = "F16.4"
    MISSING_CONSTANT        = "-1.000"
    DESCRIPTION             = "
      This field contains electron counts observed in
      azimuth bin 6 divided by the size of the azimuth
      bin grouping. A fill value of -1 is used when data is
      not available for this bin."
  END_OBJECT                = COLUMN
  OBJECT                    = COLUMN
    NAME                    = "AZIMUTH 7 COUNTS"
```

```
    COLUMN_NUMBER            = 14
    START_BYTE               = 222
    BYTES                    = 16
    DATA_TYPE                = ASCII_REAL
    FORMAT                   = "F16.4"
    MISSING_CONSTANT         = "-1.000"
    DESCRIPTION              = "
      This field contains electron counts observed in
      azimuth bin 7 divided by the size of the azimuth
      bin grouping. A fill value of -1 is used when data is
      not available for this bin."
  END_OBJECT                 = COLUMN
  OBJECT                     = COLUMN
    NAME                     = "AZIMUTH 8 COUNTS"
    COLUMN_NUMBER            = 15
    START_BYTE               = 239
    BYTES                    = 16
    DATA_TYPE                = ASCII_REAL
    FORMAT                   = "F16.4"
    MISSING_CONSTANT         = "-1.000"
    DESCRIPTION              = "
      This field contains electron counts observed in
      azimuth bin 8 divided by the size of the azimuth
      bin grouping. A fill value of -1 is used when data is
      not available for this bin."
  END_OBJECT                 = COLUMN
  OBJECT                     = COLUMN
    NAME                     = "AZIMUTH 9 COUNTS"
    COLUMN_NUMBER            = 16
    START_BYTE               = 256
    BYTES                    = 16
    DATA_TYPE                = ASCII_REAL
```

```
    FORMAT                  = "F16.4"
    MISSING_CONSTANT        = "-1.000"
    DESCRIPTION             = "
      This field contains electron counts observed in
      azimuth bin 9 divided by the size of the azimuth
      bin grouping. A fill value of -1 is used when data is
      not available for this bin."
  END_OBJECT                = COLUMN
  OBJECT                    = COLUMN
    NAME                    = "AZIMUTH 10 COUNTS"
    COLUMN_NUMBER           = 17
    START_BYTE              = 273
    BYTES                   = 16
    DATA_TYPE               = ASCII_REAL
    FORMAT                  = "F16.4"
    MISSING_CONSTANT        = "-1.000"
    DESCRIPTION             = "
      This field contains electron counts observed in
      azimuth bin 10 divided by the size of the azimuth
      bin grouping. A fill value of -1 is used when data is
      not available for this bin."
  END_OBJECT                = COLUMN
  OBJECT                    = COLUMN
    NAME                    = "AZIMUTH 11 COUNTS"
    COLUMN_NUMBER           = 18
    START_BYTE              = 290
    BYTES                   = 16
    DATA_TYPE               = ASCII_REAL
    FORMAT                  = "F16.4"
    MISSING_CONSTANT        = "-1.000"
    DESCRIPTION             = "
      This field contains electron counts observed in
```

```
    azimuth bin 11 divided by the size of the azimuth
    bin grouping. A fill value of -1 is used when data is
    not available for this bin."
  END_OBJECT                = COLUMN
  OBJECT                    = COLUMN
    NAME                    = "AZIMUTH 12 COUNTS"
    COLUMN_NUMBER           = 19
    START_BYTE              = 307
    BYTES                   = 16
    DATA_TYPE               = ASCII_REAL
    FORMAT                  = "F16.4"
    MISSING_CONSTANT        = "-1.000"
    DESCRIPTION             = "
      This field contains electron counts observed in
      azimuth bin 12 divided by the size of the azimuth
      bin grouping. A fill value of -1 is used when data is
      not available for this bin."
  END_OBJECT                = COLUMN
  OBJECT                    = COLUMN
    NAME                    = "AZIMUTH 13 COUNTS"
    COLUMN_NUMBER           = 20
    START_BYTE              = 324
    BYTES                   = 16
    DATA_TYPE               = ASCII_REAL
    FORMAT                  = "F16.4"
    MISSING_CONSTANT        = "-1.000"
    DESCRIPTION             = "
      This field contains electron counts observed in
      azimuth bin 13 divided by the size of the azimuth
      bin grouping. A fill value of -1 is used when data is
      not available for this bin."
  END_OBJECT                = COLUMN
```

```
OBJECT                      = COLUMN
  NAME                      = "AZIMUTH 14 COUNTS"
  COLUMN_NUMBER             = 21
  START_BYTE                = 341
  BYTES                     = 16
  DATA_TYPE                 = ASCII_REAL
  FORMAT                    = "F16.4"
  MISSING_CONSTANT          = "-1.000"
  DESCRIPTION               = "
    This field contains electron counts observed in
    azimuth bin 14 divided by the size of the azimuth
    bin grouping. A fill value of -1 is used when data is
    not available for this bin."
END_OBJECT                  = COLUMN
OBJECT                      = COLUMN
  NAME                      = "AZIMUTH 15 COUNTS"
  COLUMN_NUMBER             = 22
  START_BYTE                = 358
  BYTES                     = 16
  DATA_TYPE                 = ASCII_REAL
  FORMAT                    = "F16.4"
  MISSING_CONSTANT          = "-1.000"
  DESCRIPTION               = "
    This field contains electron counts observed in
    azimuth bin 15 divided by the size of the azimuth
    bin grouping. A fill value of -1 is used when data is
    not available for this bin."
END_OBJECT                  = COLUMN
OBJECT                      = COLUMN
  NAME                      = "QUALITY FLAGS"
  COLUMN_NUMBER             = 23
  START_BYTE                = 375
```

```
    DATA_TYPE                = CHARACTER
    BYTES                    = 11
    FORMAT                   = "A11"
    DESCRIPTION              = "
     These flags describe the quality of the data.
     The quality is coded in a 8 byte string. Each character can have
     the following values:
      VALUE:    MEANING:
        x       property described by flag is still unknown
        0       no disturbance, good quality
        1..9    specific disturbance/problems, see below

     Description of the specific flags:

     FLAG-STRING FLAG DESCRIPTION
     87654321
     ::::::::------ 1 OVERALL QUALITY:
     :::::::          x = overall quality not assessed
     :::::::          0 = quality good without any processing
     :::::::          1 = quality good after data processing
     :::::::            2 = quality improved by data processing, still
not good
     :::::::          3 = data disturbed by unknown source
     :::::::          4 = TBD
     :::::::          5 = TBD
     :::::::          6 = TBD
     :::::::          7 = TBD
     :::::::          8 = TBD
     :::::::          9 = quality bad
     :::::::
     :::::::------- 2 HIGH BACKGROUND PRESSURE
     ::::::           x = impact not assessed
```

```
     ::::::           0 = no disturbance
     ::::::           1 = disturbance eliminated during data analysis
     ::::::           2 = data disturbed
     ::::::
     ::::::-------- 3 HIGH DUST FLUX
     :::::            x = disturbance not assessed
     :::::            0 = no disturbance
     :::::            1 = disturbance eliminated during data analysis
     :::::            2 = data disturbed
     :::::
     :::::--------- 4 TBD
     ::::             x = no assessment
     ::::
     ::::---------- 5 TBD
     :::              x = no assessement
     :::
     :::----------- 6 TBD
     ::               x = no assessment
     ::
     ::------------ 7 TBD
     :                x = no assessment
     :
     :------------- 8 TBD
                      x = no assessment"
  END_OBJECT                = COLUMN
END_OBJECT                  = TABLE

END
```