

# SPIRE

**SUBJECT:** SPIRE Structure Subsystem Specification Document

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# SPIRE Structure Subsystem Specification Document

## Contents:

This SPIRE structure subsystem description gives an outline of the design of the instrument together with the specification of its capabilities. Furthermore, the compliance of this design with the various requirements as stated in the instrument requirements documents and the instrument interface document part A and B are compared. All interfacing subsystems are listed together with the applicable ICD's. This is for downward compatibility and traceability.

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1.1	1 October 2000	Draft revision, based upon Review of SPIRE Sub-system Specification Documents by B. Swinyard, 10 August 2000
1.2	05-November-2000	Continued revising the document conform following generic sub system specification



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## 1. Scope of the document

This specification defines the requirements applied to the performances, the design and the qualification of the SPIRE structure. It is applicable to the STM, CQM, PFM and the FS.

## 2. Documents

### 2.1. Reference documents

ID	Title	Author	Reference	Date
RD1	Instrument Requirements Document	B. M. Swinyard	SPIRE-RAL-PRJ-000034, issue 0.30	May 2000
RD2	Instrument Interface Document part A	H. Schaap/ A. Heske	PT-IID-A-04624, issue 0-2	15/02/2000
RD3	Instrument Interface Document part B	C. Cunningham B. Swinyard	PT-SPIRE-02124, issue 0-4	15/02/2000
RD4	SPIRE Product Tree	K. King	SPIRE-RAL-PRJ-00030, issue 1	September 1999
RD5	FIRST SPIRE Optical Alignment Plan	Origne, K. Dohlen	LOOM.KD.SPIRE.2000.001-1	3/1/2000
RD6	SPIRE Sub-system Budget allocations	-	-	-
RD7	Conceptual design for the 300 mK thermal strap	B. Swinyard	Technote, issue .00	24/05/2000
RD8	A stray-light baffle design for thermal strap entry ports	A. Richards	SPIRE-RAL-NOT-000344, issue 1	2 March 2000
RD9	SPIRE Instrument CQM requirements	B. Swinyard	SPIRE-RAL-NOT-000389, issue 0.1	18/04/2000
RD10	ESA fax	B. Guillaume	SCI-PT/IFI/07222	05/11/99
RD11	SPIRE Subsystem Block Diagram	B. Swinyard K. King	SPIRE-RAL-NOT-000391	26/04/2000

### 2.2. Applicable documents

ID	Title	Author	Reference	Date
AD1	Structure Design and Development Plan	B. Winter	SPIRE-MSS-PRJ-000426, issue 0.1	02-06-2000
AD2	SPIRE structure/Optics ICD	D. Pouliquen B. Winter	SPIRE-MSS-PRJ-000293, issue 0.3	06-06-200
AD3	ICD 1.1/1.5.2 structure/FTS mechanism	D. Pouliquen B. Winter	SPIRE-MSS-PRJ-000298, issue 0.2	-
AD4	ICD 1.1/1.2.1 structure/Filters	P. Hargrave B. Winter	SPIRE-MSS-PRJ-000331, issue 0.-	-
AD5	ICD 1.1/1.3 structure/cooler	L. Dubant B. Winter		
AD6				

### 2.3. Applicable drawings

ID	Title	Author	Reference	Date
DR1	Provisional SPIRE/FIRST Interface	J. Coker	A1-5264-300, issue 6, Snapshot in appendix A	30/05/00
DR2				

**2.4. Glossary**

AD	Applicable Document	Level 0	Cryostat level 0 temperature is ~ 2 K (RD1)
ATC	Astronomy Technology Centre	Level 1	Cryostat level 1 temperature is ~ 4 K (RD1)
BSM	Beam Steering Mechanism	Level 2	Cryostat level 2 temperature is ~9-11 K (RD1)
CDR	Critical Design Review	MGSE	Mechanical Ground Support Equipment
CEA	Commissariat à l' Energie Atomique	MSSL	Mullard Space Science Laboratory
CNES	Centre National des Etudes Spatiales	NA	Not Applicable
CoG	Centre of Gravity	OGSE	Optical Ground Support Equipment
CQM	Cryogenic Qualification Model	PFM	ProtoFlight Model
FIRST	Far Infra Red Space Telescope	QMW	Queen Mary and Westfield college
FOB	FIRST optical bench	RAL	Rutherford Appleton Laboratories
		SCAL	Spectrometer Calibration source
FS	Flight Spare	SMEC	Spectrometer MECHANISM
FTS	Fourier Transform Spectrometer	SOB	SPIRE Optical Bench
FOR	FIRST optical reference	SOR	SPIRE optical reference
GSFC	Goddard Space Flight Center		
I/F	Interface	SPIRE	Spectro and Photometric Image REceiver
JFET	Junction Field Effect Transistor	TBC	To Be Confirmed
JPL	Jet Propulsion Laboratory	TBD	To Be Defined
LAM	Laboratoire d'Astronomie Marseille	TBI	To Be Issued

### 3. Subsystem description

Hereafter the contents of the various subsections is explained.

- 3.1. General Overview: Explaining how the structure looks like and how it works. This includes how it fits in the overall system design, using block diagrams
- 3.2. Sub-system Design: Gives a more detailed description of the design of each of the components within the sub-system.
- 3.3. Mission profile: Gives the expected lifetime of the structure including numbers of operations (tests) it will go through for each of the models. This includes transportation.
- 3.4. Product Tree: This tree lists all the components that are delivered and by whom for each instrument model.

#### 3.1. General Overview

The SPIRE structure consists of a monocoque shell with a central optical bench. The structure is mounted on a fixed point (cone-shaped), suppressing translation in all directions. This fixed point is located on the corner of the optical bench. On the other side of the instrument (+Z direction) two A-frames are mounted suppressing each translation in the plane of the frame itself. All in all this results in a kinematic suspension with as a fixed reference with regard to the FIRST optical bench, the cone.

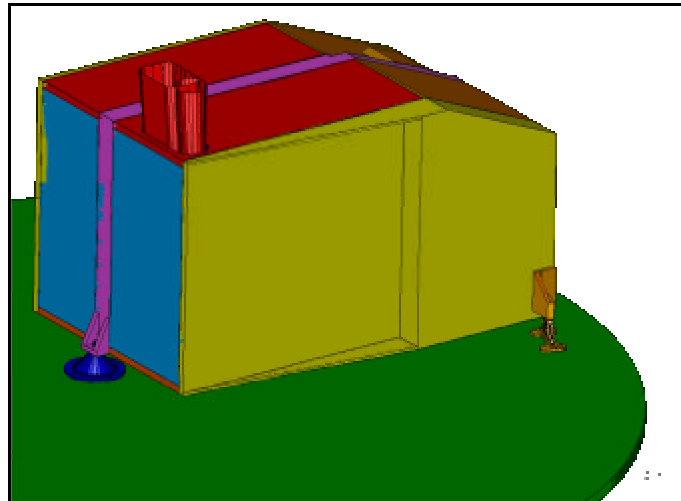


Figure 3.1-1: View of the outside of the instrument – Common Structure + Mounting

As said before the SPIRE instrument consists of a monocoque shell holding a bending stiff, internal, SPIRE optical bench. This optical bench supports a photometer and a spectrometer. All parts of these two sub-instruments are mounted on the SPIRE optical bench. See figure 3.1-2 for the photometer side of the optical bench, including IR-beams.

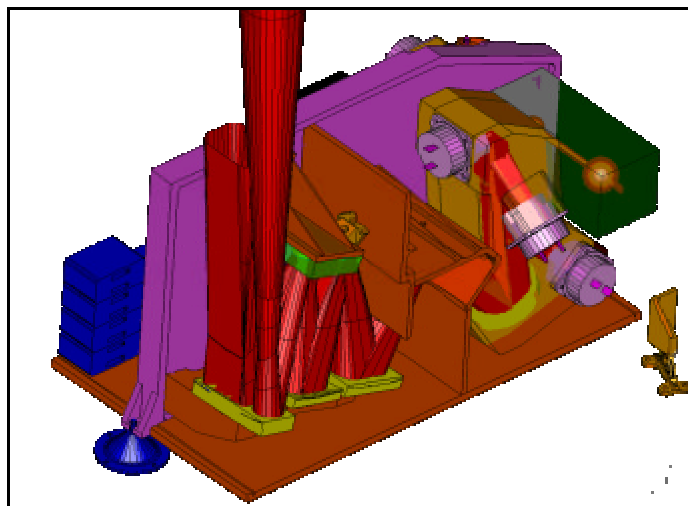


Figure 3.1-2: View of the inside of the instrument – photometer side, cover taken off



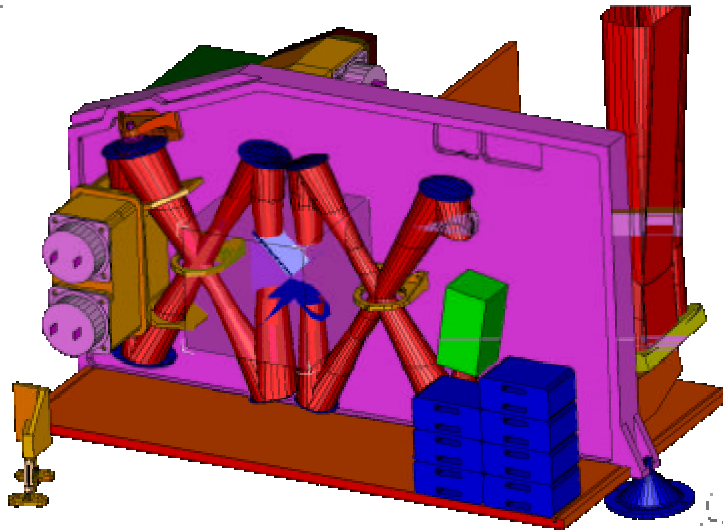
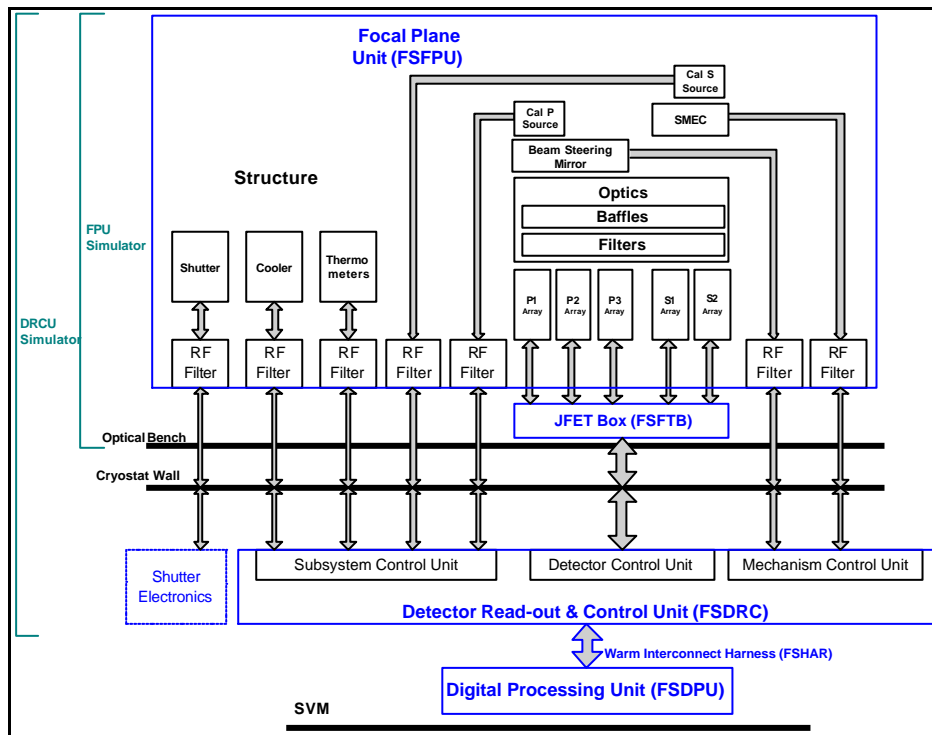


Figure 3.1-3: View of the inside of the instrument – spectrometer side, cover taken off

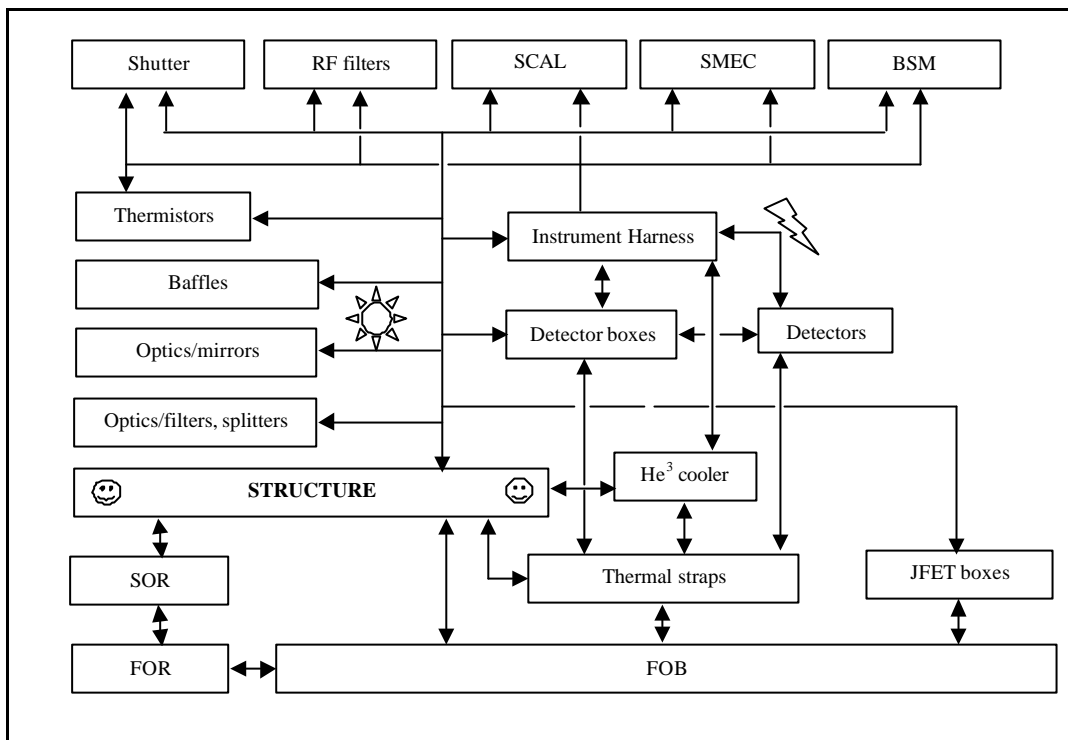
The instrument is divided into different temperature zones. The reason for this is the relative high interface temperature of the FIRST optical bench and the low operating temperature of the detectors inside the instrument. The interface temperature is 6-10 Kelvin where the operating temperature of the detectors is ~0.3 Kelvin. The temperature zones in between are the temperature of the monocoque structure with the optical bench at ~4 Kelvin and the boxes holding the detectors, filters and dichroics at ~2 Kelvin.

Hereafter the block diagram is given, showing the structure as part of the whole system.



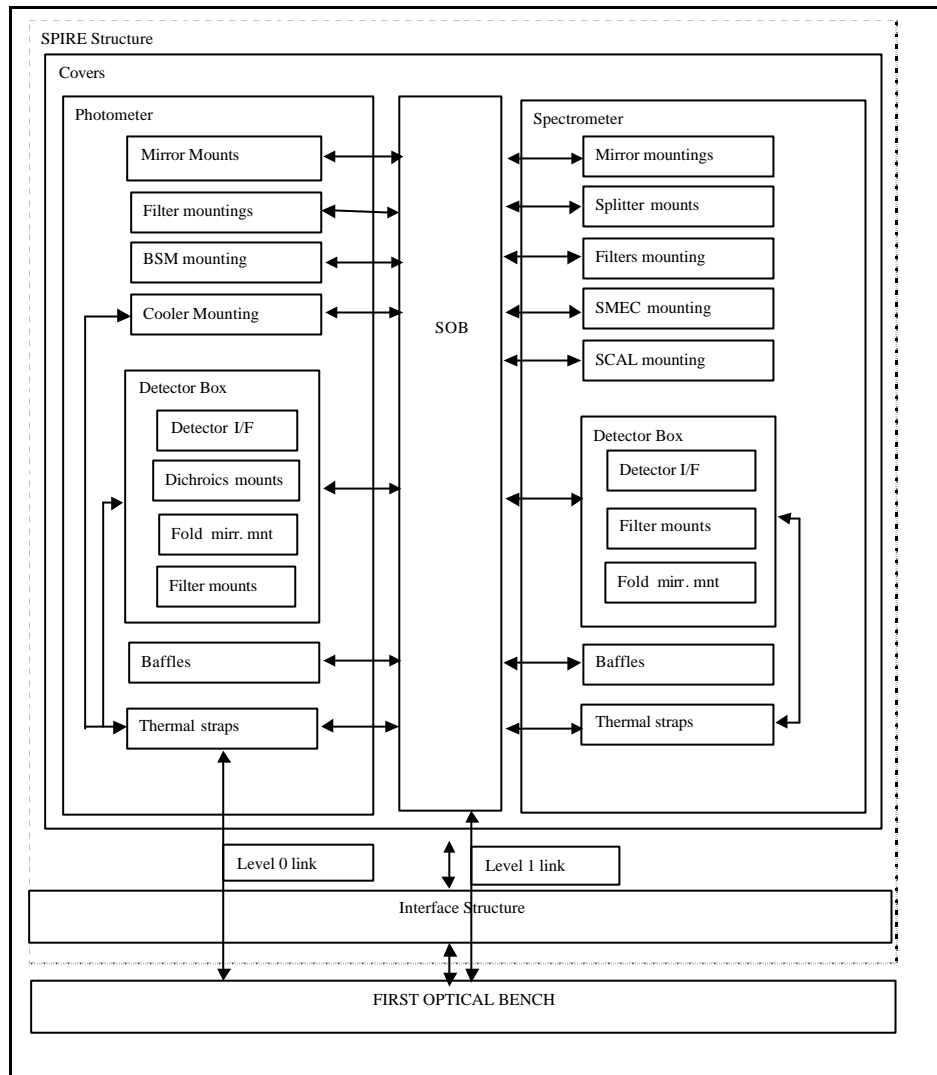
Block Diagram 3.1-1: SPIRE system

In the block diagram listed below the, many, interface with the various sub-systems is shown.



Block Diagram 3.1-2: SPIRE structure interfacing with other subsystems

In the block diagram listed below, the various components making up the structure are listed with the internal relationships.



Block Diagram 3.1-3: SPIRE structure, relations between structural components

### 3.2. Sub-system Design

The SPIRE structure is part of the SPIRE FPU, its place within the FPU is outlined in the block diagram 3.1-1. Here after the various components, together forming the design of the structure are listed and briefly addressed.

#### 3.2.1. SPIRE structure - RF sealing

As stated in section 3.1 the structure uses a monocoque shell and an optical bench to hold all different parts of the instrument. The monocoque shell also serves as RF-attenuator/shield. For this all openings/seams in the instrument need to be closed such that the instrument works as a Faraday cage attenuating the RF radiation sufficiently in the specified frequency bands. For the signal and control wiring filters will be used to ensure proper operation/functionality of the electronics. An example of an RF-seal to be used in between the joints of the common structure panels is shown in figure 3.2.1-1. MSSL is responsible for implementing the RF-attenuation with regard to the structure. In principle the consortium as a whole is responsible for the RF-shielding of the instrument

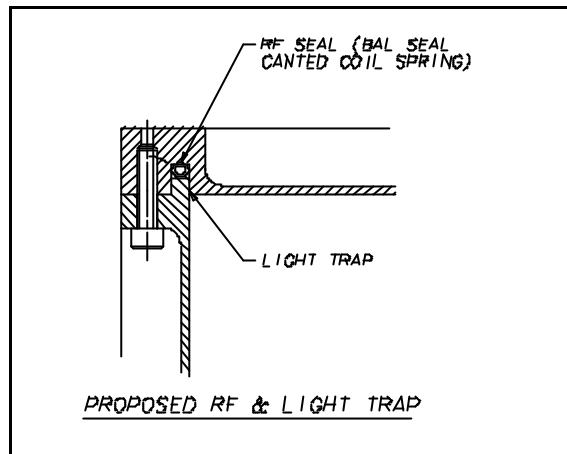


Figure 3.2.1-1

### 3.2.2. Straylight

The straylight will be attenuated as much as possible by utilising filters, black material, labyrinths and baffles where needed. The joints will have a small tight labyrinth as shown in figure 3.2.1-1. At the entry of the instrument the incoming IR beam will be filtered, excluding as much as possible the unwanted wave lengths. Inside the instrument (4 K environment) the beam will pass baffles, filters and finally, after being reflected via mirrors, end inside the detector boxes falling onto the detector noses. The proposed stray-light baffles around the thermal straps entering the common structure or the detector boxes will look like the one outlined in figure 3.2.2-1. (See for more details RD7 and RD8). MSSL is responsible for implementing the straylight shielding. RAL is responsible for the straylight modelling and advises MSSL on the implementation.

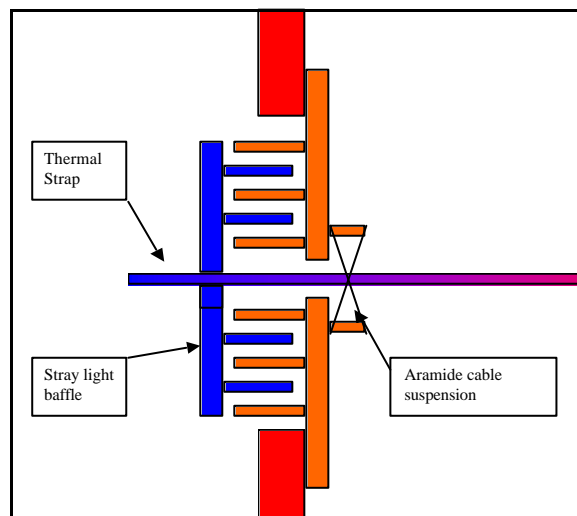
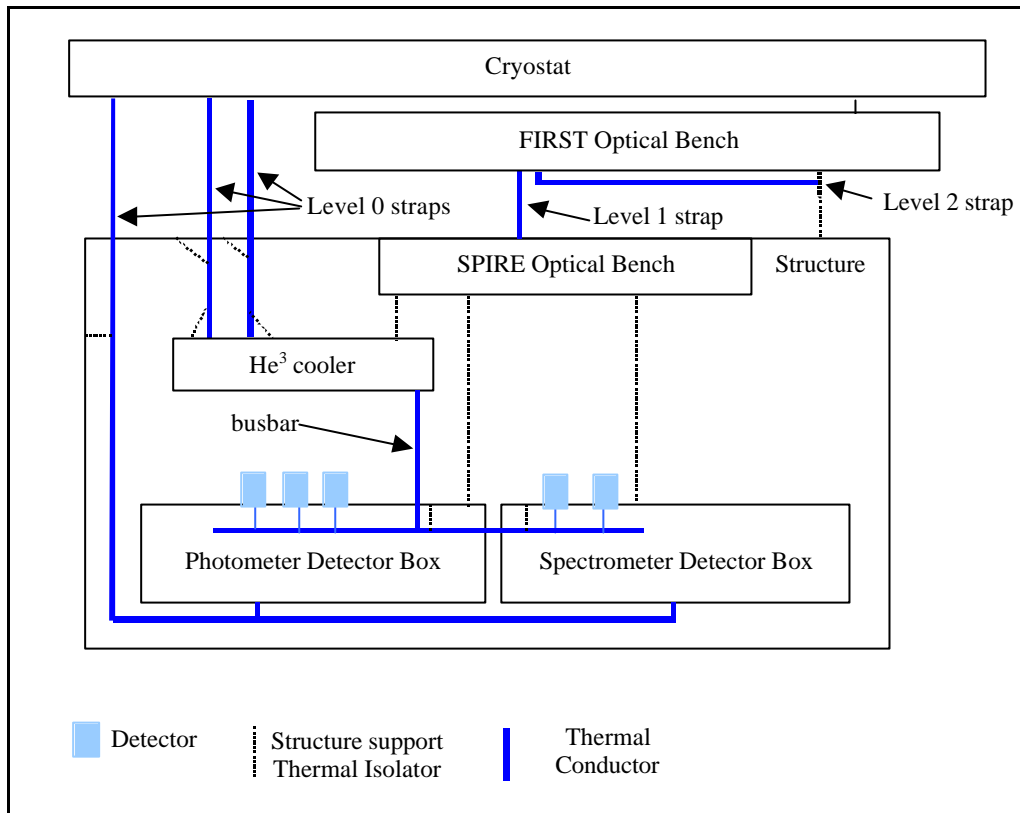


Figure 3.2.2-1: Outline of the stray-light baffle concept for thermal straps.

### 3.2.3. Thermal design

The various parts of the instrument will be connected with thermal straps to heat sinks inside the cryostat (~2 K and ~4 K). The detectors, mounted on the detector boxes, will be cooled using a thermal busbar connected to a helium-3 fridge. MSSL responsibility is manufacture and integration of the straps. The design and development responsibility is shared between CEA, RAL, QMW and MSSL.

The block diagram, listing the various thermal components in more detail than 3.1-3 does is listed hereafter in 3.2.3-1.



Block diagram 3.1-3: All thermal control hardware (except surface finishes)

### 3.2.4. Grounding

The instrument will be electrically isolated from the FIRST optical bench. The thermal straps will be electrically isolated from the cryostat. See for an outline of the baseline grounding scheme figure 3.2.4-1.

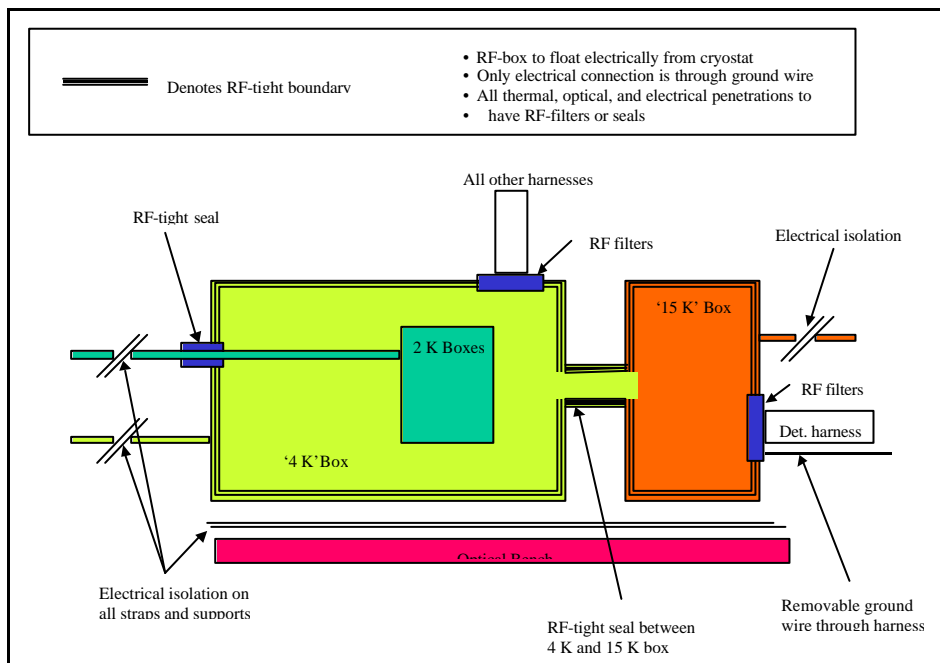


Figure 3.2.4-1: SPIRE baseline grounding scheme.

The grounding scheme is a consortium responsibility. The electrical isolation requirements will be defined by the consortium (RAL, ATC, QMW, JPL, LAM leading). And implementation by the sub-systems.

**3.2.5. SPIRE optical design**

Hereafter the topology of the optical lay-out of the instrument is given. First, the photometer side of the instrument is outlined (Figure 3.2.5-1), followed by the spectrometer side (Figure 3.2.5-2). The optical design is not a MSSL responsibility its design however defines in many ways the structure. Therefore all optical components are pictured in the two below listed figures. All relative positions of the components (following the light path) are correct.

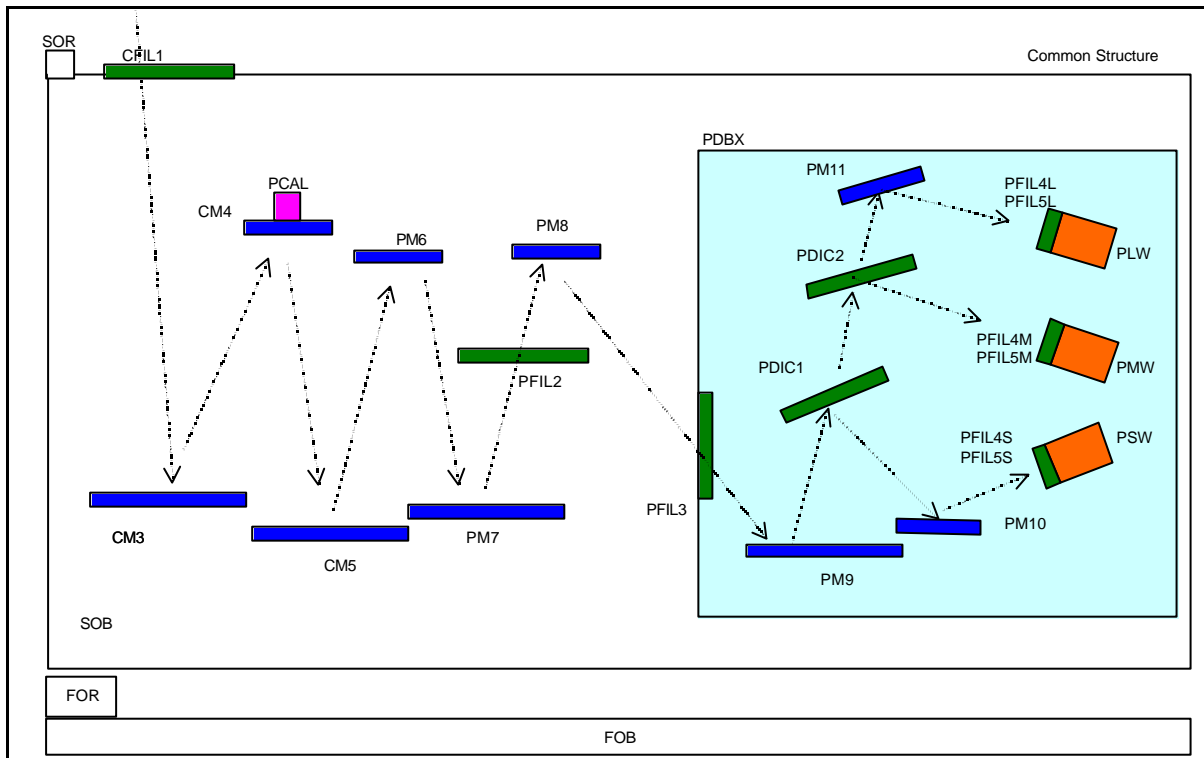


Figure 3.2.5-1: Photometer topology

Where :

- |           |   |
|-----------|---|
| CFIL1     | Common Filter 1 (entry 4 K enclosure)   |
| CM3-5     | Common Mirror 3-5                       |
| FOB       | First Optical Bench                     |
| FOR       | First Optical Reference                 |
| PCAL      | Photometer CALibrationsource            |
| PDBX      | Photometer Detector BoX                 |
| PDIC1     | Photometer DICHroic 1                   |
| PDIC2     | Photometer DICHroic 2                   |
| PFIL2     | Photometer FILTer (entrance PDBX)       |
| PFIL3     | Photometer FILTer 2 (4 K-2 K enclosure) |
| PFIL4L/5L | Photometer FILTer 4 and 5 at nose PLW   |
| PFIL4M/5M | Photometer FILTer 4 and 5 at nose PMW   |
| PFIL4S/5S | Photometer FILTer 4 and 5 at nose PSW   |
| PLW       | Photometer Long Wave detector           |
| PM6-11    | Photometer Mirror 6 to 11               |
| PMW       | Photometer Medium Wave detector         |
| PSW       | Photometer Short Wave detector          |
| SOB       | Spire Optical Bench Panel               |
| SOR       | Spire Optical Reference                 |

Not included in the overview are the shutter, baffles and the cooler.

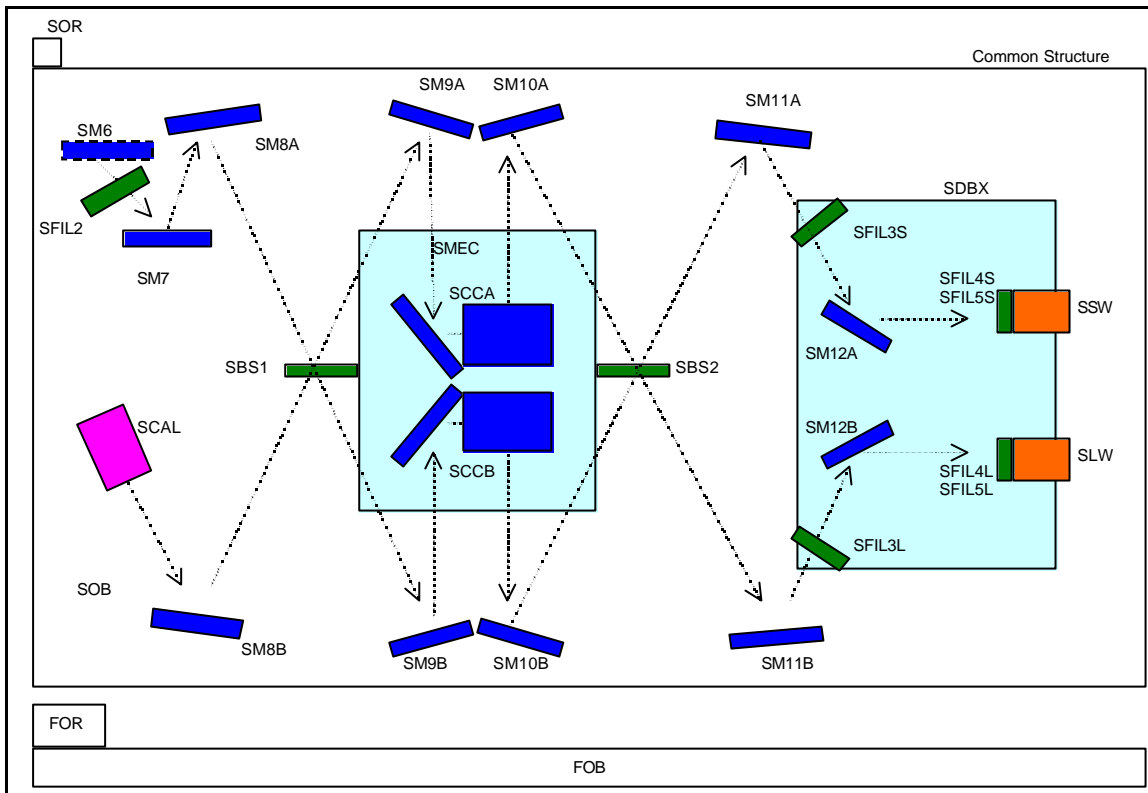


Figure 3.2.5-2: Spectrometer topology

Where:

FOB	First Optical Bench
FOR	Firs Optical Reference
SBS1	Spectrometer Beam Splitter 1
SBS2	Spectrometer Beam Splitter 2
SCCA	Spectrometer Corner Cube +X
SCCB	Spectrometer Corner Cube -X
SCAL	Spectrometer CALibration source
SDBX	Spectrometer Detector BoX
SFIL2	Spectrometer FILter 2 (4 K - 2 K enclosure)
SFIL3L	Spectrometer FILter 3 (long wave)
SFIL3S	Spectrometer FILter 3 (short wave)
SFIL4L/5L	Spectrometer FILter at nose SLW
SFIL4S/5S	Spectrometer FILter at nose SSW
SM6-7	Spectrometer Mirror6-7
SM8A-12A	Spectrometer Mirror 8-12 +X chain
SM8B-12B	Spectrometer Mirror 8-12 -X chain
SMEC	Spectrometer MECHANism
SOB	Spire Optical Bench Panel
SOR	Spire Optical Reference

Hereafter each individual subsystem mounted on or inside the SPiRE structure and therefore interfacing with it is discussed in separate sections.

### 3.2.6. SPiRE structure

In order to prevent deformation of the structure due to thermal contraction of the interface with the spacecraft, it will be suspended kinematically on the FIRST optical bench (FOB). Furthermore in order to prevent excess deformation of the instrument due to it being cooled down, it will consist of one type of material for all structural parts. The material used to construct the instrument is aluminium 6082-T6. The only deformation present after cooling down from room temperature



down to 4 Kelvin (nominally) is due to thermal contraction. Possible deformation due to thermal gradients over the structure (with all temperatures well below 6 Kelvin) will be negligible.

### **3.2.7. Structure suspension**

The structure is mounted on a fixed point (cone shape support) and two A-frames. These items should ensure sufficient mounting stiffness, such that the overall dynamics of the instrument complies with the eigenfrequency requirements. Besides that they should provide for sufficient thermal insulation such that the total heat load running through the supports is minimised. The material used for these supports is stainless steel (321). The arrangement of the supports is such that the suspension of the structure is kinematic. The fixed point suspension serves as a reference with respect to the FOB. It is the only part of the instrument that will not move relative to this bench due to possible differential thermal contraction.

### **3.2.8. SPIRE optical bench**

The SPIRE optical bench (SOB) is the part of the instrument where in principle all subsystems inside the instrument will be mounted on. The optical bench should be stiff. All items mounted on this bench will be aligned within 0.1 mm for each interface point.

### **3.2.9. Straylight baffles**

Currently the following straylight baffles are foreseen:

1. At beam entry of the instrument
2. Between PM7 and PM8, dividing the photometer in two parts. One with the common optics and one with the photometer optics, including cooler and detector box.
3. Two baffles within the spectrometer section of the instrument. Each one shielding the optics at each beam splitter

The surface finish will have an emissivity of less than 0.2. Except at the apertures, where possibly black material will be used to absorb IR-light.

### **3.2.10. Detector boxes**

On both sides of the SOB detector boxes will be mounted to provide for a level 0 enclosure. The detectors will be mounted on the outside of these boxes with the detector nose, pointing inside the box. At the apertures possibly black material will be used to absorb IR-light.

### **3.2.11. Mirror supports**

The mirrors will be mounted, baseline, on mirror mounts. These mirror mounts will be bolted on the optical bench. The exceptions are for mirrors mounted inside the detector boxes, some mirrors mounted on the central photometer straylight baffle and CM3, CM5 and CM7 mounted on a secondary optical bench. Mirror mounting is MSSL's responsibility.

### **3.2.12. Beam splitter supports**

Within the spectrometer two beam splitters are mounted on separate splitter mounts. Care will be taken to ensure the (co)planarity of these splitters.

### **3.2.13. Mounting provisions for detectors**

The detectors will be mounted on the outside of the detector boxes with the nose pointing inwards. The detector harness will exit the detectors from the back, outside the detector boxes. The 0.3 K thermal strap will exit the detectors close to the nose of the detector, inside the detector box. The thermal strap will be attached to the 0.3 K thermal busbar. The harness of the detectors will be routed via the structure outside the instrument. This routing will be RF-shielded outside the instrument.

### **3.2.14. Mounting provisions for dichroics and filters**

Two dichroics will be mounted inside the photometer detector box. Care will be taken to ensure the planarity of the dichroics. Filters will be mounted in various locations within the instrument. The first one will be located at the beam entry of the instrument. After that for each part of the instrument (photometer and spectrometer) two filters will be located in the optical train. For the spectrometer there will be one located at the optical bench beam entry. The other two at each beam entry of the





spectrometer detector box. For the photometer there will be one located at the beam passing through the central photometer straylight baffle and one at the beam entry of the photometer detector box.

### **3.2.15. Mounting provision for the SMEC**

The SMEC will be mounted on the optical bench between the two spectrometer straylight baffles (and beam splitters). There are 4 interface bolt holes. No provision for mounting a locking device (that is a SMEC internal interface). Harness will be routed via the structure to the RF-filter boxes.

### **3.2.16. Mounting provision for the <sup>3</sup>He cooler**

The cooler will be mounted next to the photometer detector box. It needs to be connected with two level 0 thermal straps to the cryostat dewar (outside the structure). The other thermal interface is with the 0.3 K thermal busbar. Harness will be routed via the structure to the RF-filter boxes.

### **3.2.17. Mounting provision for BSM**

The BSM will be mounted on the optical bench within the photometer section of the instrument. The BSM will hold the CM4 mirror and the photometer calibration source. Harness will be routed via the structure to the RF-filter boxes. The BSM will provide for its own support.

### **3.2.18. Mounting provision for spectrometer calibration source**

The spectrometer calibration source will be mounted with its own mount on the optical bench. Harness will be routed via the structure to the RF-filter boxes.

### **3.2.19. Mounting provision for shutter**

The shutter will be mounted on the optical bench. The location is before the straylight entry baffle. Effectively between the entry filter and the secondary mirror. Harness will be routed via the structure to the RF-filter boxes.

### **3.2.20. Mounting provisions for thermal straps**

The structure will provide for mounting provisions for the thermal straps and the thermal busbars. (see also section 3.2.3)

1. Level 0 thermal strap to the evaporator interface with the cooler (including straylight baffle at the entry of the structure)
  2. Level 0 thermal strap to the pump interface with the cooler (including straylight baffle at the entry of the structure)
  3. Level 0 thermal strap to the photometer detector box, continuing to the spectrometer detector box (including straylight baffle at the entry of the structure)
  4. Level 1 thermal strap, attached to the SOB on the outside of the instrument.
  5. Thermal busbar, routed from the cooler to the inside of the photometer detector box and the spectrometer detector box
- Sizing and conceptual design will be performed together with RAL and QMW.

The level 2 thermal strap has no direct interface with the instrument. The link between the level 2 vent line, and the instrument, is via the FIRST optical bench. The area around the mounting will be cooled using thermal links between the level 2 vent lines.

### **3.2.21. Mounting provisions for filters**

The various filters will be mounted in the appropriate locations.

### **3.2.22. RF-seal (attenuation) for Common Structure**

The various panels, comprising the outside of the structure will be RF-sealed. Such to attenuate incoming RF radiation sufficiently

### **3.2.23. I/F to JFET box**

The structure will interface with the two JFET boxes on the outside of the instrument. The main concerns there are RF shielding of the detector harness outside the structure, thermal insulation from the (possibly hot) JFET box and grounding.

### **3.2.24. Application of thermal finishes**

The instrument will have a low emissive finish on the inside and outside of the instrument. This to provide for a cold background inside the instrument and for thermal radiation shielding on the outside of the instrument.



**3.2.25. Application of black material**

At the various apertures within the instrument where the IR-beam will pass through possibly black material will be applied to absorb excess/stray IR-light.

**3.2.26. Mounting provision for RF-filter box**

All harness inside the instrument (except for the detector harness) will be routed via the structure to the RF-filter boxes, These are located within the spectrometer section of the instrument. The RF-filter boxes will interface with the cryo-harness inside the cryostat. This harness comprises the harness from the cooler, the various mechanisms, calibrators and the thermistors.

**3.2.27. Thermometers**

There is a requirement to have thermometers to monitor the temperature of the various structure parts. Such as the common structure, SPIRE optical bench and the detector boxes. In principle the harness of the sub-systems mounted inside the structure will be used as much as possible to route the harness of the thermometers. The actual number of thermometers that have to be mounted and routed by MSSL is TBD.

**3.3. Mission profile**

All parts will be baked out before integration (TBC).

The SPIRE CQM structure will be subjected to several tests. These tests consist of an STM test, which is a structural test verifying the dynamic properties of the structure (stiffness, strength, mass), and the CQM test programme (requirements outlined in RD9). After that a PFM model will be built, tested and mounted on the spacecraft. On the spacecraft it will go through the spacecraft proto-flight programme. The CQM will be refurbished (TBC) to serve as flight spare.

In the table hereafter (3.3-1) the allowable times a certain handling/test is performed or the allowable amount of time

Model:	STM	CQM	PFM	FS
Integration	2	2	5	1
Cold cycles (down to 2-4 K)	5 (TBC)	15 (TBC)	5 (TBC)	5 (TBC)
Mounting and dismounting detector boxes	15 (TBC)	15 (TBC)	5 (TBC)	5 (TBC)
Mounting and dismounting mirror mounts	15 (TBC)	5 (TBC)	3 (TBC)	3 (TBC)
Mounting and dismounting instrument suspension	15 (TBC)	15 (TBC)	5 (TBC)	5 (TBC)
General use <sup>1</sup>	1 Yr	2 Yr	2 Yr	1 Yr
Storage	1 Yr	2 Yr	2 Yr	2 Yr
In orbit	-	-	4.25 Yr	4.25 Yr

Table: 3.3-1: Overview life of the structure (split in the different models)

<sup>1</sup> General use encompasses 1 qualification test or two proto-flight test programmes. If the structure is needed for more tests, fatigue critical components need to be replaced.

After successfully, withstanding the qualification tests, the PFM will be constructed. The PFM will be subjected to the same tests as the CQM (except STM testing). After successfully withstanding the proto-qualification tests the PFM will be integrated in the spacecraft. Integrated in the spacecraft it will be subjected to the spacecraft proto-qualification programme. If these tests prove successful the spacecraft will be launched. More details can be found in AD1. The CQM model will be refurbished to act as a flight spare. The FS will be interchangeable with the PFM unit.

If the FS goes through a proto-qualification programme or a qualification programme it needs to be refurbished

**3.4. Product tree**

Hereafter the parts as listed in the product tree (RD4) are listed.

ID	Product Item	Descriptor	STM	AVM	CQM	PFM	FS	RESP
1.1	Structure	STRC	x		X	x	x	MSSL
1.1	Common Structure (incl. Interface wit JFET and RF filter box and FIRST optical bench)	C4K	x		X	x	x	MSSL
1.1	Photometer detector box	PDBX	x		X	x	x	MSSL



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1.1	Spectrometer detector box	SDBX	x		X	x	x	MSSL
1.1	Mounting provisions for thermal straps		x		X	x	x	MSSL
1.1	Mounting provision for thermal interface with level 0		x		X	x	x	MSSL
1.1	Mounting provision for thermal interface with level 1		x		X	x	x	MSSL
1.1	Mounting provisions for thermal straps from cooler to detectors		x		X	x	x	MSSL
1.1	Alignment reference mirror	SOR	x		X	x	x	MSSL
1.1.1.1	Structure thermistors	T_PDBX_1 T_PDBX_2 T_SDBX_1 T_SDBX_1 T_C4K_1 T_C4K_2	x		X	x	x	MSSL
1.2	Photometer mirror mounts		x		X	x	x	MSSL
1.2	Spectrometer mirror mounts		x		X	x	x	MSSL
1.2.1	Filters, Dichroics and beam splitters mountings		x		X	x	x	MSSL
1.2.2	Baffles	BAFF	x		X	x	x	MSSL
1.3	Cooler I/F	COOL	x		X	x	x	MSSL
1.4.1	Detector I/F		x		X	x	x	MSSL
1.4.2	Detector I/F		x		X	x	X	MSSL
1.5.2	SMEC I/F		x		X	x	X	MSSL
1.5.3	Shutter I/F		x		X	x	X	MSSL
1.6.2	Spectrometer calibration I/F		x		X	x	X	MSSL
1.7	JFET I/F		x		X	x	X	JPL/MSS L
1.7	RF I/F		x		X	x	X	JPL/MSS L
5.3	Mechanical ground support equipment	MGSE	x		X	x	X	MSSL
	Mathematical model structure	FEM						MSSL

A more detailed list of deliverables is part of the development plan (AD1)

## 4. Specification

### 4.1. Functional

#### 4.1.1. Performance specification

The focal plane unit structure will hold all cold sub-systems in the focal unit. This includes all thermometers necessary to monitor the instrument during cool down and operation.

Hereafter the performance of the subsystem is specified and compared with the required performances and general requirements as set out in RD1.

The mass requirements are listed in the budget allocations (RD6). The mass of the current baseline design meets the strict requirements as set out in section 3.4.3 of RD1.

IRD-SUBS-R03: The mass of the structure, including 20% contingency instrument is 32.6 kg



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Included in this mass are the straylight baffles, mirror mounts, filter splitter and dichroic clamps/mounts. The straylight baffles are not strictly a part of the structure, nor are the various subsystem supports. However they will be deliverables and are therefore added to this mass budget. The budget does not include the fasteners required to mount the various subsystems in/on the common structure or one of the detector boxes. See for a more detailed mass allocation section 4.1.2, with the technical (implementation) requirements.

The subsystem structure is further specified hereafter. The specifications are compared with the requirements.

The alignment of the structure with respect to the FOB depends on the machining tolerance of the FOB. At present the machining tolerance is expected to be within +/- 0.05 mm for each interface bolt.

ID	Description	Value	Source
IRD-STRC-R01	Alignment of the instrument w.r.t. FIRST optical axis	The common structure shall allow the alignment of the instrument and the telescope axis to within +/- 2.5 mm (TBC) lateral, TBD arcmin rotational about any axis.	RD1/table 3.5.-1

The structure consists of aluminium panels, milled down to a nominal thickness of 1.5 mm. At the connection between the panels an RF-attenuating seal is provided. See figure 3.1-4.

ID	Description	Value	Source
IRD-STRC-R02	Attenuation of RF by Common Structure	The covers as fitted on the instrument will attenuate all frequencies lower than 8 GHz by TBD dB.	RD1/table 3.5.-1

The photometer and spectrometer will be mounted on the SOB.

ID	Description	Value	Source
IRD-STRC-R03	Items required support from the Common Structure	Photometer and common subsystems, Spectrometer	RD1/table 3.5.-1

The structure will be aligned as specified in RD5.

ID	Description	Value	Source
IRD-STRC-R04	Optics and associated sub-system alignment	Specified in RD5	RD1/table 3.5.-1

The structure will have a surface finish with an  $\epsilon < 0.2$

ID	Description	Value	Source
IRD-STRC-R05	Surface finish of the Common Structure cover	The inside and the outside of the box shall have a finish with a low emissivity. At least less than $\epsilon = 0.2$ . Parts may be blackened as part of stray light control.	RD1/table 3.5.-1

The total effective pumping capacity has to be analysed.

ID	Description	Value	Source
IRD-STRC-R06	Pumping port	The total effective pumping conductance of the common structure shall be $\geq 7.8$ l/s	RD1/table 3.5.-1

The structure will be equipped with the specified number of thermometers

ID	Description	Value	Source
IRD-STRC-R07	Thermometry	See 3.3/1.1.1.1	RD1/table 3.5.-1

The surface finish on the outside of the instrument will have an emissivity with an  $\epsilon < 0.2$

ID	Description	Value	Source
IRD-STRC-R08	Attenuation of radiation from cryostat	Requirement $< 2 \times 10^{-5}$	RD1/table 3.5.-1

The structure, supported on its mounts, will have an eigenfrequency  $> 120$  Hz



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ID	Description	Value	Source
IRD-STRC-R09	First natural frequency	The structures eigenfrequency shall be above 100 Hz (req.) and preferably above 120 Hz (goal)	RD1/table 3.5-2

The instrument will be mounted on the FOB with electrically insulating washers.

ID	Description	Value	Source
IRD-STRC-R10	Instr. mechanical interface	The I/F will be directly to the FIRST optical bench. The instrument will be in direct thermal contact	RD1/table 3.5-2

All parts of the structure will be electrically connected, one to the other, with an resistance less than than 0.1 W.

ID	Description	Value	Source
IRD-STRC-R12	Grounding	All parts of the SPIRE structure shall be electrically connected one to another. Resistance to be no more than 0.1 W (TBC)	RD1/table 3.5-2

The instrument will be mounted on the FOB with electrically insulating washers.

ID	Description	Value	Source
IRD-STRC-R13	Electrical isolation from FIRST	All parts of the SPIRE structure shall be electrically isolated from the FIRST optical bench and cryostat. Resistance to be greater than TBD W	RD1/table 3.5-2

The conductance from level 2 to level 1, via the supports of the instrument, will be less than 6 mW with boundary temperatures of 9.0 K and 4.0 K respectively.

ID	Description	Value	Source
IRD-STRC-R14	Thermal isolation	The conductance from the level 2 to level 1 stage is required to be no more than 6 mW (TBC) assuming level 2 is 9 K and level 1 is 4 K.	RD1/table 3.5-2

The photometer detector box shall support the level 0 optics, dichroics, filter, detectors, straylight baffles and thermal strap supports for the detectors

ID	Description	Value	Source
IRD-STRP-R01	Items supporting	The photometer detector box shall support the level 0 optics, dichroics, filter, detectors, and thermal strap for detectors	RD1/table 3.5-3

The structure will be aligned as specified in RD5.

ID	Description	Value	Source
IRD-STRP-R02	Optics and filter alignment	See RD5	RD1/table 3.5-3

The structure will be aligned as specified in RD5.

ID	Description	Value	Source
IRD-STRP-R03	Array module alignment	See RD5	RD1/table 3.5-3

The structure will have a surface finish with an  $\epsilon < 0.2$

ID	Description	Value	Source
IRD-STRP-R04	Surface finish	The outside of the box shall have a finish with a low emissivity. At least $\epsilon = 0.2$ . The inside shall have a low reflective finish on all non-optical surfaces..	RD1/table 3.5-4

The effective pumping capacity has yet to be analysed.

ID	Description	Value	Source
IRD-STRP-R05	Pumping port	The total effective pumping conductance of the detector box shall be $\geq 5.6$ l/s	RD1/table 3.5-4



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The structure will have a surface finish with an  $\epsilon < 0.2$

ID	Description	Value	Source
IRD-STRP-R06	Attenuation from common structure	Requirement: $5 \times 10^{-7}$ with a goal of $5 \times 10^{-8}$	RD1/table 3.5-4

The first eigenfrequency of the photometer detector box on its mounts will be greater than 150 Hz

ID	Description	Value	Source
IRD-STRP-R07	First natural frequency	The first eigenfrequency of the photometer detector box on its mounts shall be greater than 100 Hz, with a goal of 150 Hz	RD1/table 3.5-4

The thermal conductance from the structure to the photometer detector box will be less than 0.75 mW via the supports, with a boundary condition of 4.0 K and 2.0 K respectively.

ID	Description	Value	Source
IRD-STRP-R09	Thermal isolation	The conductance from the common structure to the detector box shall be less than 0.75 mW. (TBC)	RD1/table 3.5-4

The spectrometer detector box shall support the level 0 optics, filter, detectors, straylight baffles and thermal strap supports for the detectors.

ID	Description	Value	Source
IRD-STRS-R01	Items supporting	The photometer detector box shall support the level 0 optics, filter, detectors, and thermal strap for detectors	RD1/table 3.5-4

The structure will be aligned as specified in RD5.

ID	Description	Value	Source
IRD-STRS-R02	Optics and filter alignment	See RD5	RD1/table 3.5-4

The structure will be aligned as specified in RD5.

ID	Description	Value	Source
IRD-STRS-R03	Array module alignment	See RD5	RD1/table 3.5-4

The structure will have a surface finish with an  $\epsilon < 0.2$

ID	Description	Value	Source
IRD-STRS-R04	Surface finish	The outside of the box shall have a finish with a low emissivity. At least $\epsilon = 0.2$ . The inside shall have a low reflective finish on all non-optical surfaces..	RD1/table 3.5-4

The effective pumping conductance of the structure has yet to be analysed

ID	Description	Value	Source
IRD-STRS-R05	Pumping port	The total effective pumping conductance of the detector box shall be $\geq 5.6$ l/s	RD1/table 3.5-4

The structure will have a surface finish with an  $\epsilon < 0.2$

ID	Description	Value	Source
IRD-STRS-R06	Attenuation from common structure	Requirement: $5 \times 10^{-7}$ with a goal of $5 \times 10^{-8}$	RD1/table 3.5-4

The first eigenfrequency of the spectrometer detector box on its mounts will be greater than 150 Hz

ID	Description	Value	Source
IRD-STRS-R07	First natural frequency	The first eigenfrequency of the spectrometer detector box on its mounts shall be greater than 100 Hz, with a goal of 150 Hz	RD1/table 3.5-4

The thermal conductance from the structure to the photometer detector box will be less than 0.75 mW via the supports, with a boundary condition of 4.0 K and 2.0 K respectively.

ID	Description	Value	Source
IRD-STRS-R09	Thermal isolation	The conductance from the common structure to the detector box shall be less than 0.75 mW. (TBC)	RD1/table 3.5-4

**4.1.2. Technical requirements**

The following co-ordinate system as reference to the cryostat shall be taken into account (RD3). See also DR1 and appendix A.

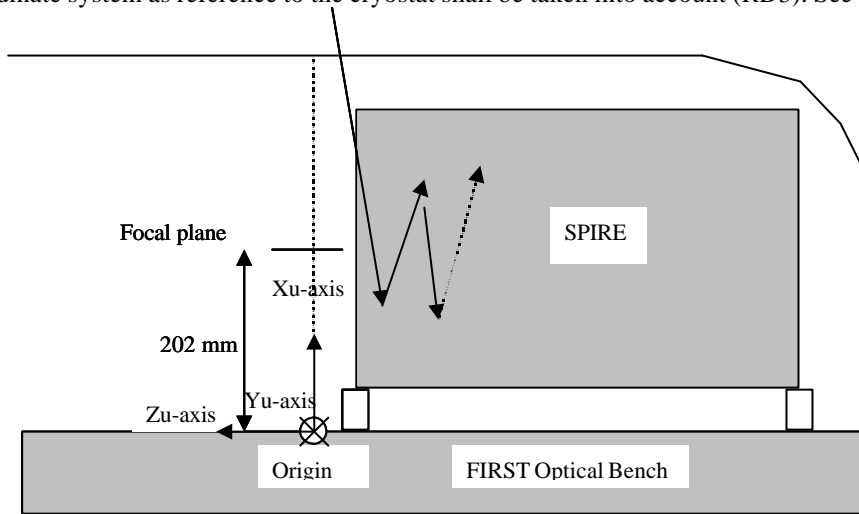


Figure 4.1.2-1 Cryostat reference co-ordinate system.

The following local co-ordinate system shall be taken into account for the SOB mounted subsystems. See also DR1 and appendix A. The origin lies on the optical bench datum point. This is visible and reachable both before, during and after integration. It will however not be visible from outside the spacecraft cryostat. As optical reference an optical alignment cube will be used, mounted on top of the optical bench, outside the common structure, closest to the spacecraft optical axis and visible from outside the cryostat. The advantage of the datum point on the optical bench is that all co-ordinates inside the instrument are positive. Both the co-ordinate systems, are co-aligned with the S/C co-ordinate system, only the origins differ.

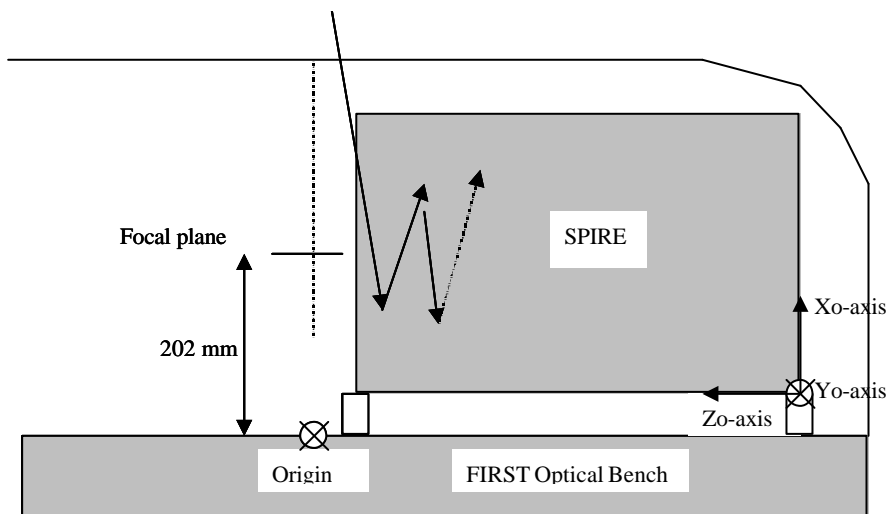


Figure 4.1.2-2 Spire optical bench reference co-ordinate system.

For the current baseline design the following mass distribution holds:



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SPIRE – structure				
Contingency taken into account: 20%				
	Nett	Contg.	Total	
Photometer cover	7.43	1.49	8.92	[kg]
Spectrometer cover	5.69	1.14	6.82	[kg]
Optical bench	7.10	1.42	8.52	[kg]
Mounting common structure	0.57	0.11	0.69	[kg]
RF-attenuation	0.25	0.05	0.30	[kg]
Phot. det. box	1.58	0.32	1.89	[kg]
Spect. det. box	1.10	0.22	1.32	[kg]
Mounts, Clamps Phot.	1.04	0.21	1.17	[kg]
Mounts, Clamps Spect.	1.06	0.21	1.42	[kg]
Straylight Baffles	1.38	0.28	1.66	[kg]
total	27.20	5.44	32.64	[kg]
Excluding:	Thermal straps			
	Harness wiring			
	RF-filter boxes			
	JFET I/F structure			
	Spect. Cal.			
	Shutter			
	Cooler I/F			
	Straylight baffles (around straps)			

Table 4.1.2-1: Mass distribution SPIRE structure and related

Implementation requirements still to be defined:

- Thermal flow characteristic through A-frame and fixe d point suspension. (RD3, section 5.9.1.2, total 6 mW)
- Thermal flow characteristic through detector box suspension. (RD3, section 5.9.1.2, total 1.5 mW)

**4.2. Operational**

**4.2.1. Operational Safety**

N.A.

**4.2.2. Lifetime**

Expected integrated lifetime is 9 years, including on orbit life, TBC

**4.2.3. Operating modes**

N.A.

**4.2.4. Telemetry**

NA

**4.2.5. Telecommands**

NA





**4.3. Interface requirements**

The interface with the FIRST optical bench is given by ja1-5264-300, pictured in appendix A.

Hereafter all subsystem interfaces and the applicable ICD's are listed. The ICD's contain the interface drawings

ICD	Part name	ICD	Other party
1.1 - 1.2	Mirrors/Optics	AD2	LAS
1.1 - 1.5.2	FTS mechanism	AD3	LAS
1.1 - 1.2.2	Straylight attenuation	TBI	RAL
1.1 - 1.2.1	Filters, splitters & dichroics	AD4	QMW
1.1 - 1.4.1/2	Detectors	TBI	JPL
1.1 - 1.5.1	BSM	TBI	ATC
1.1 - 1.6.1	Calibration source Photometer	TBI	QMW?
1.1 - 1.6.2	Calibration source Spectrometer	TBI	QMW?
1.1 - 1.5.3	Shutter mechanism	TBI	UoS
1.1 - 1.3	He <sup>3</sup> Cooler	AD5	CEA-Grenoble
1.1 - ?	Thermal hardware	TBI	?

Table 4.3-1: ICD list

Hereafter the general requirements for the subsystems are given.

The sine and random input at the interfaces considered in this document will be updated after the coupled analysis. The input spectra hold for all interfaces with subsystems mounted of the optical bench. The spectra given here are a guideline and are subject to changes in the IID-A. At all times the spectra specified in the individual ICDs will supersede the ones specified below. Again, the spectra defined in the individual ICDs are subject to changes in the IID-A which has not been issued officially at the time this document was written.

**Quasi-Static**

The qualification levels are **provisional**, copied from the qualification loads of ISO. Since they are less severe than the specified sine loads, the requirement is covered by these sine vibration loads. (RD1)

Quasi Static levels	Case 1	Case 2	Case 3	Case 4
x-direction	22.5 g (TBC)	22.5 g (TBC)	-	-
y-direction	3 g (TBC)	-	6 (TBC)	-
z-direction	-	3 g (TBC)	-	6 (TBC)

Table 4.3-1: Qualification levels for quasi static vibration

**Sine**

The qualification levels are derived from the coupled analysis with the instrument structure with the input at base of the instrument following the requirements stated in AD\*\*. These levels include the required qualification margin (factor 1.5). They do not include any further margin with respect to design loads. These will have to be added by the subsystems.

Sine vibration levels	Frequency range	Input at base (QUAL)
X-direction	5-18 Hz	22 mm (peak-peak)
	18-100 Hz	40 g
Y-direction	5-18 Hz	22 mm (peak-peak)
	18-100 Hz	50 g
Z-direction	5-18 Hz	22 mm (peak-peak)
	18-100 Hz	50 g

Table 4.3-2: Qualification levels for sine vibration (sweep rate 2.0 oct/min)

**Random**

The qualification levels are **provisional**, copied from the qualification loads of ISO.



Random vibration levels	Frequency range	Input at base (QUAL.)
X-direction	5 – 150 Hz	+6 dB
	150-700 Hz	0.15 g <sup>2</sup> /Hz
	700 – 2000 Hz	-3 dB
Y-direction	5 – 150 Hz	+6 dB
	150-700 Hz	0.15 g <sup>2</sup> /Hz
	700 – 2000 Hz	-3 dB
Z-direction	5 – 150 Hz	+6 dB
	150-700 Hz	0.15 g <sup>2</sup> /Hz
	700 – 2000 Hz	-3 dB

Table 4.3-3: Qualification levels for random vibration (20 g<sub>rms</sub>)

#### **4.4. Design and manufacture requirements**

##### **4.4.1. Design requirements**

The key baseline design requirements and implementation are the following

##### 1 -Stiffness

In order to provide for the required stiffness the baseline design of the structure consists of a monocoque structure with a central stiff optical bench panel. In principle all sub-systems are mounted on the optical bench panel. In order to comply with the sine vibration input specification the design goal set for the structure is to reach an first natural frequency of 120 Hz or higher. Because of the uncertainty taken into account for the frequency prediction of 10% the baseline design should have its predicted first natural frequency of at least 133 Hz. Furthermore the various sub-systems need to be decoupled (dynamically) as much as possible. The eigenfrequency of the SOB should be at least 20 % higher than the first overall natural frequency (160 Hz). The minimal eigenfrequency of all subsystems and detector boxes should be at least 25% higher (200 Hz). Ideally the eigenfrequency of a subsystem mounted on an other should be a factor 1.41 higher in order to avoid any amplification due to the coupling. This is in practice not possible.

##### 2 - Thermal

The structure will be thermally isolated from the FIRST optical bench as much as possible. Assuming a 6.0 K interface temperature and a 4.0 K structure temperature the maximum heat flow through the mounting of the structure should be less than 2 mW.

##### 3 - Electrical

The structure will be electrical isolated from the FIRST optical bench. This to allow for controlled grounding of the structure and its various subsystems. The thermal straps routed along various locations within the structure will be electrically isolated from the structure.

##### **4.4.2. Design rules**

No specific design rules as yet.

##### **4.4.3. Manufacture requirements**

TBD

#### **4.5. Logistic requirements**

RD2, 5.15.1.1: For all deliverable units, transport containers shall be provided.



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1. Each model will have its transportation container and handling frame. Only one set of integration tools is required. (If applicable) Since the FS will be a refurbished CQM model, only two sets of transportation containers and handling frames are required. One for the CQM/FS and one for the PFM.
2. The subsystem will be transported from MSSL to RAL, and from RAL to ESTEC (TBC)
3. The container shall be vacuum tight, be purged and slightly overpressured with dry nitrogen gas.
4. The instrument will be mounted using shock absorber supports
5. Shock recorders shall be mounted at TBD location
6. IID-B shall list size and mass of the container as well as the overall mass including the instrument package

### 4.6. Environment requirements

These requirements describe the environment the structure needs to be able to sustain during its life. Life includes all ground operations and tests, launch and all on orbit operations. For the structure it is assumed that the warm vibration test, followed by the CQM qualification programme envelopes the worst case environment for the structure. Therefore the structure

#### 4.6.1. Natural environment

This is the description of the natural environment around the subsystem during its life.

#	Parameter	Value	IRD	Note
EN1	Vacuum	Less than 10 <sup>-4</sup> Pa		In operation
EN2	Operating temperature	during system qualif and on orbit = 4K		
EN3	Storage and handling temperature	-20 to +30 °C		Overall, on ground
	Humidity	Less than 45%		In clean room
	Cleanliness	TBD		In clean room
EN4	Radiations	Less than 3.5 kRAD		On orbit

#### 4.6.2. Operating environment

N.A.

### 4.7. Verification requirements

RD1 lists in section 2.6 the verification requirement for the instrument. They are repeated hereafter and the compliance is indicated.

ID	Description	Response	Compliant
IRD-VER-01	STM verification: stiffness and strength of the structure	STM sine vibration test (warm)	yes
IRD-VER-R02	AVM verification	Not applicable	-
IRD-VER-R03	CQM verification		
	1 correct operation (all operating modes) at cryogenic temperatures	1 Cryogenic test	1 yes
	3 thermal dissipation	3 Cryogenic test	3 yes
	6 Straylight	6 Cryogenic test	6 yes
IRD-VER-R04	PFM and FS same as for CQM verification (structure)	Cryogenic test	yes

Prior to the STM and CQM test, mathematical models will be used to verify the strength and stiffness of the proposed structure. The mechanical tests will be simulated for as far as the software allows doing that. The principle loads will be derived from these analyses and forwarded to all subsystems, if required, via the ICD.

At present the draft mechanical environment requirements as issued by ESA (RD10) are listed hereafter. These apply for the structure subsystems as a whole. Not for the subsystems individual.



**Mechanical environment**

**Quasi-Static**

The qualification levels are **provisional**, copied from the qualification loads of ISO.

Quasi Static levels	Case 1	Case 2	Case 3	Case 4
x-direction	22.5 g (TBC)	22.5 g (TBC)	-	-
y-direction	3 g (TBC)	-	6 (TBC)	-
z-direction	-	3 g (TBC)	-	6 (TBC)

Table 2.4-1: Qualification levels for quasi static vibration

**Sine**

The sine vibration environment qualification loads taken from ESA fax, SCI-PT/IFI/07222, d.d. 05/11/99. These are the input acceleration at the base of the SPIRE instrument. The sweep rate is 2.0 oct/min (TBC) RD2 echo's this fax and adds the low level sweep to the specification.

Sine vibration levels	Frequency range	Input at base (QUAL)
X-direction	5-18 Hz	22 mm (peak-peak)
	18-70 Hz	15 g
	70-100 Hz	7.5 g
Y-direction	5-18 Hz	22 mm (peak-peak)
	18-50 Hz	15 g
	50-100 Hz	7.5 g
Z-direction	5-18 Hz	22 mm (peak-peak)
	18-50 Hz	15 g
	50-100 Hz	7.5 g

Table 2.4-2: Qualification levels for sine vibration

For all axis the a signature run of 0.5 g from 5.0 to 2000 Hz will be performed as the first and the last sweep. RD3, table 9.5.3.3.2-2.

**Random**

The qualification levels are **provisional**, copied from the qualification loads of ISO.

Random vibration levels	Frequency range	Input at base (QUAL.)
X-direction	5 – 150 Hz	+6 dB Hz
	150-700 Hz	0.04 g <sup>2</sup> /Hz
	700 – 2000 Hz	-3 dB
Y-direction	5 – 150 Hz	+6 dB Hz
	150-700 Hz	0.04 g <sup>2</sup> /Hz
	700 – 2000 Hz	-3 dB
Z-direction	5 – 150 Hz	+6 dB Hz
	150-700 Hz	0.04 g <sup>2</sup> /Hz
	700 – 2000 Hz	-3 dB

Table 2.4-3: Qualification levels for random vibration



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Appendix A

Interface drawing SPIRE-FIRST

