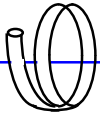


CURRENT SYSTEM DESIGN FOR THE CEA FILLED ARRAYS OPTION

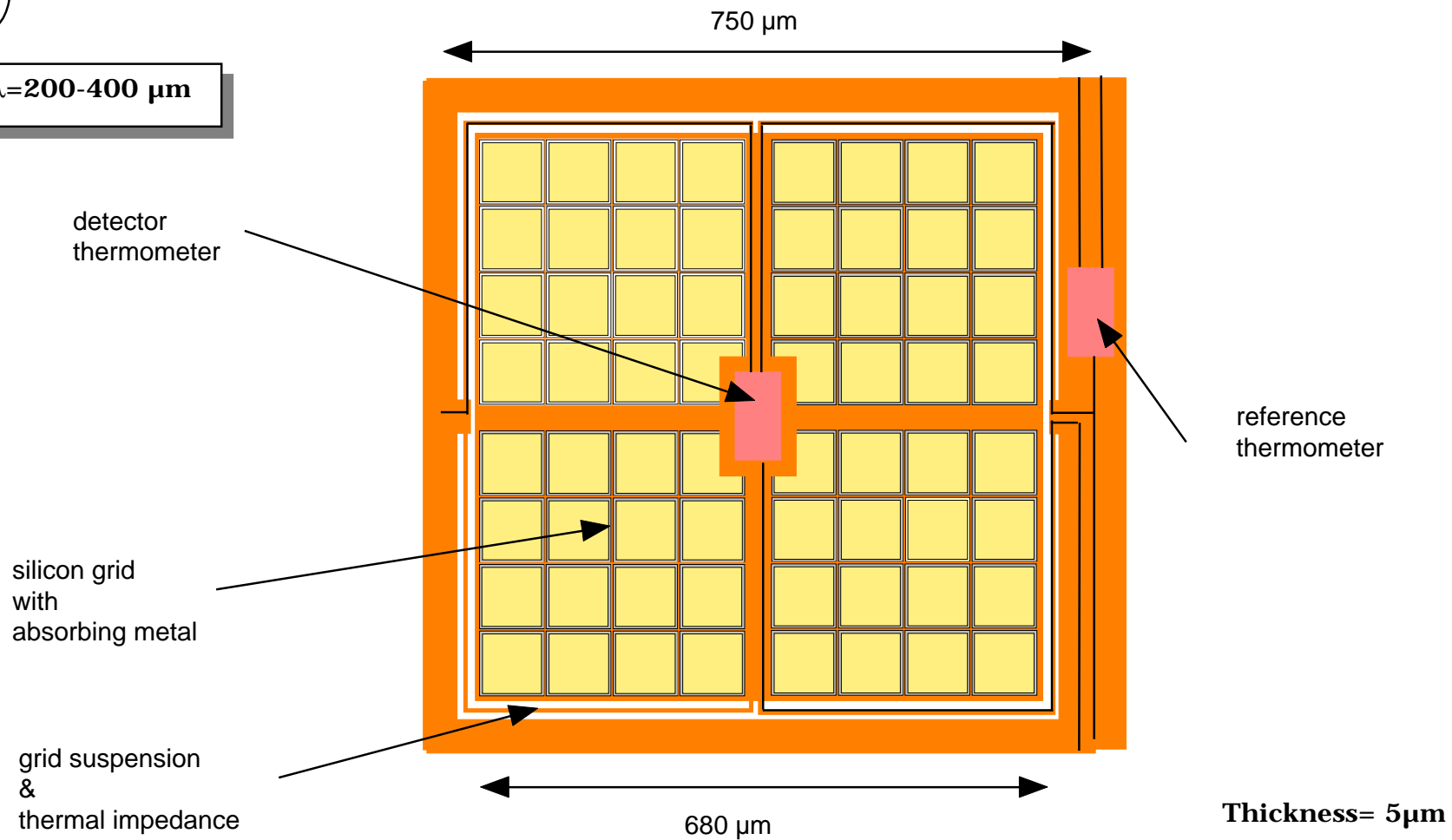
L. RODRIGUEZ
SAp
cea-Saclay
lrodriguez@cea.fr

March 5 th 1999

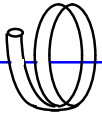


THE NEW PIXEL DESIGN

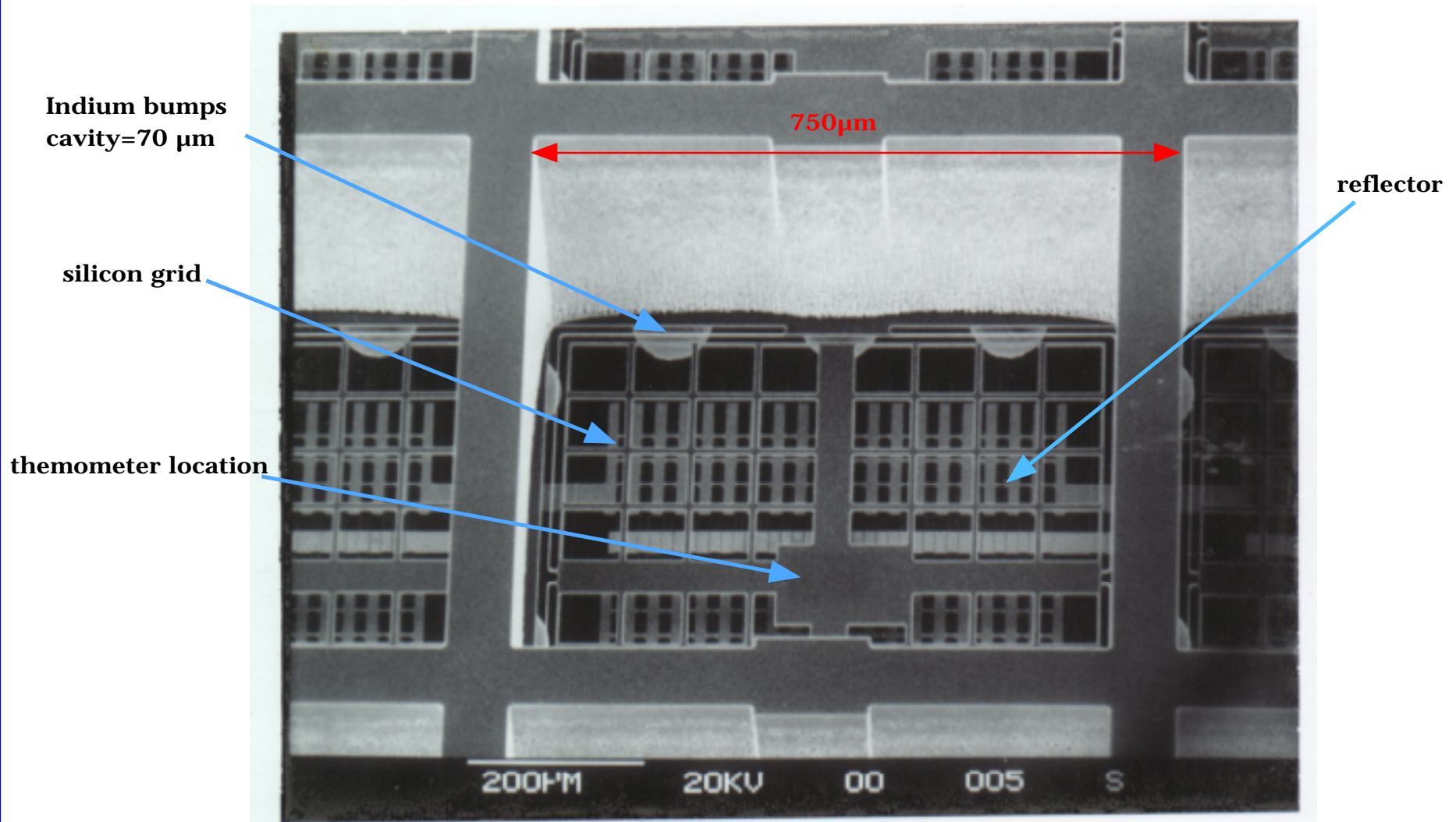
$\lambda=200-400 \mu\text{m}$



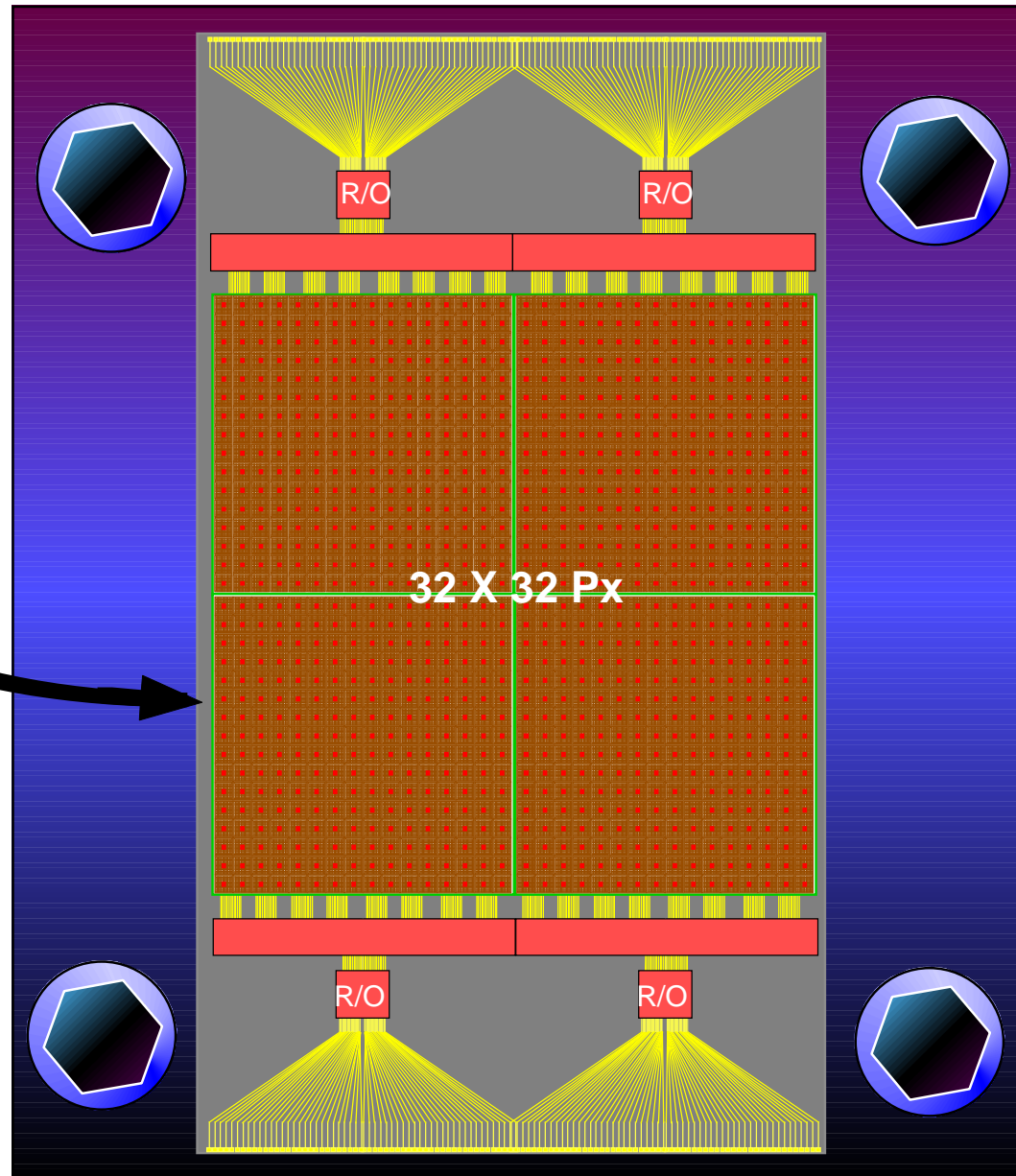
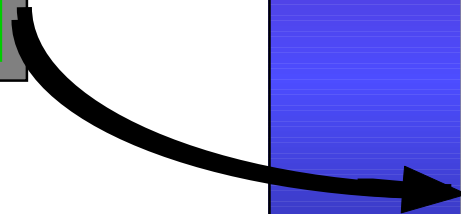
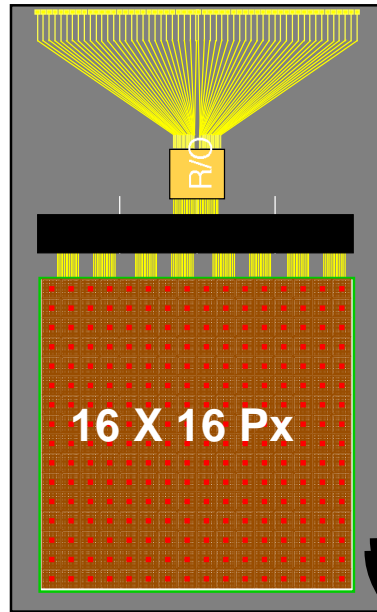
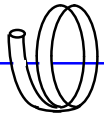
- Main changes:**
- ➔ Pattern step size reduction 900->750 μm
 - ➔ Slim grid to lower heat capacity
 - ➔ Smaller suspension beams to decrease τ_{ther}
 - ➔ Stripped reflector to lower electrical Capacitance

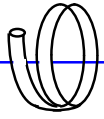


THE NEW PIXEL DESIGN

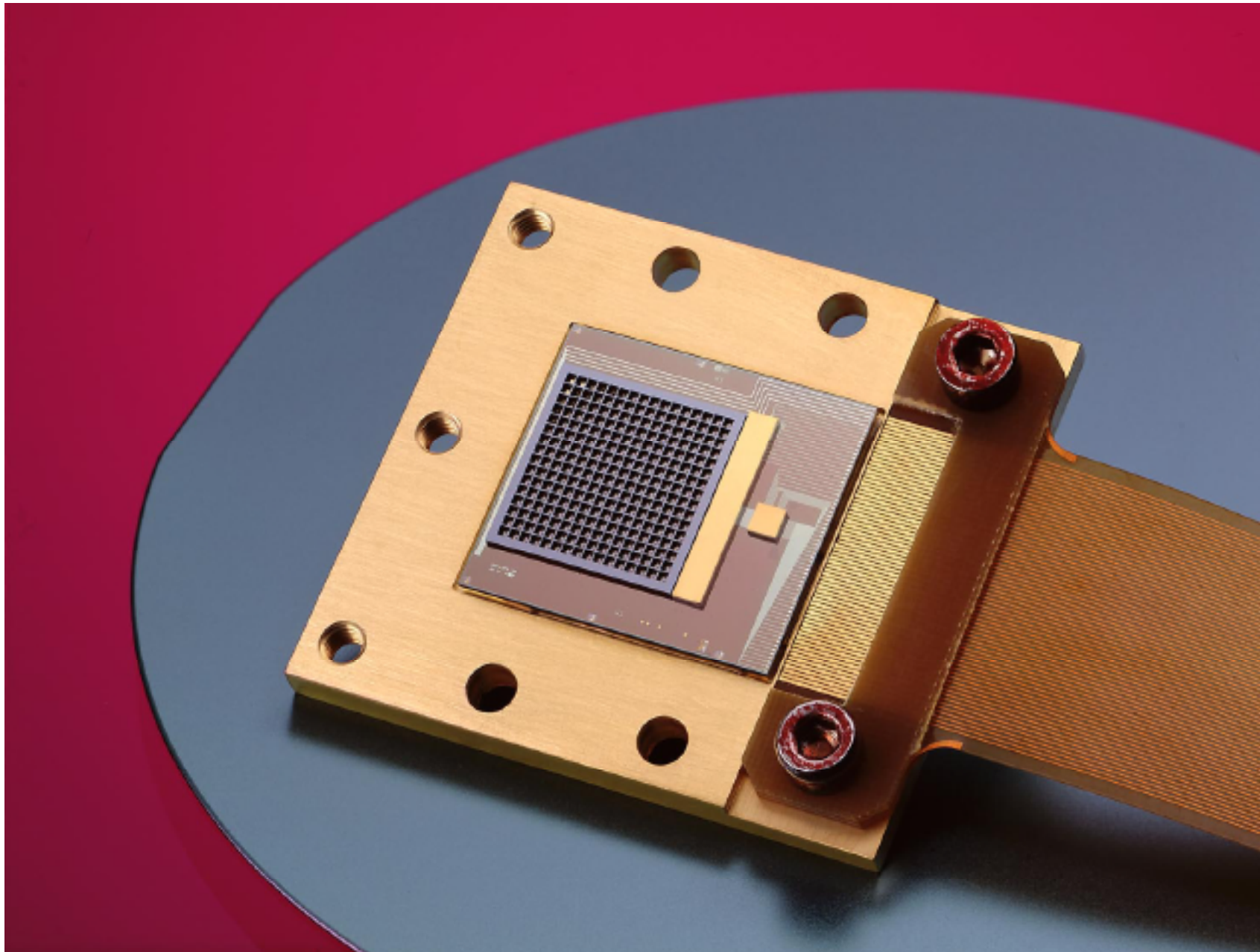


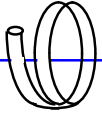
THE ARRAY CONCEPT



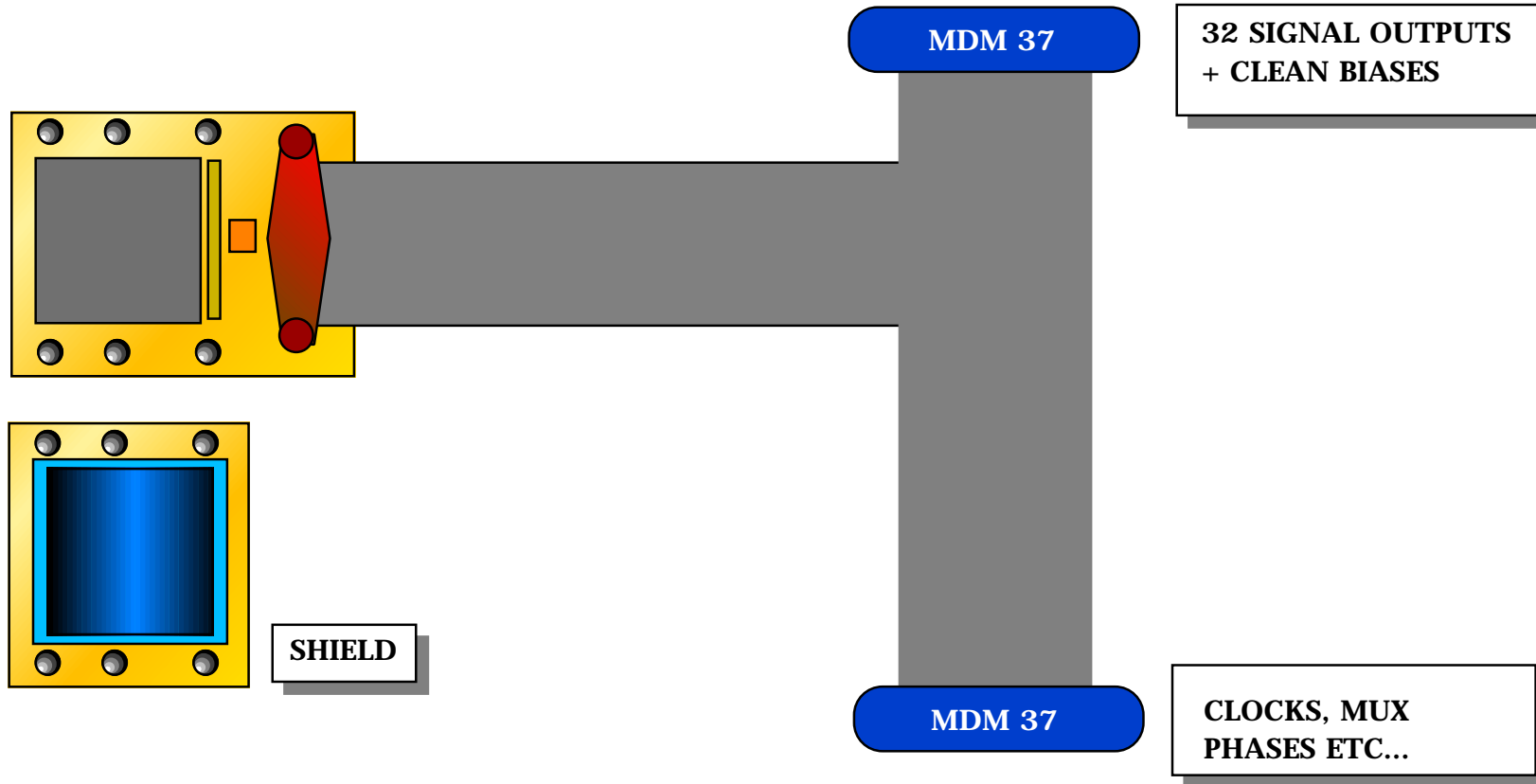


THE ARRAY CONCEPT

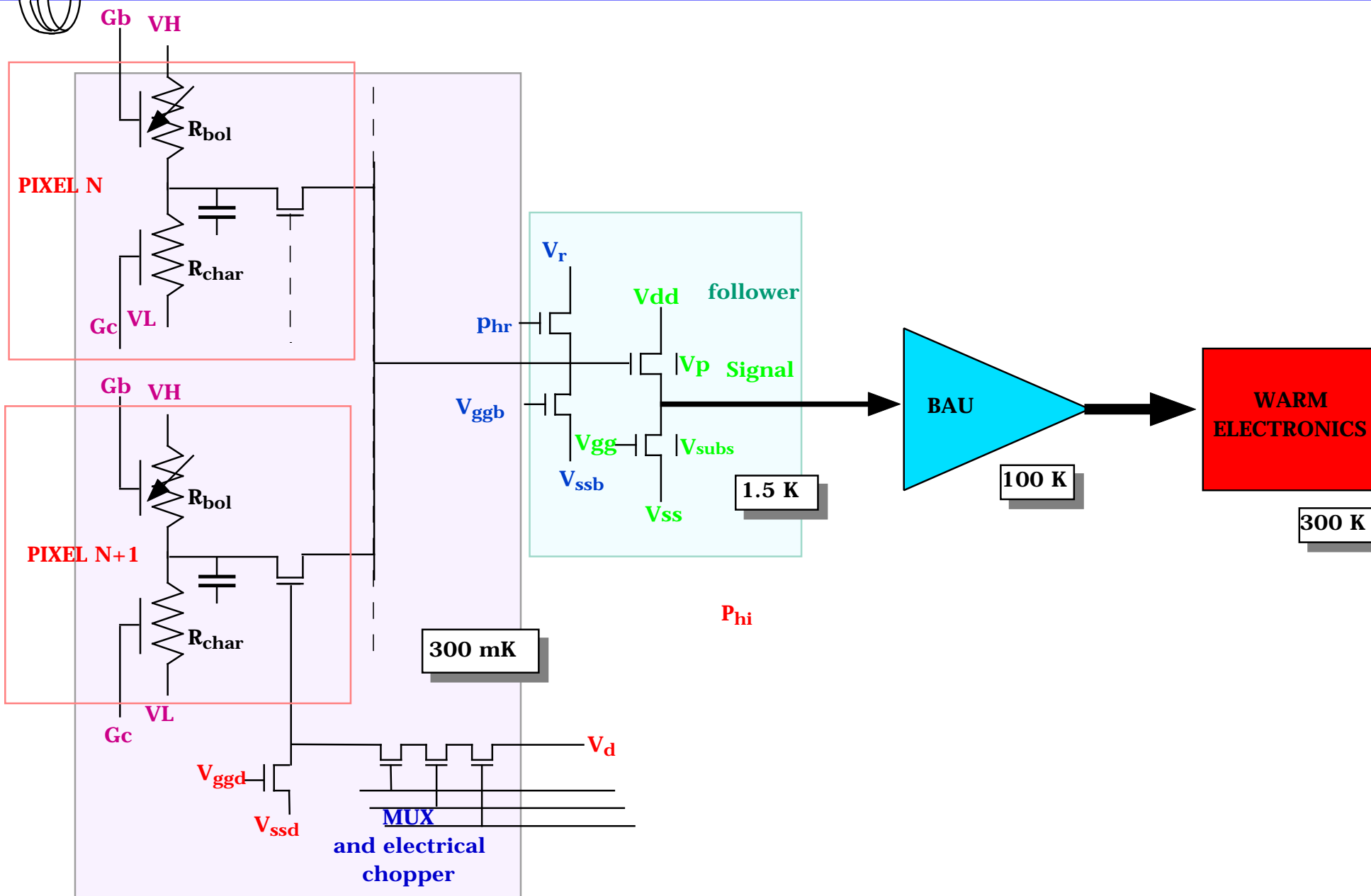
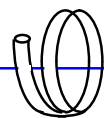


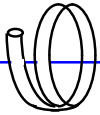


WIRING



WIRING





WIRING 4 X 4 ARCMIN 8->1 MUX

IMAGER

3 ARRAYS

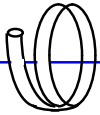
	32 X 32	24 X 24	16 X 16
ADDRESS	5 wires	5 wires	5 wires
DC i/o	2 “	2 “	2 “
bias	1 “	1 “	1 “
bus bias	1 “	1 “	1 “
DC bias	1 “	1 “	1 “
ref bias	1 “	1 “	1 “
bias	1 “	1 “	1 “
outputs +2 ref	132 ”	132 “	36 “
main bias	2 “	2 “	2 “
followers bias	1 “	1 “	1 “
substrate bias	1 “	1 “	1 “
frame bias	2 “	2 “	2 “
ground ref	1 “	1 “	1 “
bridge bias	2 “	2 “	2 “
heater bias	2 “	2 “	2 “
TOTAL	155	155	59 wires

SPECTROMETER

2 ARRAYS

	16 X 16	12 X 12
ADDRESS	5 wires	5 wires
DC i/o	2 “	2 “
bias	1 “	1 “
bus bias	1 “	1 “
DC bias	1 “	1 “
ref bias	1 “	1 “
bias	1 “	1 “
outputs +2 ref	36 “	36 “
main bias	2 “	2 “
followers bias	1 “	1 “
substrate bias	1 “	1 “
frame bias	2 “	2 “
ground ref	1 “	1 “
bridge bias	2 “	2 “
heater bias	2 “	2 “
TOTAL	59	59 wires

GRAND TOTAL = 487 WIRES



WIRING FOR 8 X 4 ARCMIN 8->1 MUX

IMAGER

3 ARRAYS

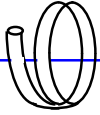
	64 X32	48 X 24	32 X 16
ADDRESS	6 wires	6 wires	5 wires
DC i/o	2 “	2 “	2 “
bias	1 “	1 “	1 “
bus bias	1 “	1 “	1 “
DC bias	1 “	1 “	1 “
ref bias	1 “	1 “	1 “
bias	1 “	1 “	1 “
outputs +2 ref	262 ”	262 “	70 “
main bias	2 “	2 “	2 “
followers bias	1 “	1 “	1 “
substrate bias	1 “	1 “	1 “
frame bias	2 “	2 “	2 “
ground ref	1 “	1 “	1 “
bridge bias	2 “	2 “	2 “
heater bias	2 “	2 “	2 “
TOTAL	286	286	93 wires

SPECTROMETER

2 ARRAYS

	32 X 16	24 X 12
ADDRESS	5 wires	5 wires
DC i/o	2 “	2 “
bias	1 “	1 “
bus bias	1 “	1 “
DC bias	1 “	1 “
ref bias	1 “	1 “
bias	1 “	1 “
outputs +2 ref	70 “	70 “
main bias	2 “	2 “
followers bias	1 “	1 “
substrate bias	1 “	1 “
frame bias	2 “	2 “
ground ref	1 “	1 “
bridge bias	2 “	2 “
heater bias	2 “	2 “
TOTAL	93	93 wires

GRAND TOTAL = 851 WIRES



OPERATION

OPERATING TEMPERATURE :

Designed to work at 0.3 K under a 1 pW optical flux.

POWER DISSIPATION AT 2 K STAGE (per array) :

-Heat load on ^3He Fridge by conduction $5\mu\text{W}/\text{array} \rightarrow 1.25 \text{ mW total}$
(assuming a 1/50 ^3He fridge efficiency).

-Bias on read out circuit $\rightarrow 1\text{-}2 \text{ mW}$,

IONISING RADIATION EFFECT :

To be determined on a accelerator device on prototype bolometer. From ISOCAM a thumb rule gives 1 proton/minute /pixel and the deposited energy is around 150 eV ($2.4 \text{ E-}11 \mu\text{J}$).

PIXEL ANGULAR RESPONSE :

Large $\pm 40^\circ$. Necessity of efficient baffling, but small size array.

CROSS TALK :

N/A to the nearest pixel if sampling the PSF. Important for electrical crosstalk between distant pixels.

ACHIEVABLE ARRAY SIZE :

64 X 32 Pixels ($3 \times 3 \text{ cm}^2$).

TELEMETRY :

TBD

Important inputs for requirements are still missing :

→ FTS operation

→ Chopper operation

• Scan Mode with AOCS

• On-board datation : resolution/accuracy

→ On-board processing : instrument data rate / telemetry data rate

• Operating modes

• Degraded modes

• Temperature regulation (He³)

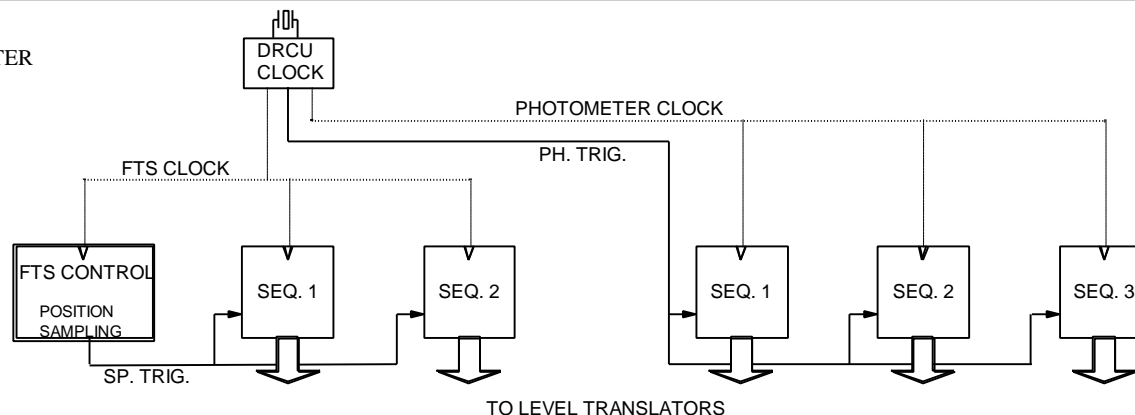
• H/K specification (temperature, else) : how many/accuracy/rate

FIRST/SPIRE

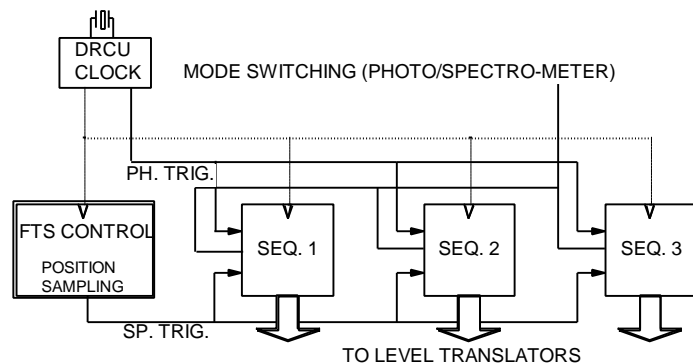
DRCU SYNCHRONIZATION TREE vs FPU CONFIGURATIONS

- 1 BOLOMETER ARRAYS FOR PHOTOMETER AND SPECTROMETER ARE DIFFERENT (BASELINE)

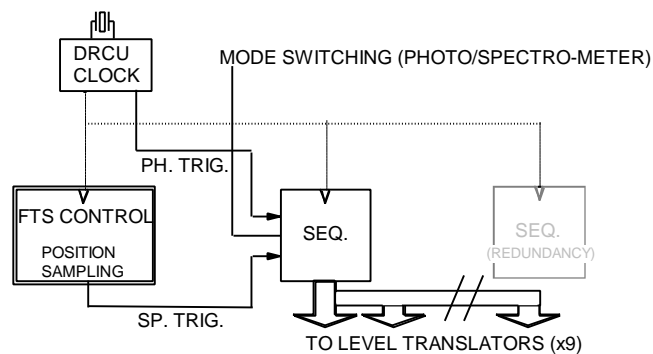
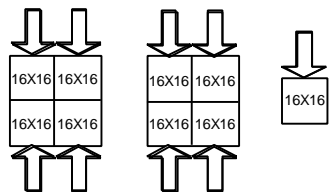
NOTE : FTS & PHOTOMETER CLOCKS MAY BE IDENTICAL



- 2 PHOTOMETER AND SPECTROMETER SHARE THE 3 BOLOMETER ARRAYS (DIFFERENT ARRAY SIZES)

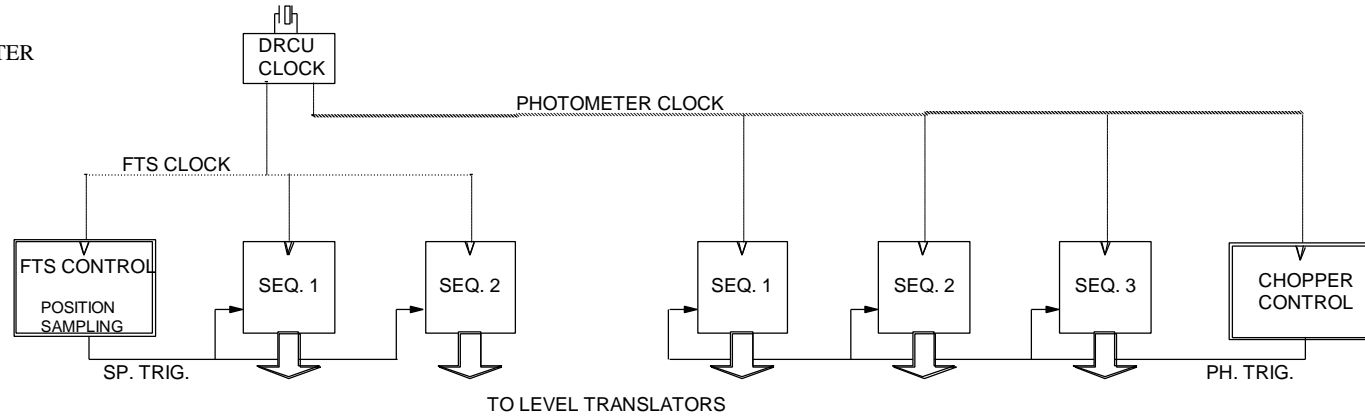


- 3 PHOTOMETER AND SPECTROMETER SHARE THE 3 BOLOMETER ARRAYS (BASED ON MULTIPLE 16x16 ELEMENTARY ARRAYS).
CEA + TES (?) OPTIONS

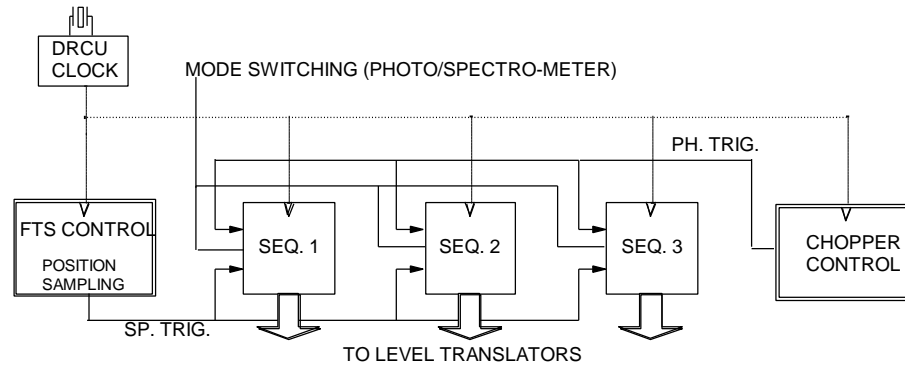


- 1 BOLOMETER ARRAYS FOR PHOTOMETER AND SPECTROMETER ARE DIFFERENT (BASELINE)

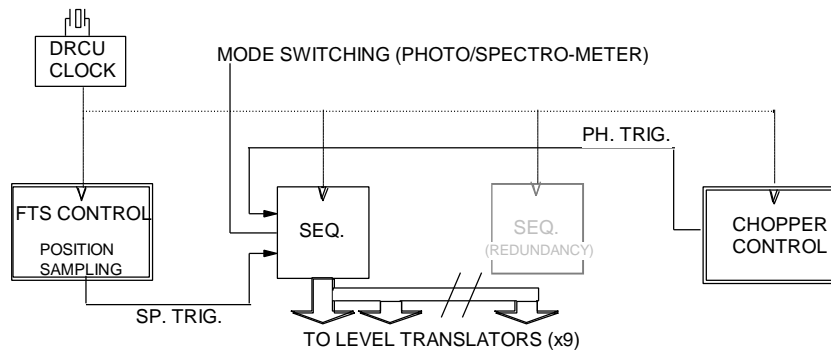
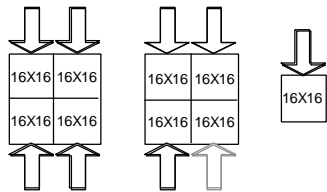
NOTE : FTS & PHOTOMETER CLOCKS MAY BE IDENTICAL



- 2 PHOTOMETER AND SPECTROMETER SHARE THE 3 BOLOMETER ARRAYS (DIFFERENT ARRAY SIZES)



- 3 PHOTOMETER AND SPECTROMETER SHARE THE 3 BOLOMETER ARRAYS (BASED ON MULTIPLE 16x16 ELEMENTARY ARRAYS). CEA + TES (?) OPTIONS



SPIRE Acquisition Rates

Photometer

Central Wavelength μm	Theoretical array sizes	Practical array sizes	Number of pixel	Acquisition rate rate Hz^{-1}	Number of acq. /s	Number of bits	Data rate bits/s
250	32x32	32x32	1024	40	40960	14	573440
350	24x24	32x32	576	40	23040	14	322560
500	16x16	16x16	256	40	10240	14	143360
Total (average) :							1039360

1 : Assuming a 2 time oversampling

Minimum compression factor compared to **40 kbits/s** (200 kbits/s-TBC) : **26 (5)** \rightarrow **Image rate = 1.5 /s**

Real compression factor will take into account data format (i.e. 3 bytes / pixel)

Spectrometer

Wavelength μm	Theoretical array sizes	Practical array sizes	Number of pixel	Acquisition rate rate Hz^{-1}	Number of acq. /s	Number of bits	Data rate bits/s
200-300	16x16	16x16	256	40	10240	14	143360
300-600	12x12	16x16	144	40	5760	14	80640
Total (average) :							224000

1 : Assuming a 2 time oversampling

Minimum compression factor compared to **40 kbits/s** (200 kbits/s-TBC) : **6 (1.12 !)** \rightarrow **Interferogram rate = 1 / 240 s**

Real compression factor will take into account data format (i.e. 3 bytes / pixel)

- Dominated by the SPECTROMETER high resolution mode memory requirements (if full spectrum co-addition).
- Assuming a 40 second scanning duration the amount required to store one scan is :

Number of arrays	Number of acq. /s	Quantization in bytes	Memory size in bytes
2	10240+5760	2	1280000

- Total Memory required includes :

- currently acquired spectrum : 2 bytes
- sum of the spectra (6) : 3 bytes
- deglitching table : 1 bytes

Total : 3.84 Mbytes

- Oversampling (factor of 1.5) may be required → Total with margin : 6 Mbytes

→ Memory could be saved if data reduction is based on error-free data compression algorithm instead of performing co-additions : no more need for “sum of spectro” & “deglitching” buffers.

Assumption : microprocessor running at 20 MHz (clock period = 50 ns)

• **CEA Option : 125824 16-bit words /s or one pixel every 8 μ s**

→ 1 data acquisition every 160 microprocessor cycles

→ DMA in not mandatory

• **JPL (Feed-Horn) Option : modulated and oversampled data gives 4 Mwords/s or one pixel every 238 ns**

→ 1 data acquisition every 4-5 microprocessor cycles

→ requires DMA in the Signal Processing Unit

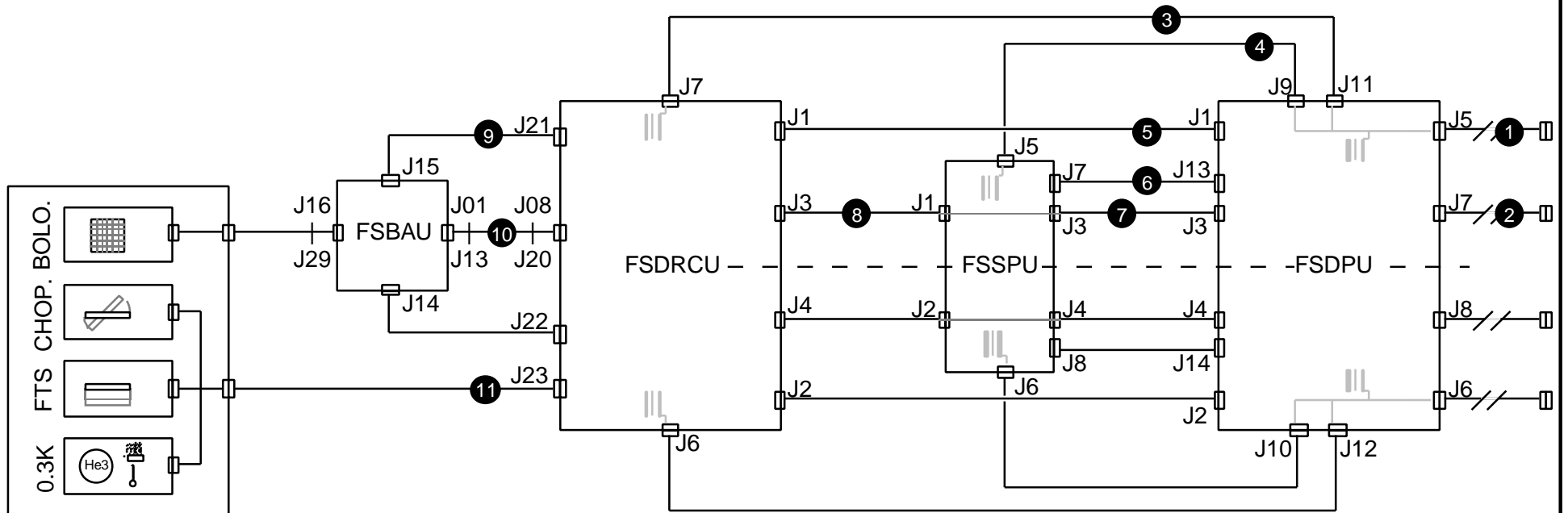
• **GSFC (T.E.S.) Option : oversampled data gives 64 Mwords/s or one pixel every 15 ns**

→ incompatible with 50 ns clock period.

→ requires data pre-processing (using hardwired electronics) in the DRCU to reduce Signal Processing input data rate (i.e. factor of 10).

→ requires DMA in the Signal Processing Unit.

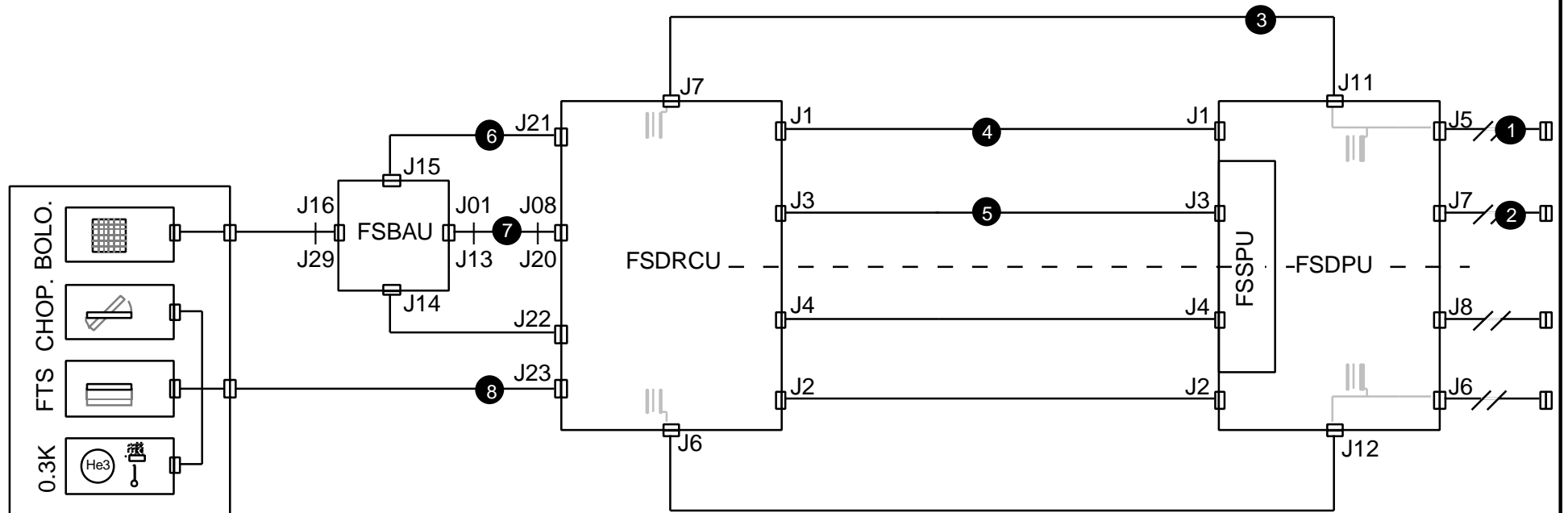
SPIRE ELECTRICAL CONFIGURATION SPU H/W & S/W OPTION



HARNESS TYPES :

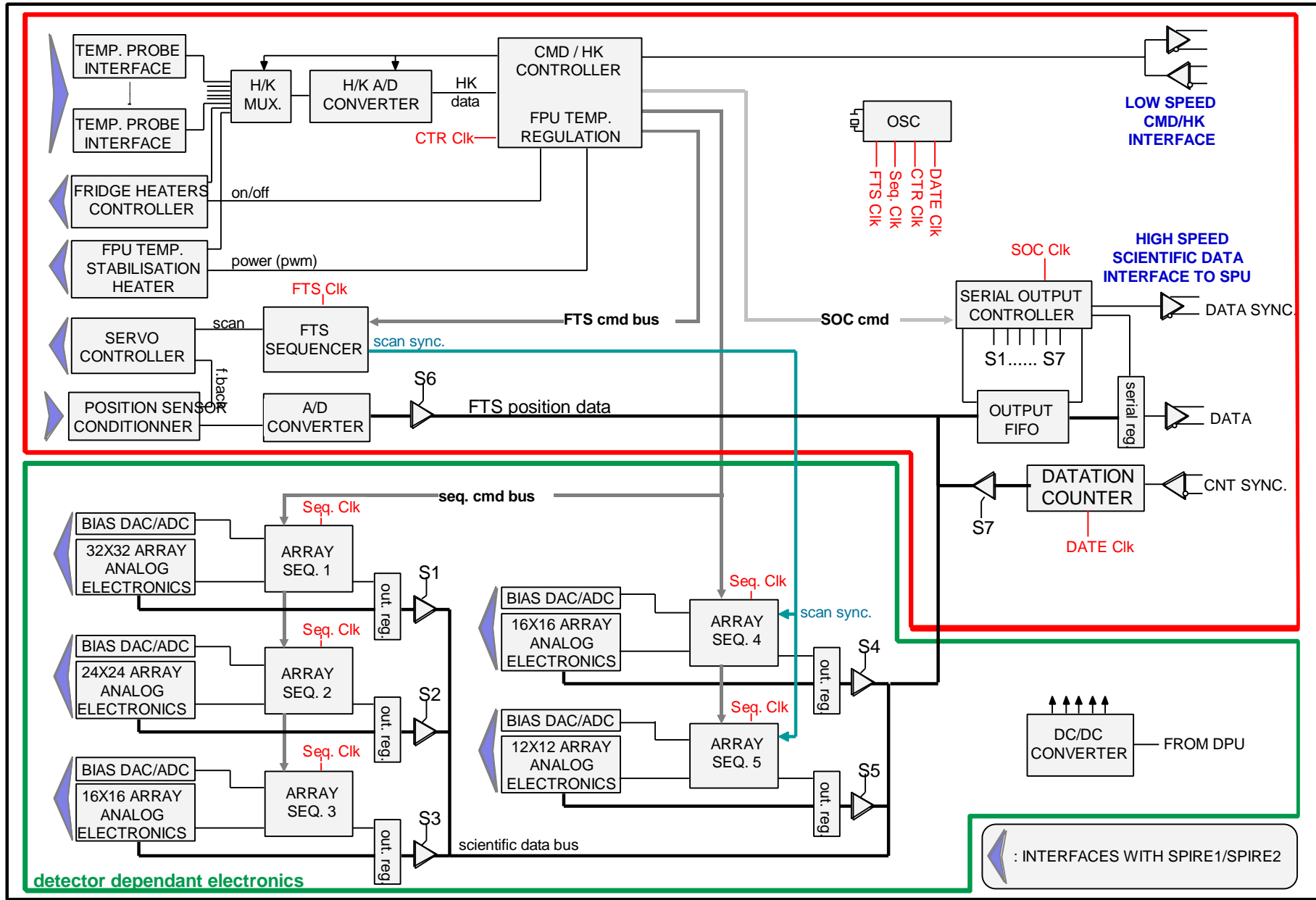
- | | | |
|----------------|---------------|---------------------------|
| ① SPIRE CMD/HK | ⑤ DRCU CMD/HK | ⑨ BAU SPWR |
| ② SPIRE PPWR | ⑥ SPU CMD/HK | ⑩ BAU SIGNAL |
| ③ DRCU PPWR | ⑦ SPU DATA | ⑪ MECHANISMS CONTROL + HK |
| ④ SPU PPWR | ⑧ DRCU DATA | |

SPIRE ELECTRICAL CONFIGURATION SPU SW ONLY OPTION



- | | |
|----------------|--------------------------|
| ① SPIRE CMD/HK | ⑤ DRCU DATA |
| ② SPIRE PPWR | ⑥ BAU SPWR |
| ③ DRCU PPWR | ⑦ BAU SIGNAL |
| ④ DRCU CMD/HK | ⑧ MECHANISM CONTROL + HK |

DRCU Block Diagram



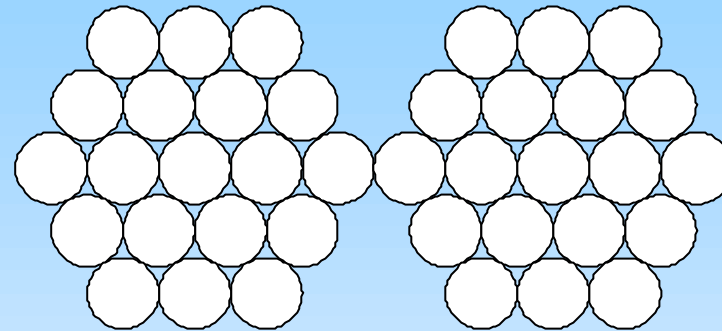
System Trade-offs for number of detectors

What are the limiting factors which determine how far we could extend the effective field of view of SPIRE?

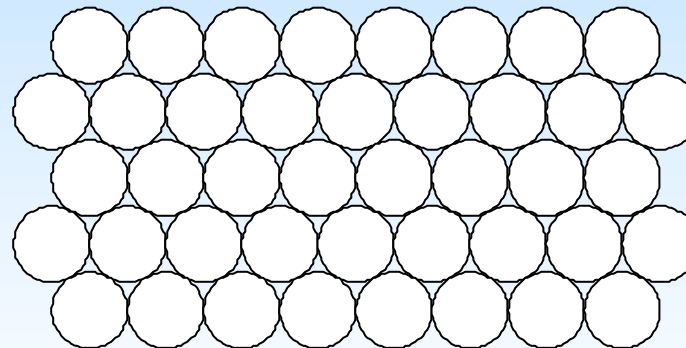
See scientific requirements -

- in summary, a 4'x8' field of view using $2F\lambda$ horns should be given serious consideration for optimum mapping speed, if the horn option is chosen
- Same applies with filled arrays, so that ideally 2 sets of 32x32, 16x16 and 12x12 arrays would be needed

Horn arrays: filling 4' x 8' field:



2x19 arrays= 38 detectors

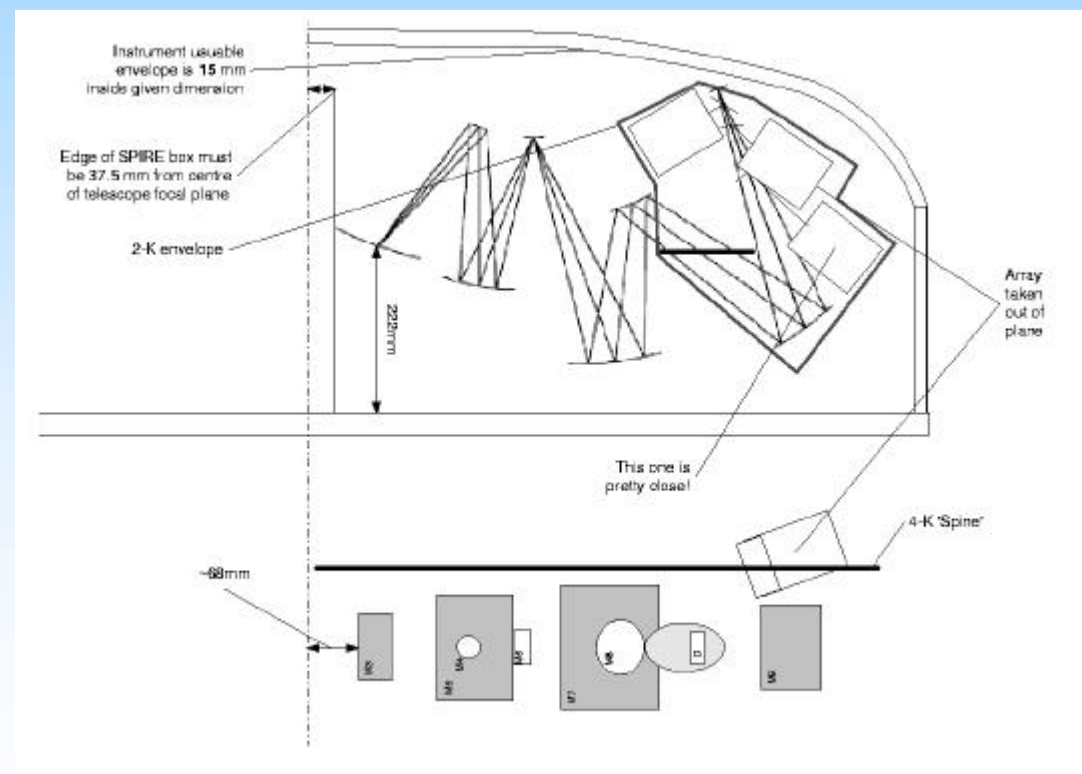


42 detectors

We will need just over twice as many detectors, and conductors, amplifiers....

Optics:

It looks like 4 by 8 arcminutes is possible in the space envelope without the mirrors becoming too large, but there may be problems with space for the photometer arrays



Signal wiring issues:

- **Inside FPU**
 - Connectors- 37, 51 or 100 way ?
 - Ribbon cables - Kapton or Woven?

- **Cold Interconnect Harness: FPU to BAU, BAU to DRC and FPU to DRC (ESA responsible)**
 - Connectors- 37, 51 or 100 way ?
 - Conventional wire bundles or Ribbon cables - Kapton or Woven?

Inside FPU

0.3-4K	4'x4' Option	4'x8' Option		
Description	Conductors	Conductors	Shields	Type
P250Signal	195	390	0/TBC	Kapton
P350Signal	123	246	0/TBC	Kapton
P500Signal	69	138	0/TBC	Kapton
S350Signal	72	144	0/TBC	Kapton
S600Signal	48	96	0/TBC	Kapton
TCBIAS	8	8	4/TBC	SST
TCSIG	8	8	4/TBC	SST
TOTAL	523	1030		

**Using twisted pairs separated by ground wires.
Detector bias on Kapton ribbons**

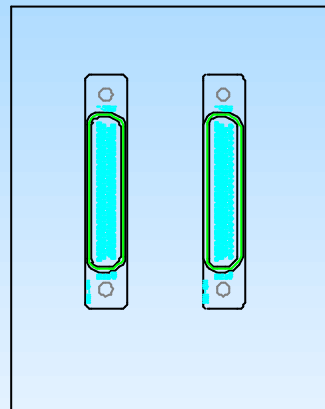
Cold Interconnect Harness

15-300K	4'x4' Option	4'x8' Option		
Description	Conductors	Conductors	Shields	Type
P250Signal	183	366	0/TBC	Kapton
P350Signal	111	222	0/TBC	Kapton
P500Signal	57	114	0/TBC	Kapton
S350Signal	60	120	0/TBC	Kapton
S600Signal	36	72	0/TBC	Kapton
Det Bias	20	40	10	SST
TCBIAS	8	8	4/TBC	SST
TCSIG	8	8	4/TBC	SST
TOTAL	483	950		

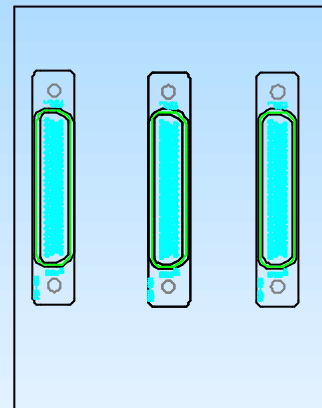
Using twisted pairs separated by ground wires, which may not be necessary, post JFET

Connector options:

61 detector array: 130 pins
2 x 100w MDM



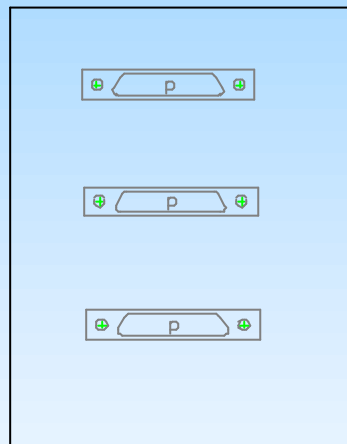
122 detector array: 255 pins
3 x 100w MDM



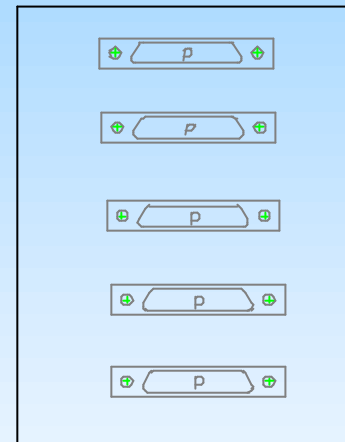
Typical space available: 95 x 75 mm

Connector options:

61 detector array: 130 pins
3 x 51w MDM



122 detector array: 255 pins
5 x 51w MDM



Using ESA-approved 37w MDMs will be even worse!

Data Rate

- **Whether we can use a higher data rate depends on the availability of x-band transponders**
- **Moving to 4'x 8' may be possible even with the current down-link: at present we assume factor 3 compression to get 25 kbps data rate.**
- **PACS assume factor 80 compression to give 40 kbps!**

Mass of array assemblies

- Array Mass will double
- Optics mass will increase
- BAU mass will increase
- JFET or cold electronics will increase in mass

Thermal loads

- 0.3K load due to supports will double:

$$Q(F, M, E_y, k_i) \equiv \frac{3 \cdot (M \cdot k_i) \cdot (2 \cdot \pi \cdot F)^2}{E_y}$$

- loads due to wiring will also double

Power consumption of cold & warm electronics

SPIRE

- **Will double**
- **DRC 25 W budget is already tight for TES and horn options**
- **-but ESA say there is some room for manoeuvre - we could push towards 100W?**

Increased Costs:

- **Manufacture:**
 - **Horns**
 - **Arrays**
 - **Cables**
 - **Load resistors**
 - **Read-out electronics**
- **Integration Time**
- **Testing Time**