

**FIRST**

Update of Instrument Interfaces



DaimlerChrysler Aerospace  
Dornier

Dornier Satellitensysteme GmbH

SPIRE - REF - MHO - 000318

**FIRST**

**Cryostat Study**

**Update of Instrument Interfaces**



## AGENDA

1. Introduction
2. Status of Action Items
3. FIRST Optical Bench Design
4. Optical Interfaces
5. Optical Bench Analysis
6. Instrument Harness evaluation status
7. Local Oscillator Unit
8. Input to IID A update
9. AOB



## 1. Introduction

### Main Objectives / Tasks

- Implementation of present FIRST instrument design
- Check of cryostat performance after implementation
- Detail of instrument interfaces



## 2. Action Item Status

### Action Items of Kick-off Meeting (10.09.1999)

Action Item #	Actionee	Subject / Description	Due Date	Status	Reference
01	ESTEC	ESTEC to verify date for PM#2	01.10.1999	closed	SCI-PT/ISI/07119, 27.10.1999
02	PACS	PACS to provide beam envelope at the level of the cryostat aperture as input to the baffle optimisation	01.10.1999	closed	Email, N. Geis, 08.11.1999
03	HIFI	HIFI to provide beam envelope at the level of the cryostat aperture as input to the baffle optimisation	01.10.1999	closed	Email, N. Nhyborn, 03.11.1999
04	HIFI	HIFI to provide update of the cryo-harness list for 15K to 300K	01.10.1999	open	
05	HIFI	HIFI to confirm diameter of LOU windows (30 mm) and distance between the LO-beams of 50 mm	01.10.1999	closed	Verbally in HIFI Meeting, 21.10.1999
06	HIFI	HIFI to provide dimensions and masses of waveguides between LOU and SVM	01.10.1999	open	
07	HIFI	HIFI to provide a detailed mass breakdown of LOU	01.10.1999	open	



**Action Items of Progress Meeting # 1 (20.10.1999)**

Action Item #	Actionee	Subject / Description	Due Date	Status	Reference
01	SPIRE	SPIRE to provide updated thermal budget for GSFC detector array	n.a.	closed	During progress meeting by Annex 1 to MoM
02	HIFI/ PACS/ SPIRE	Instrument teams to agree on / base instruments design on: 1. Mounting and dismounting of one instrument can be performed independently of all other ones 2. The mounting interfaces of the optical bench itself shall be accessible during all integration levels (for definition see Annex 3 to MoM of PM#1) 3. Instrument mounting feet shall compensate for thermal displacements between instrument and optical bench	PM#2 23.11.1999	open	
03	SPIRE	SPIRE to update location of SPIRE Filter Box to avoid conflict with optical bench mounting interface	03.11.1999	open	
04	ESTEC	For arrangement of the instruments on the optical bench a clearance requirement of 10 mm between instruments shall be implemented in the IID-A		open	



**Action Items of Progress Meeting # 1 (20.10.1999), cont'd**

Action Item #	Actionee	Subject / Description	Due Date	Status	Reference
05	ESTEC	ESTEC to check the design limit loads applicable to the optical bench	05.11.1999	closed	SCI-PT/IFI/07222, 05.11.1999
06	DSS	DSS to define thermal design and interface data required from the instruments to allow update of the thermal mathematical model of the optical bench	26.10.1999	closed	Email, DSS, 27.10.1999
07	HIFI/ PACS/ SPIRE	Instruments to provide design/interface data required by DSS to update the thermal mathematical model of the optical bench (see AI 06 of PM#1)	05.11.1999	closed	Email (HIFI), 15.11.1999; PACS-ME-TN-007, 10.11.1999; Email (SPIRE), 08.11.1999
08	ESTEC	ESTEC to reconsider the necessity of the presently defined very stringent instrument fixation accuracies (see Annex 6 to MoM)	PM#2 23.11.1999	open	
09	DSS	DSS to sketch necessary harness interface connector arrangement when applying ISO harness technology for FIRST PLM	PM#2 23.11.1999	open	



**Action Items of Progress Meeting # 1 (20.10.1999), cont'd**

Action Item #	Actionee	Subject / Description	Due Date	Status	Reference
10	DSS	DSS to sketch necessary harness interface connector arrangement required for BAU when applying ISO harness technology for FIRST PLM	PM#2 23.11.1999	open	
11	HIFI	HIFI to define the position of the outlet of the HIFI instrument harness	05.11.1999	open	
12	ESTEC	ESTEC will check what is the sine vibration specification applicable for the instruments on the PLM optical bench and whether the levels can be reduced at higher frequency	05.11.1999	closed	SCI-PT/IFI/07222, 05.11.1999
13	HIFI/ PACS/ SPIRE	Instruments to comment/confirm definition of temperature level-0, level-1 and level-2 applicable to the instruments as defined in PM#1-MoM, page 9/9	PM#2 23.11.1999	open	

# **FIRST**

Update of Instrument Interfaces



**DaimlerChrysler Aerospace**  
Dornier

Dornier Satellitensysteme GmbH

---

## **3. FIRST Optical Bench Design**

- Arrangement of FPU's on OB
- Mechanical Interfaces
- Thermal Interfaces (level 0, 1, 2)
- OB connectors, Harness routing
- LO baffle between HIFI and Common Instrument Shield

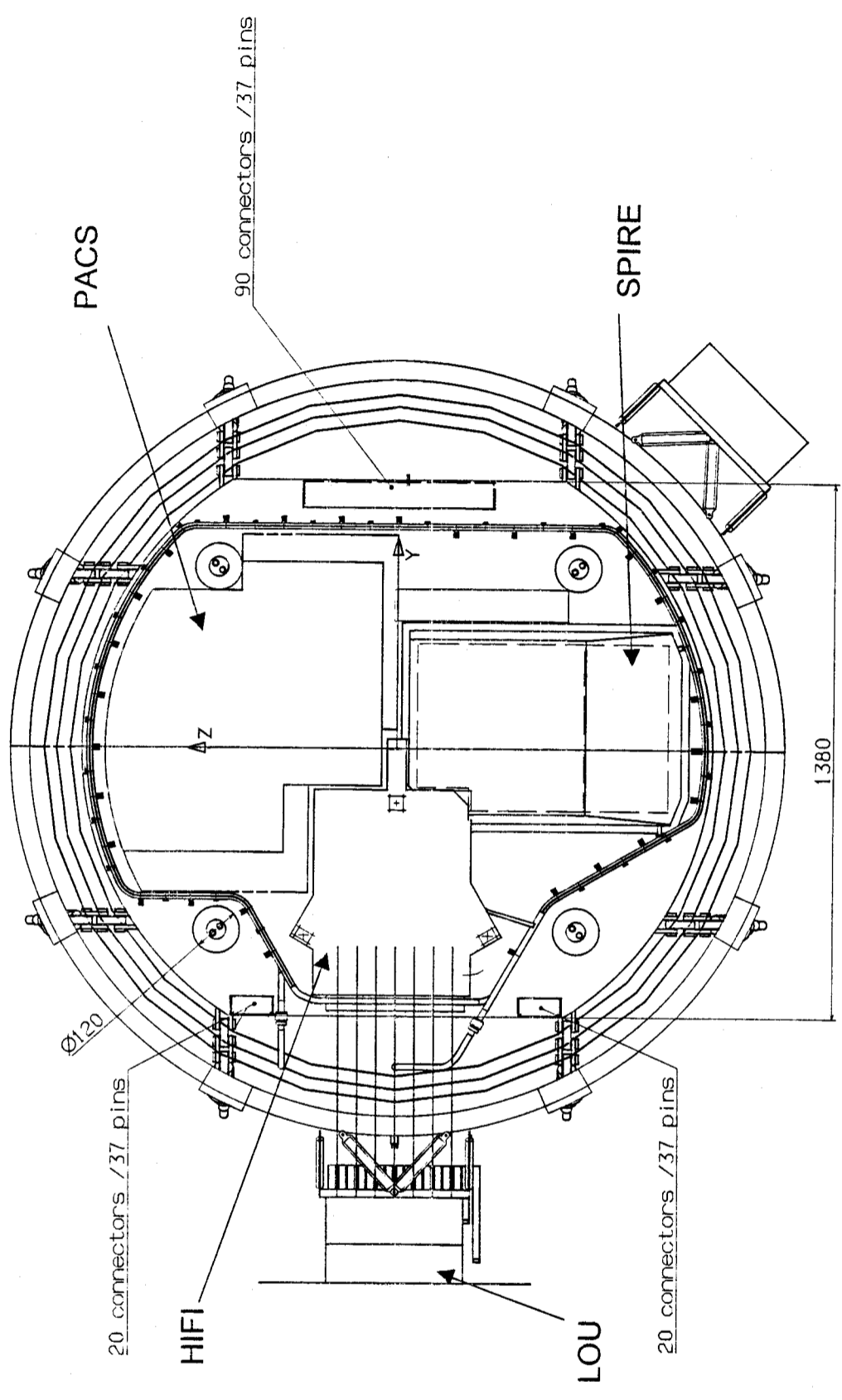




# FIRST

Update of Instrument Interfaces

## First Optical Bench Design – Arrangement of FPU's



# FIRST

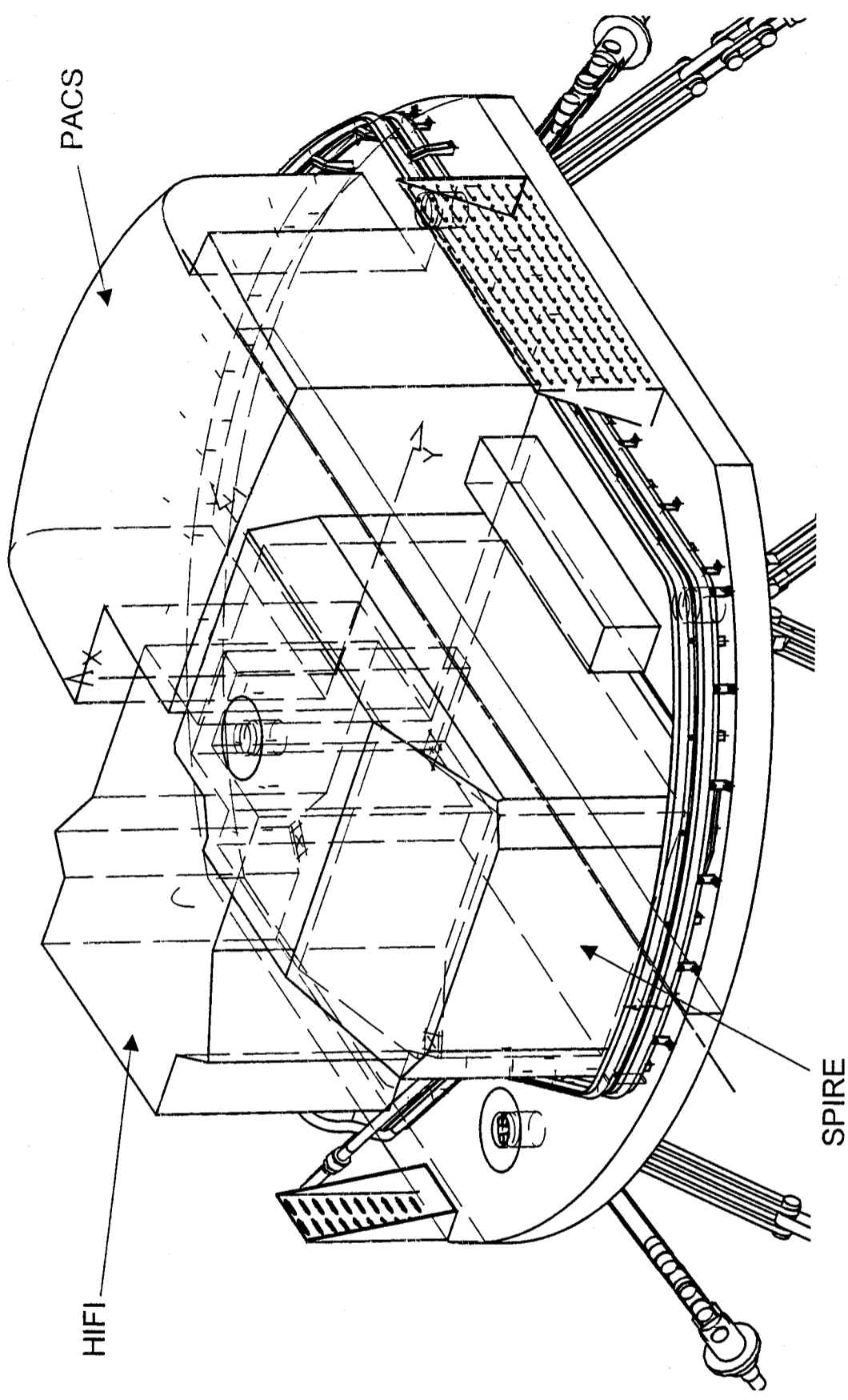
Update of Instrument Interfaces

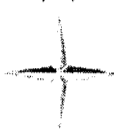


DaimlerChrysler Aerospace

Space Infrastructure

## First Optical Bench Design – Arrangement of FPU's - 3D



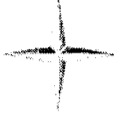


## First Optical Bench Design – CATIA Layout Status

- accessibility to OB fixation points taken into account (dia. 120mm)
- OB enlarged in  $\pm Y$  direction, thickness kept at 60mm
- heat exchanger below OB deleted
- thermal interface level 1 routed above level 2 on top of the OB
- connector brackets added on  $\pm Y$  side
- experiment shield adapted
- LOU baffle implemented
- experiment volumes no change since last meeting
- SPIRE JFET Box position open

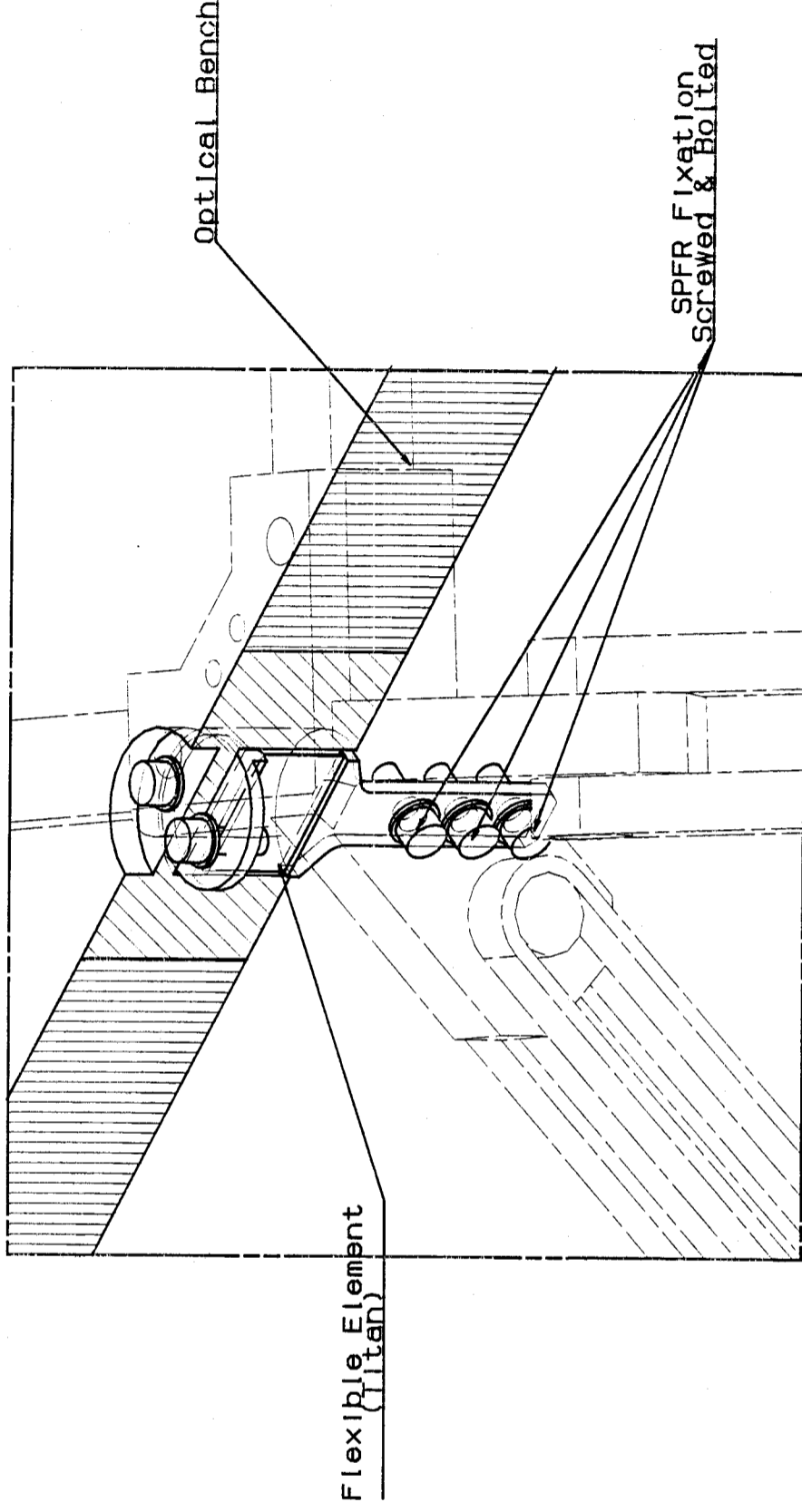
### PACS checked with larger volume:

- enlargement of volume critical in terms of clearance to SPIRE filter box and mounting area between experiments (to be checked in detail)
- no additional activities planned



## First Optical Bench Design – Mech. IF to Cryostat

- Previous Design kept – update not included in present study



# **FIRST**

Update of Instrument Interfaces



DaimlerChrysler Aerospace

Space Infrastructure

---

## **First Optical Bench Design – Mech. IF to Instruments**

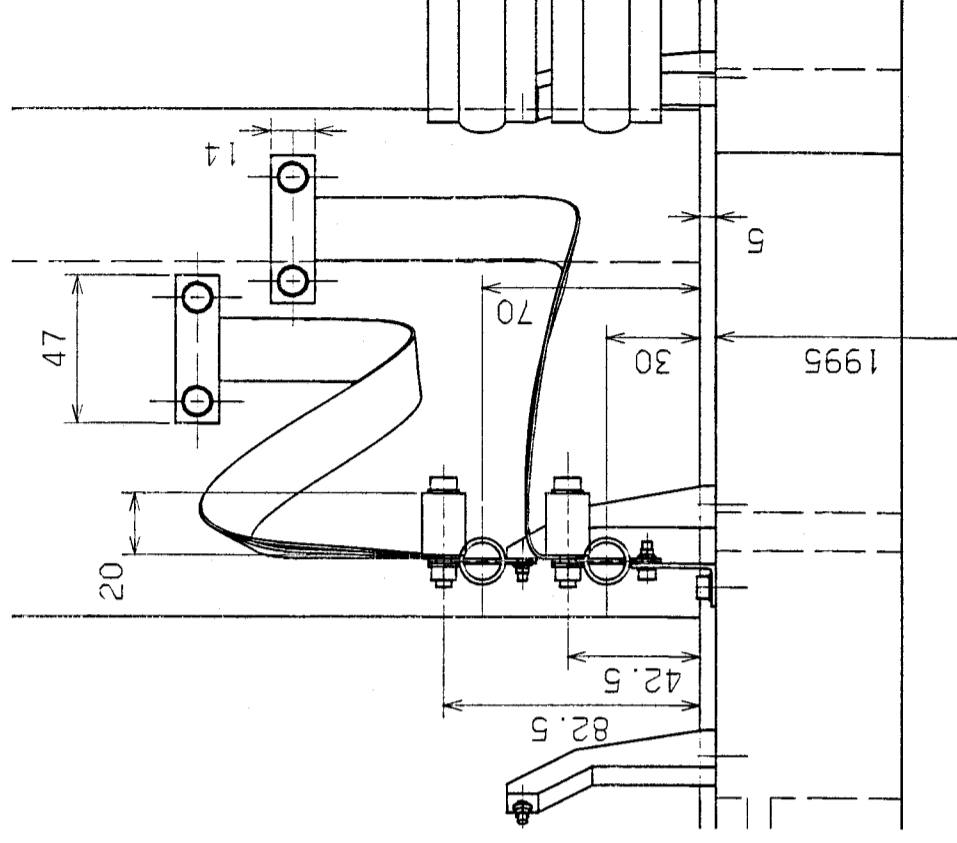
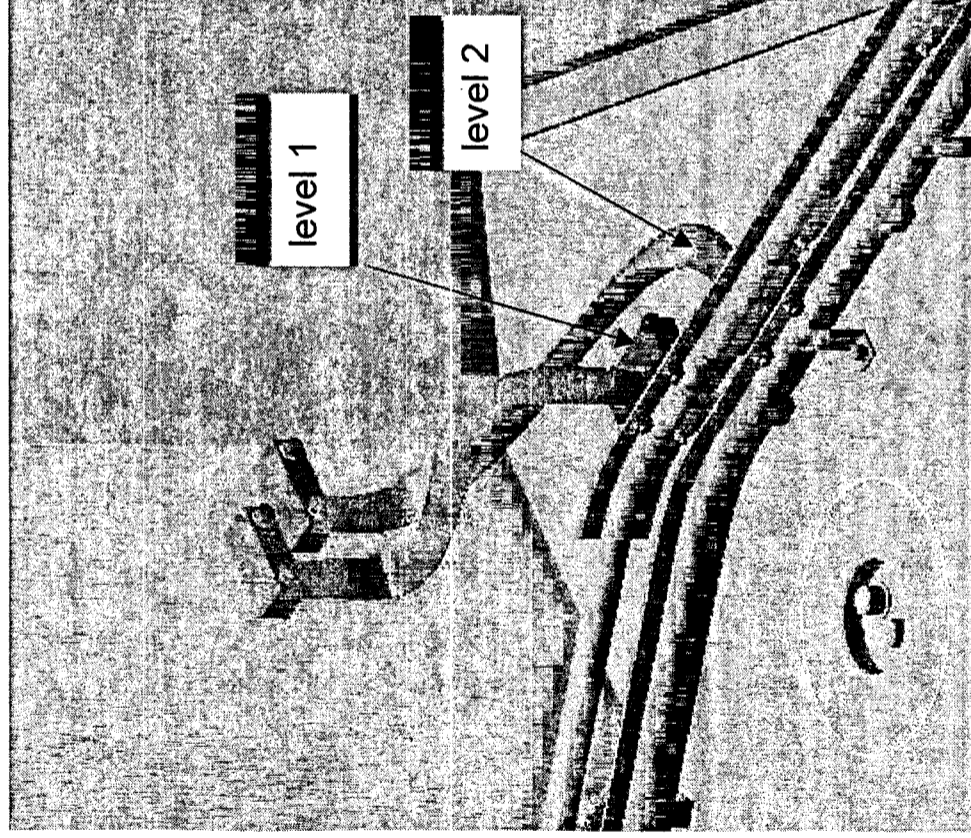
### **Summary of last Meeting (see MOM 20.10.99)**

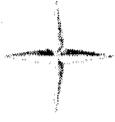
- number of fixations point different for each experiment  
PACS 3, HIFI 3, SPIRE 8
- reference plane STA 2000, one reference hole for each experiment
- fixation holes: Dia. 12H7 location tbd. and 4 screws M6 size around each hole
- all holes on OB are fixed - no provisions for compensation of thermal displacements
- provisions for thermal alignment and thermal displacements shall be done on experiment side

**FIRST**

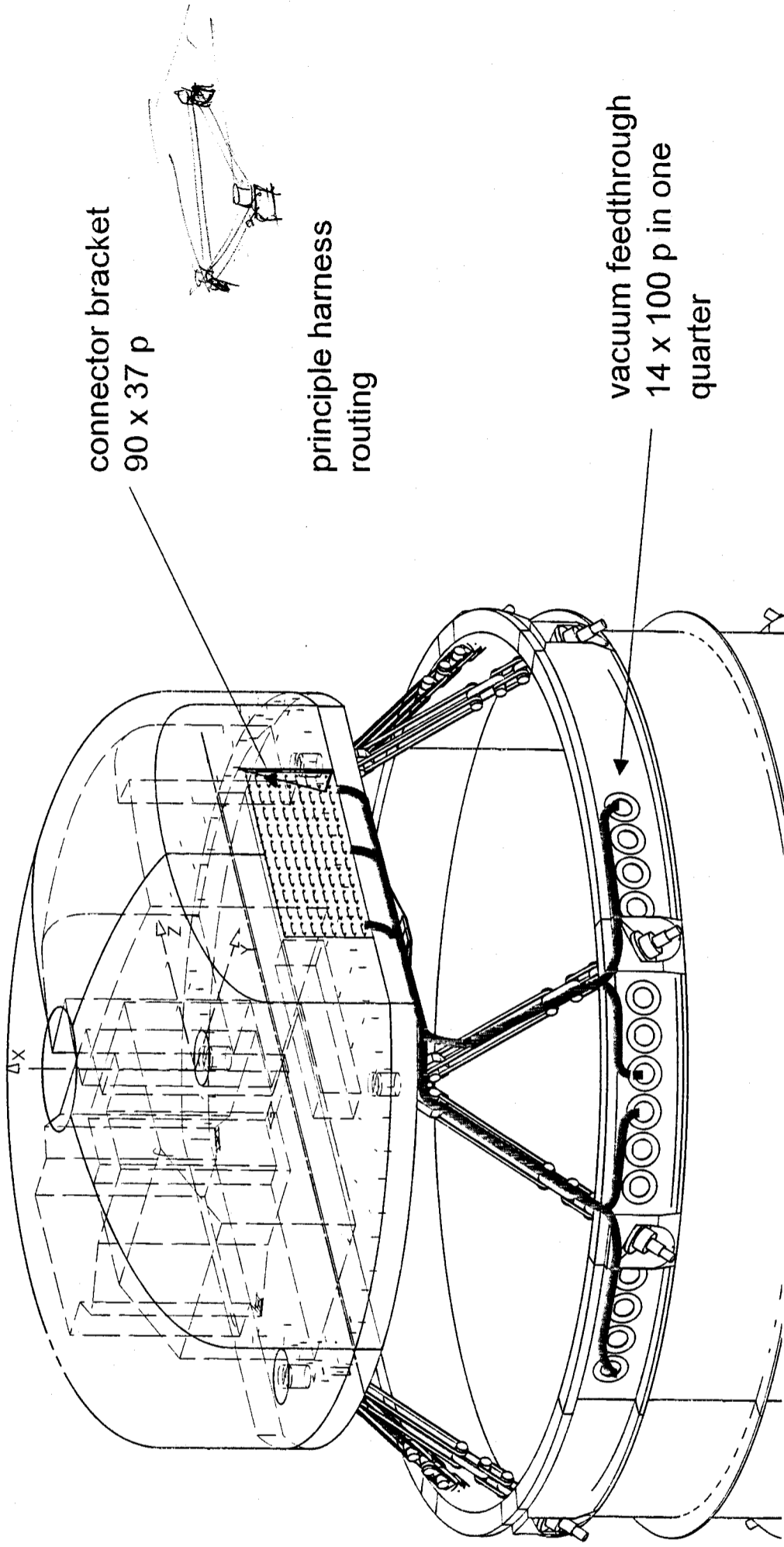
Update of Instrument Interfaces

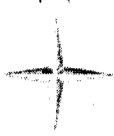
- First Optical Bench Design – Thermal Interfaces (level 0, 1, 2)





# First Optical Bench Design – Harness Routing, Connector Interfaces

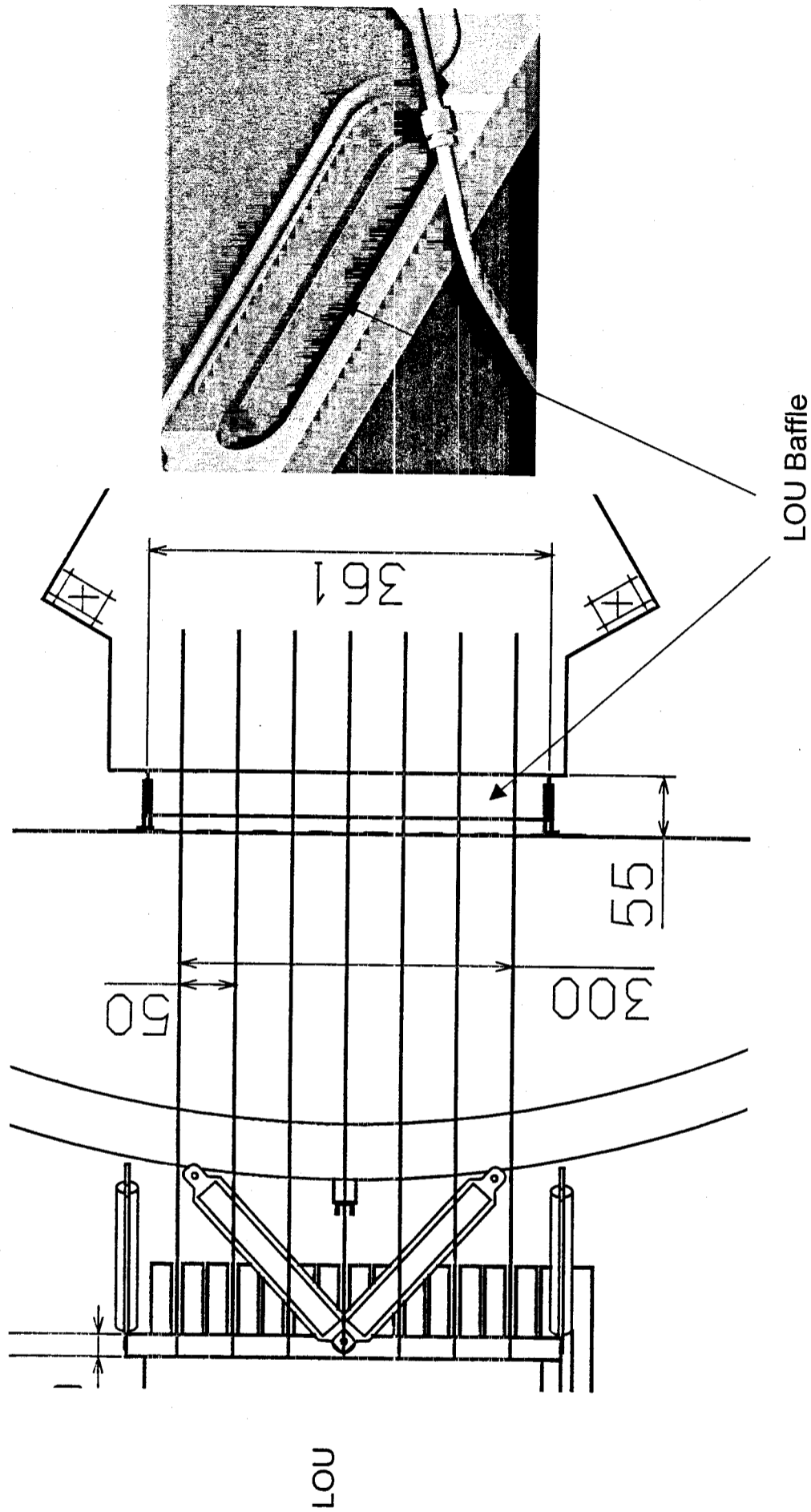




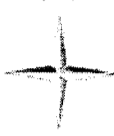
**FIRST**

Update of Instrument Interfaces

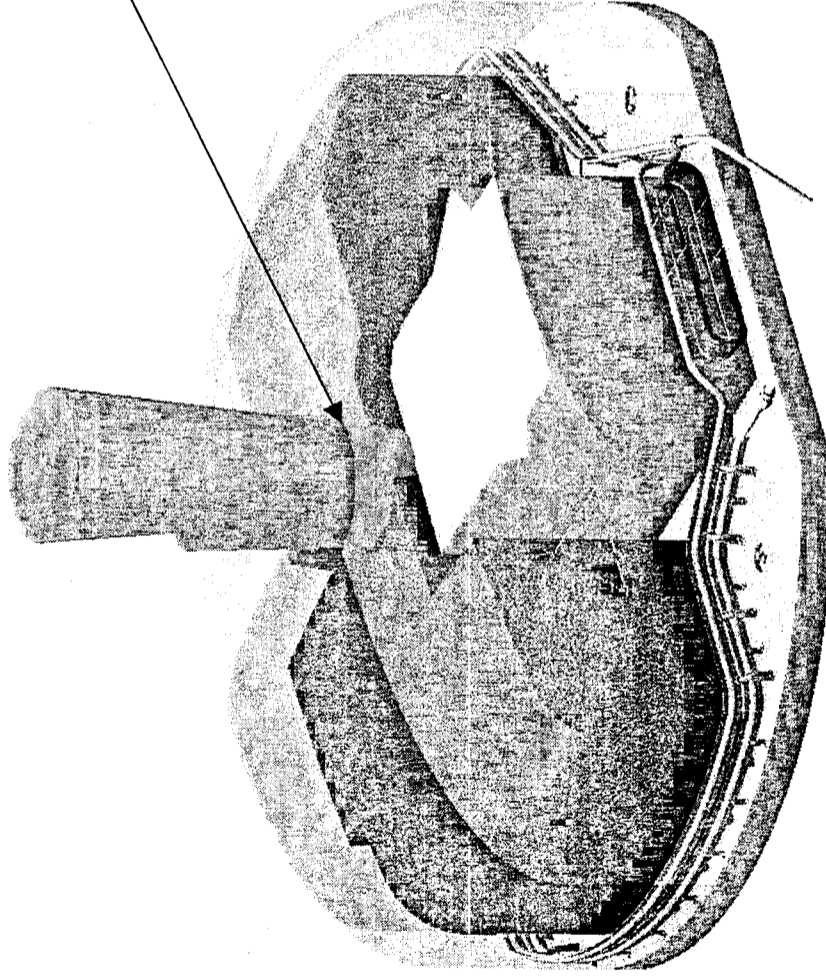
**First Optical Bench Design – LOU Baffle**



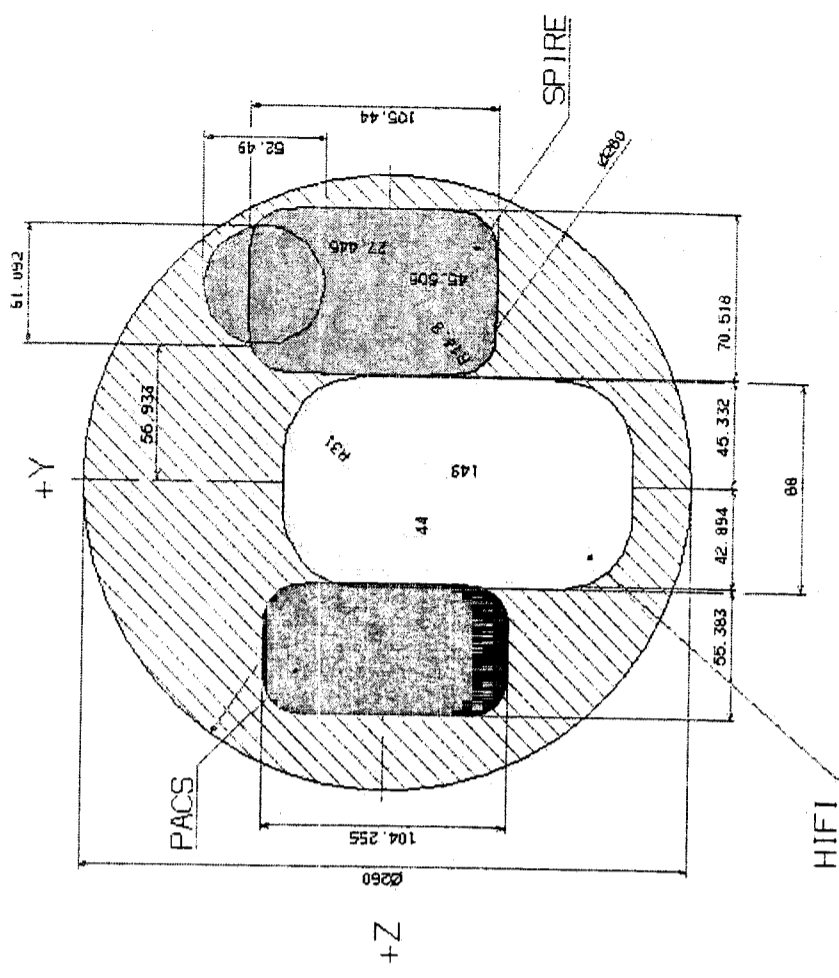




### 4. Optical Interfaces (Beam Pattern/Baffle Shaping)



SECTION X=304 above Focus



# **FIRST**

Update of Instrument Interfaces



DaimlerChrysler Aerospace  
Dornier

Dornier Satellitensysteme GmbH

---

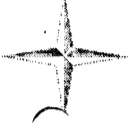
## 5. Optical Bench Analysis

### Mechanical Analysis

A structural model of the OB has been established; different configurations of the OB have been analysed

### Thermal Analysis

With the inputs from the instruments a thermal model has been build up, first calculations performed;



**FIRST**  
**Optical Bench**  
**Mechanical Analyses**

**Bernd Gröger RIA54**



Version	design	total mass	1. eigenfreq.
9	<p>100 mm alu sandwich plate</p> <ul style="list-style-type: none"> <li>- 96 mm alu core <math>G=283 \text{ N/mm}^2</math></li> <li>e.g. HEXCEL 1/8-5052-.0015-6.1</li> <li>- 2mm alu face sheets</li> </ul> <p>1mm steel blades (46x29.5mm)</p> <p>alu inserts</p> <p>instruments 156.1 kg</p> <p>sandwich plate + inserts 65.4 kg</p>	247.5 kg	81 Hz
10	<p>60 mm cfc sandwich plate</p> <ul style="list-style-type: none"> <li>- 58 mm cfc core <math>G=331 \text{ N/mm}^2</math></li> <li>e.g. HEXCEL HFT-G-327-3/16-6.0</li> <li>- 1mm cfc face sheets quasi isotropic</li> <li>M60 fibers <math>E=120000 \text{ N/mm}^2</math></li> </ul> <p>1 mm steel blades (46x29.5mm)</p> <p>cfc inserts</p> <p>instruments 156.1 kg</p> <p>sandwich plate + inserts 25.26 kg</p>	206.7 kg	57 Hz



Version	design	total mass	1. eigenfreq.
11	60 mm cfc sandwich plate - 56 mm cfc core $G = 331 \text{ N/mm}^2$ e.g. HFT-G-327-3/16-6.0 - 2mm cfc face sheets quasi isotropic M60 fibers $E = 120000 \text{ N/mm}^2$  1 mm steel blades  cfc inserts  instruments 156.1 kg  sandwich plate + inserts 30.4 kg	211.9 kg	71 Hz
12	70 mm cfc sandwich plate - 66 mm cfc core $G = 331 \text{ N/mm}^2$ e.g. HEXCEL HFT-G-327-3/16-6.0 - 2 mm cfc face sheets quasi isotropic M60 fibers $E = 120000 \text{ N/mm}^2$  cfc inserts  1mm steel blades  instruments 156.1 kg  sandwich plate + inserts 33.8 kg	215.3 kg	78.2 Hz

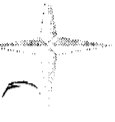


Version	design	total mass	1. eigenfreq.
13	70 mm cfc sandwich plate - 65 mm cfc core $G = 331 \text{ N/mm}^2$ e.g. HEXCEL HFT-G-327-3/16-6.0 2.5 mm cfc face sheets quasi isotropic M60 fibers $E = 120000 \text{ N/mm}^2$  1 mm steel blades cfc inserts instruments 156.1 kg sandwich plate + inserts 36.2 kg	217.7 kg	81.0 Hz
14	70 mm cfc sandwich plate - 65 mm cfc core $G = 331 \text{ N/mm}^2$ e.g. HEXCEL HFT-G-327-3/16-6.0 - 2.5 mm cfc face sheets quasi isotropic M60 fibers $E = 120000 \text{ N/mm}^2$  1 mm cfc blades ( $E = 120000 \text{ N/mm}^2$ )  cfc inserts  instruments 156.1 kg  sandwich plate + inserts 36.2 kg	217.4 kg	79.3 Hz



**Dimensioning Loadcases  
Face Sheets**

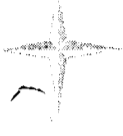
loadcase	min. princ. stress +Z fiber in [N/mm <sup>2</sup> ]	max. princ. stress -Z fiber in [N/mm <sup>2</sup> ]	min. princ. stress +Z fiber in [N/mm <sup>2</sup> ]	max. princ. stress -Z fiber in [N/mm <sup>2</sup> ]
10 g x + 5g y	-13.0	6.8	-8.8	9.5
10 g x + -5g y	-10.9	5.8	-5.1	10.4
10 g x + 5 g z	-10.5	7.7	-6.2	9.7
10 g x + -5g z	-10.6	8.2	-7.3	10.1



### Dimensioning Loadcases Face Sheets

loadcase	min. princ. stress +Z fiber in [N/mm <sup>2</sup> ]	max. princ. stress -Z fiber in [N/mm <sup>2</sup> ]	min. princ. stress +Z fiber in [N/mm <sup>2</sup> ]	max. princ. stress -Z fiber in [N/mm <sup>2</sup> ]
-10gx + 5gy	-5.8	10.9	-10.4	5.1
-10gx + -5gy	-6.8	13.9	-9.5	8.8
-10gx + 5gz	-7.7	10.5	-9.7	6.2
-10gx + -5gy	-8.2	10.6	-10.1	7.3





Allowables for quasi isotropic M60 laminate with alpha 0.02 e-6:

pressure:  $\text{Sig}_{\text{ult}} = -130 \text{ N/mm}^2$        $\text{MOS} > 7.6$

tensile:  $\text{Sig}_{\text{ult}} > 200 \text{ N/mm}^2$        $\text{MOS} > 11.5$

wrinkling:  $\text{Sig}_{\text{wrink}} = 480 \text{ N/mm}^2$        $\text{MOS} > 31$



### Quasi static dimensioning loadcases

maximum stress of blades (46x29.5x1mm,  $E=120000 \text{ N/mm}^2$ ):

- max. shear  $104 \text{ N/mm}^2$
- max. principle stresses  $360 \text{ N/mm}^2$
- min. principle stresses  $-337 \text{ N/mm}^2$
- max. von Mises  $353 \text{ N/mm}^2$

for Ti-blades:  $MOS_{ult} > 0.8$      $MOS_{shear} > 2.4$

for CFC-blades: ?

# FIRST

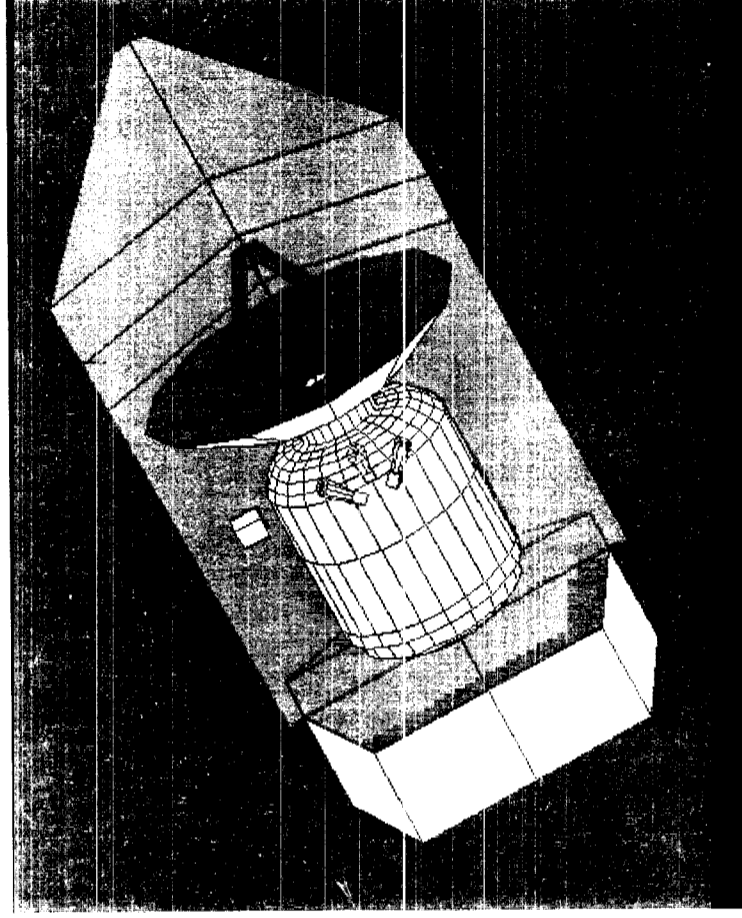
Update of Instrument Interfaces



DaimlerChrysler Aerospace  
Dornier

Dornier Satellitensysteme GmbH

## Thermal Analysis - Update of Instrument Interfaces



A. Hauser ST41

# **FIRST**

Update of Instrument Interfaces



DaimlerChrysler Aerospace  
Dornier

Dornier Satellitensysteme GmbH

---

## **Status of Updated FIRST Thermal Model**

- Instruments (PACS, SPIRE, HIFI) included
- Wheel shaped H/X replaced by GHe Level 1
- Instrument I/F to LHe, GHe Level 1 and GHe Level 2 introduced
- Instrument harness updated and disconnected from HS2 and HS3
- Cross section of tank suspension straps increased from 106 mm<sup>2</sup> to 130 mm<sup>2</sup> per strap (tbc)
- Optical Bench support changed from Ti blades to CFRP blades

# FIRST

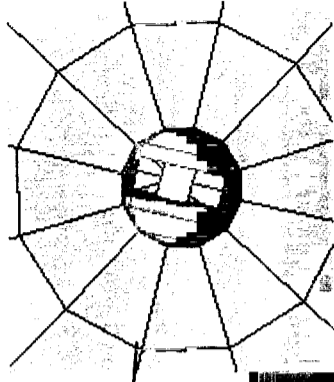
Update of Instrument Interfaces



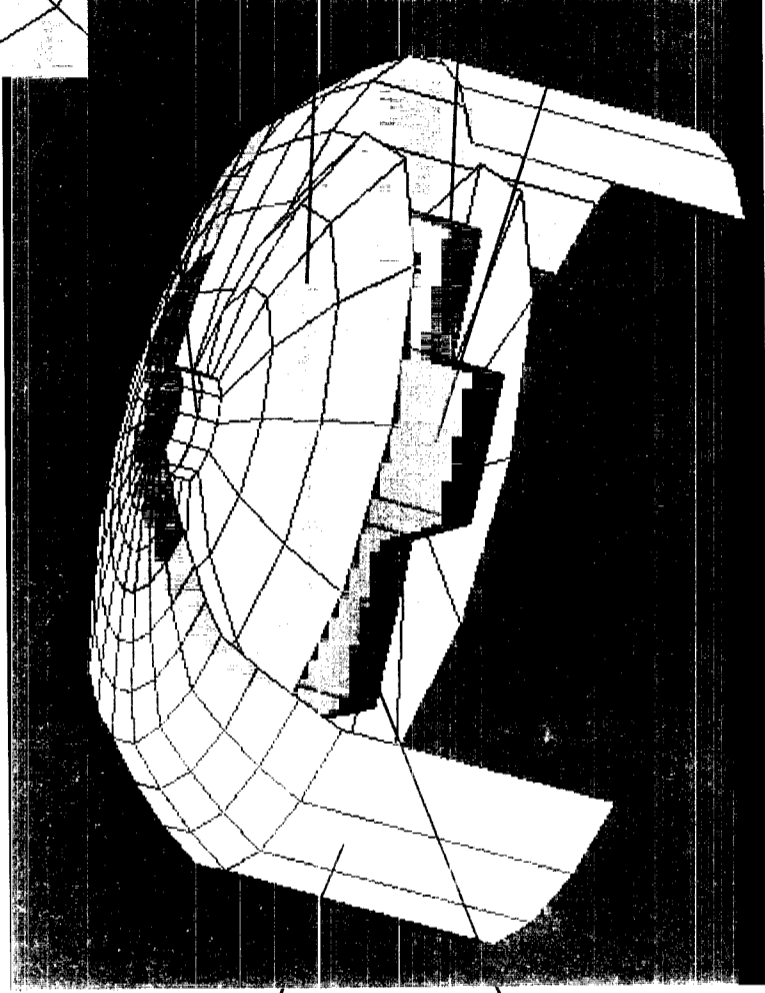
DaimlerChrysler Aerospace  
Dornier

Dornier Satellitensysteme GmbH

## Geometry Thermal Model of Optical Bench



Top View



Heat Shield Baffle

Instrument Shield

SPIRE

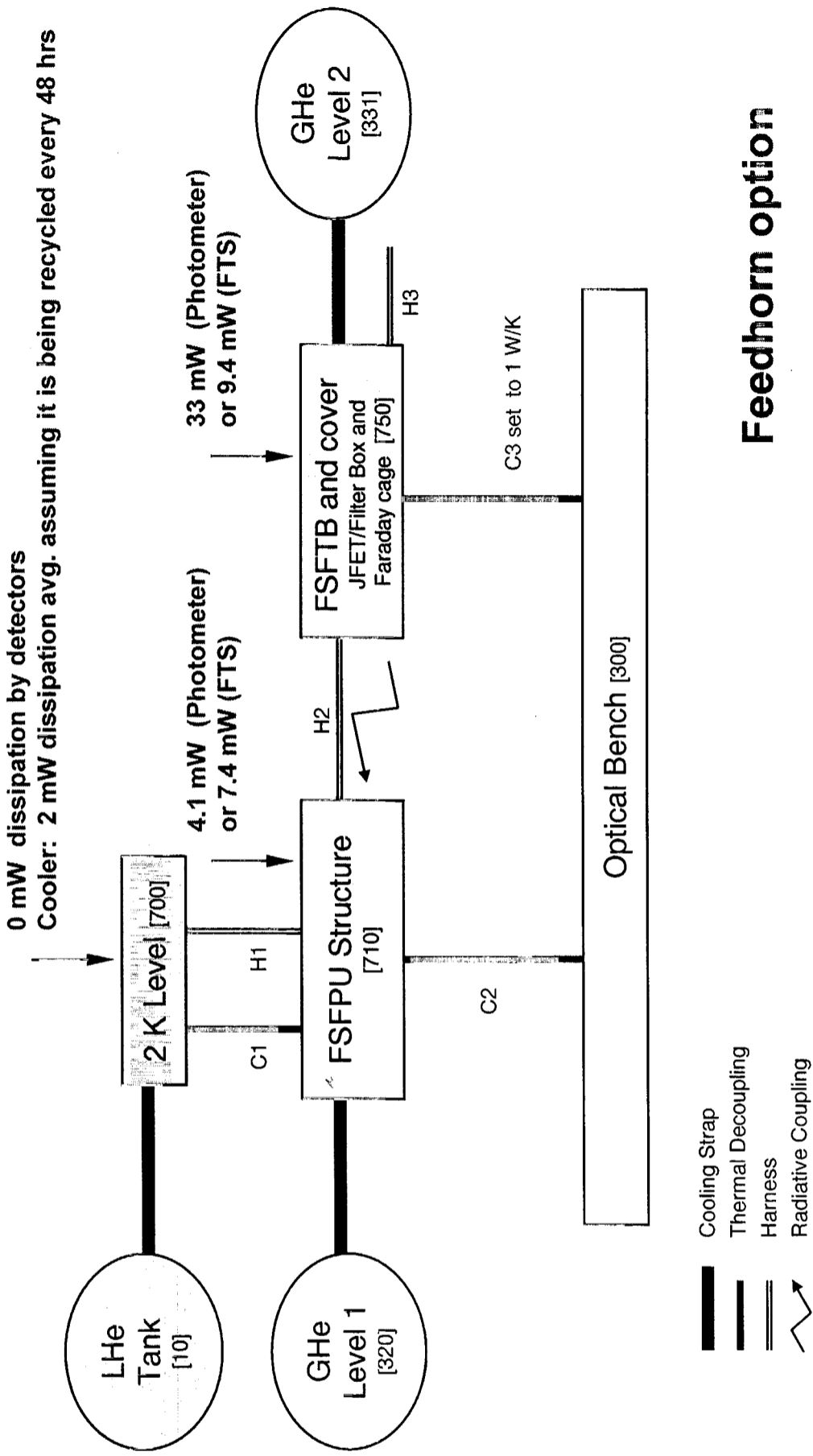
HIFI

Cryostat Vacuum Vessel

PACS



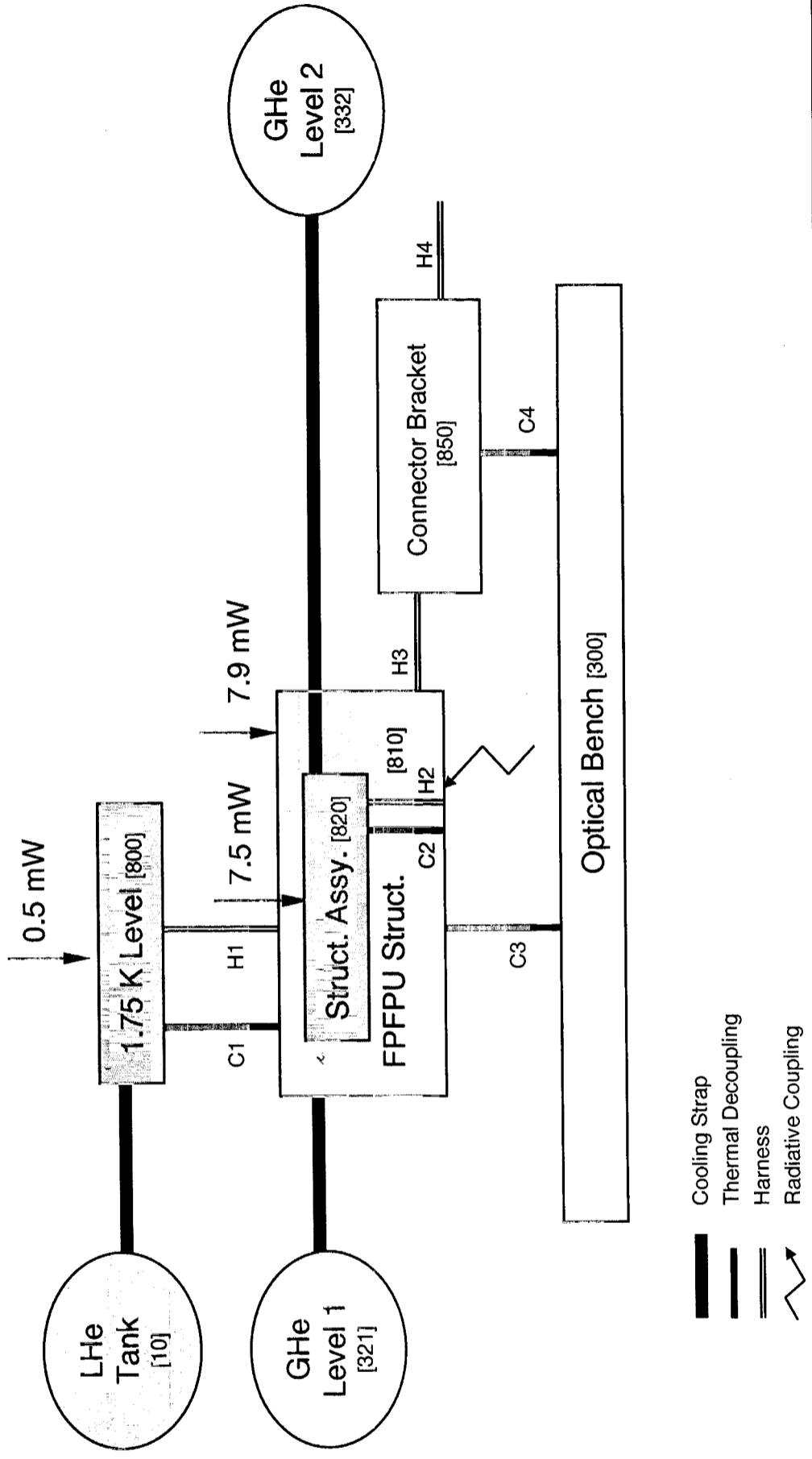
**SPIRE - FPU Thermal Sub Model for FIRST TMM**





# FIRST

## PACS - FPU Thermal Sub Model for FIRST TMM

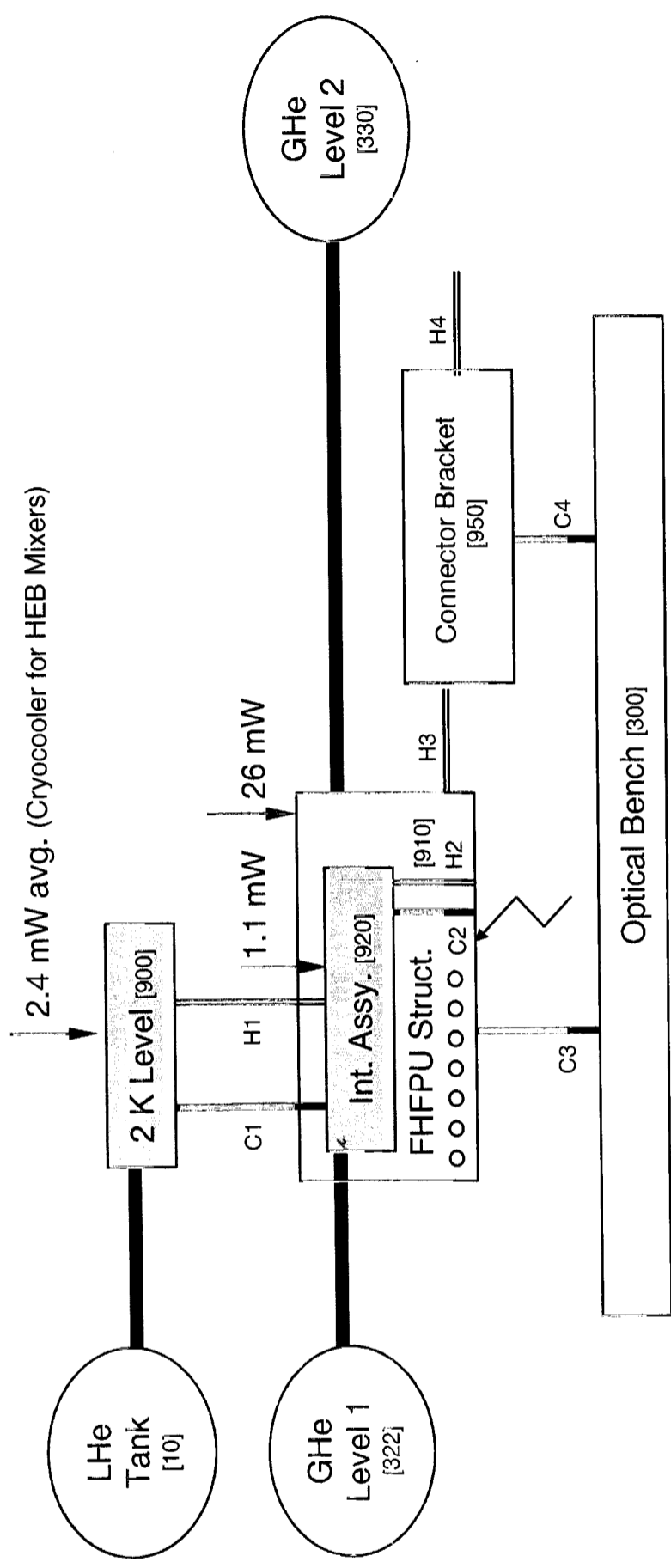




# FIRST

Update of Instrument Interfaces

## HIFI - FPU Thermal Sub Model for FIRST TMM



- Cooling Strap
- Thermal Decoupling
- Harness
- Radiative Coupling





### Input Data for FIRST TMM Update (1)

Item	Node	Material / Component	Mass	Size	Dissipation		Emiss.	Temp.
					On	Off		
LHe Tank	10							
Spatial Framework	200-213	CFRP struts						
Optical Bench	300	CFRP honeycomb panel (tbc) 4 CFRP blades (tbc)	40 kg	D=1.63m, H=70mm, 2.5mm skins cross sect.: 4x46mm <sup>2</sup> , l=30mm thickness: 0.8 mm			1.0	
Instrument Shield	310	Al cover					0.05	
GHe Ventline Level 1	320-322	Al tube CFRP brackets	0.5 kg	D=14 mm, d=12 mm cross sect.: 52mm <sup>2</sup> , l= 57.5 mm				
GHe Ventline Level 2	330-332	Al tube CFRP brackets	0.5 kg	D=14 mm, d=12 mm cross sect.: 40mm <sup>2</sup> , l= 17.5 mm				
Heat Shield Baffles	1050 2050 3050			260 mm diameter			0.78	
SPIRE 2K Level	700	thermally equivalent to Al 3 Stainless steel blades (mech. supp.) Harness to FSFPU Cooling strap to LHe tank <sup>a</sup>	7.0 kg	cross sect.: 60 mm <sup>2</sup> , l=0.03m SPIRE H1, see next table cross sect.: (20x3)mm, l=0.4m	2 mW negligible	2 mW 0		
SPIRE FSFPU	710	Al casing/structure 4 CFRP blades (mech. supp.) Harness to FSFTB Cooling strap to GHe ventline Level-1 <sup>a</sup>	23.5 kg	simplified box: 690x410x410mm cross sect.: 400 mm <sup>2</sup> , l=0.04m SPIRE H2, see next table cross sect.: (20x1)mm, l=0.3m	4.1/7.4mW <sup>b</sup> negligible	<sup>b</sup> 0	0.2	
SPIRE FSFTB & Cov.	750	JFET/ Filter Box (Al) Faraday Cage (Al) Thermal Coupl. to OB (tbd) Harness to shield 1 Cooling strap to GHe ventline Level-2 <sup>a</sup>	2.5 kg 1.0 kg	simplified box: 300x100x100mm	33/9.4 mW <sup>b</sup>	<sup>b</sup>	0.2	
PACS 2K level	800	thermally equivalent to Al 6 Kevlar struts (mech. supp.) 50 Kapton supports Harness to FPFPU Cooling strap to LHe tank	6 kg	cross sect.: 240mm <sup>2</sup> , l=60mm cross sect.: 96mm <sup>2</sup> , l= 5mm PACS H1, see next table cross sect.: (20x1)mm, l=0.4m	0.5 mW			



### Input Data for FIRST TMM Update (2)

PACS FPFPU	810	thermally equivalent to Al 6 CFRP struts to OB Harness to PACS connect. bracket Cooling strap to GHe ventline Level-1 <sup>a</sup>	50 kg	simplified box: 690x410x410mm Apertures with filters: 0.002 m <sup>2</sup> cross sect.: 2460mm <sup>2</sup> , l=420 mm PACS H3, see next table cross sect.: (20x1)mm, l=0.3 m	7.9 mW	0.05 0.1
PACS Struct. Assy.	820	thermally equivalent to Cu 6 Kevlar struts (mech.supp. to FPFPU) Harness to FPFPU	1 kg	cross sect.: 22mm <sup>2</sup> , l= 20mm PACS H2, see next table	7.5 mW 5mW	0
PACS connector bracket	850	Cooling strap to GHe ventline Level-2 <sup>a</sup> harness to shield 1	1.5 kg	cross sect.: (20x1)mm, l=0.3m PACS H4, see next table	8.81 mW	0
HIFI 2K level	900	thermally equivalent to Cu Vespe support & harness Cooling strap to LHe tank <sup>a</sup>	0.26 kg	1.1 mW/K cross sect.: (20x3)mm, l=0.4m	2.4 mW	2.4mW
HIFI FHFP	910	Al casing/structure Harness to HIFI connector bracket Cooling strap to GHe ventline Level-2 <sup>a</sup>	50 kg	box: 400x310x372mm* 0.009 m <sup>2</sup> Apertures (LO port) 0.005 m <sup>2</sup> Aperture (M3 at +Y) HIFI H3, see next table cross sect.: (20x1)mm, l=0.3m	26 mW	0.05 1.0 1.0
HIFI Internal Assy	920	Al structure Vespe support & harness Cooling strap to GHe ventline Level-1 <sup>a</sup>	1 kg	0.55 mW/K cross sect.: (20x1)mm, l=0.3m	1.1 mW	
HIFI connector bracket	950	harness to shield 1		HIFI H4, see next table PTFE, l=1.2m	11.4 mW	0

a) Cu 99.99 as baseline material, Al99.999 as option

b) Photometer / FTS for Feed Horn Option

Note: Instrument internal harness w/o margin. All other harness with 20% margin (cross section and dissipation)

# FIRST

Update of Instrument Interfaces



DaimlerChrysler Aerospace  
Dornier

Dornier Satellitensysteme GmbH

## Summary of Harness Data for FIRST TMM Update

Harness	Length m	Dissipation mW	Stainl. St. mm <sup>2</sup>	Brass mm <sup>2</sup>	CuBe mm <sup>2</sup>	PTFE mm <sup>2</sup>	Kapton mm <sup>2</sup>	Remark
SPIRE H1	0.15	~0	6.72				106	
SPIRE H2	0.15	~0	6.72				106	
SPIRE H3	1.2	0.26	18.5	8.78		222		
PACS H1	0.3 0.005	8μW ~0	0.064 0.58			0.17		
PACS H2	0.5	0.5		0.4		2		
PACS H3	0.3	5.62	27.5	4.08		262		
PACS H4	1.2	8.81	27.5	4.08		262		
HIFI H1								1.1 mW/K @ 3.0 K including mechanical supports
HIFI H2								0.55 mW/K @ 9.5 K including mechanical supports
HIFI H3	0.3	2.85	15.2	0.54	0.34	58.7		L=0.3 m assumed
HIFI H4	1.2	11.4	15.2	0.54	0.34	58.7		
CVV-HS1 X2	0.25	1.41	53.7	13.4	0.98	542.3		

Note: 20% Margin (cross section and dissipation) included for X2 and all H3 and H4

# **FIRST**

Update of Instrument Interfaces



**DaimlerChrysler Aerospace**  
Dornier

Dornier Satellitensysteme GmbH

---

## **Preliminary Results and Next Steps**

- **Lower Optical Bench temperature using CFRP blade coupling to spatial framework**
- **Level 1 Temperature: 3.1 K - 4.1 K**
- **Level 2 Temperature: 6.2 K - 9.1 K**
- **Lifetime reduced to about 4.5 years**

### **Next Steps:**

- **Update of cross sections of tank suspension and Optical Bench fixation, if necessary**
- **Implementation of BAU and update of LOU**
- **Transient analysis:**  
**SPIRE on → SPIRE off, PACS on → PACS off, HIFI on → HIFI off, SPIRE on**



## 6. Instrument Harness evaluation status

Requirements for Harness OB-Bracket to CVV feedthroughs:

### SPIRE:

number of connections (Horn Option): 1116 +180 conductors, with 77 +39 shields  
cross-sections (GSFC): 18.5 mm<sup>2</sup> SST, 8.8 mm<sup>2</sup> brass, 222 mm<sup>2</sup>teflon (incl.marg.)  
ohmic dissipation 0.22 mW/m (incl. margin)

### PACS:

number of connections : 686 conductors with 34 shields  
cross-sections: 27.5 mm<sup>2</sup> SST, 4.1 mm<sup>2</sup> brass, 261 mm<sup>2</sup> teflon (incl. margin)  
ohmic dissipation 7.34 mW/m (incl. margin)

### HIFI:

number of connections : 554 conductors with 9 shields  
cross-sections: 15.2 mm<sup>2</sup> SST, 0.54 mm<sup>2</sup> brass, 0.34 mm<sup>2</sup> BeCu , 58.7 mm<sup>2</sup> teflon  
(incl. margin); ohmic dissipation: 9.5 mW/m (incl. margin)

# **FIRST**

## Update of Instrument Interfaces



DaimlerChrysler Aerospace  
Dornier

Dornier Satellitensysteme GmbH

---

- All instruments have connector brackets on OB (see chapter OB design)
- Routing via OB lower side, to SFWK and via the tank straps to the vacuum connectors
- Wires of instruments distributed to tank straps which are closest
- Required number of connectors on OB and CVV possible to accommodate
- Wires only thermally coupled to shield 1
- Spire-Harness is considered most critical w.r.t number of connections
- Harness length from SPIRE connector bracket to vacuum feedthroughs between 870 and 1400 mm
- About 1200 wires have to be routed via two tank suspension for SPIRE to BAU (46 bundles diameter 2.2 mm with 26 wires each)

### Open point:

- Vacuum feedthrough technology w.r.t ISO



## 7. Local Oscillator Unit

### Updated LO requirements (tbc):

Mass	13 kg + radiator mass (?)
Dimension (x, z, y)	440 x 357 x 220 mm
Dissipation	1 W multiplier, 6 & 12 W amplifier
Temperature (goal)	120 K multiplier, 200K amplifier
Temperature stability	?
Alignment w.r.t FPU	$\pm 0.8$ mm, $\pm 70$ mm, $\pm 0.8$ mm, $\pm 0.027^\circ$ , $\pm 0.38^\circ$ , $\pm 0.027^\circ$
Align. stability w.r.t FPU / 100s	$\pm 0.08$ mm, $\pm 0.002$ mm, $\pm 0.08$ mm, $\pm 0.003^\circ$ , $\pm 0.04^\circ$ , $\pm 0.003^\circ$ ,
Window diameter	Free inner diameter 30 mm, distance 50mm
Transmission	$\geq 90\%$ for each window and filter
Radiator area	2 x 0.25 m <sup>2</sup>
Waveguides to SVM	tbd

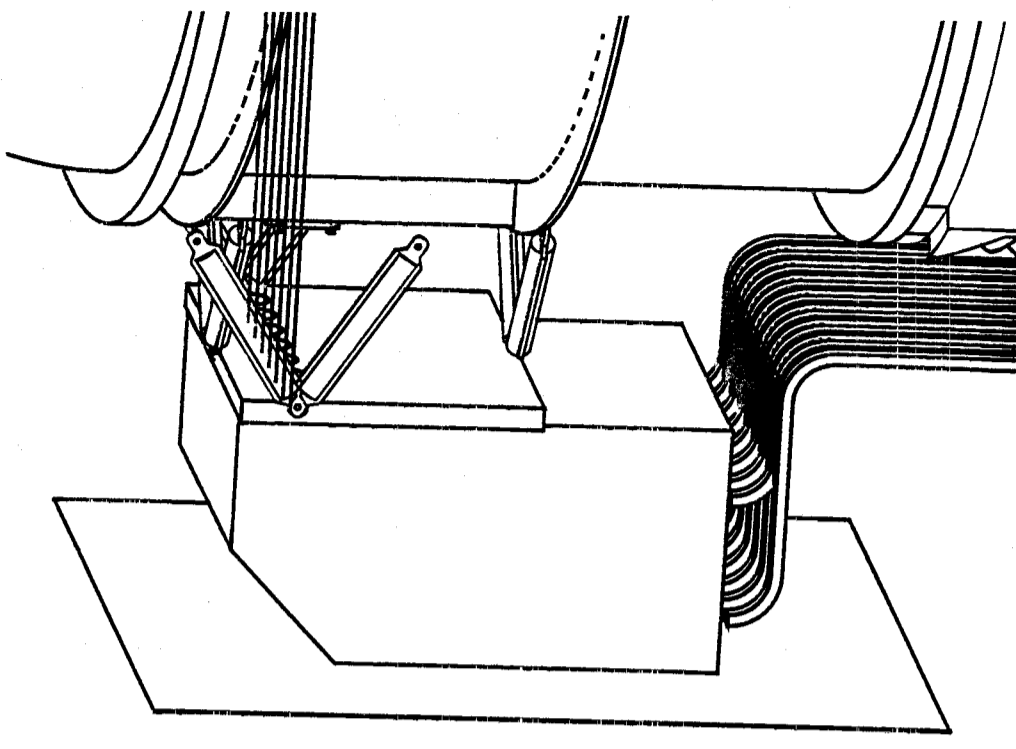
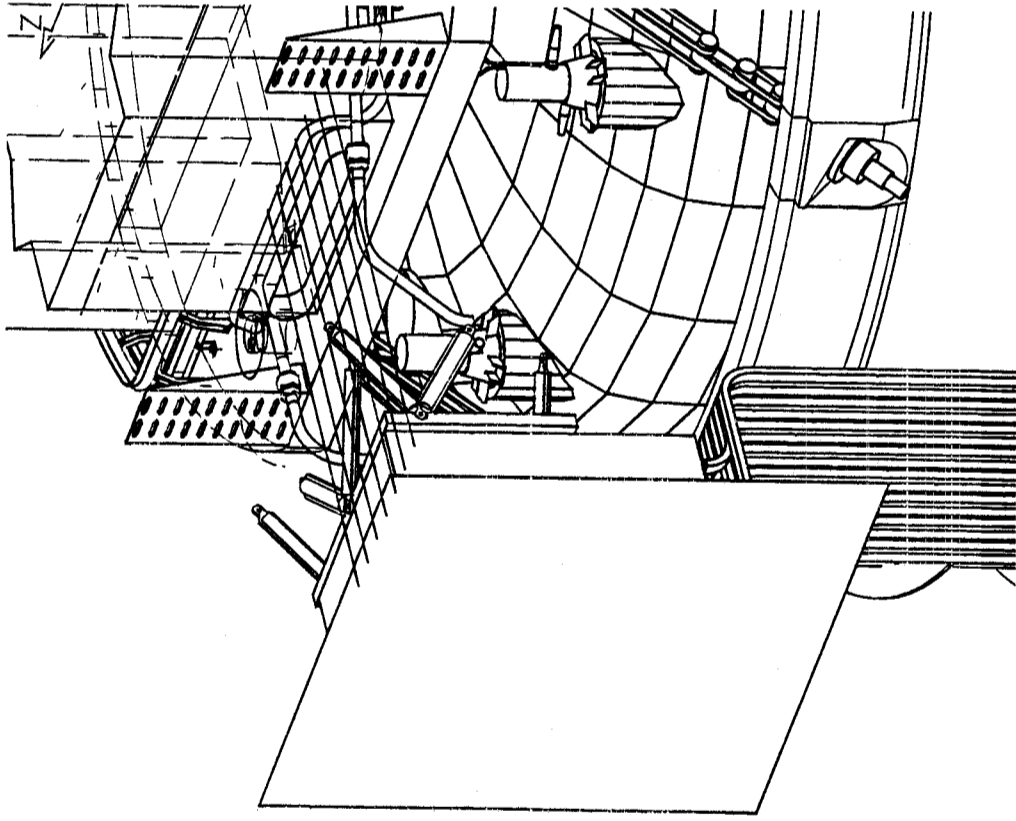
**FIRST**

Update of Instrument Interfaces

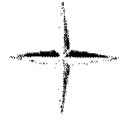
DaimlerChrysler Aerospace

Space Infrastructure

## Local Oscillator Unit – Installation Layout



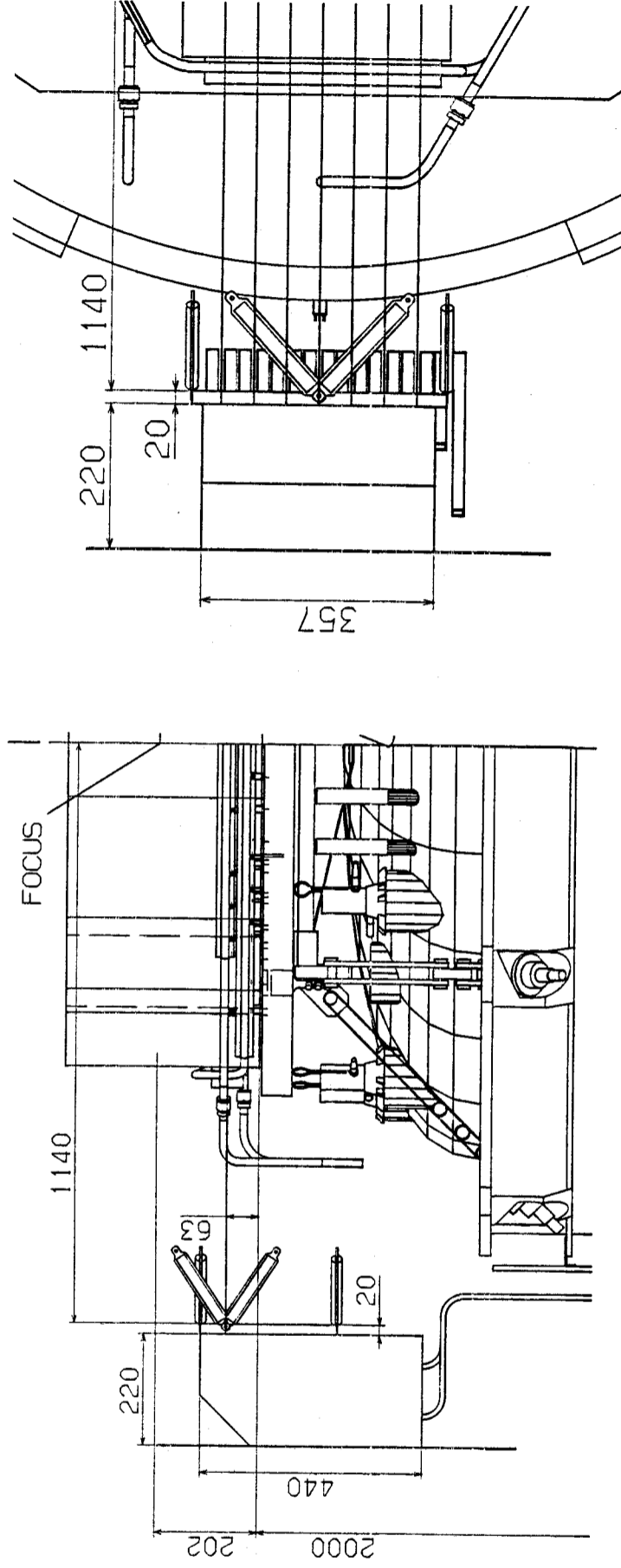




**FIRST**

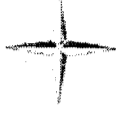
Update of Instrument Interfaces

**Local Oscillator Unit – Installation Layout**



**FIRST**

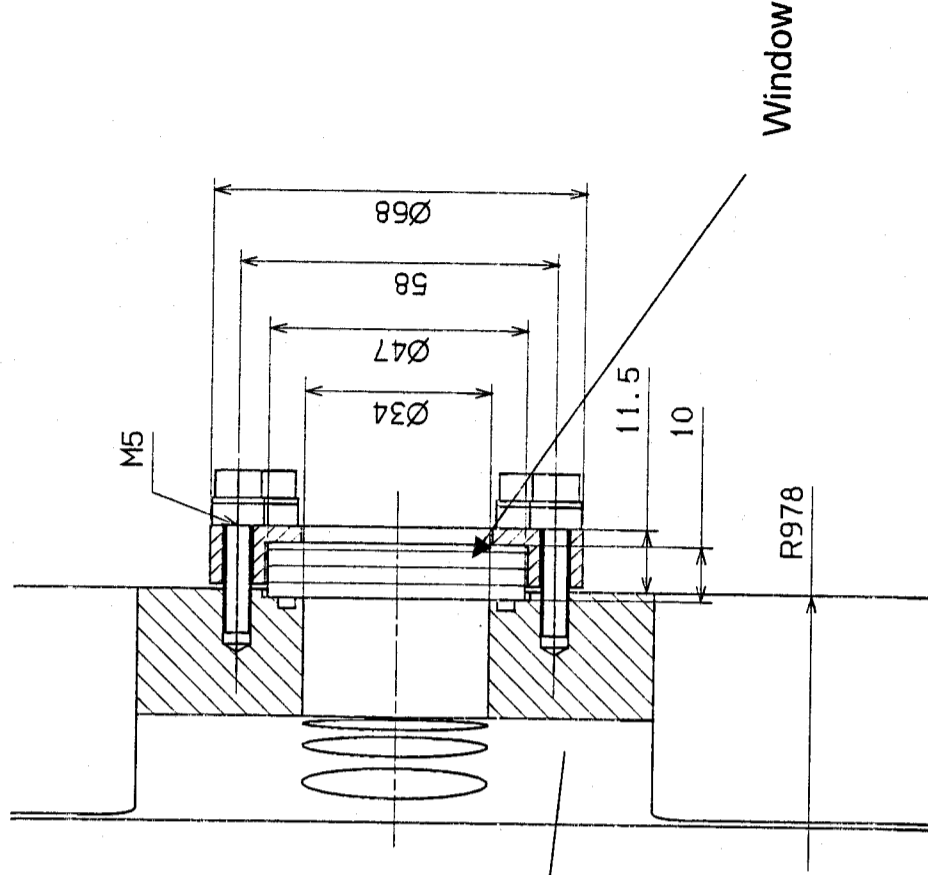
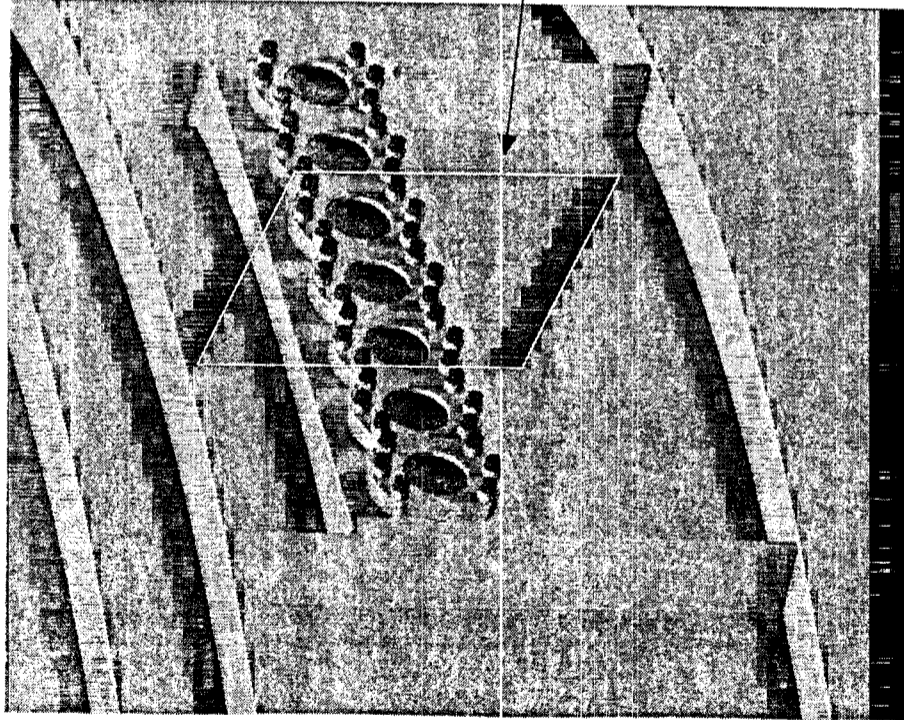
Update of Instrument Interfaces



DaimlerChrysler Aerospace

Space Infrastructure

## Local Oscillator Unit – Design Concept Optical Windows



# **FIRST**

Update of Instrument Interfaces



**DaimlerChrysler Aerospace**  
Dornier

Dornier Satellitensysteme GmbH

---

## **LOU Optical Channel**

### **Basic requirements**

- outer window has to withstand atmospheric pressure
- One heat filter per channel at cryostat shields
- transmission  $> 90\%$  per window and per heat filter



## LOU Optical Channel

### Possibility 1

- **one quartz window** in outer cryostat vessel (about 80K in orbit) and **another quartz window** at one inner shield (30K in orbit)
- use of poly-ethylene anti-reflection coatings on quartz
- inner quartz window with heat filter
- usable diameter = 30 mm
- windows are wedged and tilted by 2° (according to FIRST IID, PT-HHF1-02125)

### Possibility 2

- **one quartz window** in outer cryostat vessel (about 80K in orbit) plus **one mesh filter\*** on the inner side (30K in orbit)
- use of poly-ethylene anti-reflection coatings on quartz window
- mesh filter as a heat filter
- usable diameter = 30 mm
- quartz window and mesh filter are tilted by 2°, quartz window is wedged (according to FIRST IID, PT-HHF1-02125)

\* technique by Peter Ade, Queen Mary and Westfield College, London

# FIRST

Update of Instrument Interfaces



DaimlerChrysler Aerospace  
Dornier

Dornier Satellitensysteme GmbH

## LOU Optical Channel

### Estimation of window + heat filter transmissions (channel 7)

	Possibility 1	Possibility 2
Window in cryostat outer vessel	2 quartz windows 88%	quartz window + mesh filter 88%
Heat filter	88%	95%
<b>Total</b>	<b>77%</b>	<b>83%</b>

Remark: informations about mesh filters from Peter Ade

**LOU Optical Window****Transmission of window in outer cryostat vessel**

quartz window with polypropylene anti-reflection coating

Channel	lower frequency (GHz)	upper frequency (GHz)	required minimum transmission (tbc)	achievable minimum transmission (tbc)
1	480	636	90%	90%
2	640	792	90%	91%
3	792	948	90%	91%
4	960	1120	90%	91%
5	1104	1273	90%	91%
6a	1408	1584	90%	89%
6b	1704	1896	90%	89%
7	2400	2700	90%	88%

Temperature: 80 K

Remark: achievable minimum transmissions tbc by manufacturer!



## LOU Optical Window

### Transmission budget of window in outer cryostat vessel

quartz window with polypropylene anti-reflection coating

Channel	1	2	3	4	5	6	7
1 transmission loss of quartz window	2%	2%	2%	2%	2%	3%	5%
2 theoretical losses at edges of $\lambda/4$ coating	2%	1%	1%	1%	1%	2%	1%
3 -not ideal coating material -change of refractive index over band -change of refractive index (ambient versus 80 K) total							
4 manufacturing uncertainty	2%	2%	2%	2%	2%	2%	2%
5 degradation during life-time	2%	2%	2%	2%	2%	2%	2%
<b>total loss of transmission</b>	<b>10%</b>	<b>9%</b>	<b>9%</b>	<b>9%</b>	<b>9%</b>	<b>11%</b>	<b>12%</b>

Temperature: 80 K

**LOU Optical Window****Crystal-quartz - Transmission**

In literature the following measurement values for the transmission of 6 mm quartz at 2700 GHz

have been found:(quartz is a bire-fringent material):

temperature	T (ordinary direction)	T (extra-ordinary)
at 300 K	0.6	0.89
at 10 K	0.985	0.999

Measurements between different literature sources show differences.

No measurements for 80 K have been found yet.



# FIRST

Update of Instrument Interfaces

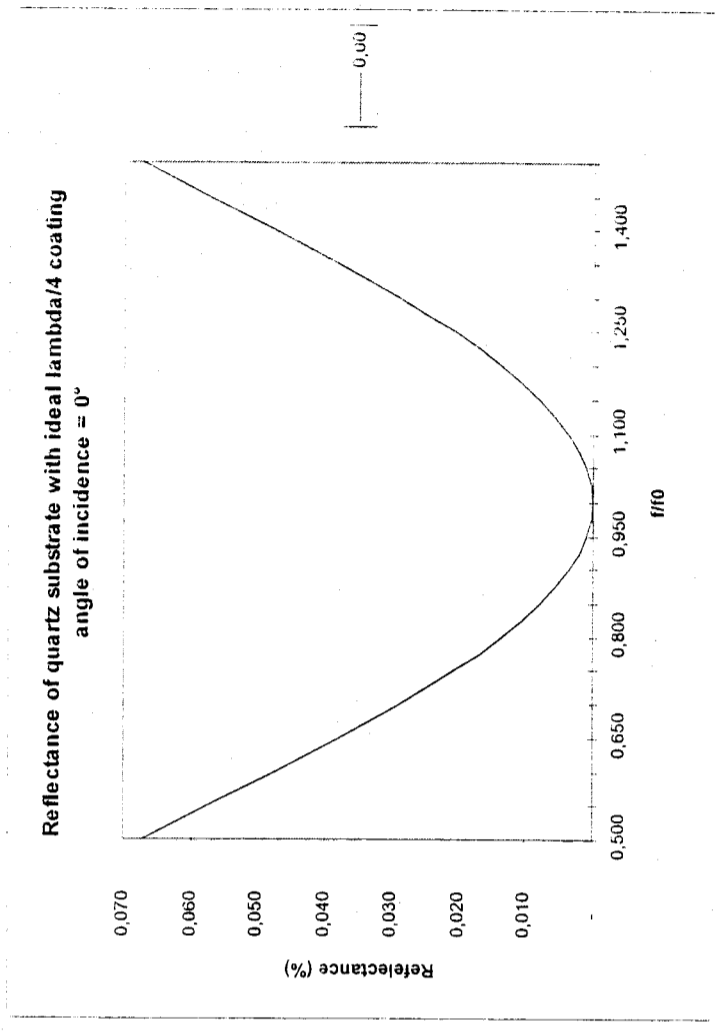


DaimlerChrysler Aerospace  
Dornier

Dornier Satellitensysteme GmbH

## LOU Optical Window

### Anti-reflection coatings



The reflectance of a single  $\lambda/4$  anti-reflection layer on a quartz surface ( $n = 2.1$ ) with ideal refractive index ( $n = 1.45$ ) as a function of frequency divided by center frequency. Candidate coating materials: polyethylene ( $n = 1.52$ ), polypropylene ( $n = 1.49$ ) or teflon ( $n = 1.44$ ).

# FIRST

Update of Instrument Interfaces



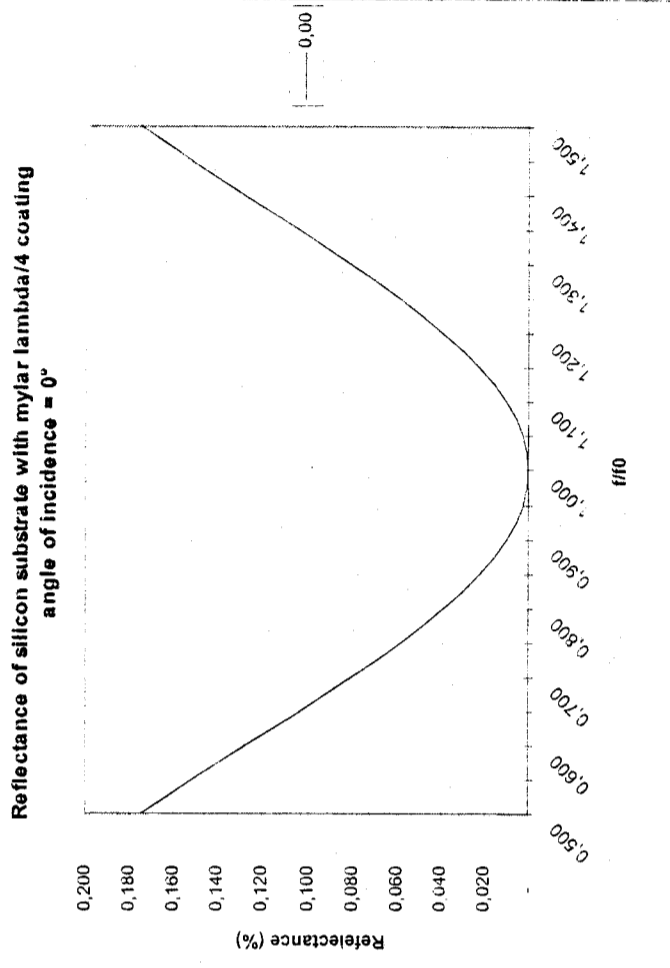
DaimlerChrysler Aerospace  
Dortmunder

Dornier Satellitensysteme GmbH

## LOU Optical Window

### Anti-reflection coatings

Example: Mylar on silicon  
substrate



The reflectance of a single  $\lambda/4$  anti-reflection layer of Mylar ( $n = 1.85$ ) on a silicon substrate ( $n = 3.4$ ) as a function of frequency divided by center frequency.

# FIRST

Update of Instrument Interfaces



DaimlerChrysler Aerospace  
Dortmund

Dornier Satellitensysteme GmbH

## LOU Optical Window

### Anti-reflection coatings

The following materials are candidates for AR-coatings:

- paraffine (n = 1.40)
- teflon (n = 1.44)
- TPX (n = 1.47)
- Polyethylene (n = 1.52)
- Polypropylene (n = 1.49)
- Polystyrene (n = 1.60)
- Parylene (n = 1.63)
- Mylar (n = 1.85)

The coating must be mechanically and optically qualified for operating temperatures

# FIRST

Update of Instrument Interfaces



DaimlerChrysler Aerospace  
Dornier

Dornier Satellitensysteme GmbH

## **FPU to CVV Integration and Alignment Sequence**

- Integration of the **3 FPU**s on **Optical Bench (OB)** and Alignment to OB reference cube.
- Integration of OB into CVV (w/o upper dome) and **coarse alignment of OB to CVV**.
- Mounting of CVV upper dome and **fine alignment of OB/HIFI wrt**
  - **the 7 LOU-windows** - using LOU/HIFI alignment channels - and
  - the CVV reference cube (when closed, cover window required)
- **Integration of LOU** on lateral CVV framework and alignment to HIFI - using LOU/HIFI alignment channels
- **Alignment check in cold conditions** using cover window and LOU/HIFI alignment channels



## LOU to HIFI Alignment

The following considerations are made to:

- 1.) establish the achievable instrument alignment budget (driver: LOU to HIFI alignment accuracy)
- 2.) derive the requirement on the optically free LOU-windows size based on the requirement to guarantee an optically free 30mm diameter for the LOU beams in operating conditions

Allowed tolerances of LOU to HIFI alignment:

Xrot	Yrot	Zrot	Xpos	Ypos:	Zpos
$\pm 0.027\text{deg}$	$\pm 0.38\text{deg}$	$\pm 0.027\text{deg}$	$\pm 0.8\text{mm}$	$\pm 70\text{mm}$ (uncritical)	$\pm 0.8\text{mm}$

Contributions to alignment budget:

1. Alignment accuracy of HIFI to LOU-windows (through LOU-alignment windows)
2. Alignment accuracy of LOU to HIFI (through the LOU-alignment windows)
3. Structural deviations (ISO experience)



## Alignment accuracy of HIFI to LOU- windows

(Rotations are measured by autocollimating theodolite to the normals on the 2 LOU-alignment windows, located left and right of the 7 LOU-IR windows. Positions are measured relative to reticles on the same windows).

### Xrot and Zrot uncertainties:

1	unknown deviation of HIFI cube normal from nominal position (TBC by HIFI)	$\pm 0.001\text{deg}$ *
2	unknown deviation of LOU-alignment window normal from nom. pos.	$\pm 0.001\text{deg}$
3	measurement uncertainty on HIFI cube	$\pm 0.001\text{deg}$
4	measurement uncertainty on LOU-alignment window normal	$\pm 0.001\text{deg}$
5	HIFI adjustment resolution	$\pm 0.001\text{deg}$
	<b>Total rot uncertainty 2 to 5 (worst case)</b>	<b><math>\pm 0.004\text{deg}</math></b>
	Xrot, Zrot: $\pm 0.004\text{deg}$ correspond to a <b>linear error</b> at LOU-window position of	<b><math>\pm 0.02\text{ mm}</math></b>
	Yrot calculated from position uncertainty beneath ( $\pm 0.5\text{mm}$ on 300mm)	$\pm 0.1\text{deg}$

### Xpos and Zpos uncertainties:

1	unknown deviation of HIFI marks from nominal position (TBC by HIFI)	$\pm 0.1\text{mm}$ *
2	unknown deviation of LOU-alignment window marks from nom.pos.	$\pm 0.1\text{mm}$
3	measurement uncertainty on HIFI marks	$\pm 0.1\text{mm}$
4	measurement uncertainty on LOU-alignment window marks	$\pm 0.1\text{mm}$
5	manufacturing tolerances of window positions to each other	$\pm 0.1\text{mm}$
6	HIFI adjustment resolution	$\pm 0.1\text{mm}$
	<b>Total pos uncertainty 2 to 6 (worst case)</b>	<b><math>\pm 0.5\text{mm}</math></b>

)\* not considered on LOU- window Alignment Budget

**Ypos: uncritical**



**Alignment accuracy of LOU to HIFI**

(Sequence: LOU is adjusted to HIFI through the LOU-alignment windows. Afterwards LOU alignment to LOU-alignment windows is checked).

**Xrot and Zrot uncertainties:**

1	unknown deviation of LOU cube normal from nominal position (TBC by HIFI)	$\pm 0.001\text{deg}$ )**
2	unknown deviation of HIFI cube normal from nominal position (TBC by HIFI)	$\pm 0.001\text{deg}$ )**)**
3	measurement uncertainty on LOU cube	$\pm 0.001\text{deg}$
4	measurement uncertainty on HIFI cube	$\pm 0.001\text{deg}$
5	LOU adjustment resolution	$\pm 0.001\text{deg}$
	<b>Total rot uncertainty (LOU to LOU-windows) 1, 3, 4, 5 (worst case):</b>	$\pm 0.004\text{deg}$
	<b>Total rot uncertainty (LOU to HIFI) 3, 4, 5 (worst case):</b>	$\pm 0.003\text{deg}$
	Xrot, Zrot: $\pm 0.003\text{deg}$ correspond to a linear error at LOU window position of	
	Yrot calculated from position uncertainty beneath ( $\pm 0.4\text{mm}$ on 300mm) ~	$\pm 0.1\text{deg}$

**Xpos and Zpos uncertainties:**

1	unknown deviation of LOU marks from nominal position (TBC by HIFI)	$\pm 0.1\text{mm}$ )**
2	unknown deviation of HIFI marks from nominal position (TBC by HIFI)	$\pm 0.1\text{mm}$ )**)**
3	measurement uncertainty on LOU marks:	$\pm 0.1\text{mm}$
4	measurement uncertainty on HIFI marks:	$\pm 0.1\text{mm}$
5	LOU adjustment resolution:	$\pm 0.1\text{mm}$
	<b>Total pos uncertainty (LOU to LOU-windows) 1, 3, 4, 5 (worst case):</b>	$\pm 0.4\text{mm}$
	<b>Total pos uncertainty (LOU to HIFI) 3, 4, 5 (worst case):</b>	$\pm 0.3\text{mm}$

)\*\* not considered on LOU to LOU-IR window Alignment Budget )\*\* not considered on LOU to HIFI Alignment Budget

Ypos is uncritical



**Structural experience from ISO**

No	ITEM	Rot.	Pos.
1	Expected <b>rotational</b> ground to orbit transition effect on LOU (Zrot only)	+ 0.015deg)*	
2	Settling effects derived from TV test (Xrot, Yrot, Zrot)	0.02deg	
	Sum: 1 + 2 (worst case) (Zrot only)	± 0.035deg)**	^ 0.2 mm
	Sum: 1 + 2 (worst case) (Xrot, Yrot)	± 0.02 deg	^ 0.12mm
3	Expected <b>position</b> ground to orbit transition effect on LOU (assessment)		0.2 mm)*
4	Uncertainty of shrinking effect of vacuum vessel		± 0.2 mm)**
	<b>Total position uncertainty (worst case) (Zrot)</b>		± 0.6 mm
	<b>Total position uncertainty (worst case) (Xrot, Yrot)</b>		± 0.52 mm

)\* due to in orbit change of: 1.) outer CVV temperature and straps pretension, 2.) gravity, 3.) atmospheric pressure (can be compensated to some extent by precompensation)

)\*\* leads to a **linear offset** at the LOU windows of **0.2mm** (~330mm distance)

[for determination of the required LOU window diameters the linear LOU beam offset (translation value) is the crucial figure!]

)\*\*\*)shrinking effect of Al for the expected  $\Delta T$ : 0.004 (4 ‰) is known with an accuracy of 5%; assumed length: 900mm; shrinking consequently: 3,6mm; shrinking uncertainty: 5% of 3,6mm = 0.2mm





## Alignment Budgets

### LOU to HIFI

Allowed tolerances of LOU to HIFI alignment					
Xrot	Yrot	Zrot	Xpos	Ypos:	Zpos
±0.027deg	±0.38deg	±0.027deg)*	±0.8mm	±70mm (uncritical)	±0.8mm
Calculated accuracy of LOU to HIFI alignment					
±0.003deg	±0.1deg	±0.003deg	±0.32mm		±0.32mm
Structural uncertainty					
±0.02deg	±0.02deg	±0.02deg)*	±0.52mm	±0.52mm	±0.4mm)**

)\* precompensation of 0.015 deg applied

)\*\* precompensation of 0.2 mm applied

*0,4 0,3 803 427  
CK*



## Size of LOU windows

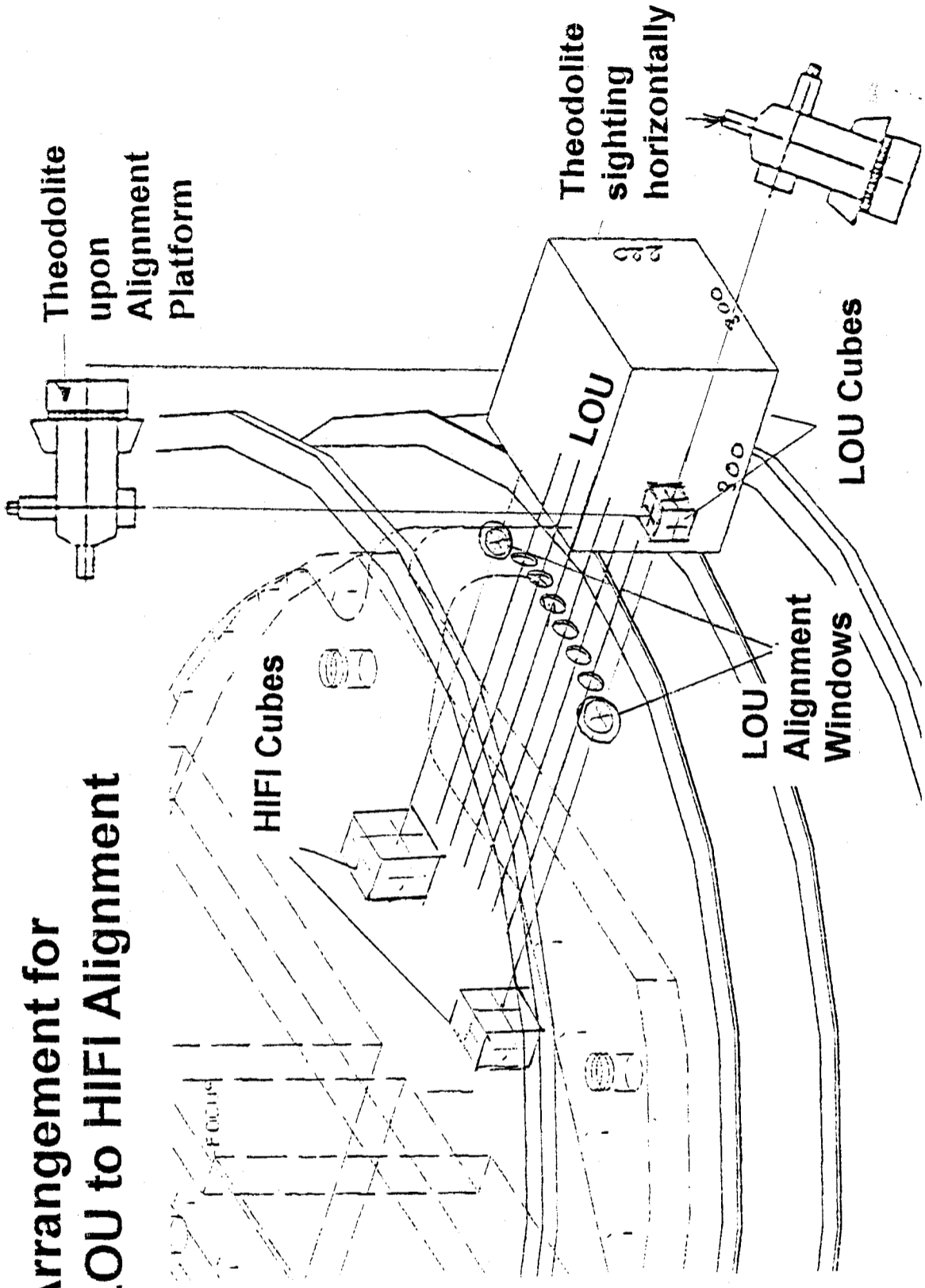
Worst case position uncertainties of LOU Beams to LOU windows (Xpos, Zpos) are derived to obtain the really needed free LOU-IR-windows diameters on the basis of a free 30mm diameter for the LOU beams taking into account possible alignment errors and structural uncertainties.

<b>1</b>	<b>HIFI alignment to LOU-alignment windows on CVV</b>	
	* worst case position error of HIFI to LOU-windows is:	<b>± 0.5 mm</b>
	* contribution of HIFI rotation error to LOU-windows is:	<b>± 0.02mm</b>
<b>2</b>	<b>LOU to HIFI alignment</b>	
	* worst case position error of LOU to HIFI is:	<b>± 0.4 mm</b>
	* contribution of LOU rotation error of LOU to HIFI is:	<b>± 0.02mm</b>
<b>3</b>	<b>Uncertainty derived from ISO structural experience</b>	<b>± 0.6 mm</b>
<b>4</b>	<b>Shift of LOU Beam due to LOU-IR-window tilt (3°)*</b>	<b>± 0.16 mm</b>
	<b>Total uncertainty (worst case)</b>	<b>± 1.7 mm</b>
	<b>Total uncertainty (worst case) inclusive margin</b>	<b>± 2 mm</b>
	To guarantee a free LOU-window diameter of 30mm at operational conditions, the actually needed LOU-window size has to be increased in diameter by 4 mm	<b>34 mm</b>

)\* Quartz window 6mm thick, 3deg tilt,  $n = 2.1$  for far IR ( $v = \epsilon \cdot d^* \cdot (n-1)/n$ ) (can be precompensated!)



## 4. Arrangement for LOU to HIFI Alignment



## **FIRST**

Update of Instrument Interfaces

---



**DaimlerChrysler Aerospace**  
Dornier

Dornier Satellitensysteme GmbH

LO waveguides:

- Have to be elastic to compensate
  - differential movement between LO and CVV and between CVV and SVM during vibration
  - different thermal expansion between waveguides and CVV, SVM
- Minimum heatinput to CVV (shall be  $< 0.1$  W) by using suited materials (e.g. SST or CFRP) and insulating from CVV

Next steps:

- Mechanical analyses of LO support
- Thermal Analyses of LO (taking into account TMM of B. Collaudin)

# FIRST

Update of Instrument Interfaces



DaimlerChrysler Aerospace  
Dornier

Dornier Satellitensysteme GmbH

## 8. Input to IID A update (proposal)

- Mounting, Dismounting of all instruments shall be independent from each other
- The mounting interfaces areas of the OB (diameter 120 mm) shall be accessible at all integration levels (i.e. not covered by instruments)
- No forces due to thermal expansions of instruments shall be applied to the optical bench

## 9. AOB