

# SPIRE

**SUBJECT: ICD Structure - Mechanical I/F**

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## Change Record

ISSUE	DATE	
0.1	November 2000	New document
0.2	February 2001	Update, updated topology, updated I/F drawings and cable lengths


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## Glossary

BSM	Beam Steering Mechanism
BSBR	Thermal BuSBaR between cooler and detectors
CBFL1	Common BaFfLe 1
CFIL1	Common FILter 1 (entry 4 K enclosure)
CM3-5	Common Mirror 3-5
CoG	Centre of Gravity
CSTR1	Cold STRap 1 (level 0, Cryostat to base evaporator heat switch)
CSTR2	Cold STRap 2 (level 0, Cryostat to base sorption pump heat switch)
CSTR3	Cold STRap 3 (level 0, Cryostat to photometer detector box)
CSTR4	Cold STRap 4 (level 1, Vent line to optical bench)
FOB	First Optical Bench
FOR	Firs Optical Reference
<sup>3</sup> He Cooler	<sup>3</sup> He cooler (located in photometer side of the instrument)
JFBX	JFET box
MoI	Moments of inertia
NA	Not Applicable
PBFL2	Photometer BaFfLe 2
PDBX	Photometer Detector Box
PCAL	Photometer CALibrationsource
PDBX	Photometer Detector BoX
PDIC1	Photometer DIChroic 1
PDIC2	Photometer DIChroic 2
PFIL2	Photometer FILter (entrance PDBX)
PFIL3	Photometer FILter 2 (4 K-2 K enclosure)
PFIL4L/5L	Photometer FILter 4 and 5 at nose PLW
PFIL4M/5M	Photometer FILter 4 and 5 at nose PMW
PFIL4S/5S	Photometer FILter 4 and 5 at nose PSW
PLW	Photometer Long Wave detector
PM6-11	Photometer Mirror 6 to 11
PMW	Photometer Medium Wave detector
SBFL1	Spectrometer BaFfLe 1
SBFL2	Spectrometer BaFfLe 2
SBS1	Spectrometer Beam Splitter 1
SBS2	Spectrometer Beam Splitter 2
SCAL	Spectrometer CALibration source
SCCA	Spectrometer Corner Cube +X
SCCB	Spectrometer Corner Cube -X
SDBX	Spectrometer Detector BoX
SFIL2	Spectrometer FILter 2 (4 K - 2 K enclosure)
SFIL3L	Spectrometer FILter 3 (long wave)
SFIL3S	Spectrometer FILter 3 (short wave)
SFIL4L/5L	Spectrometer FILter at nose SLW



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SFIL4S/5S	Spectrometer FILter at nose SSW
SHUT	Shutter mechanism
SM8A-12A	Spectrometer Mirror 8-12 +X chain
SM8B-12B	Spectrometer Mirror 8-12 -X chain
SM6-7	Spectrometer Mirror6-7
SMEC	Spectrometer MEChanism
SOB	SPIRE Optical Bench
SOR	Spire Optical Reference
SPIRE	Spectral and Photometric Imaging REceiver

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## References

### Applicable Documents

- AD1 Instrument Requirements Document, issue 0.30, May 2000 (FIRST/SPIRE)  
AD2 Systems Budgets, issue 1.1, 14 June

### REFERENCE DOCUMENTS

- RD01 SPIRE Mirrors: Mass estimate. LAM.PJT.SPI.OPT.990001 ind 3, 20 May 2000  
RD02 SPIRE filters sub-system specification – Draft 1.0 – 16<sup>th</sup> May 2000  
RD03 BSM subsystem specification, 14 June 2000  
RD04 Cable routing, e-mail J. Bock, 16-Nov-2000  
RD05 Re: Cable routing, e-mail T. Cafferty, 18-Nov-2000  
RD06 ESA fax, B. Guillaume, SCI-PT/IFI/07222, 5-Nov-1999  
RD07 Instrument Interface Document part A, SCI-PT-IIDA-04624, issue 1.0, Sept. 2000

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## 1. INTRODUCTION

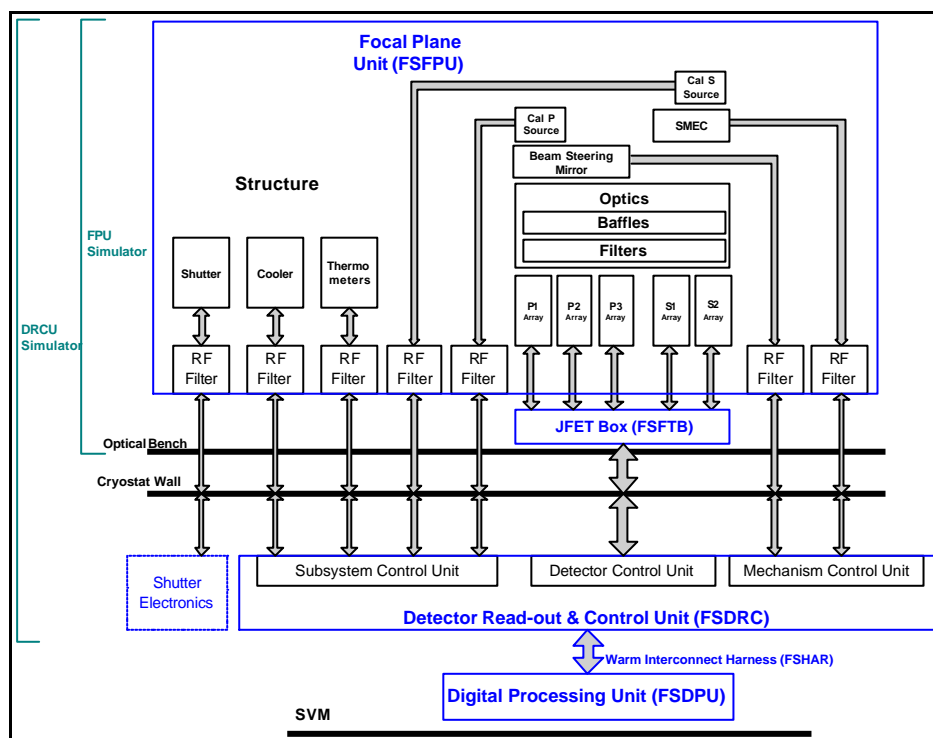
### 1.1 Purpose of Document

The purpose of this document is to define the interface between the SPIRE structure and all the sub-systems. Within this document the mechanical and thermal requirements with respect to these interfaces are given. The mechanical interface requirements are related to the actual interface (interface drawing), CoG, MOI, input spectra. The thermal interface requirements deals with contact area, surface finishes, thermal mass and cross sectional areas.


This issue is a draft issue. The purpose of this issue is to define the interface, such that an agreed baseline interface is available at the beginning of the final design process. During the final design process the definitions may be refined and possibly changed. This must be done following a formal change procedure, agreed by all parties involved in the design of a sub-system linked to that particular interface requirement.

The design of the instrument and its subsystems obeys the requirements as stated in RD1.

An overview in block diagram format of all the sub-systems and their relationships is given hereafter in block diagram: 1.1-1.

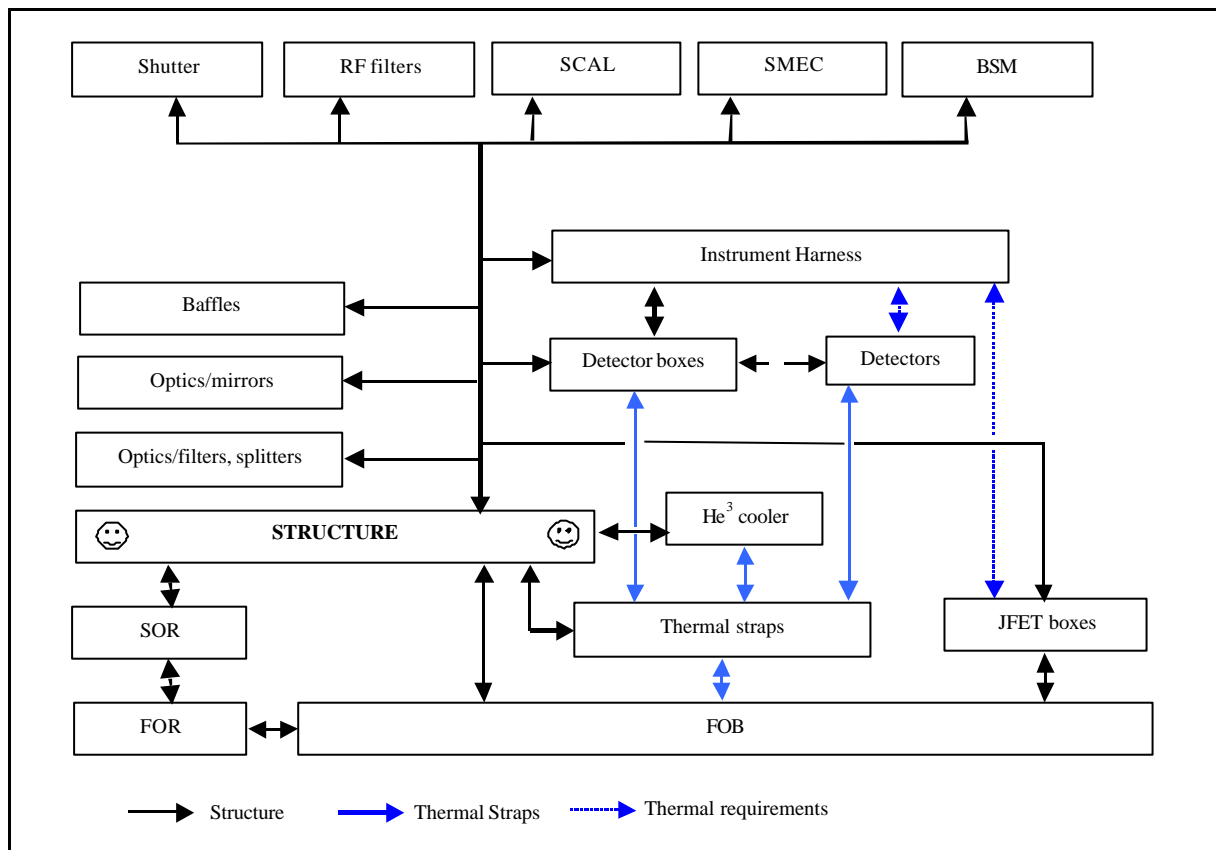


Block diagram: 1.1-1: SPIRE system

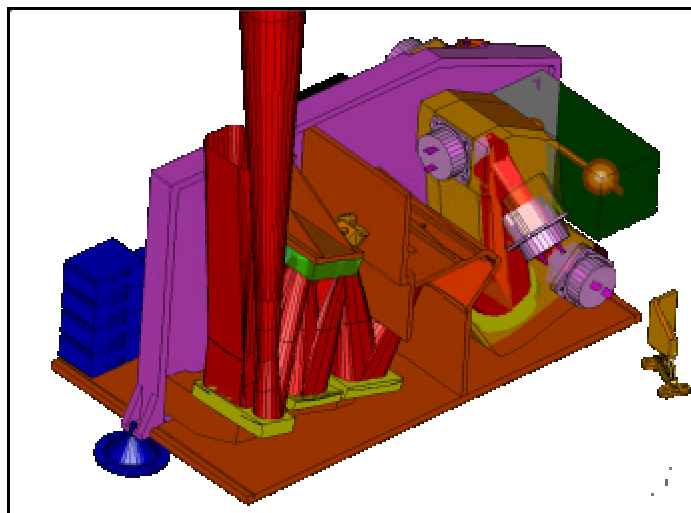
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## 1.2 SCOPE

The interfaces between the structure and the other subsystems are divided into the mechanical and thermal. For each interface the relevant requirements are listed. The interfaces considered are between the SPIRE structure and the relevant subsystems. Block diagram 1.2-1 gives a graphical interpretation of the interfaces defined, including mechanical and thermal requirements, in this document.



Block diagram: 1.2-1: Connections, mechanical and thermal, between the subsystems




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Figure 1.2-1: View of the inside of the instrument – photometer side, cover taken off

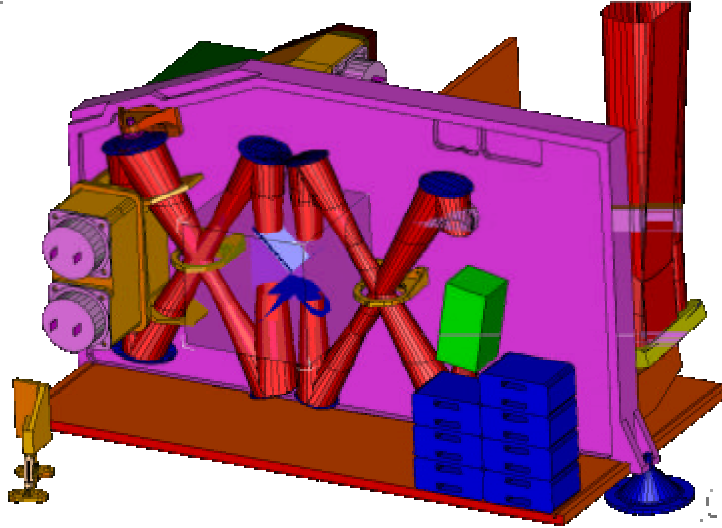


Figure 1.2-2: View of the inside of the instrument – spectrometer side, cover taken off

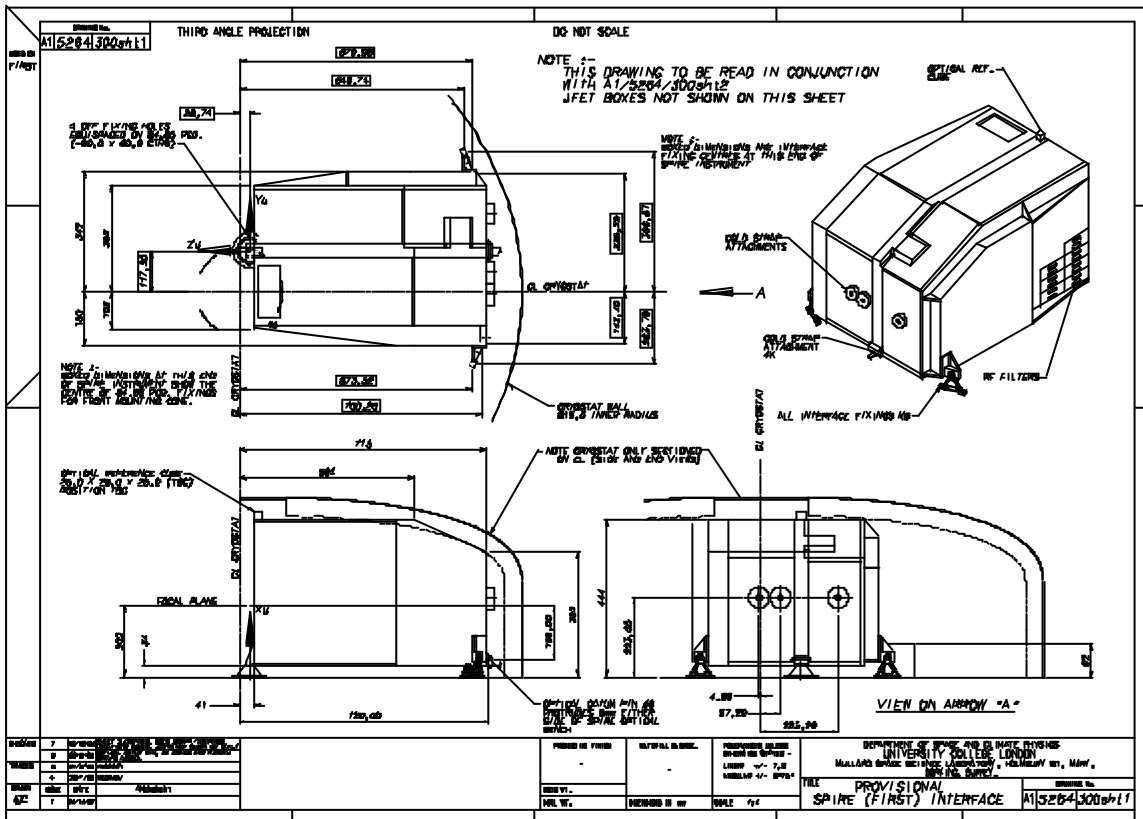


Figure 1.2-3: Interface drawing between structure and FOB-1



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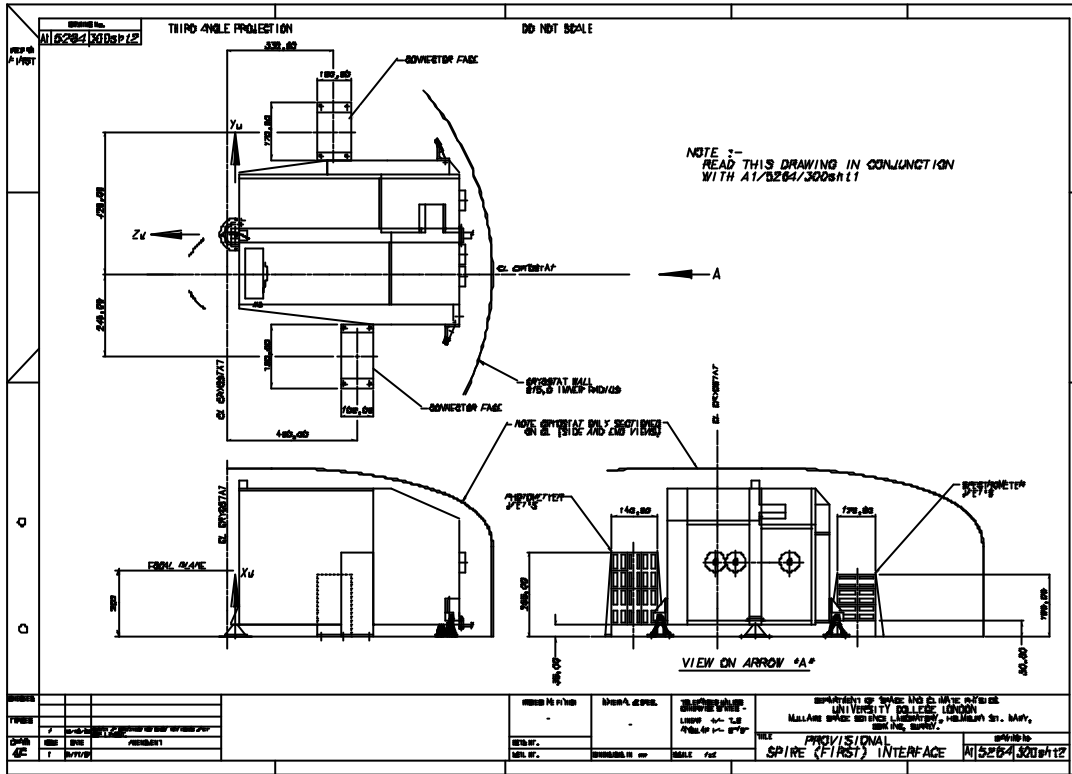
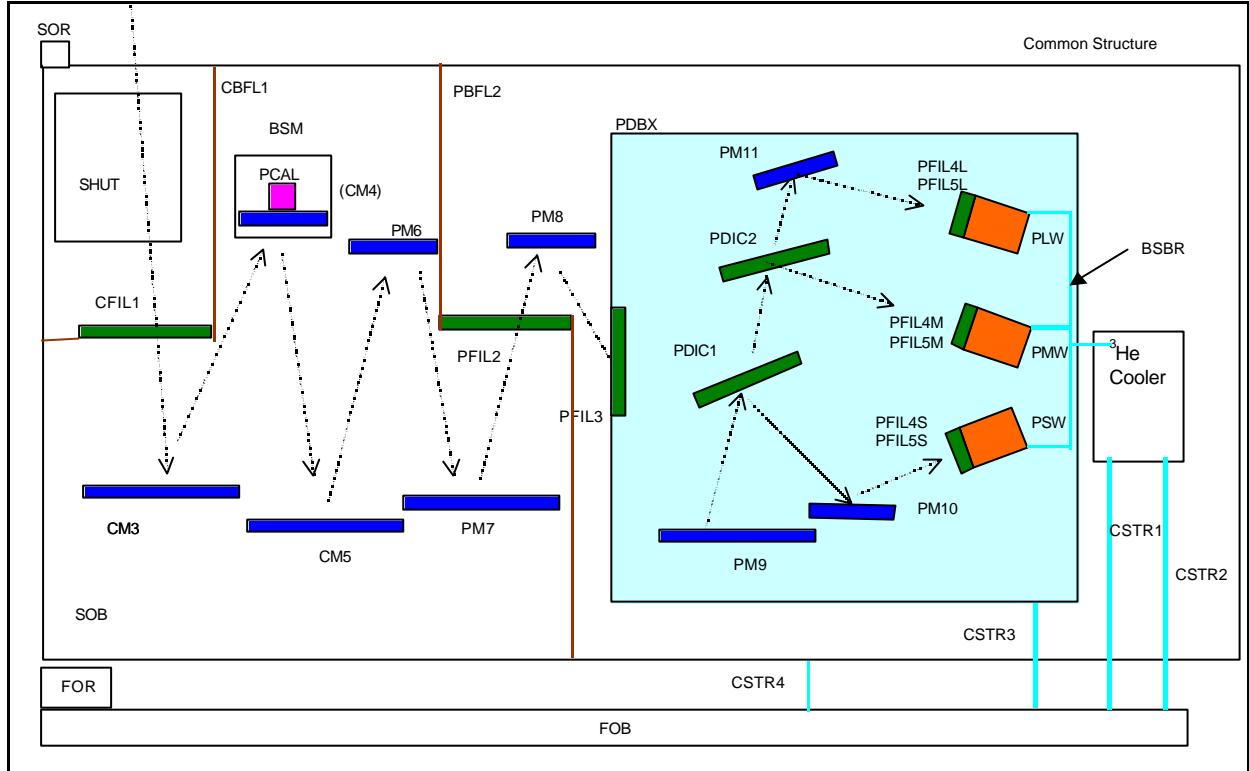


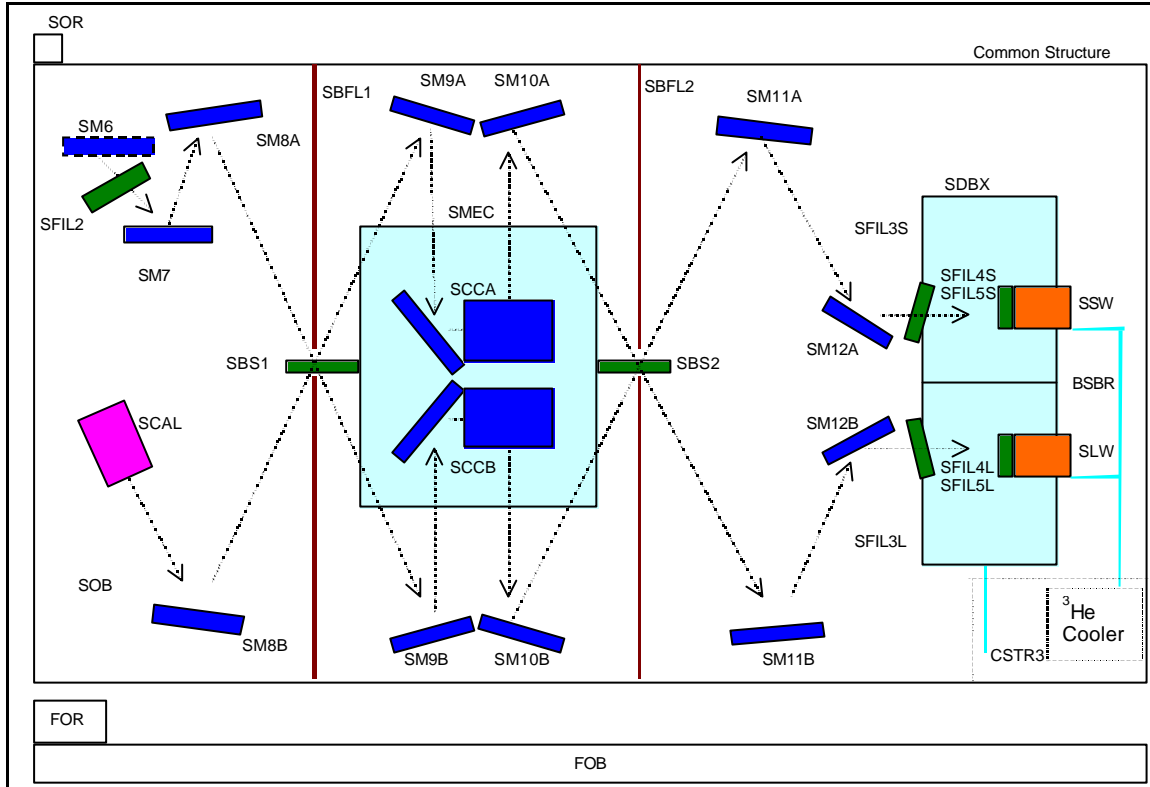
Figure 1.2-4: Interface drawing between structure and FOB-2

### 1.3 The sub-system requirements listed (mechanical and thermal)



Block diagram 1.3-1: Photometer side of the instrument

BSBR	Thermal BuSBar between cooler and detectors	PCAL	Photometer CALibrationsource
BSM	Beam Steering Mechanism	PDBX	Photometer Detector BoX
CBFL1	Common BaFfLe1	PDIC1	Photometer DIChroic 1
CFIL1	Common FILter 1 (entry 4 K enclosure)	PDIC2	Photometer DIChroic 2
CM3-5	Common Mirror 3-5	PFIL2	Photometer FILter (entrance PDBX)
CSTR1	Cold STRap 1 (level 0, Cryostat to base evaporator heat switch)	PFIL3	Photometer FILter 2 (4 K-2 K enclosure)
CSTR2	Cold STRap 2 (level 0, Cryostat to base sorption pump heat switch)	PFIL4L/5L	Photometer FILter 4 and 5 at nose PLW
CSTR3	Cold STRap 3 (level 0, Cryostat to photometer detector box)	PFIL4M/5M	Photometer FILter 4 and 5 at nose PMW
CSTR4	Cold STRap 4 (level 1, Vent line to optical bench)	PFIL4S/5S	Photometer FILter 4 and 5 at nose PSW
FOB	First Optical Bench	PLW	Photometer Long Wave detector
FOR	First Optical Reference	PM6-11	Photometer Mirror 6 to 11
<sup>3</sup> He Cooler	<sup>3</sup> He cooler (located in photometer side of the instrument)	PMW	Photometer Medium Wave detector
PBFL2	Photometer BaFfLe2	PSW	Photometer Short Wave detector
		SHUT	SHUTter mechanism
		SOB	Spire Optical Bench Panel
		SOR	SPIRE Optical Reference




Block diagram 1.3-2: Spectrometer side of the instrument

CSTR3	Cold STRap 3 (Level 0, cryostat base to SDBX, via PDBX)	SDBX	Spectrometer Detector BoX
FOB	First Optical Bench	SFIL2	Spectrometer FILter 2 (4 K - 2 K enclosure)
FOR	Firs Optical Reference	SFIL3L	Spectrometer FILter 3 (long wave)
<sup>3</sup> He Cooler	<sup>3</sup> He cooler (located in photometer side of the instrument)	SFIL3S	Spectrometer FILter 3 (short wave)
SBFL1	Spectrometer BaFfile 1	SFIL4L/5L	Spectrometer FILter at nose SLW
SBFL2	Spectrometer BaFfile 2	SFIL4S/5S	Spectrometer FILter at nose SSW
SBS1	Spectrometer Beam Splitter 1	SM6-7	Spectrometer Mirror6-7
SBS2	Spectrometer Beam Splitter 2	SM8A-12A	Spectrometer Mirror 8-12 +X chain
SCCA	Spectrometer Corner Cube +X	SM8B-12B	Spectrometer Mirror 8-12 -X chain
SCCB	Spectrometer Corner Cube -X	SMEC	Spectrometer MECHANISM
SCAL	Spectrometer CALibration source	SOB	Spire Optical Bench Panel
		SOR	Spire Optical Reference

See for the schematic location in the instrument blockdiagrams 1.3-1 and 1.3-2

Sub system ID	Description	Acronim	Section
1.2	Mirrors	CM3-CM5 PM6-PM11 SM6-SM12	2.1
1.2.1	Filters	CFIL1, PFIL2, PFIL3, SFIL2,	2.2



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		SFIL3L, SFIL3L	
1.2.1	Beam splitters	SBS1, SBS2	2.3
1.2.1	Dichroics	PDIC1, PDIC2	2.4
1.2.2	Straylight attenuation (Baffles and QMW-black)	PBFL1, PBFL2, SBFL1, SBFL2	2.5
1.3	He <sup>3</sup> Cooler		2.6
1.4.1-1	Photometer detectors	PSW,PMW, PLW	2.7
1.4.1-2	Spectrometer detectors	SSW,SLW	2.8
1.5.1	Beemsteering mechanism	BSM	2.9
1.5.2	Spectrometer mechanism	SMEC-m	2.10
1.5.3	Shutter mechanism	SHUT	2.11
1.6.1	Calibration source Photometer	PCAL	part of the BSM
1.6.2	Calibration source Spectrometer	SCAL	2.12
-	Thermal straps	CSTR1-4 BSBR	2.13
-	JFET box	JFBX	2.14
-	RF-filters		2.15

Table 1.3-1: Subsystems interfaces listed in this document

## 1.4 General requirements

### Mechanical environment

RD07 states the following with regard to the mechanical environment and testing. (General requirements for design and analysis on system level: section 9.4 RD07). The mechanical environment as defined in RD07 will change. So the values listed hereafter are provisional.

Safety factors and Sizing factors (9.4.1.2.3, RD07)

- Qualification factor  $j_q = 1.5$
- Yield safety factor  $j_y = 1.1$
- Ultimate safety factor  $j_u = 1.25$

Material properties/data should be assessed within 99% statistical probability at a confidence level of 95% (A-level) The following sizing factors shall be applied to the limit loads:

□ Load at yield failure point:


$$F_y = j_q \cdot j_y \cdot F_{limit}$$

$$F_y = 1.5 \cdot 1.1 \cdot F_{limit} = 1.65 \cdot F_{limit}$$

□ Load at ultimate failure point:

$$F_y = j_q \cdot j_u \cdot F_{limit}$$

$$F_y = 1.5 \cdot 1.25 \cdot F_{limit} = 1.875 \cdot F_{limit}$$

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For the SPIRE instrument as a whole the following flight limit load are given:

### Quasi-Static

RD07, section 9.4.1.2.4:

Quasi Static levels	Case 1	Case 2	Case 3
x-direction	25 g	-	-
y-direction	-	14 g	-
z-direction	-	-	14 g

Table 1.4-1: Instrument qualification levels for quasi-static vibration

All subsystems will be subjected to the same environment as defined above. The in table 1.4-1 defined accelerations are very important for limiting interface loads for the various sub-systems subjected to either sine, random or shock loads. At no point during these tests, the interface load between two subsystems is allowed to exceed the quasi-static interface load. This is the baseline for notching during these tests. For an accurate analysis of the interface loads reduced models of all subsystems are required. These models should give an accurate approximation of the first eigenmode and mass in all three orthogonal directions. There is one condition to be allowed a notch based upon this quasi-static interface load. The condition is that the subsystem is only allowed a notch at a given frequency band, if that system has significant effective mass in that frequency band in the excitation direction.

### Sine

RD07, section 9.5.3.3.2, gives as the requirement for the sine vibration test (instrument level):

Sweep rate: 2 oct/min


And the following input .for the SPIRE instrument mounted on the FIRST optical bench:

Sine vibration levels	Frequency range	Input at base (QUAL)
X-direction	5-40 Hz	30 g
	40-100 Hz	15 g
Y-direction	5-45 Hz	18 g
	45-100 Hz	7.5 g
Z-direction	5-45 Hz	18 g
	45-100 Hz	7.5 g

Table 1.4-2: Instrument qualification levels for sine vibration (2 oct/min)

Due to mechanical amplification of the instrument itself, subjected to these levels, the subsystems within the instrument will see a different environment. This has been analysed using FEA. The results of this analysis are translated into sine level specifications for subsystems individually. This will be done taking the following into account:

- Frequency uncertainty: The instrument is designed to meet the frequency requirement including a uncertainty of 10%. To meet the frequency uncertainty requirement the above listed levels are extended to 110 Hz.
- Analysis uncertainty: A analysis uncertainty of 20% will be take into account The analysed amplification is multiplied with the analysis uncertainty factor of 1.2. The latter is assumed to be sufficient since the influence of structural damping is minimal. The input spectrum for each subsystem is analysed (and given) at its interface plane with the instrument structure.

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### Random

RD07, section 9.5.3.4, gives the following requirements for the SPIRE instrument.


The test duration is 2 minutes and 30 seconds.

Random vibration levels	Frequency range	Input at base (QUAL.)	RMS
X-direction	20 – 80 Hz	+3 dB Hz	6.67 g RMS
	80 - 300 Hz	0.077 g <sup>2</sup> /Hz	
	300 – 2000 Hz	-6 dB	
Y-direction	20 – 80 Hz	+3 dB Hz	6.67 g RMS
	80 - 300 Hz	0.077 g <sup>2</sup> /Hz	
	300 – 2000 Hz	-6 dB	
Z-direction	20 – 80 Hz	+3 dB Hz	6.67 g RMS
	80 - 300 Hz	0.077 g <sup>2</sup> /Hz	
	300 – 2000 Hz	-6 dB	

Table 1.4-3: Instrument levels for random vibration (qualification levels)

Due to mechanical amplification of the instrument itself, subjected to these levels, the subsystems within the instrument will see a different environment. This has been analysed using FEA. The results of this analysis are translated into random level specifications for subsystems individually. This will be done taking the following into account:

- structural damping of 3% between 20 and 200 Hz
- structural damping of 4% between 200 and 400 Hz
- structural damping of 5% above 400 Hz

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## 2. MECHANICAL, THERMAL INTERFACE REQUIREMENTS PER SUBSYSTEM

In the sections here after some of the requirements that are listed are for information only. Since they are defined in higher level documents. The listed requirements, for information only, are:

### 1. Mass, reference document: AD2 Systems budgets

The sections per interface are divided into 9 different sub-sections. The first subsection deals with overall information to give something of a background (briefly) to the function of the subsystem. If needed a block diagram will be added to clarify the inter-relations between the different parts/interfaces.

The second and third sub-section deals with possible inputs and outputs. For a mechanical interface this consists, if applicable, of a definition of input micro-vibrations and output micro-vibrations.

The fourth sub-section gives the interface drawings. One set defining the interface on the instrument side and one set defining the interface on the subsystem side. The drawings in this document itself are not the official drawings. They serve as an illustration only. Proper (A1 or A2) interface drawings will be exchanged between the parties concerned. It is the responsibility of each side of the interface to provide the other side with their interface drawing. The interface drawing should contain a reference hole. This reference hole will be tightly tolerated at both sides. In case of mounting on the SOB, it is the responsibility of the subsystem, mounting on this bench, not to introduce additional stress to the optical bench during and after integration. That is no other than the loads induced through the mechanical en

The fifth sub-section gives the mass properties. As stated before, the mass is given for information only. Together with the mass the centre of gravity and the moments of inertia around the CoG are given. The mass properties listed should contain all mass of the subsystem mounted on this interface. Since the listed mass, in accordance with AD2, will serve as input for the overall CoG and MoI figures for the instrument. According to RD07, the mass properties have to be measured within the following tolerances:

- CoG +/- 1%
- MoI +/- 5%


These margins are the same for the actual mass properties and predictions of both the CQM and the PFM model.

The sixth sub-section gives the mechanical environment. The mechanical environment defines the analysed interface levels as a result of the required input for the instrument as a whole. In effect it is the translation of the ESA requirements at the base of the instrument to responses at the interfaces as outlined in section 1.4. For this analysis the qualification levels for the instrument are used. For the qualification of the subsystems, a extra qualification factor of 1.5 is needed. Unfortunately the specified levels are subject to change due to the fact that the ESA mechanical environment specification has not been frozen.

The seventh sub-section deals with the mechanical implementation of thermal interface requirements. That is, there could be a need for thermal insulation (like the structure support of thermal straps). Or there could be a need for maximised thermal conductance, as in the case of the calibration sources where the generated heat needs to be diffused inside the optical bench.

Sub-section eight deals with the harness routing. The number of connectors involved for each interface, allowable curvature of wiring, cable length, screening requirements.

Sub-section nine deals with alignment requirements if applicable.

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## 2.1 Mirrors

### 2.1.1 Functional Description and Block Diagram

The following holds for all interfaces considered in section 2.1 unless otherwise specified.

#### **Allowable interface loads**

The mirror will not add significantly to the rigid body inertia's of the mirrors themselves. That is the eigenfrequency of the mirrors shall be higher than 350 Hz against a rigid interface. With a mount, mounted against a rigid interface shall be higher than 250 Hz.

#### **Interface drawing**

Since the mirrors interface both with their mounting and the infra red beams, which they reflect, the relative position of the mounting in the instrument depends both on the interface with the mounting as well as the thickness of the mirror. Baseline are two different mirror thicknesses. The mounting spigot will be the same for all the mirrors. As a result of this there are only two different interface drawings required. Hereafter these drawings are presented and the datasheet for each mirror will refer to one of these drawings. In the drawings the dimension H-height refers to the thickness of the mirror at its mounting interface. This is either 15 or 7 mm. In principle each mirror has a rotational degree of freedom, since each mirror mount is rotationally symmetrical. This degree of freedom will be suppressed by means of a dowel pin. The location of this pin, indicated on each drawing, is between the mirror mount spigot and the mirror mount bracket interface with the structure. If in case of possible confusion, the position of the dowel pin will be defined explicitly. See section 2.1.4.

No block diagram

### 2.1.2 Inputs


N.A.

### 2.1.3 Outputs

N.A.

### 2.1.4 I/F drawing

Hereafter the interface drawing of the mirror is included, outlining the generic interface. In the drawing (Figure 2.1-1) They should be replaced by 15 (mm) for CM3, CM5, PM7 and PM9. And should read 7 (mm) for the other mirror interfaces.

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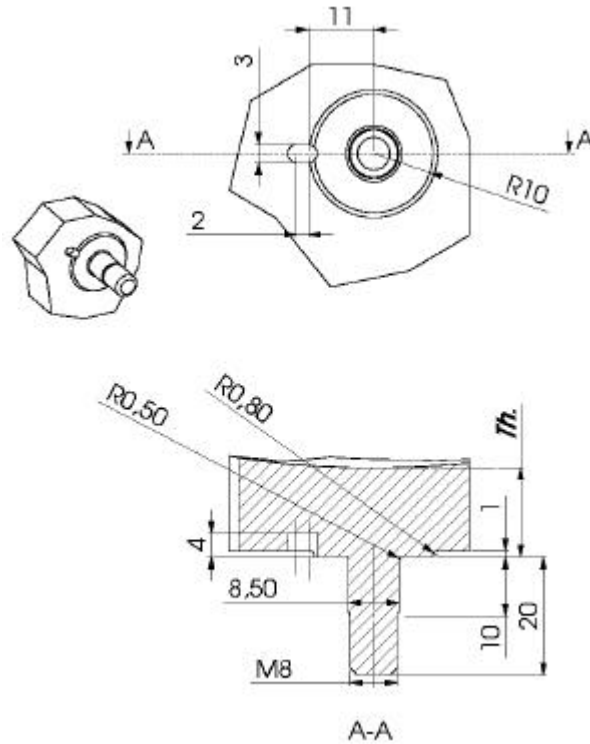


Figure 2.1.4-1: In interface drawing.

### 2.1.5 Mass Properties

Hereafter for each mirror the CoG and the MoI are listed with respect to the co-ordinate system listed in figure 2.1-2. The Z axis is pointing away from the mirror surface. The Y axis is pointing away from the SOB.

The mass budget for the mirrors is 1720 +1100 gr (AD2, excluding contingency)

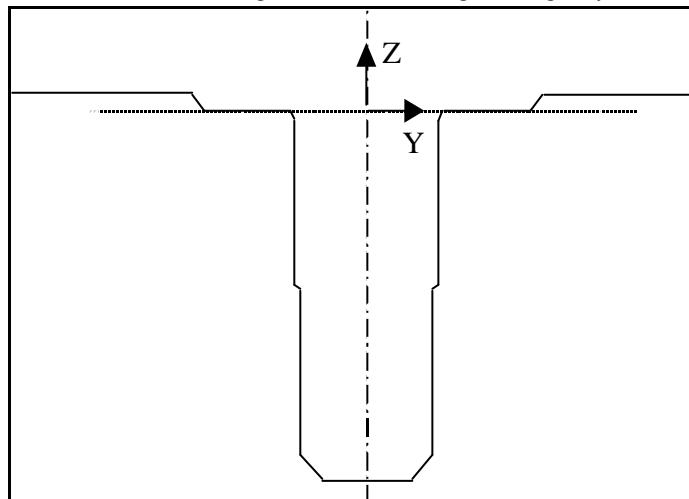



Figure 2.1-2: CoG location reference co-ordinate system for mirrors (Y pointing away form SOB)

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ID	L	W	Thick-ness	Mass nominal	CoG [m]	Ixx	Iyy	Izz
	[mm]	[mm]	[mm]	[gr]	[mm]	[grmm <sup>2</sup> ]	[grmm <sup>2</sup> ]	[grmm <sup>2</sup> ]
CM3	132	54	15	183				
CM5	161	85	15	360				
PM6	46	27	7	23+nut				
PM7	118	101	15	223+nut				
PM8	60	60	7	37+nut	2.8			
PM9	112	112	15	184+nut				
PM10	78	40	7	65+nut				
PM11	56	53	7	43+nut				
SM6	29	51	7	27				
SM7	40	57	7	40+nut				
SM8	60	60	7	37+nut				
SM9	50	50	7	37+nut				
SM10	60	60	7	55+nut				
SM11	74	74	7	79+nut				
SM12	21	16	7	20+nut				

Table 2.1.5-1: Mass properties of the mirrors (updated from IGES files)

## 2.1.6 Mechanical Environment

Assuming a first natural frequency, mounted against a rigid interface above 170 Hz.

As specified in 1.4, including the qualification factor of 1.5. The provisional sine spec is:

Sine vibration levels	Frequency range	Input at base (QUAL)
X-direction	5-30 Hz	11 mm (0-peak)
	30-100 Hz	40 g
Y-direction	5-30 Hz	11 mm (0-peak)
	30-100 Hz	25 g
Z-direction	5-30 Hz	11 mm (0-peak)
	30-100 Hz	25 g

Table 2.1.6-1 Qualification sine input (2 oct/min)


Random spec to be issued.

## 2.1.7 Thermal Interface

When cooling down the instrument, the mirrors with their overall significant thermal mass, should not lag behind in temperature significantly (time wise). Therefore a significant thermal conductance should exist between the mirrors, through the mirror mounts to the optical bench. Since the contact area between mirror and mirror mount is significant, the mirror mounts themselves are significant no problem is expected. The mirror mounts are mounted with three interface bolts on the optical bench.

Between mirror and mirror mount:

Total contact area: between 100 and 250 mm<sup>2</sup> (TBC)

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Surface roughness of contact area: Ra=1.5 (TBC).

### 2.1.8 Harness Routing



N.A.

### 2.1.9 Alignment

The following requirements hold for the alignment of the I/F plane between the mirror mounts and the structure.

- Off centre 0.1 mm all directions simultaneously
- Longitudinal +/-0.5 mm
- Tilt +/-1 arcminute.
- Roll +/-2 arcminute (TBC)



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## 2.2 Filters

### 2.2.1 Functional Description and Block Diagram

The first filters in the chain, situated on the 4-K box, should reflect back to the sky as much unwanted high frequency radiation as possible. Filters on lower temperature shields should then reject radiation from higher temperature boxes and shields.

Strategic placement of filters will enable us to:

- Define the spectral passbands.
- Minimise the thermal loading on the  $^3\text{He}$  fridge, 2-K, and 4-K stages by rejecting short wavelength thermal energy.

Minimise stray light getting to the detectors.

- Maximise the in-band spectral transmission.

The sequence of filters starts with the 2 common 4-K filters at the input to the instrument. They minimise radiation from  $T > 4\text{-K}$  entering the instrument. The filters at the entrance of the detector boxes minimise the transmission of radiation from  $T > 2\text{-K}$ . The 0.3 K filters at the nose of the detectors are not discussed in this document since they interface only with the detectors.

The filters are clamped against the structure using a light tight clamp ring. The function of the OBP (reviewing the interface between the OBP and the filters) is to support the beam splitters and their mounts. The mechanical interface should provide for a precise defined and accurately machined mounting surface. The mechanical interface should provide for sufficient stiffness, such that the mechanical loads on the mirrors are minimised. The same holds for the interfaces with the detector boxes and the dichroics. The dichroics and beam splitters are sensitive to warping, and this is considered in the design of the mounts for these components.

There is no block diagram.

### 2.2.2 Inputs

N.A.

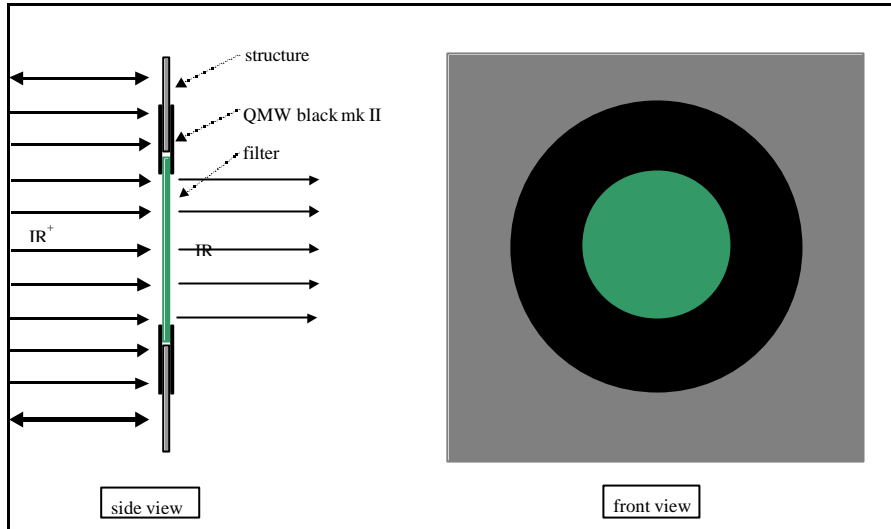
### 2.2.3 Outputs

N.A.

### 2.2.4 I/F drawing

TBD

As a guide the filters will be sized to the envelope the 20% oversized beam at the mounting location. At this location the filters closes an aperture and the area around it, if possible, will be covered with the famous QMW black mk II.



Hot pressed filters may be ring-mounted. TBC on a case by case basis

### 2.2.5 Mass Properties

The mass budget for the filters is 450 gr (AD2, excluding contingency)

ID	Mass	Ixx	Iyy	Izz
CFIL1	TBD	TBD	TBD	TBD
PFIL2	TBD	TBD	TBD	TBD
PFIL3	TBD	TBD	TBD	TBD
SFIL2	TBD	TBD	TBD	TBD
SFIL3S	TBD	TBD	TBD	TBD
SFIL3L	TBD	TBD	TBD	TBD

Table 2.2.5-1: Mass properties table, I with regard to CoG

### 2.2.6 Mechanical Environment


Assuming a first natural frequency, mounted against a rigid interface above 170 Hz.

As specified in 1.4, including the qualification factor of 1.5. The provisional sine spec is:

Sine vibration levels	Frequency range	Input at base (QUAL)
X-direction	5-30 Hz	11 mm (0-peak)
	30-100 Hz	40 g
Y-direction	5-30 Hz	11 mm (0-peak)
	30-100 Hz	25 g
Z-direction	5-30 Hz	11 mm (0-peak)
	30-100 Hz	25 g

Table 2.2.6-1 Qualification sine input (2 oct/min)

Random spec to be issued.

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### 2.2.7 Thermal Interface

TBD

### 2.2.8 Harness Routing


N.A.

Note some filters may need to be part of the Faraday filters.

### 2.2.9 Alignment

The following requirements hold for the alignment of the I/F plane between the filter mounts and the structure.  
RD2

Off centre 1 mm all directions simultaneously (TBC). The flatness of the mounting surface for the cold stops is TBD.

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## 2.3 Beam Splitters

### 2.3.1 Functional Description and Block Diagram

The beam splitters in the FTS part of the instrument reflect 50% and pass 50% of the incident radiation virtually independently of frequency across the entire SPIRE band.

### 2.3.2 Inputs

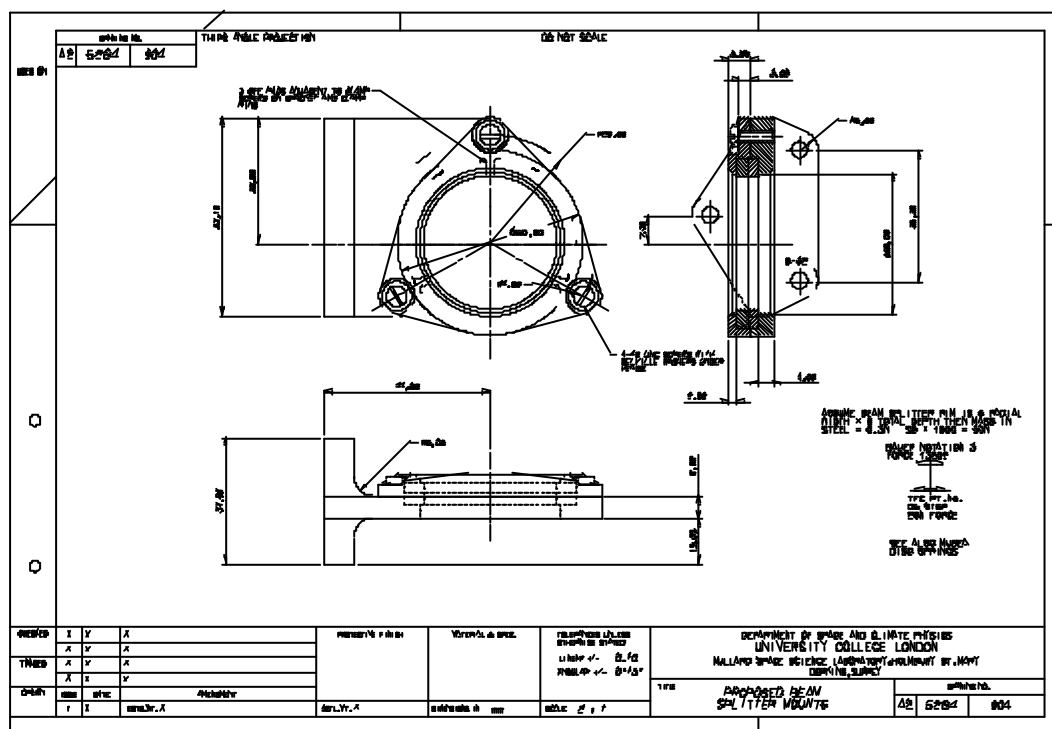
N.A.

### 2.3.3 Outputs

N.A.

### 2.3.4 I/F drawing


The the outer diameter of the beam splitter is TBD mm.



### 2.3.5 Mass Properties

The mass budget for the splitters is 100 gr (AD2, excluding contingency) (100 gr taken from mass for the spectrometer filters) This is the splitter itself, without mount.

ID	Mass	Ixx	Iyy	Izz
SBS1	TBD	TBD	TBD	TBD

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SBS2	TBD	TBD	TBD	TBD
------	-----	-----	-----	-----

Table 2.3.5-1: Mass properties table, I with regard to CoG

### 2.3.6 Mechanical Environment

Assuming a first natural frequency, mounted against a rigid interface above 170 Hz.

As specified in 1.4, including the qualification factor of 1.5. The provisional sine spec is:

Sine vibration levels	Frequency range	Input at base (QUAL)
X-direction	5-30 Hz	11 mm (0-peak)
	30-100 Hz	40 g
Y-direction	5-30 Hz	11 mm (0-peak)
	30-100 Hz	25 g
Z-direction	5-30 Hz	11 mm (0-peak)
	30-100 Hz	25 g

Table 2.3.6-1 Qualification sine input (2 oct/min)

Random spec to be issued.

### 2.3.7 Thermal Interface


N.A.

### 2.3.8 Harness Routing

N.A.

### 2.3.9 Alignment

TBD

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## 2.4 Dichroics

### 2.4.1 Functional Description and Block Diagram

The dichroics within the PDBX reflect high frequencies and pass low frequencies and hence aid definition of the SPIRE photometer channels.

The maximum warping of the dichroics induced by the structure is TBD .Tilt 1 arcminute (TBC)

### 2.4.2 Inputs

N.A.

### 2.4.3 Outputs

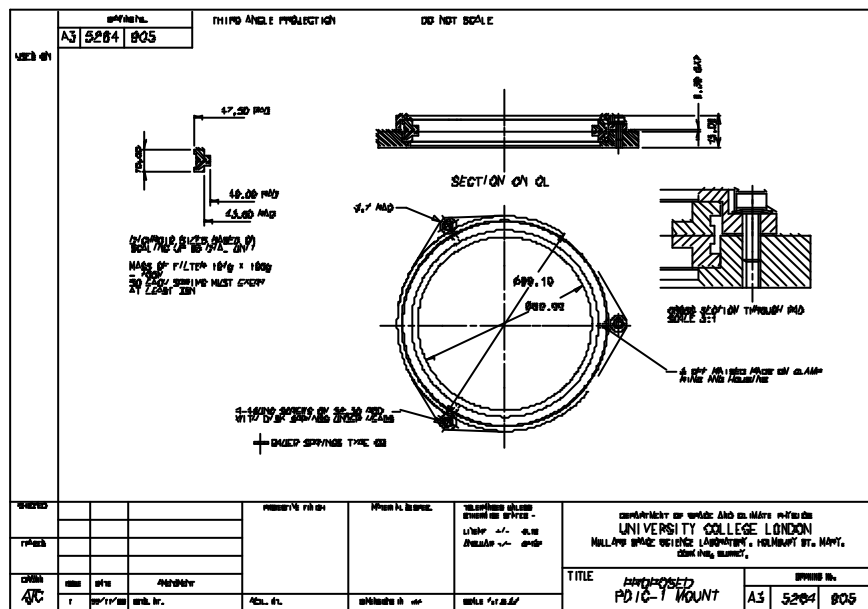
N.A.

### 2.4.4 I/F drawing

The outer dimension of the dichroics are:

PDIC1: TBD mm


PDIC2: TBD mm



### 2.4.5 Mass Properties

The mass budget for the splitters is 100 gr (AD2, excluding contingency) (100 gr taken from mass for the spectrometer filters)

ID	Mass	Ixx	Iyy	Izz
PDIC1	TBD	TBD	TBD	TBD

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PDIC2	TBD	TBD	TBD	TBD
-------	-----	-----	-----	-----

Table 2.4.5-1: Mass properties table, I with regard to CoG

## 2.4.6 Mechanical Environment

Assuming a first natural frequency, mounted against a rigid interface above 170 Hz.

As specified in 1.4, including the qualification factor of 1.5. The provisional sine spec is:

Sine vibration levels	Frequency range	Input at base (QUAL)
X-direction	5-30 Hz	11 mm (0-peak)
	30-100 Hz	40 g
Y-direction	5-30 Hz	11 mm (0-peak)
	30-100 Hz	25 g
Z-direction	5-30 Hz	11 mm (0-peak)
	30-100 Hz	25 g

Table 2.4.6-1 Qualification sine input

Random spec to be issued.

## 2.4.7 Thermal Interface


N.A.

## 2.4.8 Harness Routing

N.A.

## 2.4.9 Alignment

TBD

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## 2.5 Straylight Baffles and QMW black (Mk2)

### 2.5.1 Functional Description and Block Diagram

The straylight baffles reflect and partially attenuate straylight at various locations in the optical chain. At present 4 baffles are foreseen. The first baffle is located inside the instrument at the entrance of the IR-beam into the instrument. Effectively it is located between the secondary mirror and CM3. It encloses together with the photometer side walls the shutter mechanism. The second one is located between PM7 and PM8. Separating the common environment from the photometer and effectively enclosing the photometer detector box. The third and fourth are located inside the spectrometer enclosure. They split the spectrometer into three zones, each at the beam splitter where the two spectrometer optical chains cross.

See for a block diagram 1.3-1 and 1.3-2

### 2.5.2 Inputs

N.A.

### 2.5.3 Outputs

N.A.

### 2.5.4 I/F drawing

TBD

### 2.5.5 Mass Properties

TBD

### 2.5.6 Mechanical Environment

Assuming a first natural frequency, mounted against a rigid interface above 170 Hz.

As specified in 1.4, including the qualification factor of 1.5. The provisional sine spec is:

Sine vibration levels	Frequency range	Input at base (QUAL)
X-direction	5-30 Hz	11 mm (0-peak)
	30-100 Hz	40 g
Y-direction	5-30 Hz	11 mm (0-peak)
	30-100 Hz	25 g
Z-direction	5-30 Hz	11 mm (0-peak)
	30-100 Hz	25 g


Table 2.5.6-1 Qualification sine input (2 oct/min)

Random spec to be issued.

### 2.5.7 Thermal Interface

N.A.




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### 2.5.8 Harness Routing

N.A.

### 2.5.9 Alignment

The opening in the straylight baffles will be sufficient to allow the 20% oversized beam to cross it with a positive margin of no more than 0.5 mm.

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## 2.6 <sup>3</sup>He Cooler

### 2.6.1 Functional Description and Block Diagram

The structure of the cooler is mounted on the SPIRE optical bench panel (SOB) at ~4 K. The cooler interfaces via 2 thermal straps with the level 0 thermal sink (~1.8 K) inside the FIRST cryostat. The cold finger of the cooler interfaces via a thermal busbar with all detectors (3 photometer detectors and 2 spectrometer detectors). The cooler harness interfaces with the DRCU outside the SPIRE structure via the RF-filter box mounted inside the structure.

- SOB – Cooler (mounting)
- PSW – Cooler (thermal strap)\*
- PMW – Cooler (thermal strap)\*
- PLW – Cooler (thermal strap)\*
- SLB – Cooler (thermal strap)\*
- SHB – Cooler (thermal strap)\*
- Cooler – level 0 heat sink (two different straps)
- Harness routing (from cooler via structure to RF-filter boxes)

\*It should be noted that the interface between the detectors and the cooler, as it is currently foreseen, will run via a thermal busbar. Strictly speaking there will be in this case no (direct) interface between the cooler and the detectors.

The cooler's purpose is to provide for the cooling of the sensors within the detectors down to ~.3 K. For this the cooler needs to be able to dump excess heat during its recycling and operating phase, via 2 straps to the level 0 heat sink within the cryostat. During operation and recycling different parts within the cooler will be connected to the level 0 heat sink. This will be done using heat switches, which are located inside the cooler itself. The (support) structure of the cooler will be interfacing with the optical bench panel only (at level 1).

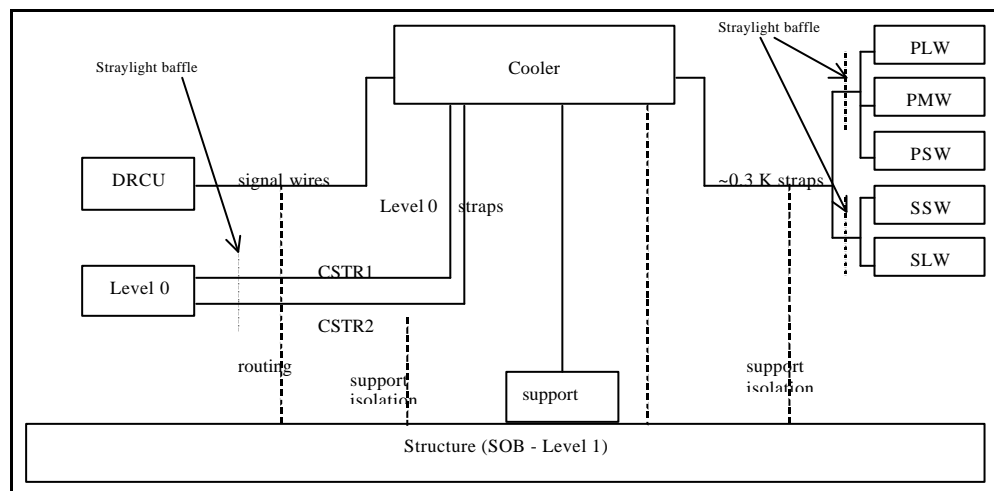



Figure 2.6.1-1: Block diagram

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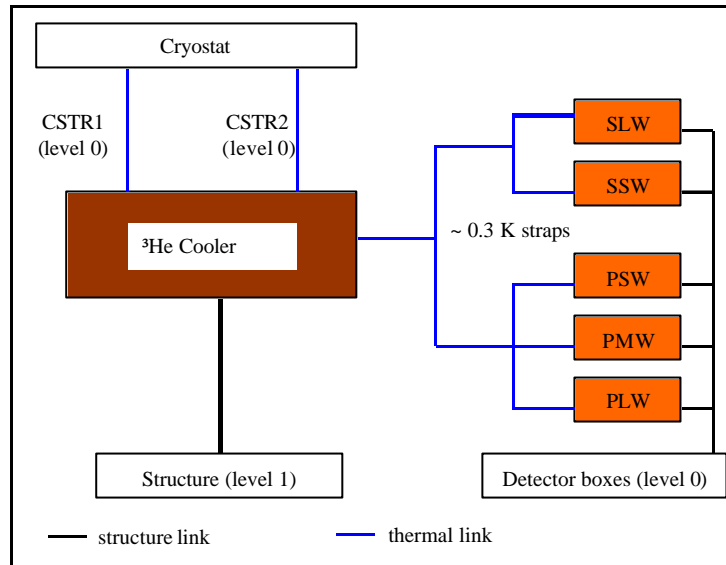


Figure 2.6.1-2: Thermal straps, conductive paths

## 2.6.2 Inputs

mounting interface structure (SOB)

Two separate straps, connected with level 0 heat sink

control wiring (routing towards RF-filter box)

Heat flow from level 1 towards cooler

Heat flow from cooler into Level 0

Micro vib. n.a.

To be specified: allowable interface load for the thermal strap-interface (CSTR1 and CSTR2)


To be specified: stiffness for the thermal strap interface.

## 2.6.3 Outputs

Strap to the thermal busbar between both detector boxes

Heat flow from detectors into cooler

Micro vib. n.a.

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## 2.6.4 I/F drawing

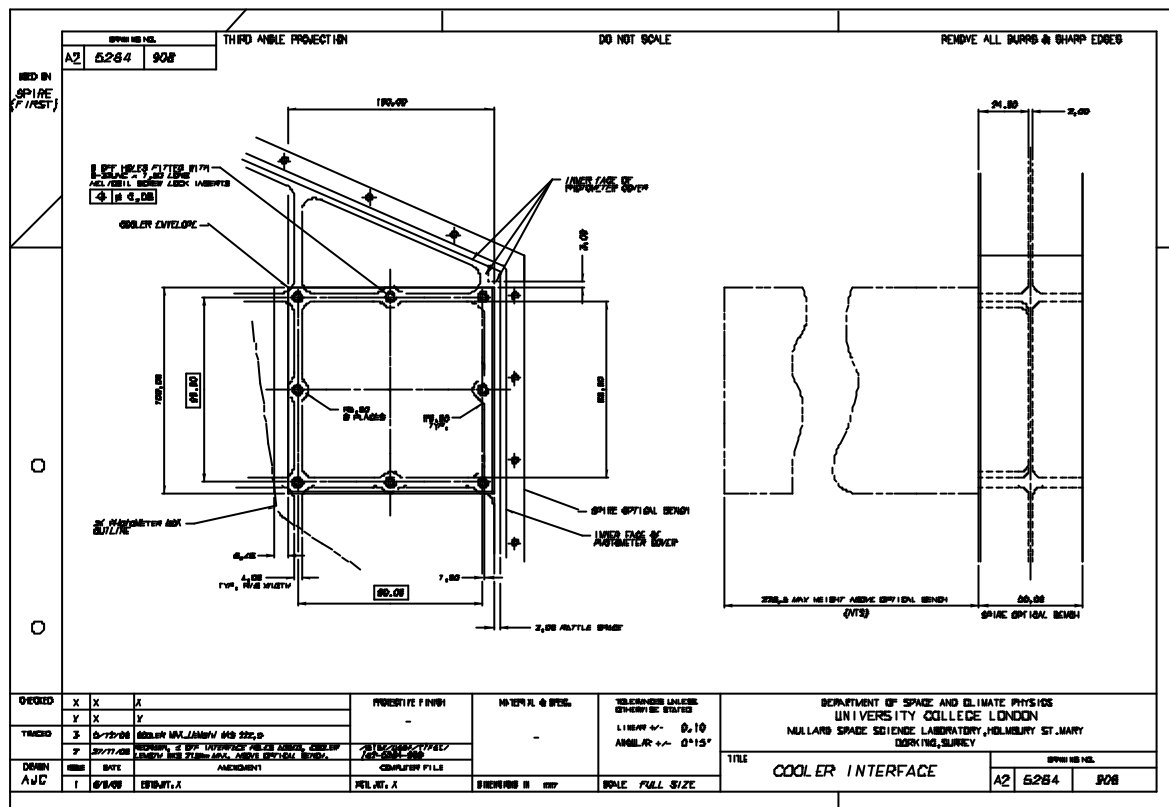


Figure 2.6.4-1: Outdated interface drawing, current I/F has 8 I/F holes, indicated with dotted circles

## 2.6.5 Mass Properties

Current mass budget for the cooler is 1.4 kg + 20% (TBC) contingency, see AD2 (to be updated). COG and MOI to be defined.


## 2.6.6 Mechanical Environment

Assuming a first natural frequency, mounted against a rigid interface above 170 Hz. As specified in 1.4, including the qualification factor of 1.5. The provisional sine spec is:

Sine vibration levels	Frequency range	Input at base (QUAL)
X-direction	5-30 Hz	11 mm (0-peak)
	30-100 Hz	40 g
Y-direction	5-30 Hz	11 mm (0-peak)
	30-100 Hz	25 g
Z-direction	5-30 Hz	11 mm (0-peak)
	30-100 Hz	25 g

Table 2.6.6-1 Qualification sine input (2 oct/min)

Random spec to be issued.

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	X	Y	Z
f	g <sup>2</sup> /hz	g <sup>2</sup> /hz	g <sup>2</sup> /hz
20	0.042	0.042	0.042
80	0.208	0.229	0.229
100	0.229	1.04	0.249
120	2.08	218.0	2.08
140	20.8	218.0	104.0
160	83.1	12.5.0	104.0
180	83.1	12.5.0	16.6.0
200	18.7	12.5.0	16.6.0
300	18.7	12.5.0	16.6.0
400	1.25	12.5.0	0.312
600	0.416	12.5.0	0.021
800	0.010	0.021	0.010
1000	0.002	0.002	0.002
<b>g rms</b>	<b>109.9</b>	<b>125.0</b>	<b>50.0</b>

Table 2.6.6-3: Cooler qualification levels for random vibration (2.5 minutes) TBC

## 2.6.7 Thermal Interface

See for budgets thermal configuration control document.

CSTR1

minimal thermal resistance between structure and CSTR1 support is TBD W/K

CSTR2

minimal thermal resistance between structure and CSTR2 support is TBD W/K

Thermal Busbar

maximal heat leak from structure into the busbar is 1 micro-W (TBC)

## 2.6.8 Harness Routing


Two connectors on cooler housing and interconnect harness between cooler and RF-filter box. The connectors are 37 pins. Two of them located at the outside of the cooler housing. Provisional location indicated on cooler design, 25 October 2000, still to be included in the I/F drawing.

Harness length: Cooler RF-filter box  $845 \pm 170$  mm

The harness length from the cooler to the RF-filter boxes is: TBD mm

## 2.6.9 Alignment

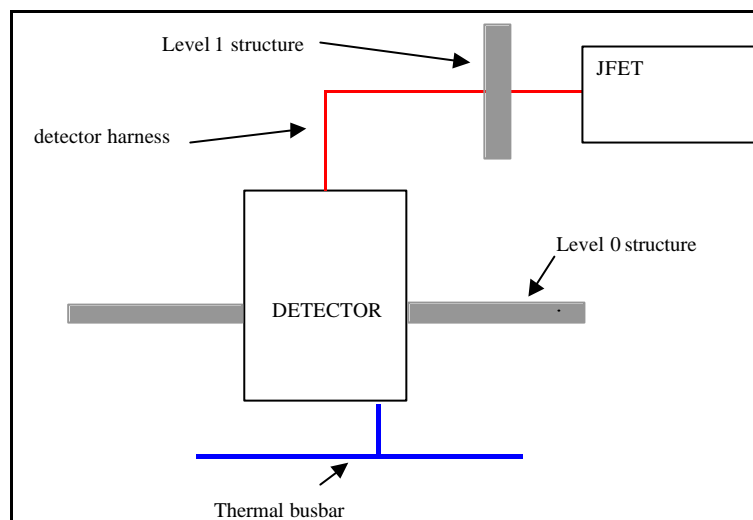
There are no strict alignment requirements on the cooler. The thermal straps attached to the cooler need to be able to sustain 0.5 mm (TBC) differential motion between the interface with the structure and the cooler. This does not take into account differential thermal contraction.

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## 2.7 Photometer Detectors

### 2.7.1 Functional Description and Block Diagram

The photometer detectors consist of a set of three detectors PSW, PMW and PLW. The detectors are mounted on the photometer detector box. Inside the box the detectors interface with the thermal busbar and outside the box with the harness routing towards the JFET modules outside the structure.



Block diagram 2.7.1-1:

### 2.7.2 Inputs

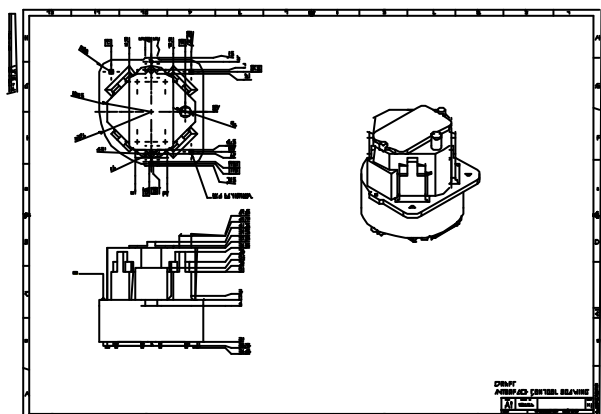
N.A.

### 2.7.3 Outputs


N.A.

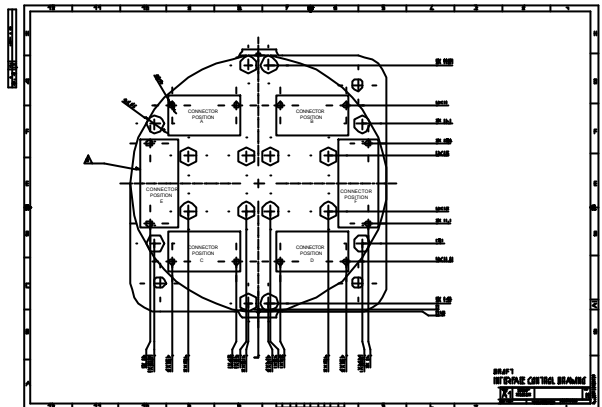
### 2.7.4 I/F drawing

The listed drawings hereafter are draft versions. The current design has 4 interface bolt-holes.

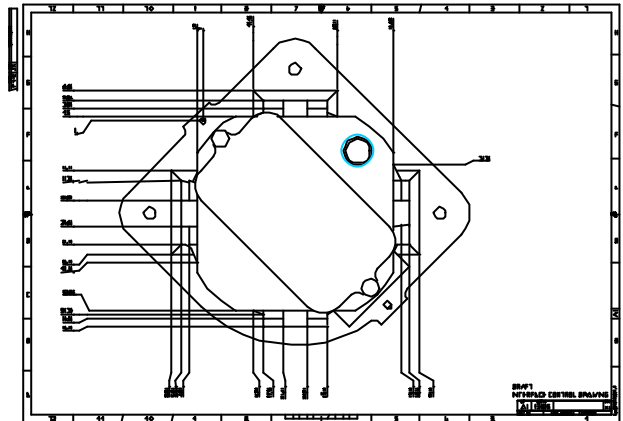


2.7.4-1: Overview of the detector

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2.7.4-2: Back of the detector with 6 possible connector positions



2.7.4.-3 Front of the detector with filter. Thermal interface indicated in blue

## 2.7.5 Mass Properties

On average the mass (for all five detectors) of the detector is 500 gr excluding contingency, see AD2.

## 2.7.6 Mechanical Environment


Assuming a first natural frequency, mounted against a rigid interface above 170 Hz.

As specified in 1.4, including the qualification factor of 1.5. The provisional sine spec is:

Sine vibration levels	Frequency range	Input at base (QUAL)
X-direction	5-30 Hz	11 mm (0-peak)
	30-100 Hz	40 g
Y-direction	5-30 Hz	11 mm (0-peak)
	30-100 Hz	25 g
Z-direction	5-30 Hz	11 mm (0-peak)
	30-100 Hz	25 g

Table 2.7.6-1 Qualification sine input (2 oct/min)

Random spec to be issued.

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### 2.7.7 Thermal Interface

The detector interfaces with the 0.3 K thermal busbar inside the photometer detector box. The connection will be made at the nose of the detector, as outlined in the draft interface drawing. The detector structure will be supported of the 2 K detector box.

There are thermal requirements for the 0.3 K busbar, they are listed in section...

- 1 Minimum cable length between detector connector and common structure is 80 mm. (RD5)
- 2 The cable loom support between the structure and the JFET rack is allowed a maximum parasitic heatload of TBD mW


The connector at the back of the detector serves also as a 2 K heat sink.

### 2.7.8 Harness Routing

The requirements for the harness routing are:

1. Firmly connected to (support) structure to minimise microphonics
2. Baseline routing is from the back of the detector down to the detector box routed along the A-frames supporting the detector box down to the SPIRE optical bench. After that the cable loom will be routed to the edge of the optical bench where it is routed through an interface plate provided for by JPL. The interface plate will be bolted to the optical bench interface bracket. Subsequently the loom is routed along the outer panel to the JFET rack. Between the SPIRE structure and the JFET rack the cable loom will be supported via a thermal insulating structure.
3. PLW to connector panel: 1200 mm TBC
4. PMW to connector panel: 1000 mm TBC
5. PSW to connector panel: 1100 mm TBC
6. Connector panel to JFET modules: 200 mm TBC



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## 2.8 Spectrometer Detectors

### 2.8.1 Functional Description and Block Diagram

For the spectrometer detectors the same hold as for the photometer detectors, unless specified otherwise.

### 2.8.2 Inputs

See 2.7.

### 2.8.3 Outputs

See 2.7.

### 2.8.4 I/F drawing

See 2.7.

### 2.8.5 Mass Properties

See 2.7.

### 2.8.6 Mechanical Environment

See 2.7.

### 2.8.7 Thermal Interface

See 2.7.

### 2.8.8 Harness routing


See 2.7.

The harness lengths are:

1. SLW to connector panel 600 mm TBC
2. SLW to connector panel 700 mm TBC
3. Connector panel to JFET modules: 200 mmTBC

### 2.8.9 Alignment

See 2.7.

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## 2.9 Beam Steering Mechanism

### 2.9.1 Functional Description and Block Diagram

The BSM is used to steer the optical beam of the SPIRE photometer channel over the detector arrays.

The BSM comprises a flat mirror mounted on a pivot system which system allows precise angular motion of the mirror over a small range of angular travel in two orthogonal axes. Electrical actuators are used to provide motion of the mirror. Electrical transducers are used to measure the mirror position to allow control of the mirror position.

The BSM also provides an aperture through which the Photometer Calibration Source is directed towards the detector arrays. The Photometer Calibration Source is mounted behind the BSM.

The BSM comprises 3 main parts; the cryogenic mechanism (BSMm), the structural support (BSMs), and the warm electronics (BSMe). The structural support may be integral to the BSMm housing, or attached as a distinct entity (TBD). This document only deals with the interface between the BSM support and the SOB.

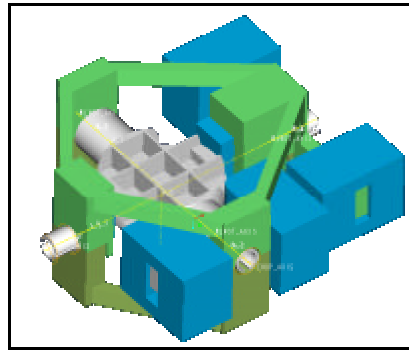


Figure 2.9.1-1: BSM mechanism, green-support structure, blue-actuators, grey-bottom view mirror RD3

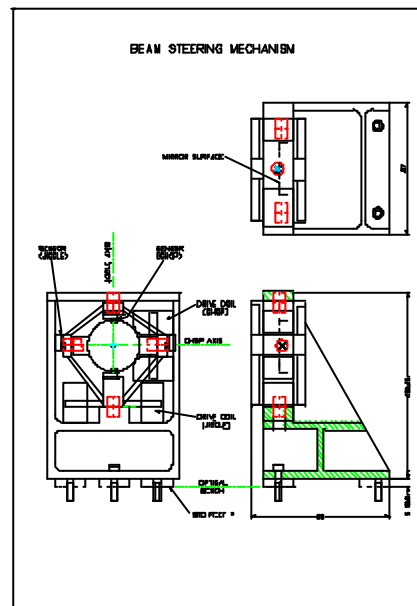



Figure 2.9.1-2: Sketch of the BSM support with BSM and Calibration source RD3

The support structure mounts the beam steering mirror mechanism from the SPIRE optical bench. Note that there will be a baffle incorporated to minimise straylight and emission into the optical beam. Electrical wiring and connectors are not shown. The photometer calibrator will also be mounted on this structure, and its wiring

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incorporated into the BSM harness. The outline design shown here is at present non-optimised, and will be further refined following FEA.

No block diagram

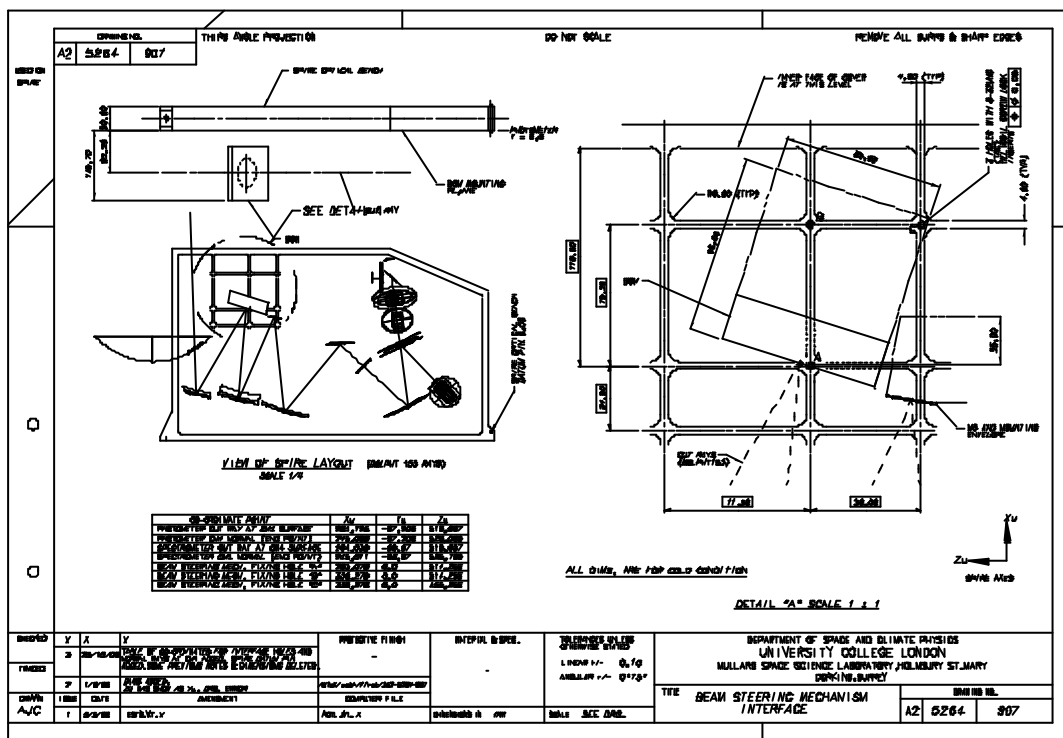
## 2.9.2 Inputs

Micro vibrations: TBD

## 2.9.3 Outputs

Micro vibrations: TBD

## 2.9.4 I/F drawing



## 2.9.5 Mass Properties


The mass budget for the BSM is 1100 gr including harness. (AD2, excluding contingency) excluding calibrator.

CoG and MoI TBD

## 2.9.6 Mechanical Environment

As specified in 1.4, including the qualification factor of 1.5. The provisional sine spec is:

Sine vibration levels	Frequency range	Input at base (QUAL)
X-direction	5-30 Hz	11 mm (0-peak)

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	30-100 Hz	40 g
Y-direction	5-30 Hz	11 mm (0-peak)
	30-100 Hz	25 g
Z-direction	5-30 Hz	11 mm (0-peak)
	30-100 Hz	25 g

Table 2.9.6-1 Qualification sine input (2 oct/min)


### 2.9.7 Random spec to be issued. Thermal Interface

This interface needs to dissipate heat generated by the photometer calibration source into the SOB. This could be via the currently defined interface or via a thermal strap. In the last case we need to specify a thermal strap bolt hole.

### 2.9.8 Harness Routing

- Harness will be connected to the RF-filter boxes and routed over the SOB.
- The harness length from the BSM connectors to the RF-filter boxes is 415 mm  $\pm$ 90 mm

### 2.9.9 Alignment

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## 2.10 Spectrometer Mechanism

### 2.10.1 Functional Description and Block Diagram

The SMEC holds the central corner cube mirror's of the FTS part of the FIRST/SPIRE instrument. These mirrors need to travel between the two ray paths within the spectromter, shortening one ray path and simultaneously lengthen the other. The total required travel is 32 mm in +X direction and 3.2 mm in -X direction (AD1, IRD-SMEC-R01) (Where +X is the launch direction). There will be an over-travel on top of this which still needs to be specified. It is expected that the overall total travel will be about something like 40 mm.

#### Mechanism Interface with other components

- Spire optical bench (SOB)
- Optics
- RF-filter boxes, via the harness

The function of the SOB (reviewing the interface between the SOB and the SMEC) is to support the mechanism. The mechanical interface should provide for a precise defined and accurately machined mounting surface. The mechanical interface should provide for sufficient stiffness, such that the mechanical loads on the SMEC are minimised.


There is no block diagram

### 2.10.2 Inputs

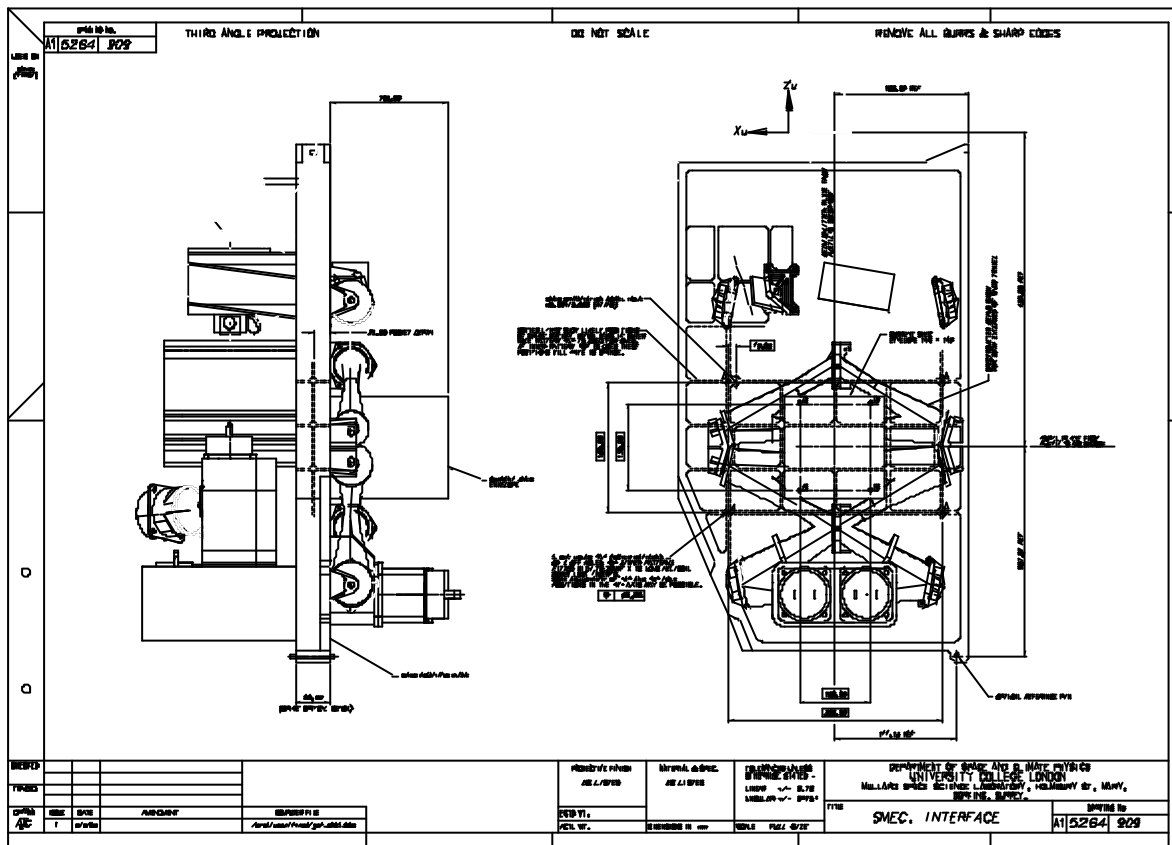
Micro-vibrations coming from the SOB TBS

### 2.10.3 Outputs

Micro-vibrations going into the SOB TBS

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### 2.10.4 I/F drawing



2.10.4-1: Interface drawing, the interface bolt hole pattern is indicated on the inside of the optical bench within the are enclosed by the collimating mirrors and the beam splitters..

### 2.10.5 Mass Properties

The current mass budget for the SMEC is 1.0 kg + 20% contingency, without mirrors and without harness? See AD2.

d.d. 28 Nov 2000


- SMEC-m: 1100 gr
- SMEC-p: 200 gr (mounted somewhere on the SOB, near the RF-filter boxes.
- SMEC-m mirrors: 200 gr (about)

### 2.10.6 Mechanical Environment

Assuming a first natural frequency, mounted against a rigid interface above 170 Hz.

As specified in 1.4, including the qualification factor of 1.5. The provisional sine spec is:

Sine vibration levels	Frequency range	Input at base (QUAL)
X-direction	5-30 Hz 30-100 Hz	11 mm (0-peak) 40 g

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Y-direction	5-30 Hz	11 mm (0-peak)
	30-100 Hz	25 g
Z-direction	5-30 Hz	11 mm (0-peak)
	30-100 Hz	25 g

Table 2.10.6-1 Qualification sine input (2 oct/min)

Random spec to be issued.

### 2.10.7 Thermal Interface

An extra interface bolt hole (on top of the 4) will be added to the interface to provide for a possible thermal strap mounting interface.


### 2.10.8 Harness Routing

Connectors (number TBD) on SMEC interface and harness between connector and RF-filter boxes.

The harness length from the SMEC connectors to the RF-filter boxes is  $425 \pm 85$  mm

### 2.10.9 Alignment

The overall alignment of the interface plane dictates that the planarity stays within 0.1 mm, the interface bolt holes are located within  $\pm 0.1$  mm. TBC.

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## 2.11 Shutter Mechanism

### 2.11.1 Functional Description and Block Diagram

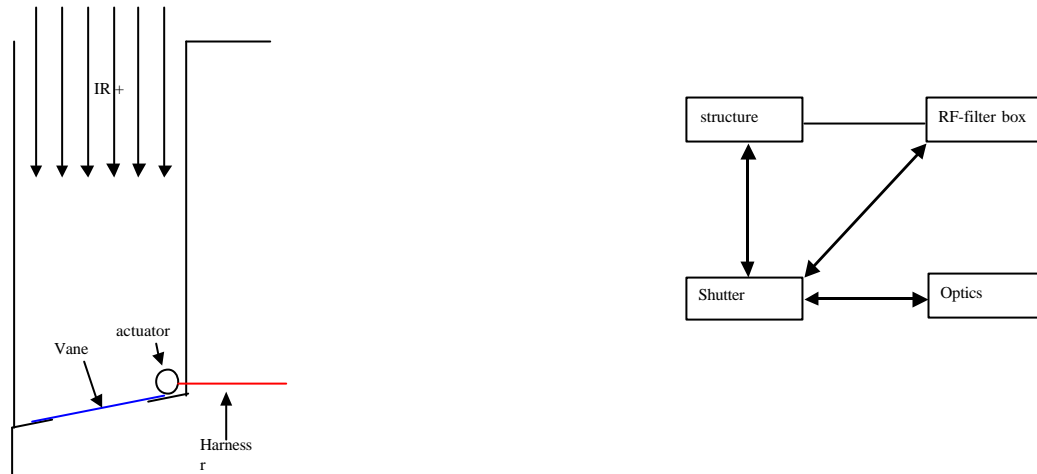


Table 2.11.1-1: Block diagram

### 2.11.2 Inputs

N.A.

### 2.11.3 Outputs

N.A.

### 2.11.4 I/F drawing

Action: MSSL to define available envelope for the shutter. Currently the volume constraint is about  $X \times Y \times Z = 80 \times 50 \times 50 \text{ mm}^3$  for the actuator (if mounted on the SOB). And the volume constraint for the vane is:  $80 \times 150 \times 50$ .

Mounting surface is either the SOB or the straylight baffle (to be decided) depends on RF-sealing of the instrument.

### 2.11.5 Mass Properties


See AD2, 200 gr without contingency.

### 2.11.6 Mechanical Environment

Assuming a first natural frequency, mounted against a rigid interface above 170 Hz.

As specified in 1.4, including the qualification factor of 1.5. The provisional sine spec is:



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Sine vibration levels	Frequency range	Input at base (QUAL)
X-direction	5-30 Hz	11 mm (0-peak)
	30-100 Hz	40 g
Y-direction	5-30 Hz	11 mm (0-peak)
	30-100 Hz	25 g
Z-direction	5-30 Hz	11 mm (0-peak)
	30-100 Hz	25 g

Table 2.11.6-1 Qualification sine input (2 oct/min)

Random spec to be issued.

### 2.11.7 Thermal Interface


Maximum allowed heat load on the SOB is 1 mW during operation of the vane (telescope broadband background simulation). TBD/TBC RAL. This does not include latching and short term events. These peak events will not exceed an influx of more than 0.01 J (?) into the instrument

### 2.11.8 Harness Routing

The harness will be routed over and through the SOB.

The harness length from the Shutter connector to the RF-filter boxes is  $410 \pm 95$  mm

The number of connectors is one at the shutter and possibly 2 at the RF-filter box. RAL needs to clarify this.

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## 2.12 Spectrometer Calibration Source

### 2.12.1 Functional Description and Block Diagram

The spectrometer calibration source provides for a reference channel for the Fourier Transform Spectrometer. It consists of a heated black body inside a cavity. The heat generated inside the housing needs to be diffused in the optical bench. It is mounted with four bolts. The available envelope is 50 x 50 x 100 mm possibly extendable to 60 x 60 x 100 mm.

No block diagram

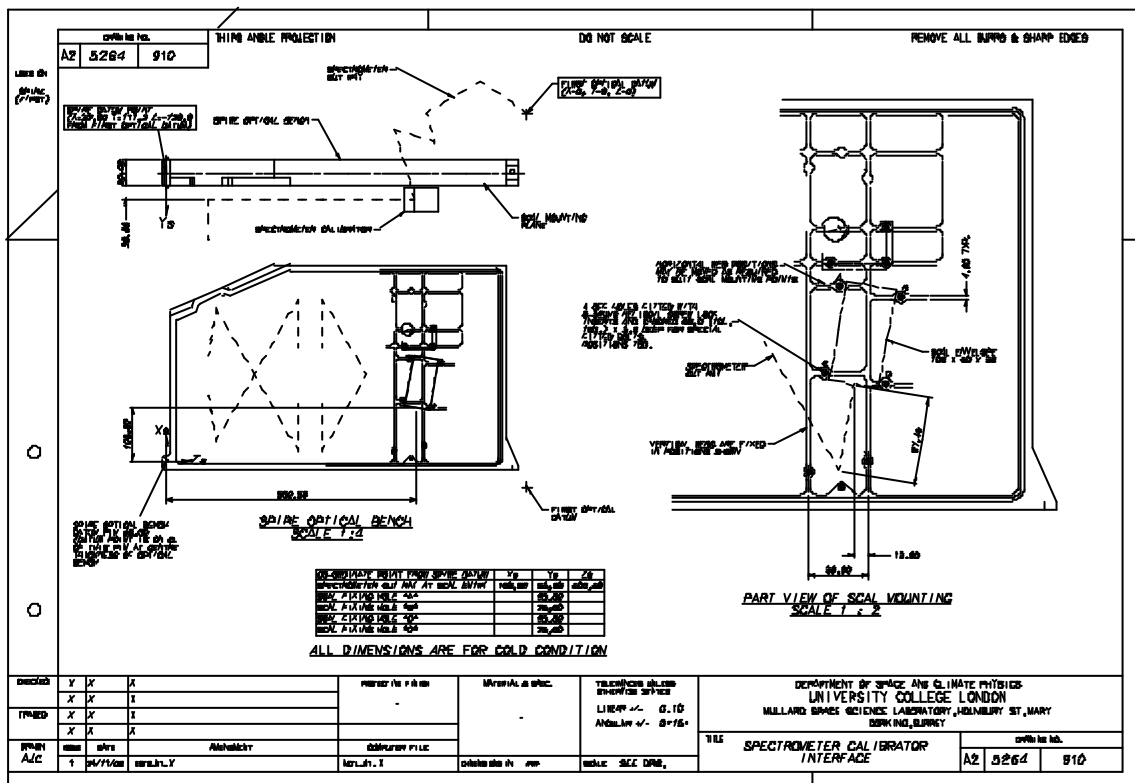
### 2.12.2 Inputs

N.A.

### 2.12.3 Outputs


N.A.

### 2.12.4 I/F drawing



### 2.12.5 Mass Properties

The mass allocated to the SCAL is 200 gr without contingency, including harness. See AD2

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### 2.12.6 Mechanical Environment

Assuming a first natural frequency, mounted against a rigid interface above 170 Hz.

As specified in 1.4, including the qualification factor of 1.5. The provisional sine spec is:

Sine vibration levels	Frequency range	Input at base (QUAL)
X-direction	5-30 Hz	11 mm (0-peak)
	30-100 Hz	40 g
Y-direction	5-30 Hz	11 mm (0-peak)
	30-100 Hz	25 g
Z-direction	5-30 Hz	11 mm (0-peak)
	30-100 Hz	25 g

Table 2.12.6-1 Qualification sine input (2 oct/min)

### 2.12.7 Random spec to be issued. Thermal Interface

The connection with the SOB should allow for thermal conductance in order to dissipate the heat generated by the SCAL sufficiently.

- 1 Minimum surface area for I/F: TBD
- 2 Minimum torque for each bolt: TBD
- 3 Surface treatment... TBD
- 4 Use of a thermal gasket is TBD/TBC


### 2.12.8 Harness Routing

The harness is routed via the SOB to the RF-filter boxes. The harness routing will be part of the I/F drawing, likely on a separate sheet(s).

The length of the harness from the SCAL to the RF-filter boxes will be  $325 \pm 65$  mm.

### 2.12.9 Alignment

The overall alignment of the interface plane dictates that the planarity stays within 0.1 mm, the interface bolt holes are located within  $\pm 0.1$  mm. TBC.

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## 2.13 Thermal Straps

### 2.13.1 Functional Description and Block Diagram

SPIRE has several different temperature stages. In order to diffuse excess heat to the cryostat, various components of the instrument are connected directly or indirectly to the cryostat. Baseline the straps are made of copper.

#### Level 0 straps.

There are three level 0 straps attached to the instrument. Two of them are connected to the cooler and one connected to the photometer detector box and the spectrometer detector box. The level 0 thermal straps are directly connected with the cryostat. Their nominal temperature is between 1.7 and 1.8 Kelvin.

#### Level 1 strap

There is one level 1 strap attached to the structure. It is routed from the vent line, running around the FOB to the optical bench.

#### Thermal busbar

The thermal busbar runs from the cold end of the cooler to the photometer detector box and the spectrometer detector box. It branches off inside these boxes and runs along the detectors. At each detector interface a copper strap branches off and is connected to the thermal interface with the detector.

### 2.13.2 Inputs

N.A.

### 2.13.3 Outputs

N.A.

### 2.13.4 I/F drawing

TBD

### 2.13.5 Mass Properties

TBD

### 2.13.6 Mechanical Environment

Assuming a first natural frequency, mounted against a rigid interface above 170 Hz.

As specified in 1.4, including the qualification factor of 1.5. The provisional sine spec is:

Sine vibration levels	Frequency range	Input at base (QUAL)
X-direction	5-30 Hz 30-100 Hz	11 mm (0-peak) 40 g
Y-direction	5-30 Hz 30-100 Hz	11 mm (0-peak) 25 g
Z-direction	5-30 Hz 30-100 Hz	11 mm (0-peak) 25 g


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Table 2.13.6-1 Qualification sine input (2 oct/min)

### 2.13.7 Random spec to be issued. Thermal Interface


ID	Maximum parasitic load from the structure	Boundary temperatures	
		CSTR1	TBD
CSTR2	TBD	4.0 K	2.0 K
CSTR3	TBD	4.0 K	2.0 K
CSTR4	N.A.	4.0 K	4.0 K
Thermal busbar	1 micro-W	1.7 K	0.3 K

Table 2.13.7:

Add bolt patterns (I/F drawing) + torques + surface finish

### 2.13.8 Harness Routing

N.A.

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## 2.14 JFET box

### 2.14.1 Functional Description and Block Diagram

The JFET modules are stored outside the instrument because the added heat load from the JFETs themselves would seriously add to the overall (parasitic) heat load of the instrument. Stored outside the common structure on the FOB the heat can be dissipated via that bench and through additional ventlines (level 2) if needed. The JFET modules run at a temperature of about 120 Kelvin.

In order to minimise the cable length between the detectors and the JFET modules the JFETs are mounted in two different enclosures. One located near the photometer detectors, 6 modules in all, and one located near the spectrometer detectors, 3 modules in all. The JFET modules interface between the detector harness coming from the instrument and the cry-harness provided for by ESA. See the pictures here after for a global view of SPIRE and the JFET boxes on the FOB.

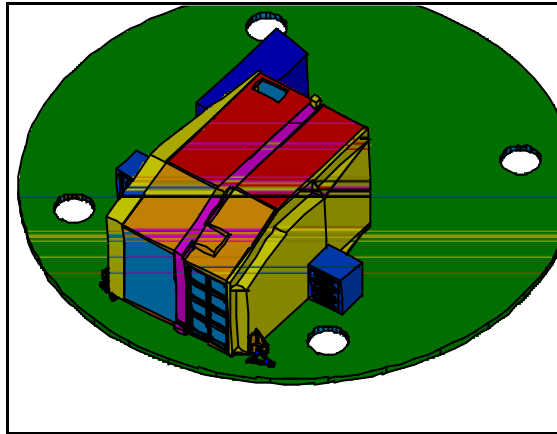


Figure 2.14.1-1: View on the spectrometer FBX

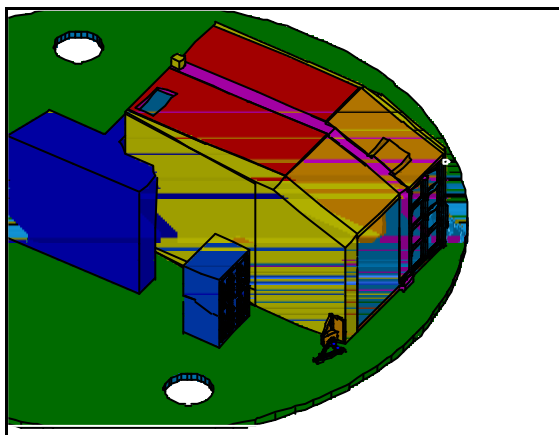



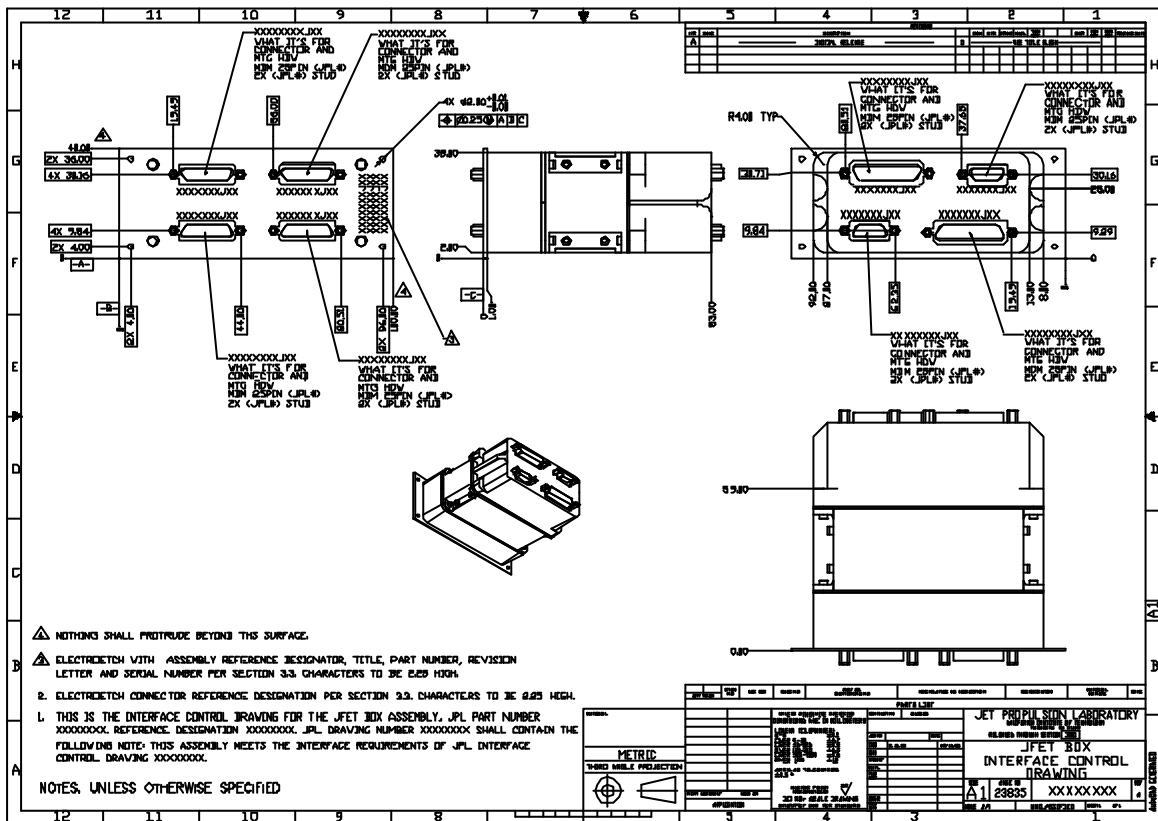
Figure 2.14.1-2: View on the photometer FBX

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### 2.14.2 Inputs

### 2.14.3 Outputs

### 2.14.4 I/F drawing



### 2.14.5 Mass Properties

TBD


### 2.14.6 Mechanical Environment

As defined by the in

### 2.14.7 Thermal Interface

### 2.14.8 Harness Routing


RD4

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1. The cables should be kept short. We wish to achieve a capacitance of 50 pF between signal lines, including all sources of capacitance. Please give us a rough length so we can investigate cable capacitance per unit length.
2. The minimum mechanical frequency of the cables should be 1 kHz. The cables must not be disturbed during servicing of the instrument.
3. Minimum bend radius is 3 mm (TBC).
4. We would like to investigate the possibility of eliminating the connectors on the 4 K shield, instead connecting a drilled aluminum plate to connect the shield to the RF box. This eliminates the connectors and simplifies the cable.

The harness will be supported of both the SPIRE structure as well of the JFET structure. This in order to minimise the parasitic heat load from the JFETs into the structure and to meet the stiffness requirement of the cables.



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## 2.15 RF-Filters

To be written.

Currently five units with 2 25 way connectors are foreseen.

## 2.16 Thermometry harness

To be written.

For the structure there will only be two flight thermometers on the detector boxes (+ redundant thermometers TBC) and possibly a extra thermometer on the secondary optical bench for ground testing purposes.